

Neutrino Masses and Flavor-Changing Radiative Corrections

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- * Brief history of radiative contributions to m_ν
- * The form of m_ν (2002 edition) and its underlying symmetry
- * Three-fold degeneracy and FCRC
- * Two-fold degeneracy and FCRC

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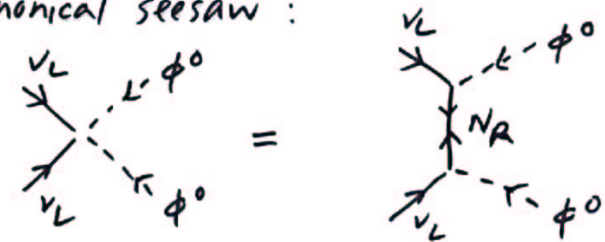
* Where does m_ν come from?

Weinberg (79):

$$\mathcal{L}_M = -\frac{f_{ij}}{2\Lambda} (\nu_i \phi^0 - e_i \phi^+) (\nu_j \phi^0 - e_j \phi^+) + \text{H.c.}$$

$$\Rightarrow (m_\nu)_{ij} = \frac{f_{ij} \langle \phi^0 \rangle^2}{\Lambda}$$

Canonical seesaw:



is just 1 of 3 treelevel realizations.
There are also 3 generic 1-loop realizations. [Ma (98)]

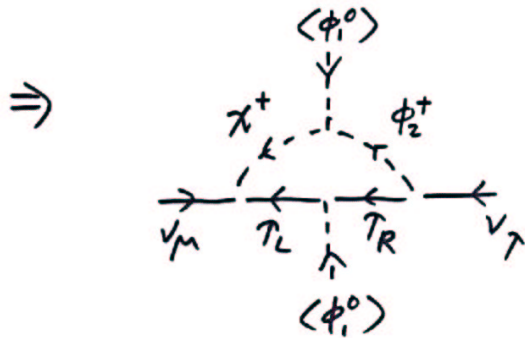
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There are 2 kinds of radiative corrections.

(I) $\begin{array}{c} \rightarrow \text{O} \leftarrow \\ \nu_L \quad \nu_L \end{array}$, i.e. mass correction

First example: Zee (80)

Add χ^+ and (ϕ_2^+, ϕ_2^0)

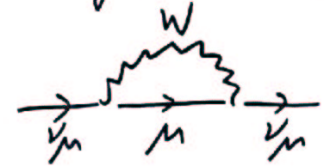


Result: $\delta m_\nu \rightarrow \delta m_\nu + \delta \delta m_\nu$

(II) $\begin{array}{c} \rightarrow \text{O} \rightarrow \\ \nu_L \quad \nu_L \end{array}$, i.e. wave-function correction

First example: Babu, Leung, Pantaleone (93)

Standard Model:



Most general such corrections:

$$R = \begin{pmatrix} r_{ee} & r_{e\mu} & r_{e\tau} \\ r_{e\mu}^* & r_{\mu\mu} & r_{\mu\tau} \\ r_{e\tau}^* & r_{\mu\tau}^* & r_{\tau\tau} \end{pmatrix}$$

Result:

$$\delta m_\nu \rightarrow (1+R) \delta m_\nu (1+R^T)$$

Atmospheric & Solar neutrino data

$$\Rightarrow \begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} \cos\theta & -\sin\theta & 0 \\ \frac{\sin\theta}{\sqrt{2}} & \frac{\cos\theta}{\sqrt{2}} & -\frac{1}{\sqrt{2}} \\ \frac{\sin\theta}{\sqrt{2}} & \frac{\cos\theta}{\sqrt{2}} & \frac{1}{\sqrt{2}} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

to a good first approximation

$\Rightarrow m_\nu$ is a function of 4 real parameters (m_1, m_2, m_3, θ) and may be written in the form

$$m_\nu = \begin{pmatrix} a+2b+c & d & d \\ d & b & a+b \\ d & a+b & b \end{pmatrix},$$

$$\text{then } \frac{d^2}{2c^2} = \frac{\tan^2\theta}{(1-\tan^2\theta)^2} = \left(\frac{\tan 2\theta}{2}\right)^2 \quad [\text{Ma (02)}]$$

$$m_{1,2} = a+2b+c \mp \sqrt{c^2+2d^2}, \quad m_3 = -a$$

Consider the most general m_ν , i.e.

