

# Self energy spectroscopy of superconductors

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with

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Ewald Schachinger

# Crystal growers:

## UBC

R. Liang  
D.A. Bonn  
W.N. Hardy

## Stanford

M. Greven  
H. Eisaki

## Brookhaven

Genda Gu

## Saclay

D. Colson  
V. Viallet-Guillen

## Chernogolovka

N.N. Kolesnikov

## McMaster

H. Dabkowska  
B.D. Gaulin  
R. Hughes  
J. Preston

## Ames

N. Ni  
P.C. Canfield  
S.L. Bud'ko

## Westinghouse

J. Talvacchio  
M.G. Forrester

# Outline

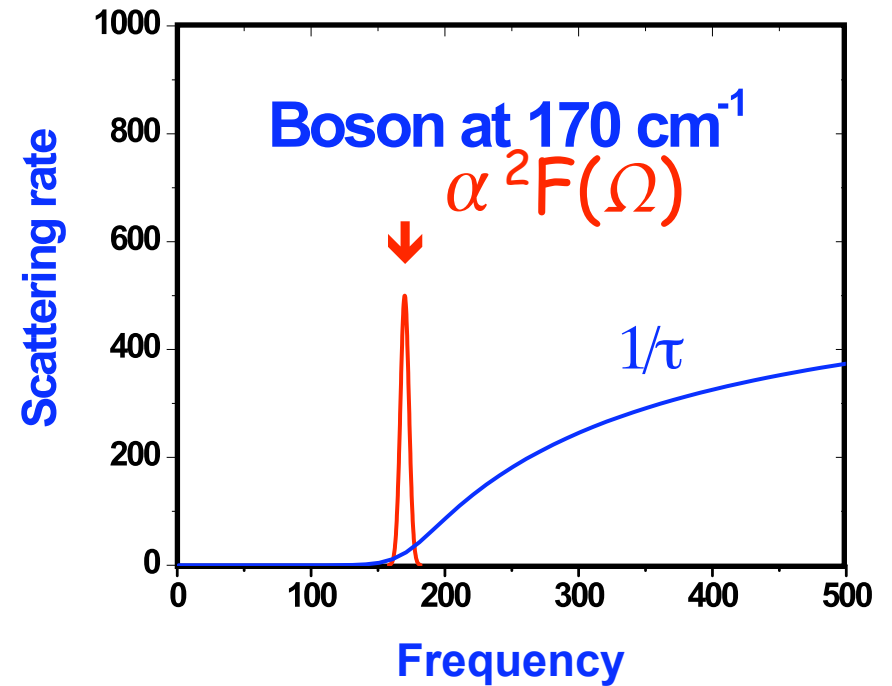
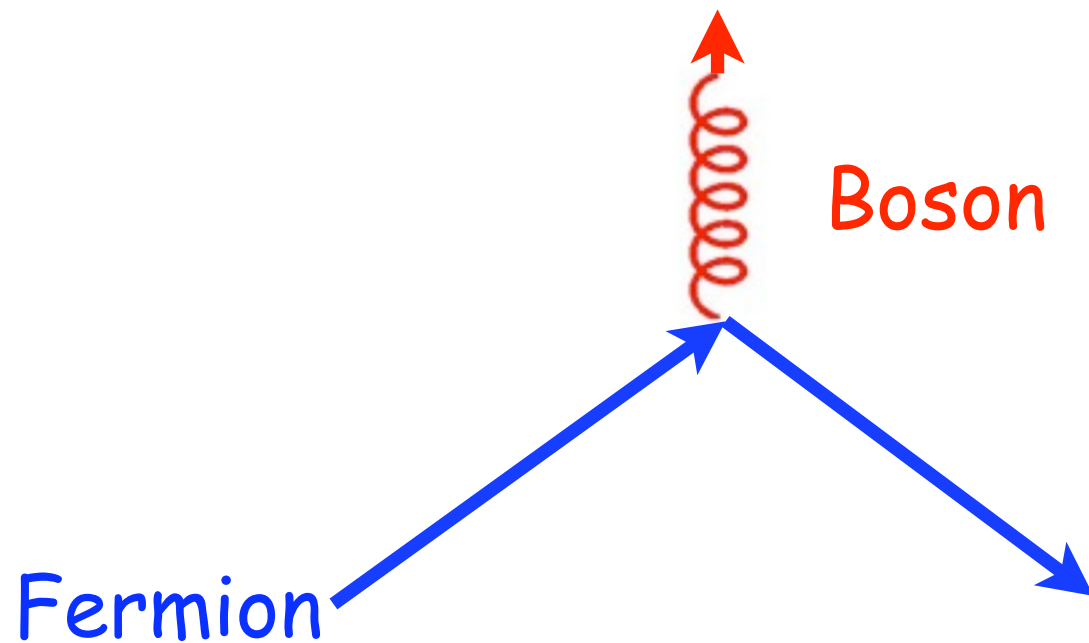
- Finding the glue in the old superconductors, tunneling -> neutrons -> phonons
- New superconductors, new spectroscopy
  - Angle resolved photo emission (ARPES)
  - Optical conductivity
- The glue function in various systems
  - Ortho II YBCO
  - $\text{Ba}_{0.35}\text{K}_{0.45}\text{Fe}_2\text{As}_2$
  - Bi-2212
  - LSCO
  - Mercury
- The isotope effect
- Conclusion: Are we there yet?

