Using numerical relativity waveforms in gravitational wave searches

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Introduction

- Two classes of searches relevant for BBH systems
 - Un-modeled bursts
 - Matched filter searches (inspiral/ringdown)
- We want to improve these pipelines using information from numerical relativity
 - How do existing pipelines perform with NR waveforms?
 - Can we use information from NR to make our vetoes and coincidence tests better? This is important for improving false alarm rates and thus the sensitivity of our searches.
 - Immediate need for ongoing S5 searches. Probably not feasible to do full NR based merger search for S5, but we need to know how well current searches perform.
 - How important are spins and higher modes for detection?
 - Should we use matched filtering to search for the merger?
- Many of these issues can be addressed before we have solved the PN-NR matching and template bank problems

Introduction

- To start addressing these questions, we need to be able to inject NR waveforms
- Injection studies (hardware and software) are a crucial ingredient in all GW searches
- Provides end-to-end test and comparisons of pipelines
- Powerful way to study complicated data analysis questions which are hard to understand analytically – non-stationarities and non-gaussianities, efficiencies, upper limits, parameter space structure etc.
- A large fraction of the time for carrying out a search is spent in injection studies

Exchanging data with NR groups

- To start injection studies, we need to exchange numerical waveforms with NR groups
- Data analysis people involved directly: D. Brown,
 S. Fairhurst, R. Kopparapu, B. Krishnan, R.A. Mercer,
 L. Santamaria, J.T. Whelan
- Document specifying data formats for the exchange: arXiv:0709.0093 [gr-qc]
- Is this a sensible data format? Comments and suggestions are welcome

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The data format

The time and the waveforms scale with M

$$h_{ij} = A_{ij} \frac{M}{r} + \mathcal{O}\left(r^{-2}\right), \qquad h_{ij} = h_+ \left(\mathbf{e}_+\right)_{ij} + h_{\times} \left(\mathbf{e}_{\times}\right)_{ij}$$

We work with the mode decomposition in spin weighted spherical harmonics:

$$h_{+} - ih_{\times} = \frac{M}{r} \sum_{\ell=2}^{\infty} \sum_{m=-\ell}^{\ell} H_{\ell m}(t)^{-2} Y_{\ell m}(\iota, \phi).$$

$$rh^{(\ell m)}_+(t) - irh^{(\ell m)}_\times(t) \equiv MH_{\ell m}(t)$$
.

Provide $rh_{+}^{(\ell m)}(t)$, $rh_{\times}^{(\ell m)}(t)$ and a metadata file for each NR simulation

What we intend to do to the data

- ASCII files are converted to "frame files" (one for each simulation) and metadata is contained in the file itself
- Intention is to distribute the frame files within the LSC by usual mechanism used for LIGO data itself ⇒ every LSC member can access the waveforms
- Use these frame files to create a population of NR signals with different masses, sky positions, orientations at different times and inject into pipelines
- Immediate aim is to set up the infrastructure to do these injections
- We want to get started with this as soon as possible first using simulated data and later with real data

A toy example with the inspiral searches

- Inject a polulation of 10 hybrid waveforms with
- $3M_{\odot} < m_1, m_2 < 35M_{\odot}$
- $6M_{\odot} < m_1 + m_2 < 40M_{\odot}$
- 1*Mpc < d <* 100*Mpc*
- Systems assumed to be optimally oriented

Try to recover them with

- TaylorT1 2-PN template bank
- 2100 templates, minimal match = 0.97
- $3M_{\odot} < m_1 < 35M_{\odot}$
- $m_1 + m_2 < 40 M_{\odot}$

Want to look at parameter estimation accuracy

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The inspiral pipeline

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The template bank used



Chirp-mass estimation



Symmetric mass ratio estimation



The effective distance



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Conclusions

- Even before solving the template construction problem, input from numrel can help improve our searches
- The ability to inject numrel waveforms is very important, and we have made progress in this direction
- We would like to start this data exchange as soon as possible
- Next....
 - NR injections in a burst pipeline Adam Mercer
 - Status of ringdown searches Jolien Creighton

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