

# The Princeton Magneto Rotational Instability Experiment

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Physics of Astrophysical Outflows and Accretion Disks, KITP, 27 May 2005

## Overview of the experiment

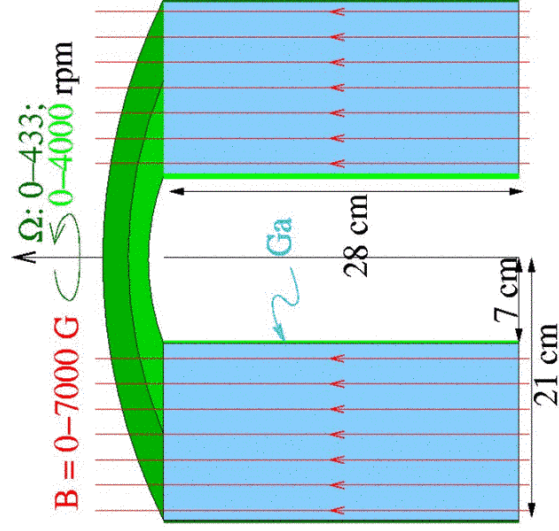
- magnetic Couette flow
  - eutectic gallium
  - axial field
- independently driven cylinders and endcaps

$$Re \equiv \frac{r_1(r_2 - r_1)\Omega_1}{\nu} \approx 10^7$$

$$Re_M \equiv \frac{r_1(r_2 - r_1)\Omega_1}{\eta} \approx 20$$

$$S \equiv \frac{(r_2 - r_1)V_A}{\eta} \approx 4$$

$$Pr_M \equiv \frac{\nu}{\eta} \approx 10^{-6}$$



**Aim to start from a centrifugally stable ( $\kappa^2 > 0$ ) magnetic steady state ( $\mathbf{B} \cdot \nabla \Omega = 0$ )**

## Motivation

- MRI largely theoretical to date
- Explore resistive, low- $Pr_M$ , high- $Re$  regime
  - protostellar disks, quiescent dwarf novae...
- Benchmark astrophysical codes
  - ZEUS, ...
- Seek nonlinear hydrodynamic instabilities
  - High- $Re$ , centrifugally stable flows are little studied
- Develop laboratory astrophysics

## Outline

- Motivation
- Couette flow
- Linear stability
- Water experiments & Ekman circulation
- ZEUS simulations
- Current status

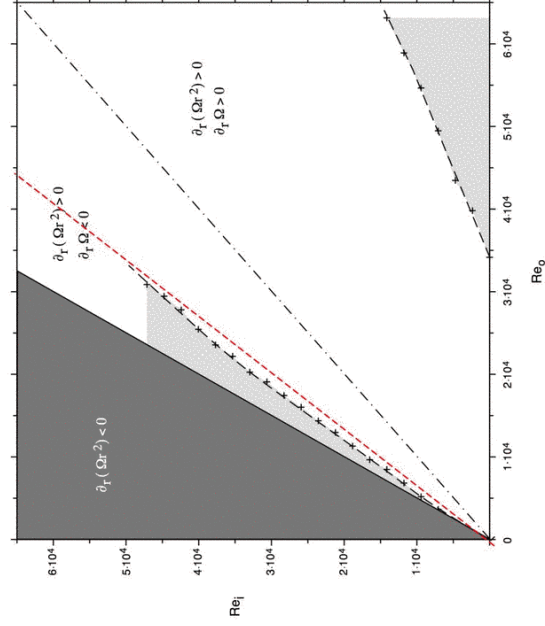
## Couette flow

- Pure rotation and  $h \rightarrow \infty$ :  $v \nabla^2 [r \Omega(r) \hat{\phi}] = 0 \Leftrightarrow \Omega = a + \frac{b}{r^2}$
- Centrifugal (Rayleigh) inviscid axisymmetric stability:
 
$$\kappa^2 \equiv r^{-3} \frac{d}{dr} (r^2 \Omega)^2 \geq 0 \Rightarrow ab \geq 0$$

*i.e.*:  $\Omega_1 \Omega_2 \geq 0$  &  $r_2^2 \Omega_2 \geq r_1^2 \Omega_1$
- Taylor number  $Ta \equiv -\kappa^2 \Delta r^4 / \nu^2$  controls viscous stability
- **MRI instability:  $\Omega_2 < \Omega_1$  and  $Re, Re_M$  &  $S$  sufficiently large**