# History of glue:

## The old superconductors

# Self energy from boson exchange

T=0, boson creation



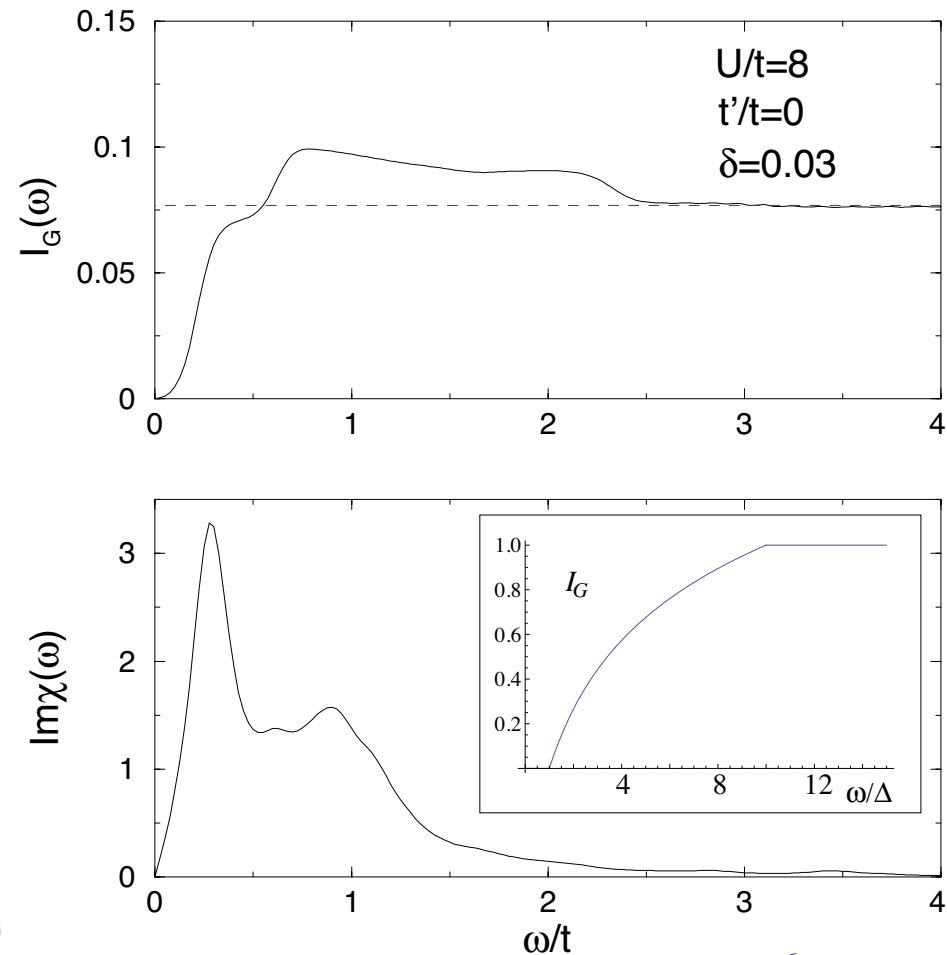
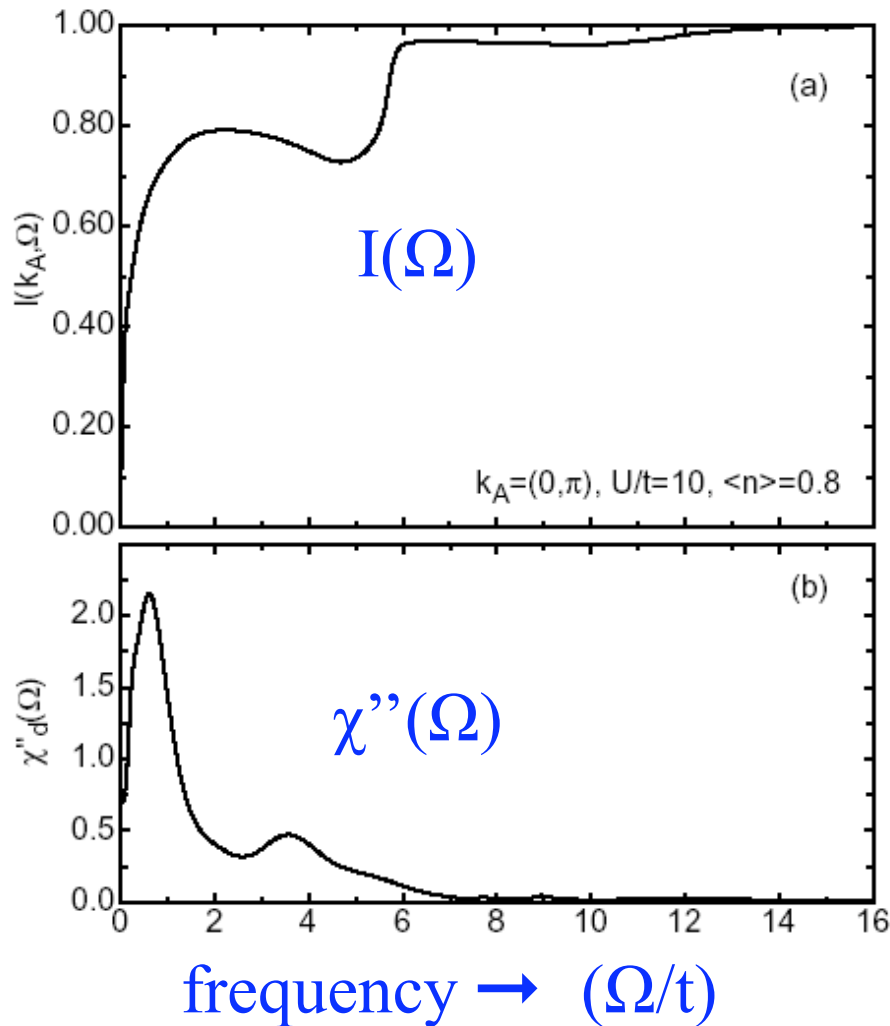
Goal: to determine the bosonic spectral function  $\alpha^2 F(\Omega)$  by spectroscopy.