$$m_\nu = \begin{bmatrix} A & D & E \\ D & B & F \\ E & F & C \end{bmatrix}$$

and the discrete symmetry

$$\nu_e \leftrightarrow \nu_e, \nu_\mu \leftrightarrow \nu_\tau, \text{ i.e. } U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{pmatrix},$$

then $U m_\nu U^T = m_\nu$

$$\Rightarrow m_\nu = \begin{bmatrix} A & D & D \\ D & B & F \\ D & F & B \end{bmatrix}$$

= m_ν of p.5 if A, B, D, F are real.

Examples :

$$A = B+F \quad [\text{Mohapatra, Nussinov (99)}]$$

$$A+D = B+F \quad [\text{Harrison, Scott (02)}]$$

$$A+B+F = 0 \quad [\text{Joshi-pura (02)}]$$

Consider $A = F$, $B = D = 0$, i.e.

$$M_\nu = m_0 \begin{bmatrix} 1 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{bmatrix}$$

This pattern is the result of the symmetry A_4 [Ma, Rajasekaran (01)] for the complete theory, with

$$\left. \begin{array}{l} (v_i, l_i)_L \sim \underline{\underline{3}}, \\ l_{1R} \sim \underline{\underline{1}}, l_{2R} \sim \underline{\underline{1}'}, l_{3R} \sim \underline{\underline{1''}} \end{array} \right\} \text{under } A_4$$

$$\Rightarrow m_\ell = \underbrace{\frac{1}{\sqrt{3}} \begin{pmatrix} 1 & 1 & 1 \\ 1 & \omega & \omega^2 \\ 1 & \omega^2 & \omega \end{pmatrix}}_{U_L} \begin{pmatrix} m_e & 0 & 0 \\ 0 & m_\mu & 0 \\ 0 & 0 & m_\tau \end{pmatrix} \quad [\omega = e^{2\pi i/3}]$$

$$\Rightarrow m_\nu = U_L^T \begin{pmatrix} m_0 & 0 & 0 \\ 0 & m_0 & 0 \\ 0 & 0 & m_0 \end{pmatrix} U_L = m_0 \begin{pmatrix} 1 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{pmatrix}$$

as desired.

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What is A_4 ?

and why is it special?

Theaetetus (~ 390 BCE) \Rightarrow
only 5 perfect geometric solids

solid	faces	vertices	Plato	Group
Tetrahedron	4	4	fire	A_4
Octahedron	8	6	air	S_4
Icosahedron	20	12	water	A_5
Hexahedron	6	8	earth	S_4
Dodecahedron	12	20	?	A_5

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A_4 = group of even permutations
of 4 objects

Character Table

Class	n	h	$\chi^{(1)}$	$\chi^{(2)}$	$\chi^{(3)}$	$\chi^{(4)}$
C_1	1	1	1	1	1	3
C_2	4	3	1	ω	ω^2	0
C_3	4	3	1	ω^2	ω	0
C_4	3	2	1	1	1	-1

n = number of elements

h = order of each element

$$\omega = e^{\frac{2\pi i}{3}}$$

$$\underline{3} \times \underline{3} = \underline{1} + \underline{1}' + \underline{1}'' + \underline{3} + \underline{3}$$