## Nonlinear hydrodynamic instability

- Exp'ts of Wendt (1933), Taylor (1936), Richard (2001) suggest anomalous torque even when  $\kappa^2 > 0$
- Richard & Zahn (1999):
 
$$v_\tau \approx -\beta r^3 \frac{\partial \Omega}{\partial r}; Re_{crit} \approx \frac{1}{\beta}; \beta \approx 3 \times 10^{-5}$$
  - should be measurable in Pct. exp't  $Re \sim 10^7, \Gamma_\tau \sim 6 \text{ N-m}$
  - nonlocal if applicable to disks
- However, the torque might be a wall or Ekman effect
  - Direct shearing-box simulations laminar for  $\zeta > 0.05$  (Hawley, Balbus, & Winters 1999)
  - Empirical closure scheme suggests laminar flow for  $\zeta > 0.33$  (Ogilvie 2005). **Keplerian:  $\zeta = 0.5$**

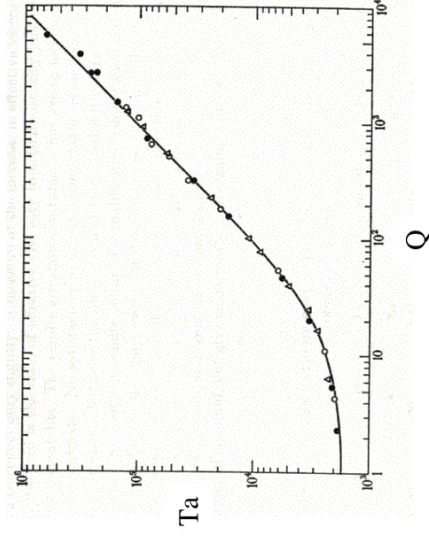


Stability in plane of  $\Omega_1$  versus  $\Omega_2$   
**red line=keplerian**  
 (Richard 2001)

# Magnetized T-C Flow

- **Velikhov (1959) & Chandrasekhar (1960)** discovered ideal MRI
- **Chandrasekhar (1961)** made first thorough analysis of magnetized Couette flow, followed by many other authors.
  - Concentrated on **magnetic stabilization** of Rayleigh-unstable flows: critical Ta increases with B or Chandrasekhar number  $Q=(V_A d)^2/\nu\eta$
  - **Term  $r\Omega^2\delta B_r$  crucial to MRI was omitted** (but OK in this case)
- **Confirming experiments** followed: Donnelly & Ozima (1960, 1962) and more carefully by Donnelly & Caldwell (1964).
- Recent non-ideal MRI analyses for infinite/periodic cylinders: Goodman & Ji (2002), Rüdiger & Shalybkov (2002)

$$\begin{aligned} \delta \dot{B}_r - B \partial_z \delta v_r &= \eta (\partial_r \partial_t^2 + \partial_z^2) \delta B_r \\ \delta \dot{B}_\theta - B \partial_z \delta v_\theta - \delta B_r r \partial_r \Omega &= \eta (\partial_r \partial_t^2 + \partial_z^2) \delta B_\theta \\ \partial_t^2 \delta B_r + \partial_z \delta B_z &= 0 \\ \delta \dot{v}_r - 2\Omega \delta v_\theta + \partial_r \delta \bar{p} - V_A \partial_z \delta B_r &= \nu (\partial_r \partial_t^2 + \partial_z^2) \delta v_r \\ \delta \dot{v}_\theta + \delta v_r \partial_r^2 (r\Omega) - V_A \partial_z \delta B_\theta &= \nu (\partial_r \partial_t^2 + \partial_z^2) \delta v_\theta \\ \partial_t^2 \delta v_r + \partial_z \delta v_z &= 0 \end{aligned}$$



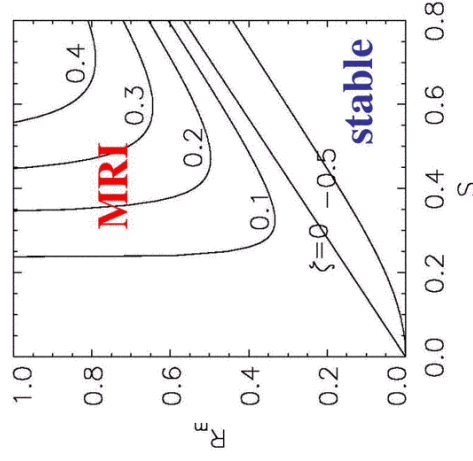
## MRI Stability Limits ( $v \ll \eta$ )

**WKB works well:**  $k^2 V_A^2 + \frac{\kappa^2 (\eta k^2)^2}{(k_z V_A)^2} + \frac{\partial \Omega^2}{\partial \ln r} \geq 0$ :  $k_z \approx \frac{\pi}{h}$ ,  $k_r \approx \frac{\pi}{d}$

