

Probing Neutrinos with BBN and the CMB

FORENSIC COSMOLOGY : PROBING NEUTRINOS WITH BBN AND THE CMB

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Abundances of D, ^3He , ^7Li

Are RATE Limited

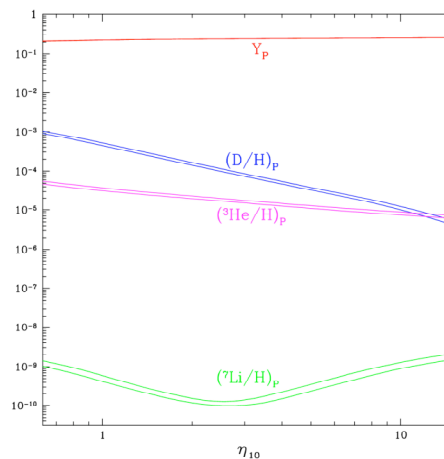
□ D, ^3He , ^7Li Are

Possible Baryometers

$$\eta \equiv n_N / n_B$$

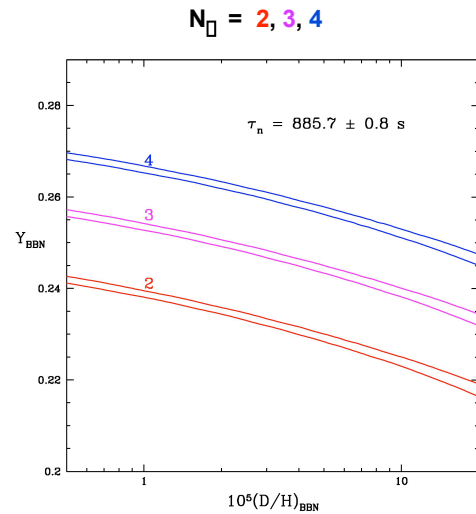
$$\eta_{10} = 10^{10} \eta = 274 \eta_B h^2$$

BBN – Predicted Primordial Abundances



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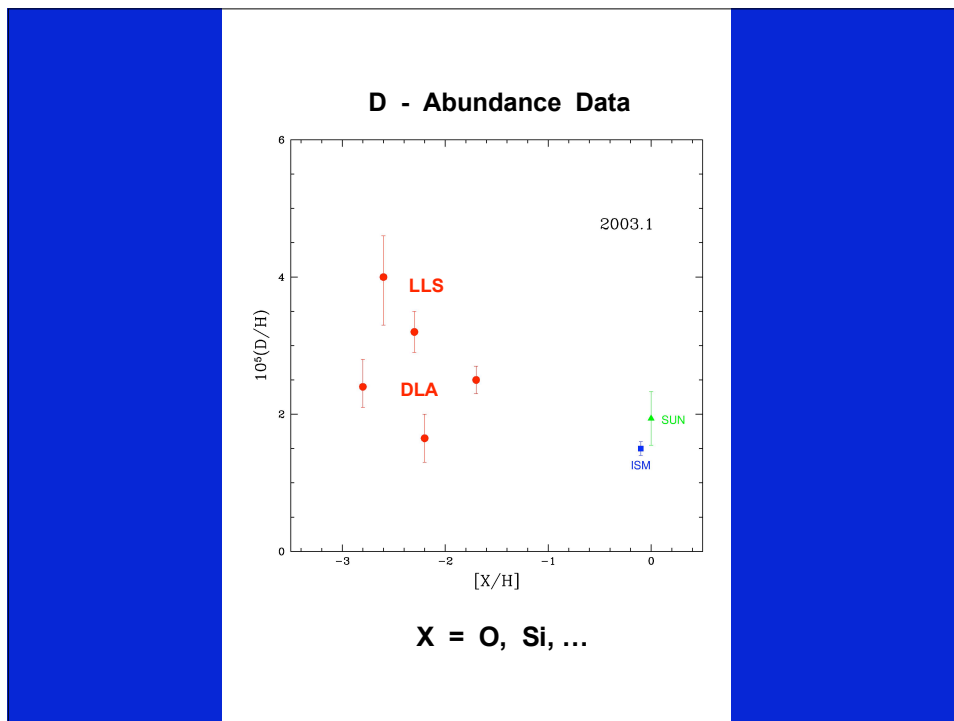
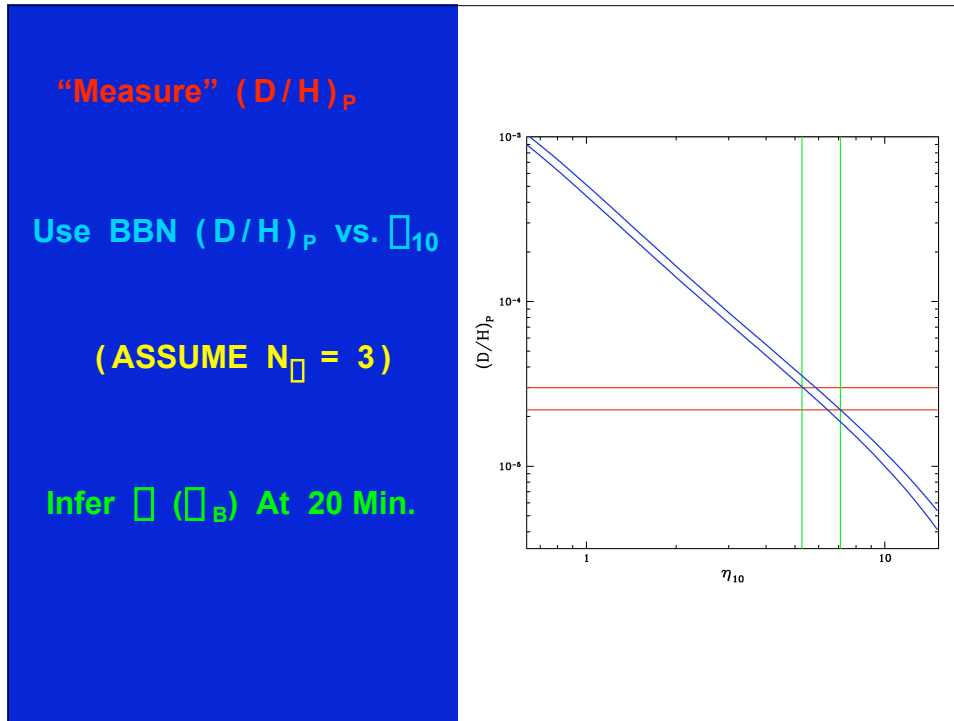
- ^4He (mass fraction Y)
Is NOT Rate Limited
- ^4He IS n/p Limited \square
 Y Is Sensitive To The
EXPANSION Rate (H)
- $S \equiv H'/H \equiv (1 + 7 \square N_\nu / 43)^{1/2}$
 $N_\nu \equiv 3 + \square N_\nu$