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Add the most general radiative
correction [Babu, Ma, Valle (02)]

$$\text{i.e. } m_\nu \rightarrow (1+R)m_\nu(1+R^T)$$

$$= m_0 \begin{pmatrix} 1+\delta_0+\delta+\delta^*+2\delta' & \delta'' & \delta''^* \\ \delta'' & \delta & 1+\delta_0+\frac{\delta+\delta^*}{2} \\ \delta''^* & 1+\delta_0+\frac{\delta+\delta^*}{2} & \delta^* \end{pmatrix}$$

$$\text{where } \delta_0 \equiv r_{\mu\mu} + r_{\tau\tau} - r_{\mu\tau} - r_{\mu\tau}^*$$

$$\delta \equiv 2r_{\mu\tau}$$

$$\delta' \equiv r_{ee} - \frac{1}{2}r_{\mu\mu} - \frac{1}{2}r_{\tau\tau} - \frac{1}{2}r_{\mu\tau} - \frac{1}{2}r_{\mu\tau}^*$$

$$\delta'' \equiv r_{e\mu}^* + r_{e\tau}$$

$\Rightarrow m_\nu$ = that of p.5 automatically
for all δ 's real, i.e. b, c, d
are generated radiatively!

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SM (no FCRC) : $S'' = S = 0$

To get a realistic m_ν , we need new flavor-changing physics.

Example: Back to the Zee model, but do not add second Higgs doublet. Instead add [Ma(02)]

$\eta_i^+ \sim \underline{3}$ under A_4 , $\overbrace{A_4 \text{ softly}}^{\text{breaks}}$

with $\mathcal{L} = f \epsilon_{ijk} (\nu_i e_j - e_i \nu_j) \eta_k^+ + \overbrace{m_{ij}^2 \eta_i^+ \eta_j^-}$

$$\Rightarrow S = -\frac{f^2}{4\pi^2} \sum_{i=1}^3 U_{mi}^* U_{\tau i} \ln m_i^2$$

$$S' = -\frac{f^2}{8\pi^2} \sum_{i=1}^3 \left(\frac{1}{2} |U_{mi} - U_{\tau i}|^2 - |U_{ei}|^2 \right) \ln m_i^2$$

$$S'' = -\frac{f^2}{8\pi^2} \sum_{i=1}^3 (U_{mi}^* U_{ei} + U_{\tau i} U_{ei}^*) \ln m_i^2$$

$$(\Delta m^2)_{atm} \simeq 4\delta m_0^2, \quad \mathcal{U}_{e3} \simeq \frac{i \text{Im} S''}{\sqrt{2} \delta},$$

$$\left[\frac{(\Delta m^2)_{sol}}{(\Delta m^2)_{atm}} \right]^2 \simeq \left[\frac{\delta'}{\delta} + |\mathcal{U}_{e3}|^2 \right]^2 + 2 \left[\frac{\text{Re} S''}{\delta} \right]^2$$

$$m_0 < 0.23 \text{ eV} \quad [\text{WMAP}(03)]$$

$$\Rightarrow \delta \gtrsim \frac{2.5 \times 10^{-3}}{4(0.23)^2} \simeq 1.2 \times 10^{-2}$$

$$[\text{compared to say QCD} : \alpha_s/2\pi \simeq 1.9 \times 10^{-2}]$$

$$\Rightarrow m_0 \text{ cannot be much below } 0.2 \text{ eV}$$

$$\Rightarrow \text{near-future } \text{O}\nu\beta\beta \text{ decay experiments should measure } m_0 !$$

Recall the condition $U m_\nu U^T = m_\nu$.

This implies $U^n m_\nu (U^T)^n = m_\nu$.

Unless $U^n = 1$ for some $n = \bar{n}$,
the only possible solution for m_ν is
 $m_\nu \propto 1$ (and U real).

Example:

$$U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{pmatrix}, \quad U^2 = 1$$

* Suppose we want $U^2 = \begin{pmatrix} 1 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{pmatrix}$,

now $U^4 = 1$, but what is U ?