# Hubbard model DCA calculations

Contribution from spin fluctuations to the gap function  $I(\Omega)$

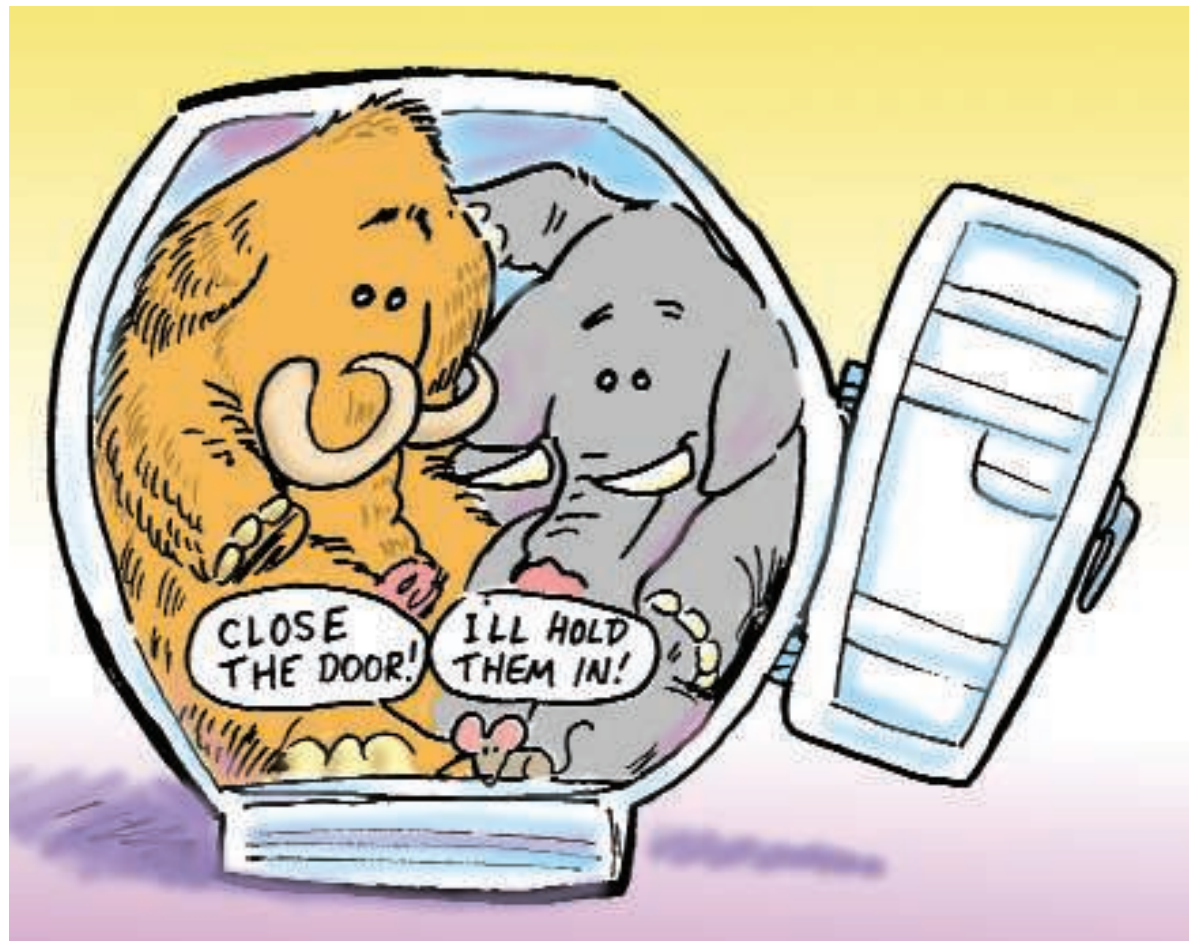
Maier *et al.* cond-mat/801.4506

Kyung *et al.* arXiv: 0812.1228



Anderson's view:  
The bosonic mouse vs RVB elephant &  
mammoth

P.W. Anderson, Science **317**, 1705 (2007)