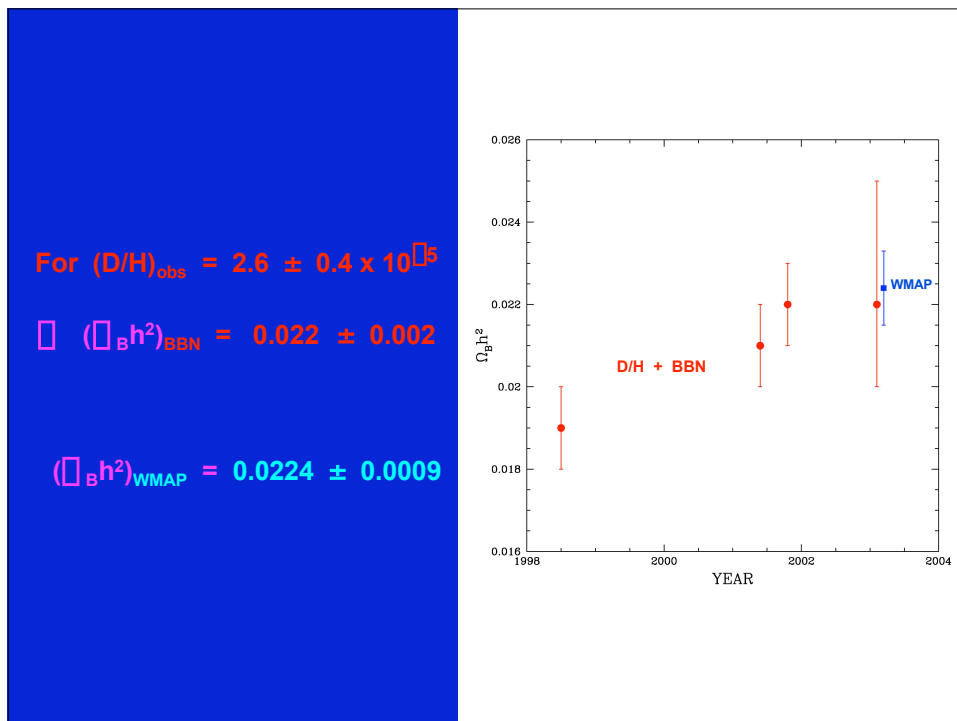
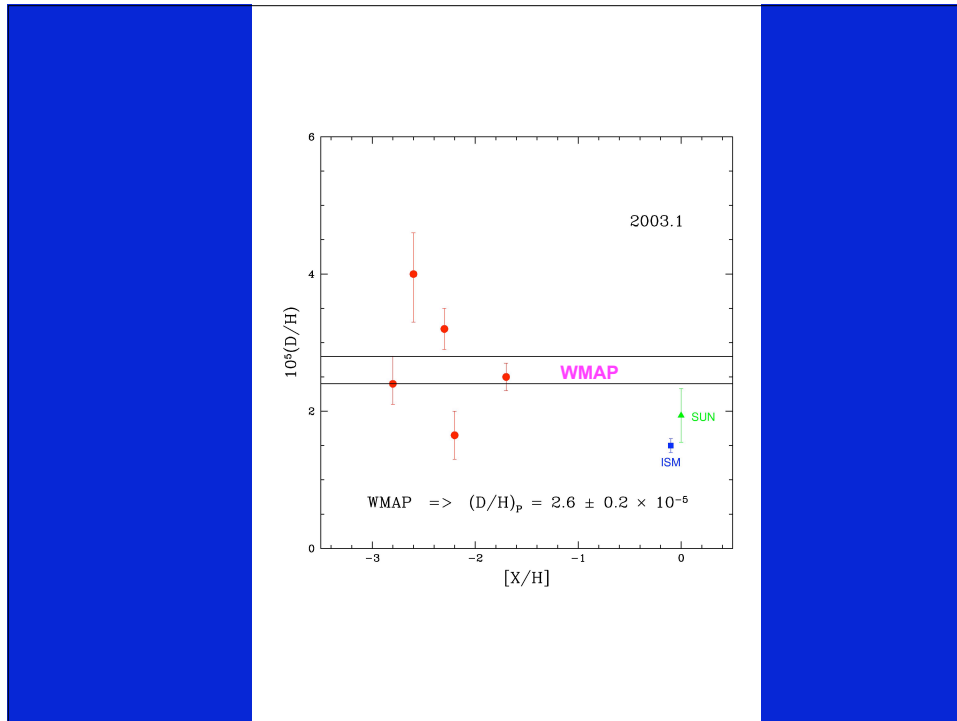
DEUTERIUM -- The Baryometer Of Choice