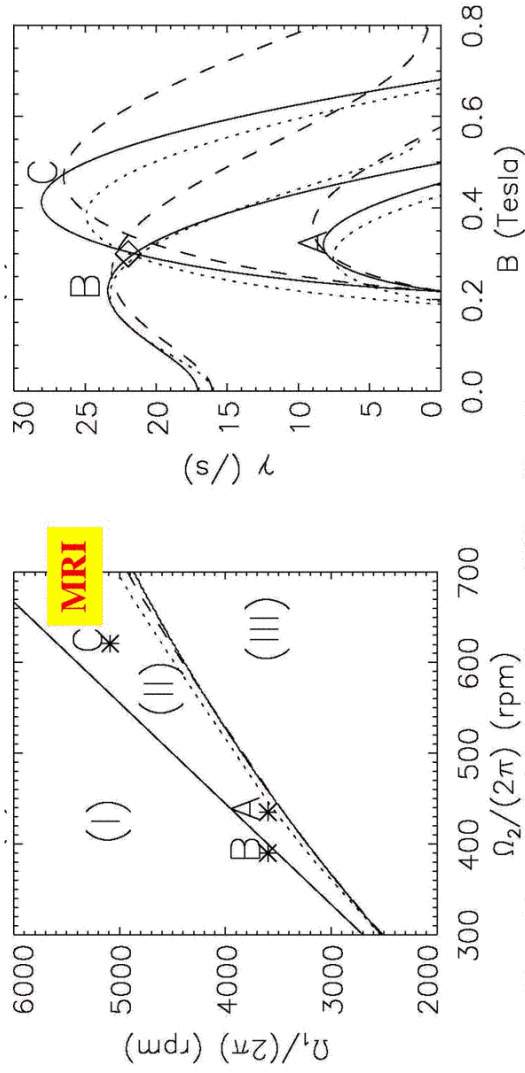
$$R_m^2 \leq \frac{S^4 [1 + (h/r)^2]}{2[(2-s)S^2 - s]}$$

$$\zeta \equiv \frac{\partial \ln(r^2 \Omega)}{\partial \ln r}$$

**minimum critical  $R_m$   
at some S when  $\zeta > 0$**



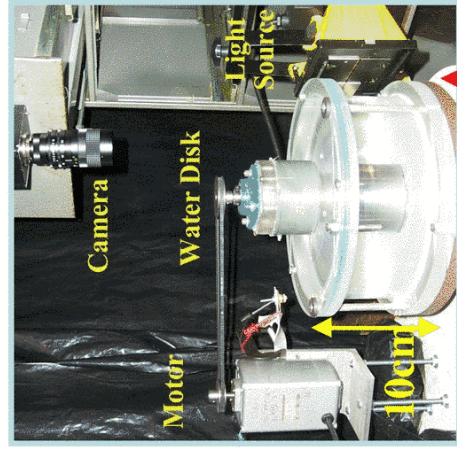
# Stability Diagram



- (I): Unstable but can be stabilized by B
  - (II): Stable but can be destabilized by B: MRI
  - (III): Always stable
- WKB agrees well with global analysis*
- $r_1 = 5 \text{ cm}$   
 $r_2 = 15 \text{ cm}$   
 $h = 10 \text{ cm}$

## Prototype water experiment (2002)

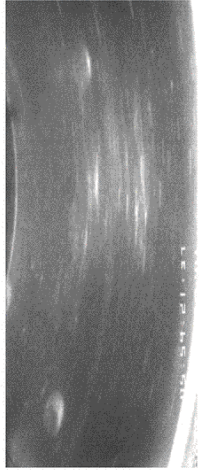
- Independent rotation of inner and outer cylinders
- Endcaps corotate with outer cylinder
- Very short geometry:  $h/r = 1$
- Seed particles to monitor stability and to measure flow



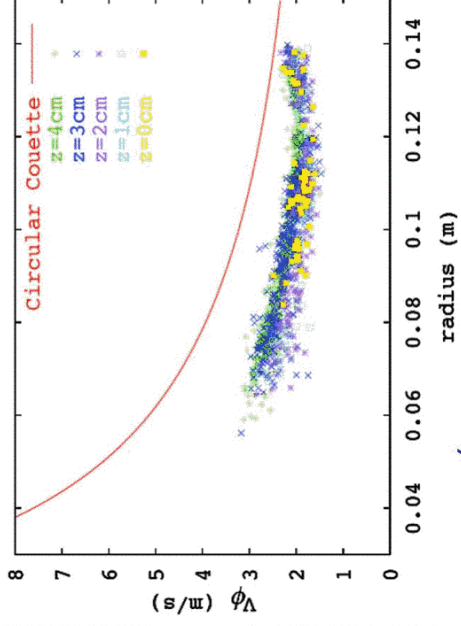
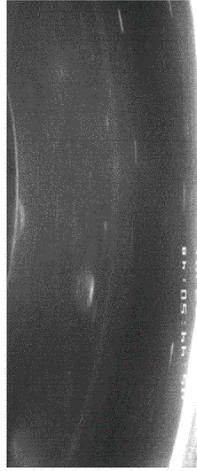
Potter's wheel