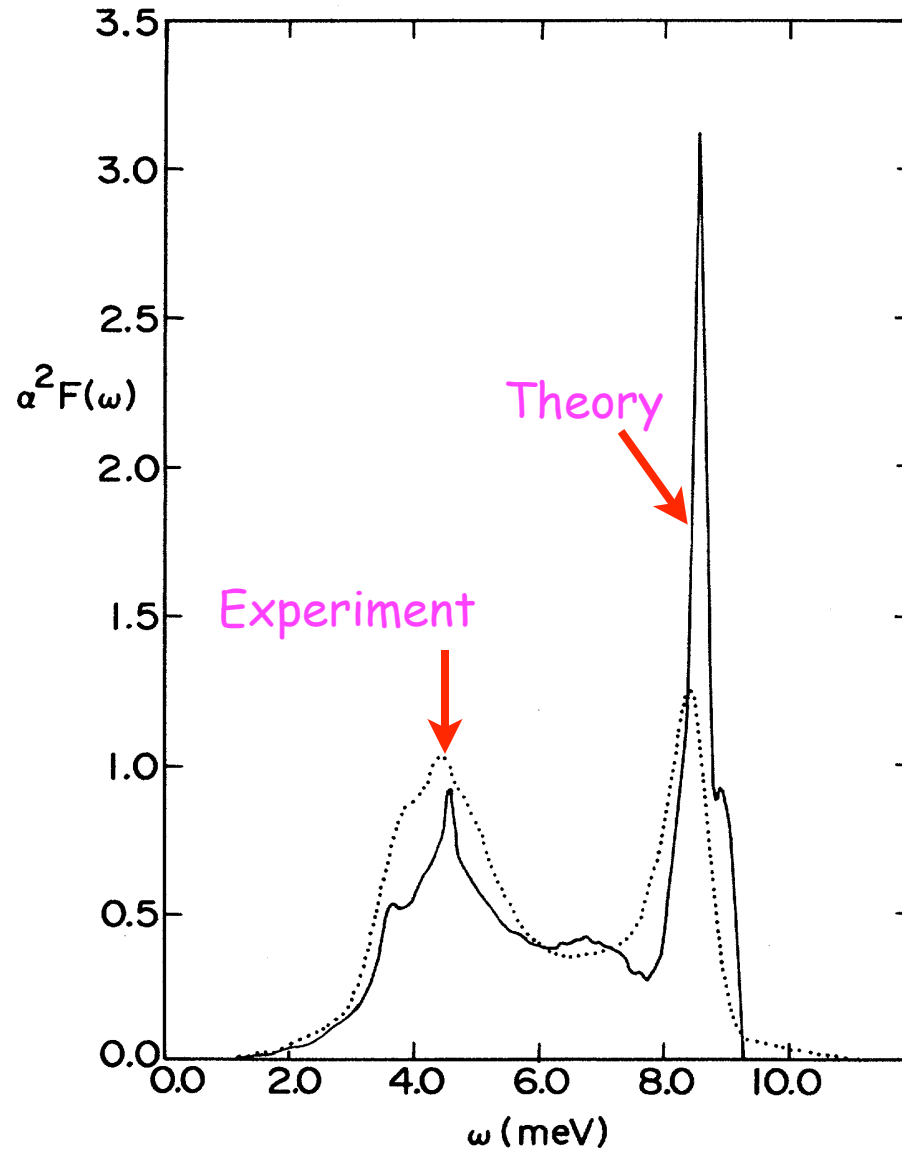


# Measuring self energy by spectroscopy

- Tunneling
- Angle resolved photo emission (ARPES)
- Optical reflectance



# Tunneling in lead



## Experiments:

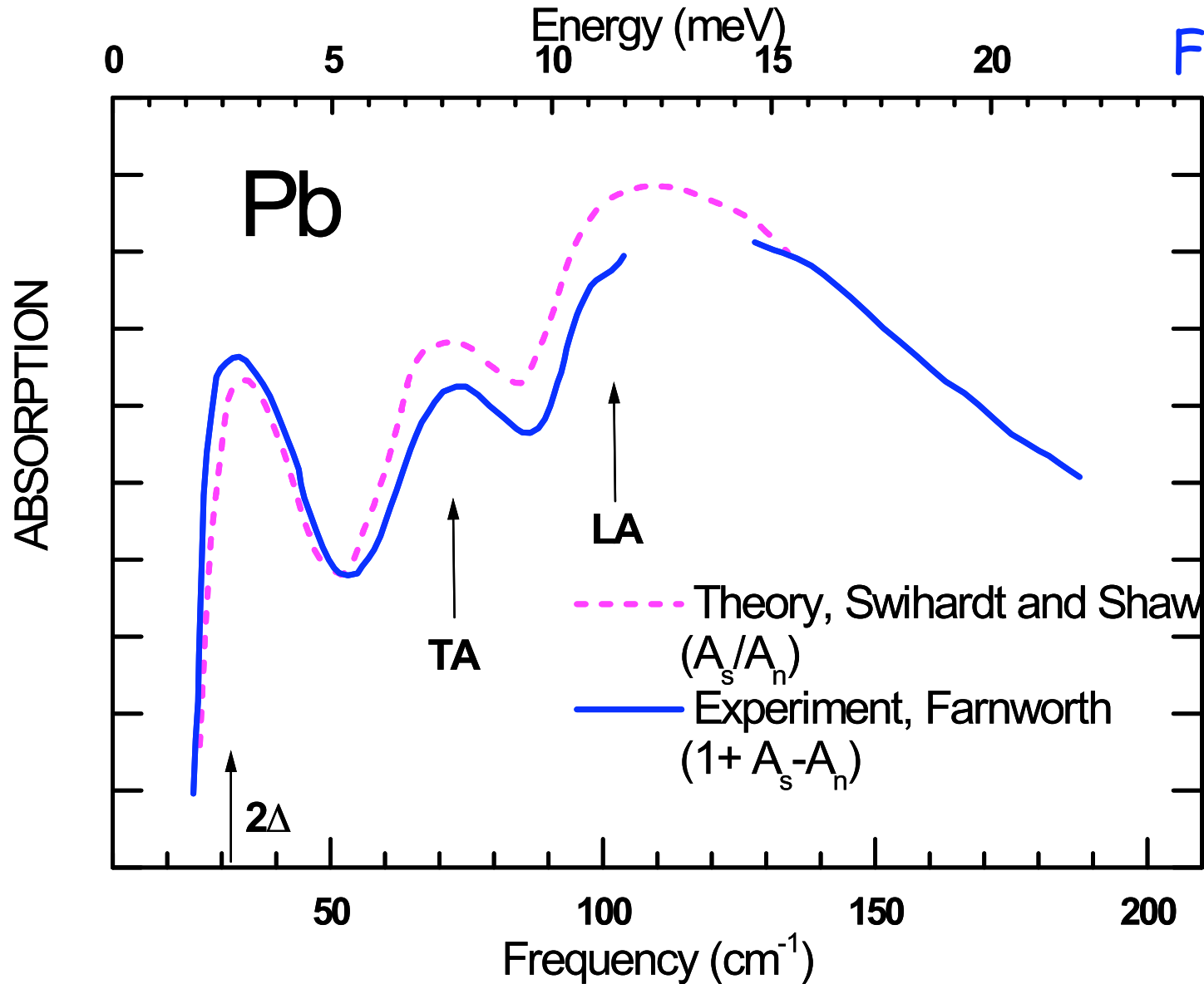
Rowell and McMillan  
tunneling  
Eliashberg inversion