- D is only DESTROYED in stars \square Anywhere, Anytime: $D/H \square (D/H)_p$
- For $t \ll t_0$ ($Z \ll Z_0$; $z \gg 0$) \square Deuterium Plateau: $(D/H)_t \square (D/H)_p$
- $(D/H)_p$ is sensitive to the baryon density $\square (D/H)_p \mu \square_{10}^{1.6}$
- BUT ...
 - * HI and DI spectra are identical (except for a wavelength shift)
 - \square HI Interlopers Masquerading As DI?
 - * Unresolved velocity structure \square Systematic errors in $N(\text{HI})$?

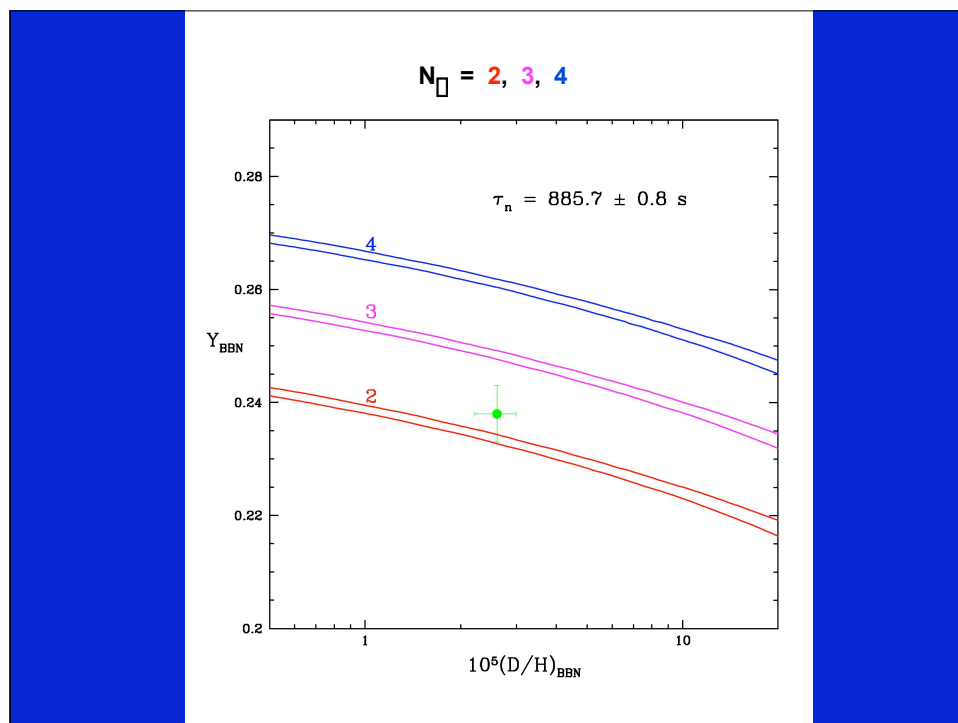
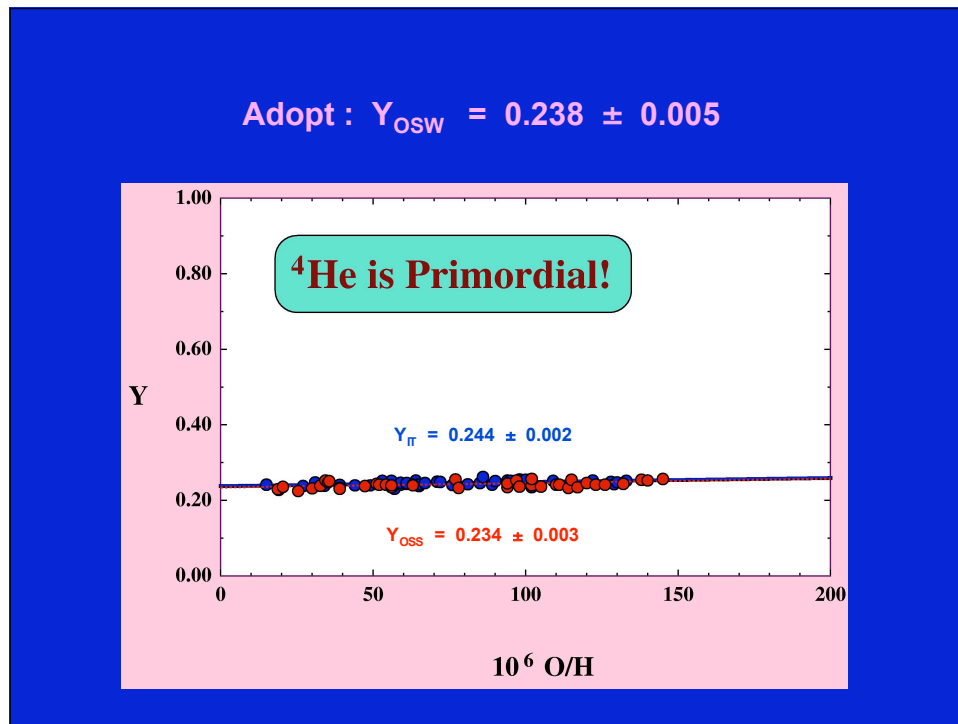
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