# Stability and Flow Measurements

**Unstable flow**

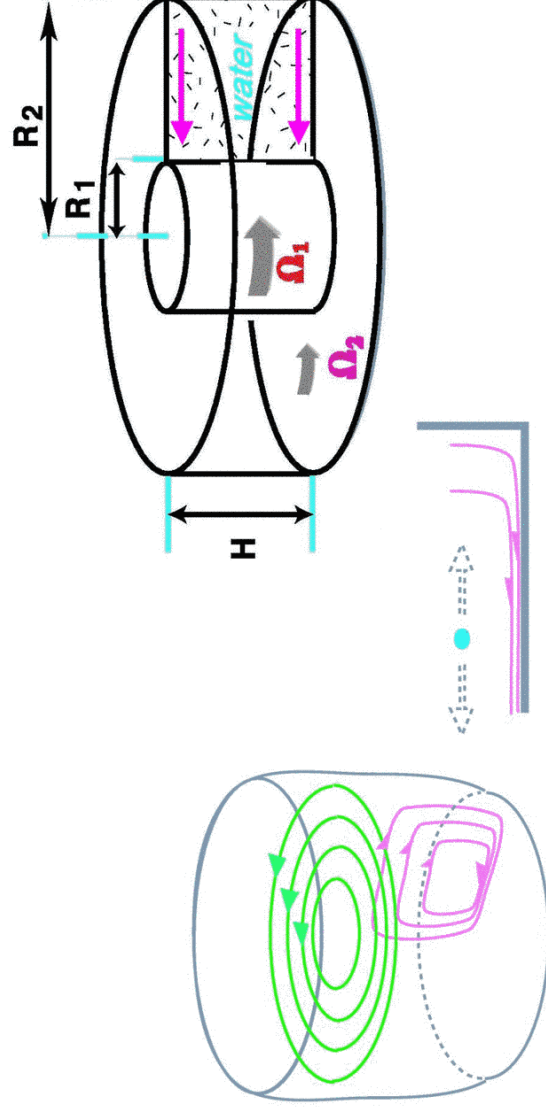


**Stable flow**



- Significant deviation from Couette profile due to **Ekman circulation**.
- Unstable to Rayleigh mode near the inner wall.

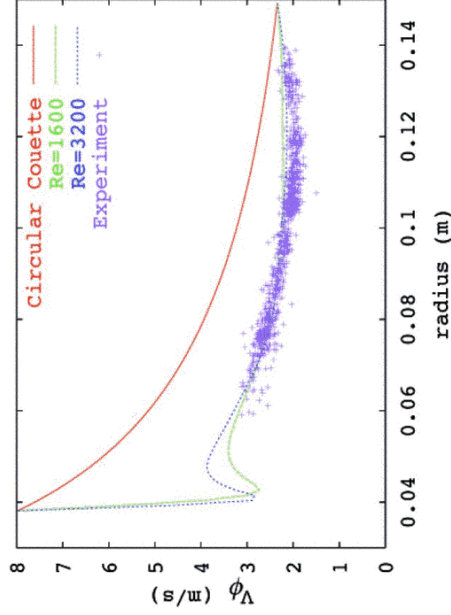
## Ekman Circulation



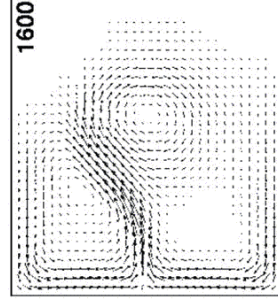
- Significant vertical transport of angular momentum
- Timescale  $\sim h/(v\Omega)^{1/2} \sim \Omega^{-1} Re^{1/2}$

## Measurements Explained by Simulations

Kageyama et al. (2004)