## Theory:

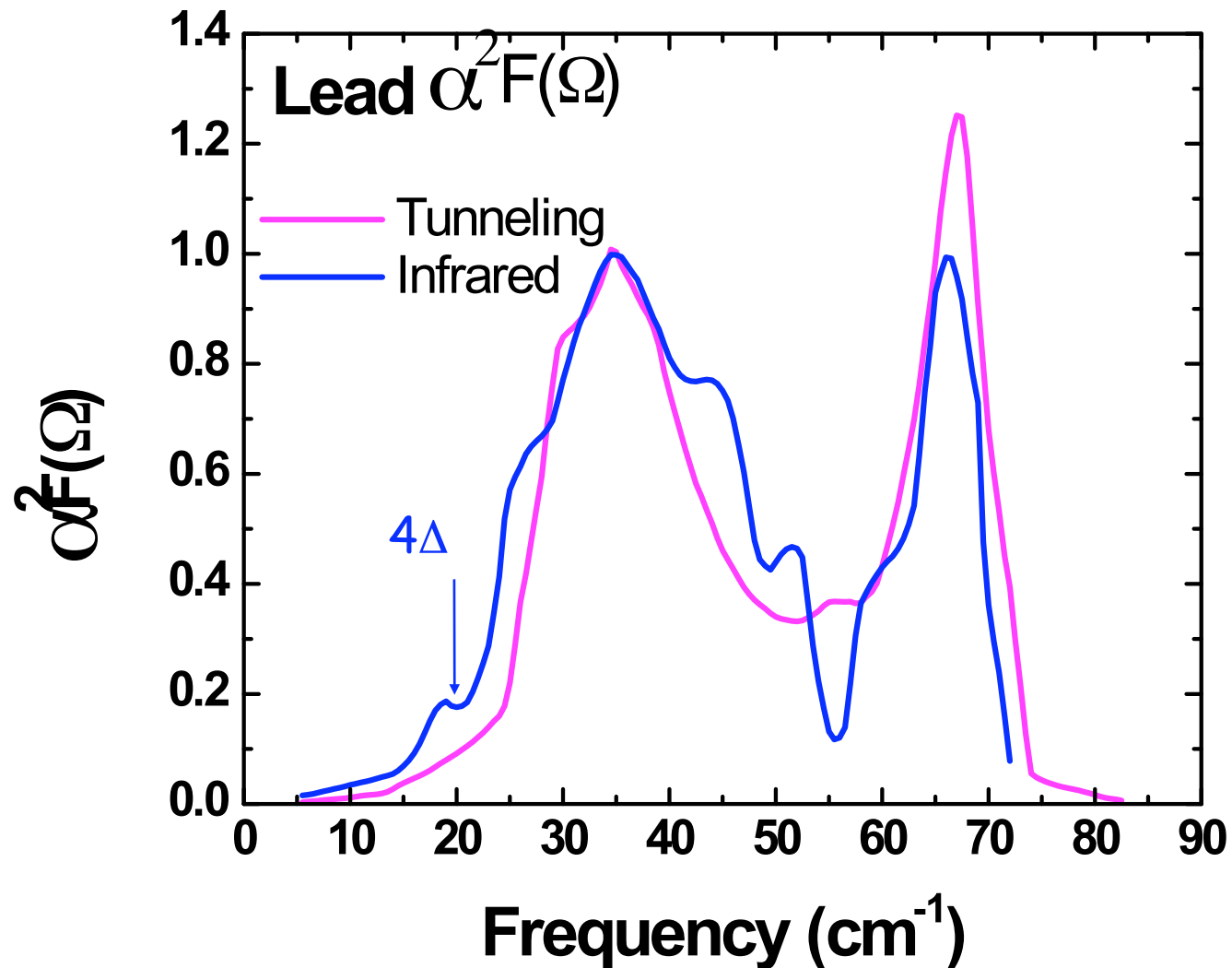
Tomlinson Carbotte  
Pseudopotential  
electron-ion potential  
Phonons from neutron  
scattering