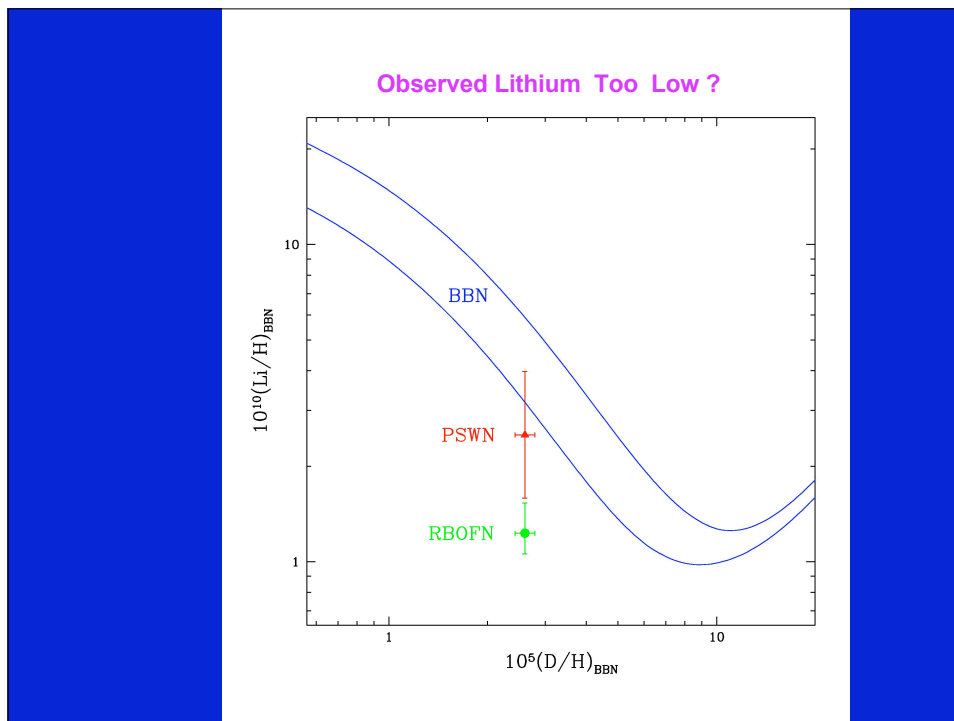
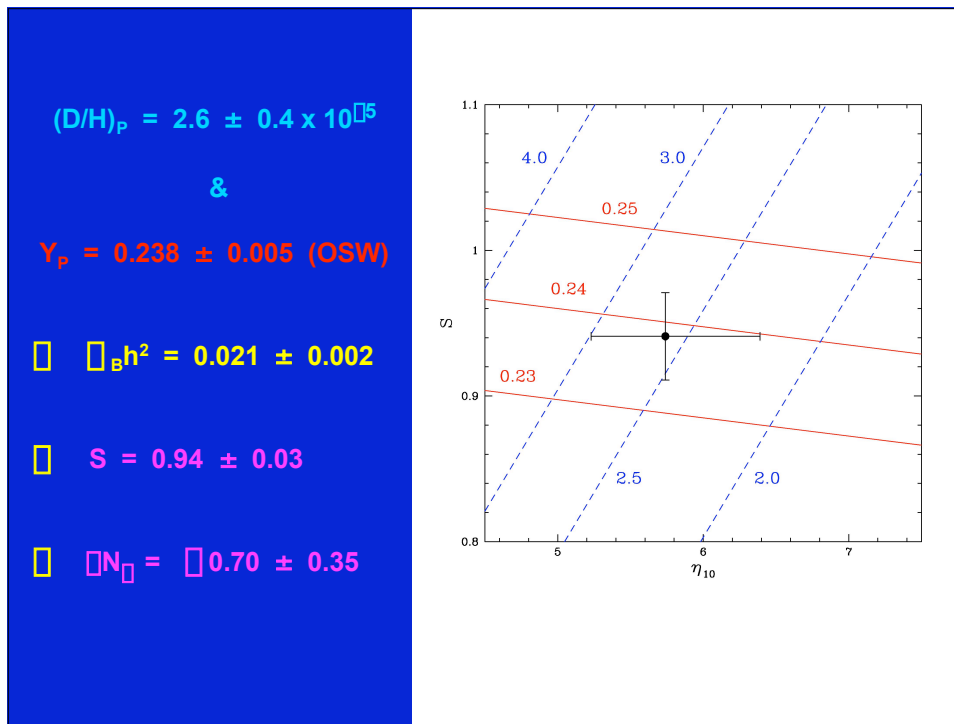
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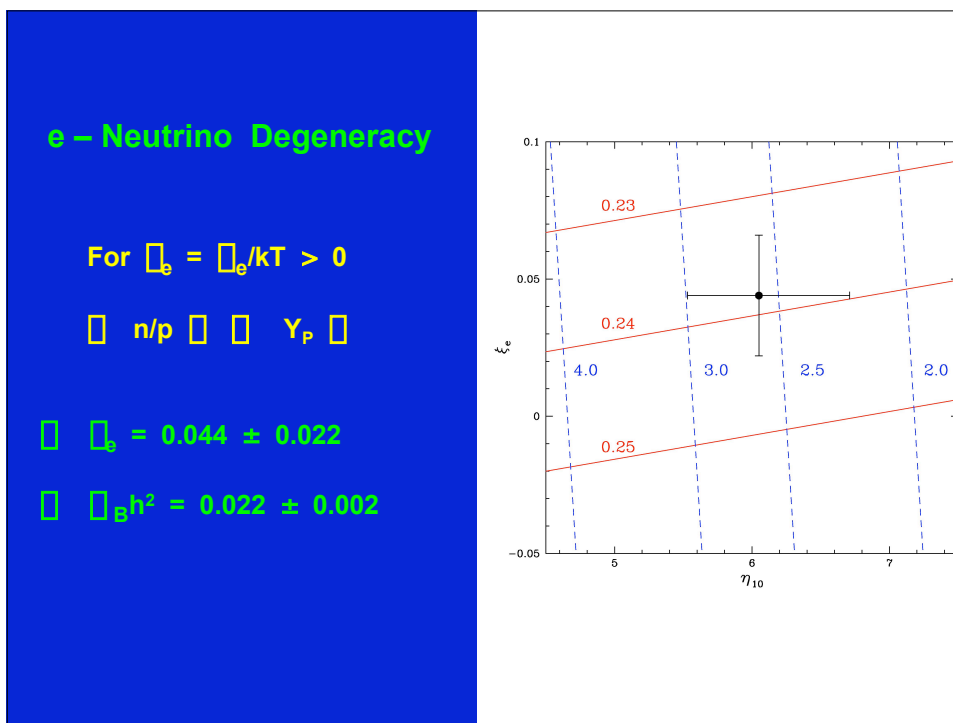
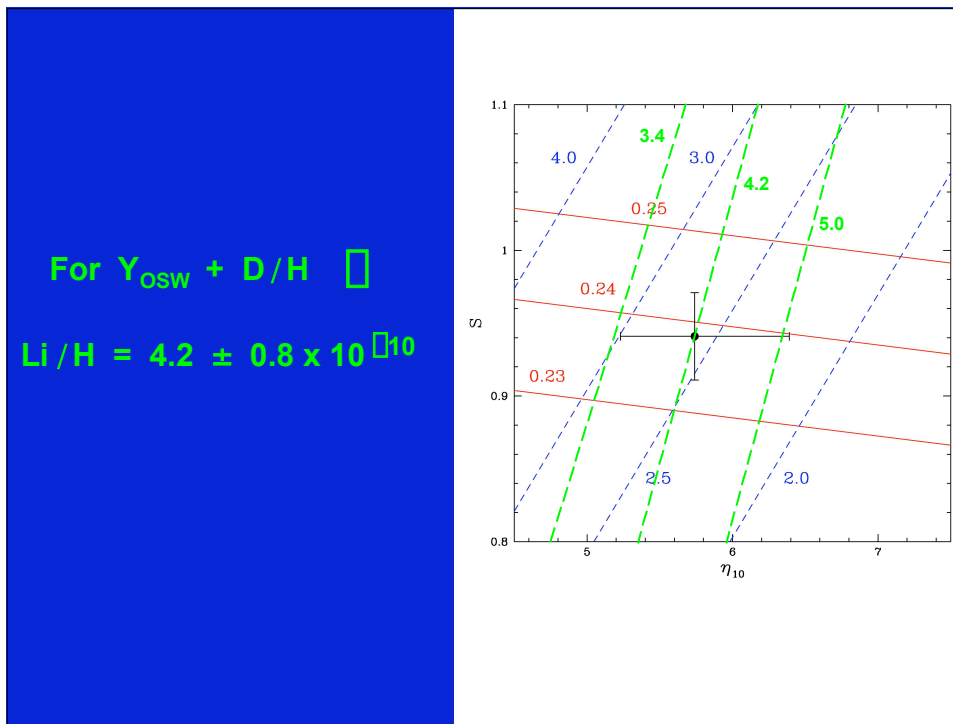
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OSW “Best Fit” (x):