**Toroidal flow**



**Poloidal flow**

**Agreements are remarkable since simulations are in 2D with  $\text{Re} \ll \text{Re in exp}$**

## Spin-down Measurements

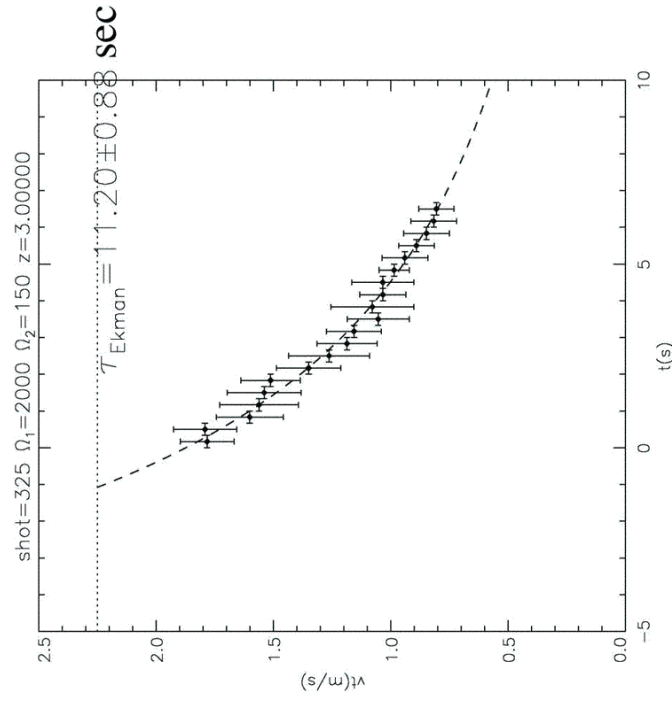
- Ekman Spin-down time:

$$\tau_E = \frac{h}{2\delta_E \bar{\Omega}} = \frac{h}{2\sqrt{\nu \bar{\Omega}}}$$

$$\delta_E = \sqrt{\frac{\nu}{\bar{\Omega}}}$$

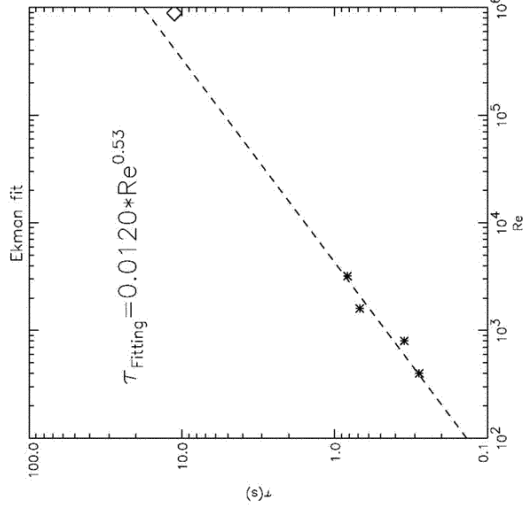
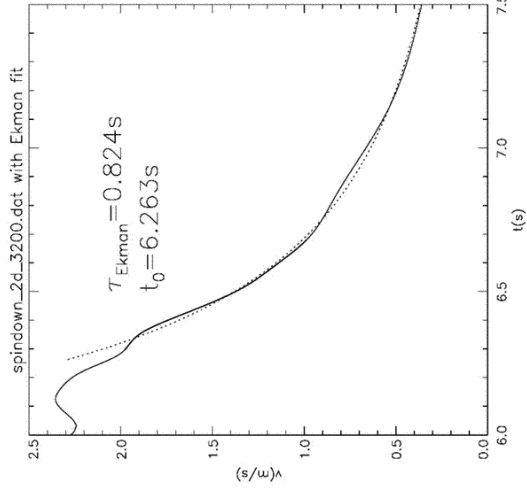
$$\frac{d\bar{\Omega}}{dt} = -\frac{\bar{\Omega}}{\tau_E} \propto -\bar{\Omega}^{3/2}$$

$$\bar{\Omega}(t) = \frac{\bar{\Omega}(t_0)}{\left(1 + \frac{t - t_0}{\tau}\right)^2}$$



# Measured Spin-down Time Consistent with Simulations

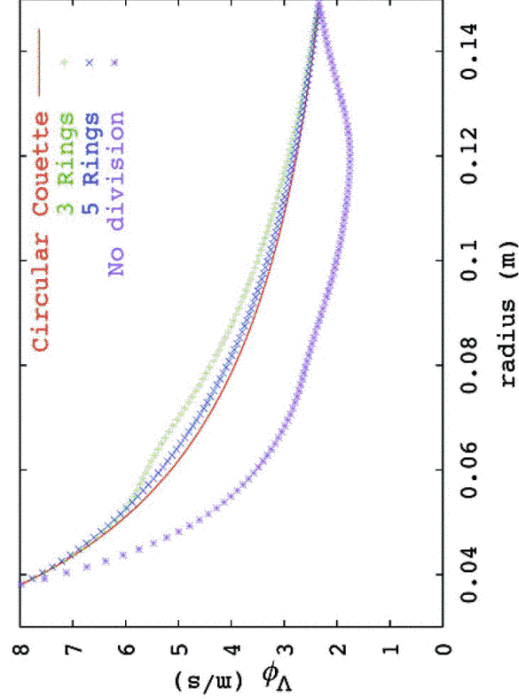
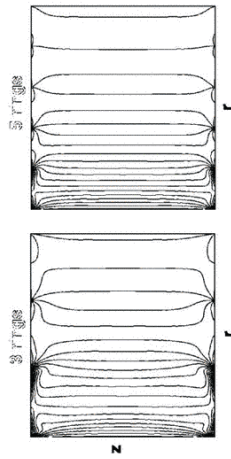
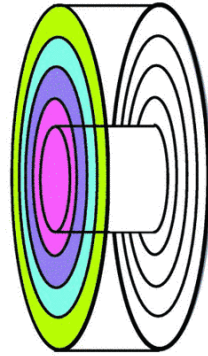
## Simulated spin-down: Scaling with Re:



$$\tau_E = \frac{h}{2\sqrt{\nu\Omega}} = \frac{h}{2\sqrt{R_1(R_2 - R_1)}\sqrt{\Omega_1\Omega}}$$

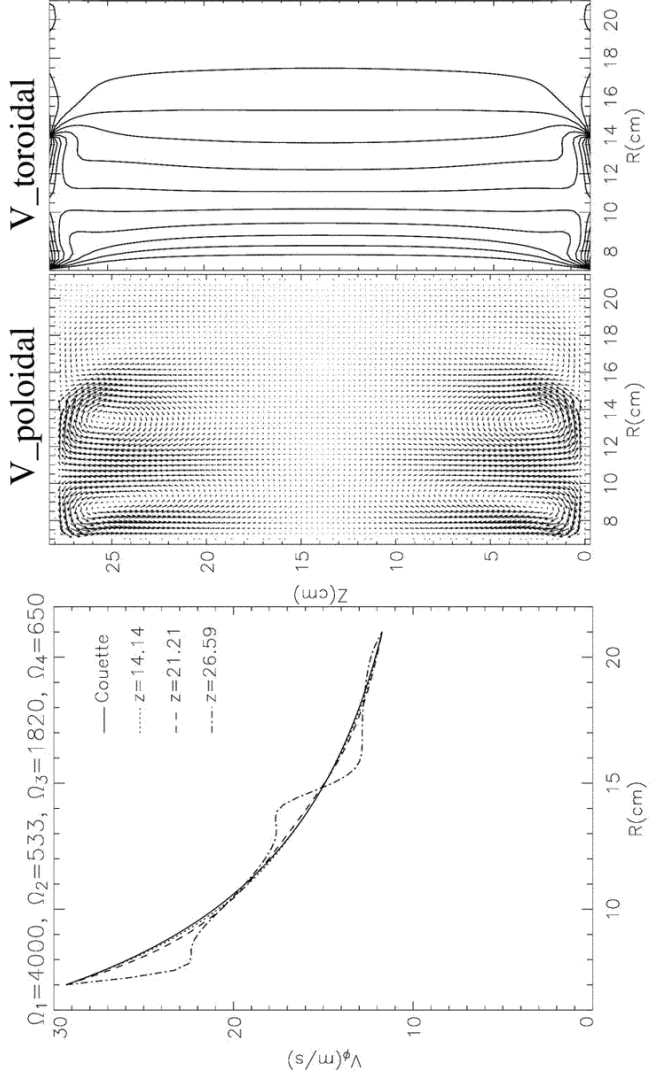
$$\sqrt{\text{Re}} = 0.0110\text{Re}^{0.5} \text{ s}$$