# Optical absorption in Pb

Farnworth et al.



# Comparing tunneling and optics in Pb



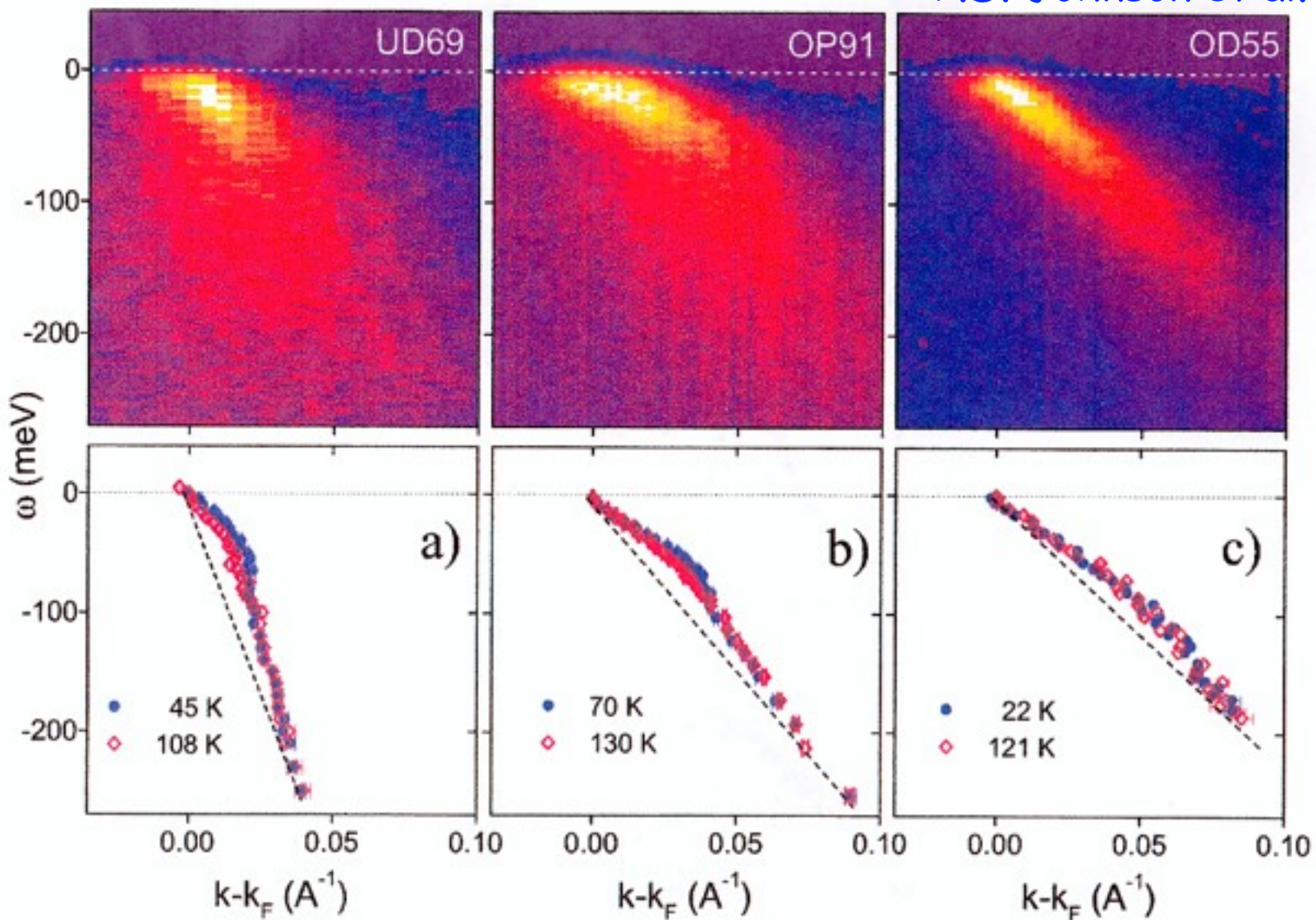
Can we do this in the new superconductors?