* Solution (i)

$$U_1 = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \frac{1-i}{2} & \frac{1+i}{2} \\ 0 & \frac{1+i}{2} & \frac{1-i}{2} \end{pmatrix}$$

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$$\Rightarrow m_\nu = \begin{pmatrix} A & D & D \\ D & B & B \\ D & B & B \end{pmatrix} = \begin{pmatrix} 2b+c & d & d \\ d & b & b \\ d & b & b \end{pmatrix}$$

$$m_{1,2} = 2b+c \mp \sqrt{c^2+2d^2}, \quad m_3 = 0,$$

i.e. inverted hierarchy with

$$(\Delta m^2)_{\text{atm}} \approx (2b+c)^2 \approx 4b^2$$

$$(\Delta m^2)_{\text{sol}} \approx 4(2b+c)\sqrt{c^2+2d^2} \approx 8b\sqrt{c^2+2d^2}$$

$$m_{\nu_e} (0\nu\beta\beta) \approx \sqrt{(\Delta m^2)_{\text{atm}}} \approx 0.05 \text{ eV}$$

* Solution (ii)

$$U_2 = \frac{1}{\sqrt{3}} \begin{pmatrix} 1 & 1 & 1 \\ 1 & \omega & \omega^2 \\ 1 & \omega^2 & \omega \end{pmatrix}, \quad \omega = e^{\frac{2\pi i}{3}}$$

$$\Rightarrow c=d, \text{ i.e. } \tan^2\theta = 2-\sqrt{3} = 0.27.$$

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$$\frac{(\Delta m^2)_{sol}}{(\Delta m^2)_{atm}} \simeq \frac{10^{-4} \text{ eV}^2}{2.5 \times 10^{-3} \text{ eV}^2} = 0.04$$

$$\Rightarrow \frac{\sqrt{c^2 + 2d^2}}{b} \simeq 0.02$$

$\therefore c, d$ may come from FCRC

$$\text{Let } m_\nu = m_0 \begin{pmatrix} 1 & 0 & 0 \\ 0 & \frac{1}{2} & \frac{1}{2} \\ 0 & \frac{1}{2} & \frac{1}{2} \end{pmatrix},$$

then $(1+R)m_\nu(1+R^T)$

$$= m_0 \begin{pmatrix} 1+2r_{ee} & \frac{3r_{e\mu}+r_{e\tau}}{2} & \frac{r_{e\mu}+3r_{e\tau}}{2} \\ \frac{3r_{e\mu}+r_{e\tau}}{2} & \frac{1}{2}+r_{\mu\mu}+r_{\mu\tau} & \frac{1+r_{\mu\mu}+r_{\tau\tau}+r_{\mu\tau}+r_{\mu\tau}^*}{2} \\ \frac{r_{e\mu}+3r_{e\tau}}{2} & \frac{1+r_{\mu\mu}+r_{\tau\tau}+r_{\mu\tau}+r_{\mu\tau}^*}{2} & \frac{1}{2}+r_{\tau\tau}+r_{\mu\tau}^* \end{pmatrix}$$

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$$\text{Example: } \mathcal{L} = f_e (\nu_e \mu - e \nu_\mu) \chi^\dagger + f_\tau (\nu_\tau \mu - \tau \nu_\mu) \chi^\dagger + \text{H.c.}$$

i.e. $L_e + L_\tau - L_\mu$ symmetry, with

$$r_{e\mu} = r_{\mu\tau} = 0, \quad r_{\mu\mu} = r_{ee} + r_{\tau\tau}, \quad r_{e\tau}^2 = r_{ee} r_{\tau\tau}$$

$$\Rightarrow m_1 = m_0 (1 + r_{ee})$$

$$m_2 = m_0 (1 + 2r_{ee} + 2r_{\tau\tau})$$

$$m_3 = 0$$

$$\tan^2 \theta = \frac{r_{ee}}{2r_{\tau\tau}} < 0.8$$

$$|U_{e3}| \simeq \sqrt{\frac{r_{ee} r_{\tau\tau}}{2}} < 0.005$$

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Conclusion

* Symmetry + FCRC


⇒ realistic m_ν

(i) 3-fold degeneracy :

$m_0(0\nu\beta\beta)$ should not be much
below 0.23 eV (WMAP bound)

(ii) 2-fold degeneracy :

inverted hierarchy with $m_0 \approx 0.05$ eV,
predicts WMAP $\Sigma m_\nu = 0.1$ eV

* FCRC ⇒ 

likely to be observable !

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