## Solution: Multiple Driven Rings at Each End

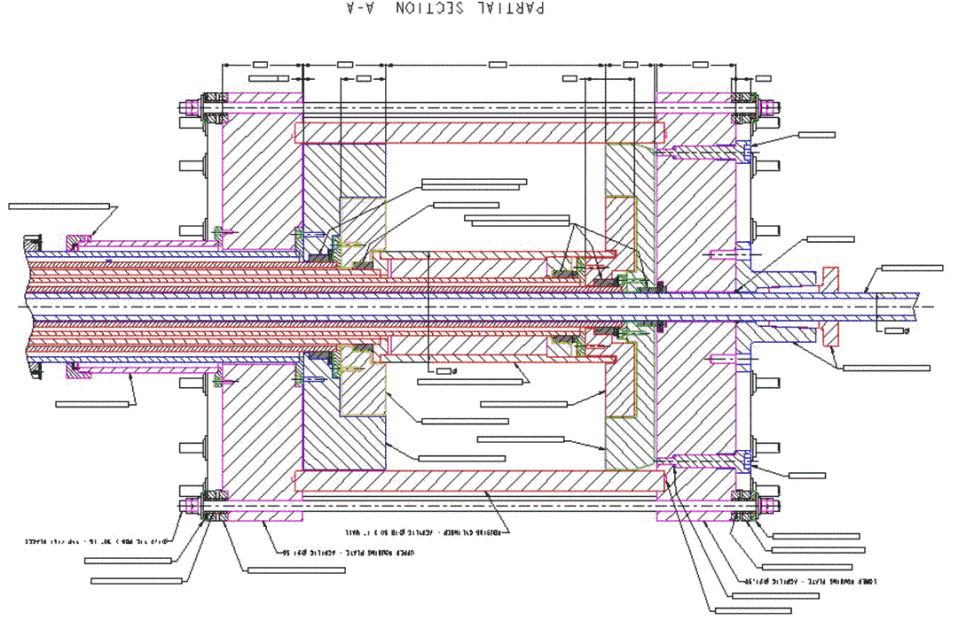
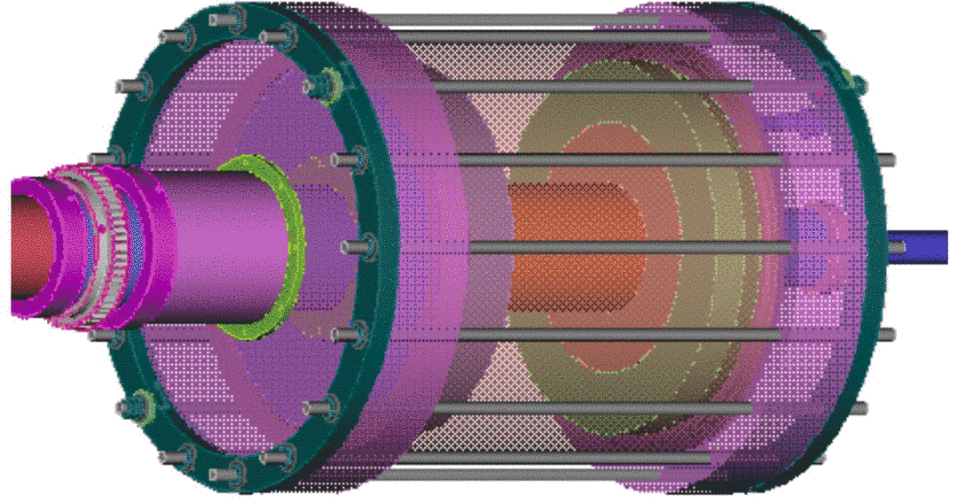