# Measuring self energy by spectroscopy

- Tunneling
- Angle resolved photo emission (ARPES)
- Optical reflectance

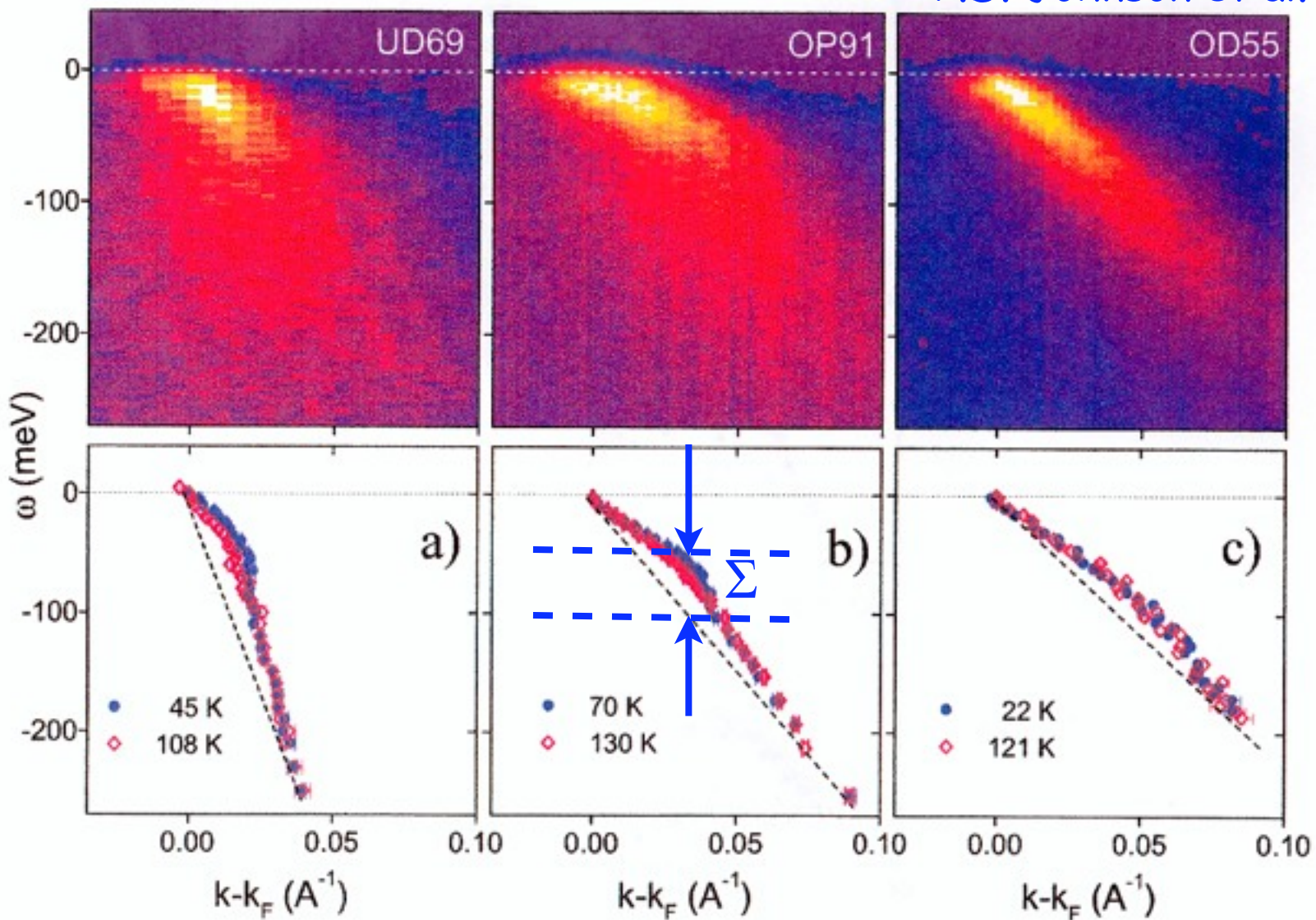
# Observing the ARPES "kink"

P.D. Johnson et al.