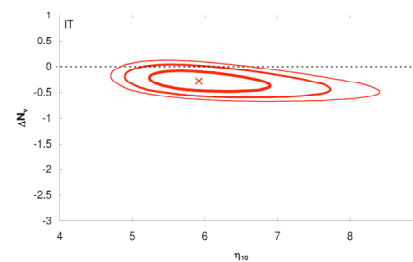
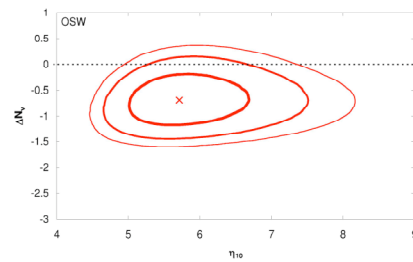
$$\Omega_{10} = 5.7; \Omega_B h^2 = 0.021$$

$$N_\nu = 0.69$$

IT “Best Fit” (x):

$$\Omega_{10} = 5.9; \Omega_B h^2 = 0.022$$

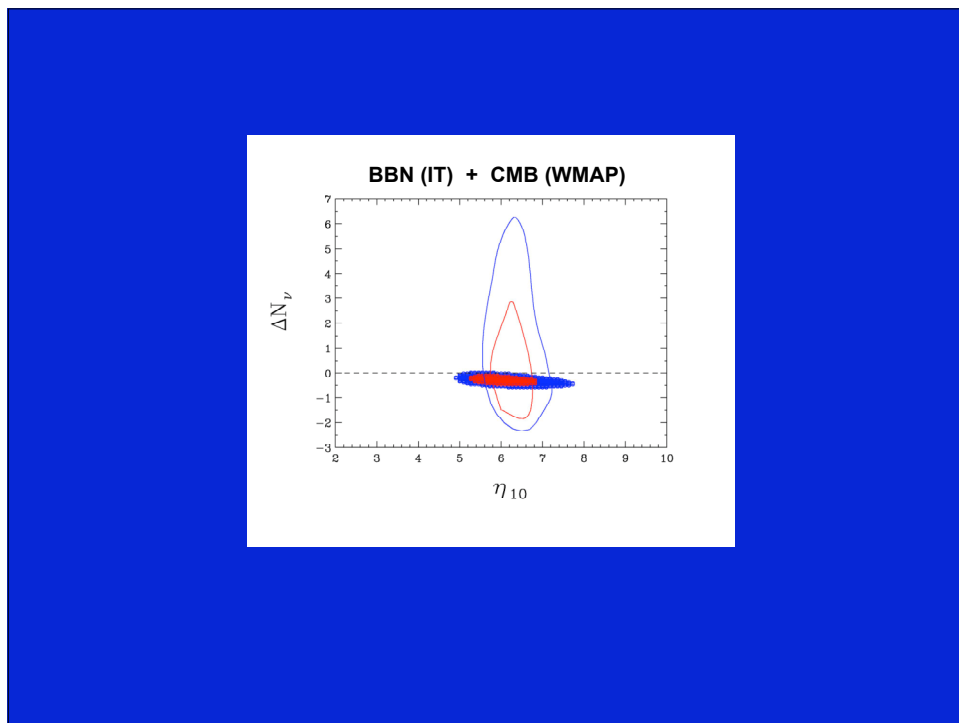
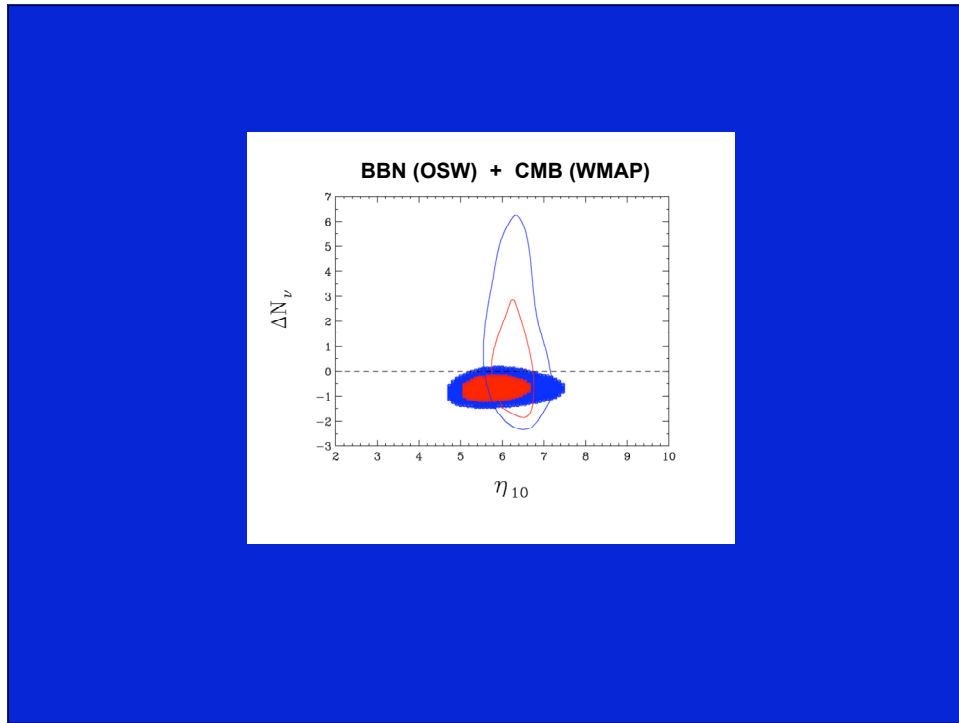
$$N_\nu = 0.29$$



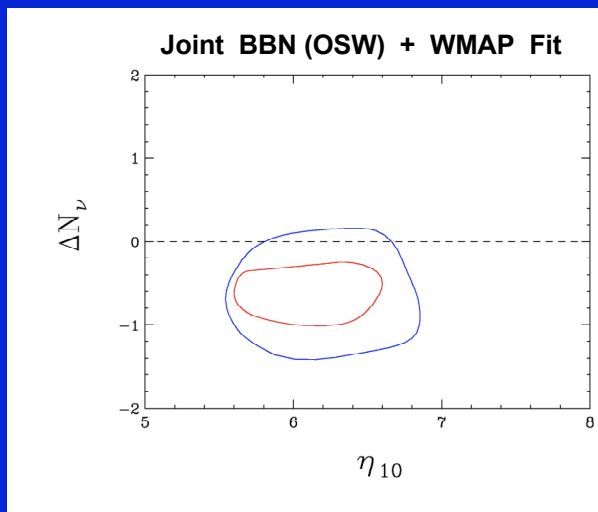
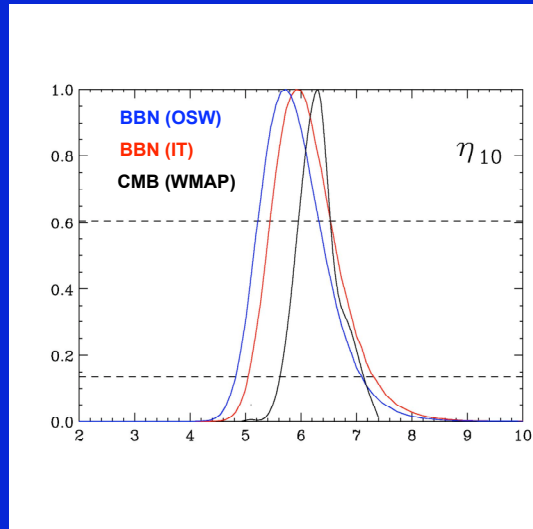
CONCLUSIONS - BBN ALONE