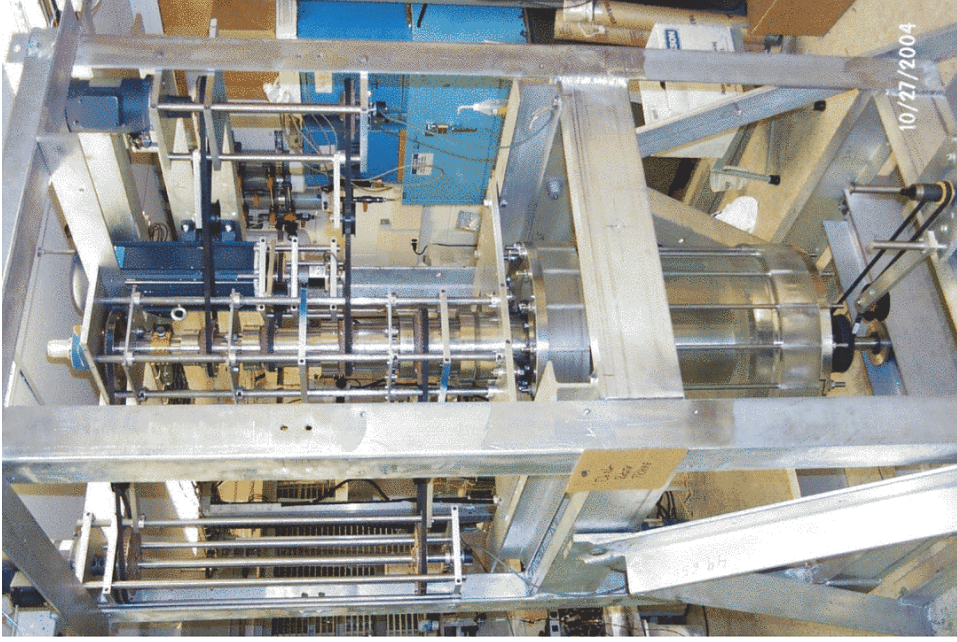


# Final 2-ring Design



2-rings,  $r1=7\text{cm}$ ,  $r2=20.3\text{cm}$ ,  $h=28\text{cm}$

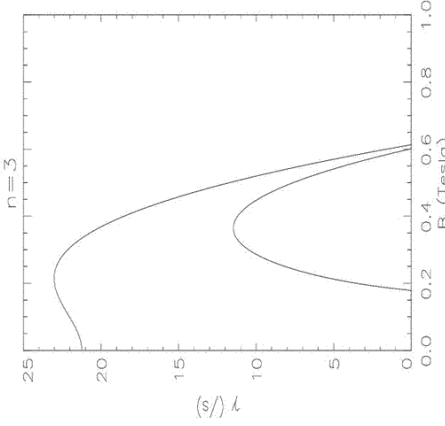




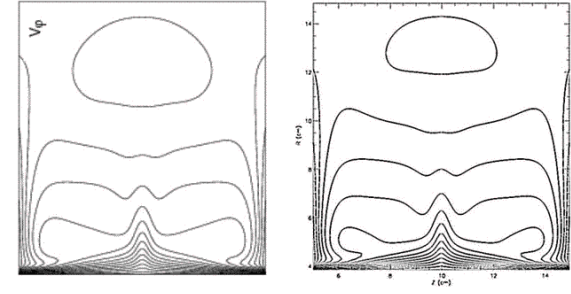
## Diagnostics

- **Torque couplings between inner cylinder and other rotating parts**
- **Surface and internal magnetic perturbations**
- **Surface pressure perturbations**
- **Internal flows by ultrasound**

## Experimental Protocols

- Time scales:  $\tau_\eta \sim \Omega^{-1} \sim 10$  ms  $\tau_A \equiv \frac{h}{2V_A} \sim 24$  ms
  - $\tau_{\text{MRI}} \sim 50$  ms  $\tau_{\text{spin-up}} \sim \tau_E \sim 20$  s
  - MRI:  $\tau_V \sim 5000$  s
  - Spin up Couette flow without B
  - Impose B on penetration time  $\tau_\eta$
  - MRI grows in  $\tau_{\text{MRI}}$
  - Magnetic Ekman circulation:
    - Setup MRI stable flows and impose B
- 

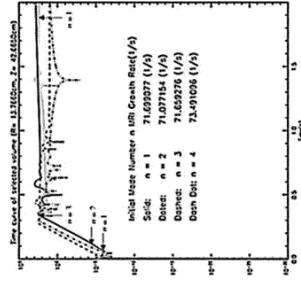
## Modelling with ZEUS2D

- **Wei Liu** added viscosity and resistivity to ZEUS2D
    - excellent agreement with incompressible hydro code if run at peak Mach number  $M=1/4$
    - MRI growth rates as expected
    - explicit resistivity is expensive, but that can be fixed
  - Currently exploring MRI saturation and magnetic Ekman circulation
- 

$V_\phi$  from incompressible code (top) versus ZEUS2D at Mach number 1/4.  $Re=1600$ , endcaps corotating with outer cylinder.

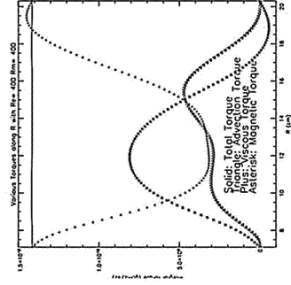
## Vertically periodic flow

$$\text{Re}=\text{Re}_M=400, B=0.5 \text{ T}$$

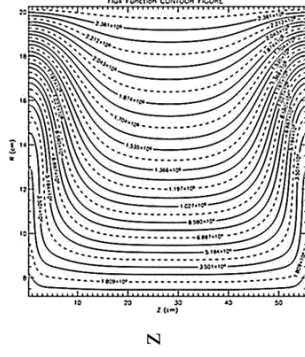


Growth and saturation.  
 $\gamma \approx 70 \text{ sec}^{-1}$   
 $n=3$  mode

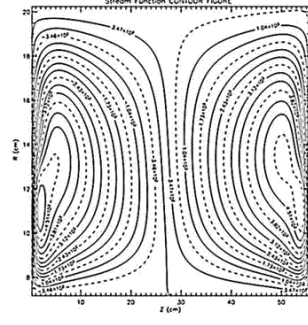
Viscous, magnetic, & advective torques in final state.



$r \rightarrow$



Poloidal field lines



Poloidal stream lines

## Current status and plans

- Couette-flow apparatus is in hand
  - balance, bearings, seals almost debugged (?)
- This summer: take data in water
  - torque measurements
  - flow visualization
- This Fall: begin experiments in gallium
- Simulations and modeling
  - implement insulating boundaries for finite height
  - speed up resistive time step
  - move to 3D