- $(D/H)_P = 2.6 \pm 0.4 \times 10^{-5}$
- $Y_P = 0.238 \pm 0.005$
- $\xi_e = 0$
 - * $\Omega_B h^2 = 0.021 \pm 0.002$ ($\eta_{10} = 5.74^{+0.65}_{-0.51}$)
 - * $\Delta N_\nu = -0.70^{+0.35}_{-0.34}$
 - * BUT: $(Li/H)_P \approx 4.5 \times 10^{-10}$
- $\xi_e \neq 0$ ($N_\nu = 3$)
 - * $\Omega_B h^2 = 0.022 \pm 0.002$ ($\eta_{10} = 6.15^{+0.68}_{-0.53}$)
 - * $\xi_e = 0.044 \pm 0.022$
 - * BUT: $(Li/H)_P \approx 4.2 \times 10^{-10}$

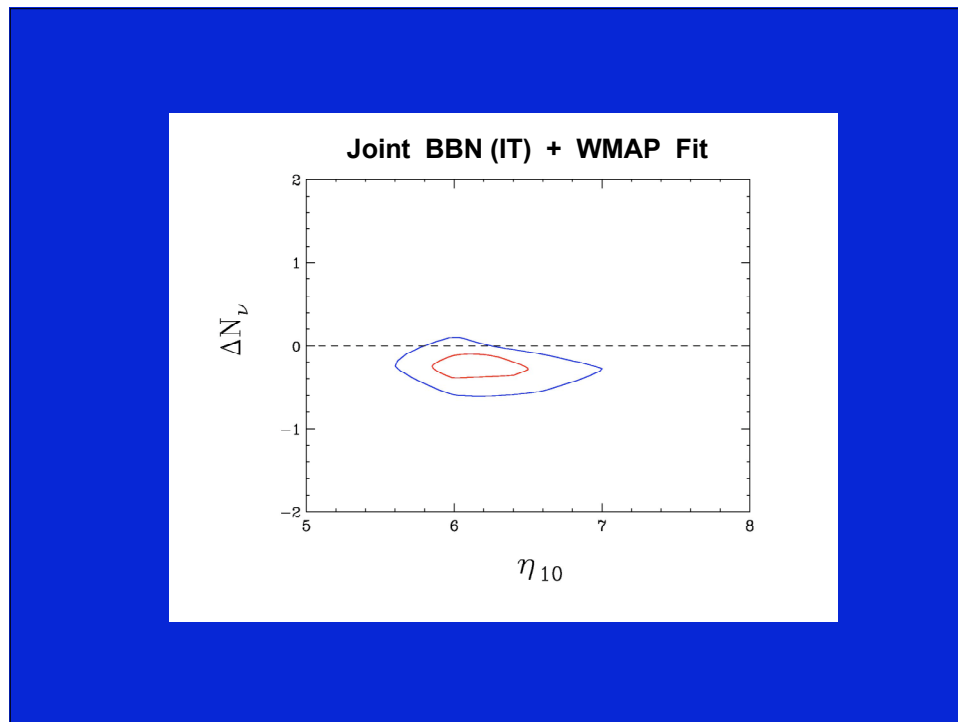
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SUMMARY AND CONCLUSIONS

BBN and CMB are consistent \square

The Universe at 20 min. and 380 kyr AGREE !

Best : $\eta_{10} = 6.16 (6.30)$; $\Omega_B h^2 = 0.0225 (0.0230)$; $N_\nu = 2.25 (2.75)$

95% Ranges : $5.57 \leq \eta_{10} \leq 6.84$; $0.0203 \leq \Omega_B h^2 \leq 0.0250$

$(5.66 \leq \eta_{10} \leq 6.84 ; 0.0207 \leq \Omega_B h^2 \leq 0.0250)$

$1.80 \leq N_\nu \leq 2.94 (2.49 \leq N_\nu \leq 2.99)$

$\Omega_b = \Omega_\nu = \Omega_\gamma = 0.044 \pm 0.022$