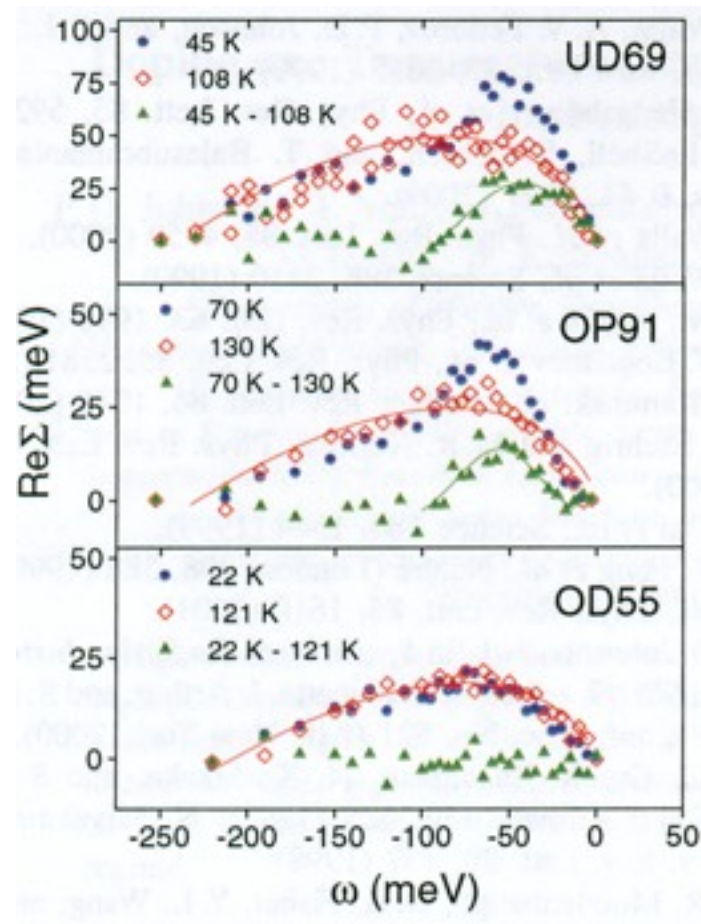
# Observing the ARPES "kink"

P.D. Johnson et al.



# Self energy $\Sigma$ observed by ARPES

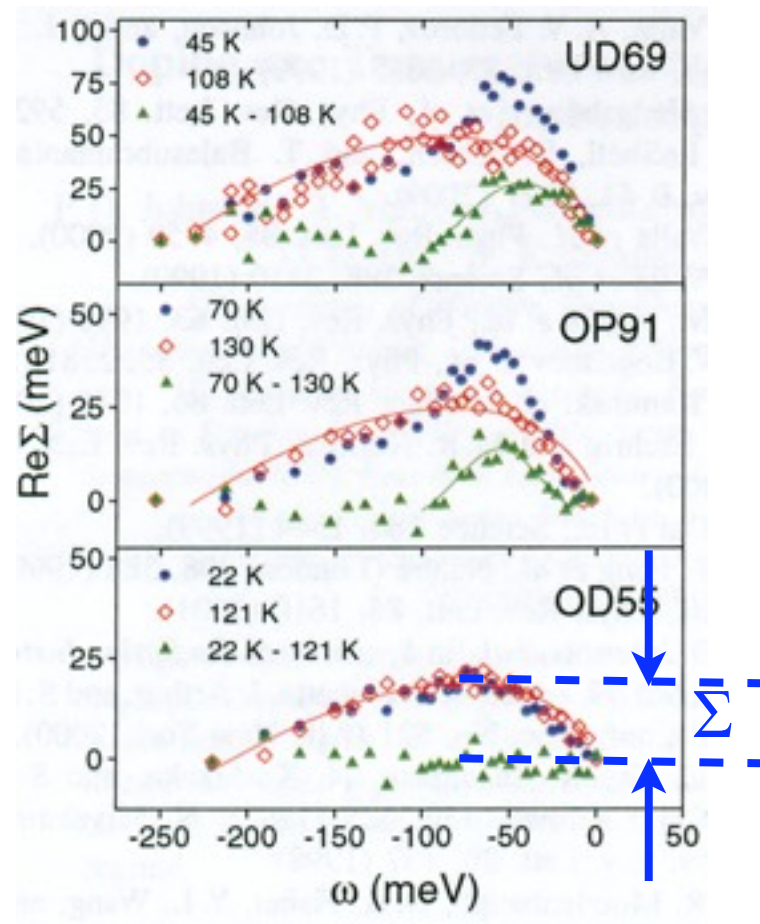
P.D. Johnson et al.





# Self energy $\Sigma$ observed by ARPES

P.D. Johnson et al.



# Measuring self energy by spectroscopy

- Tunneling
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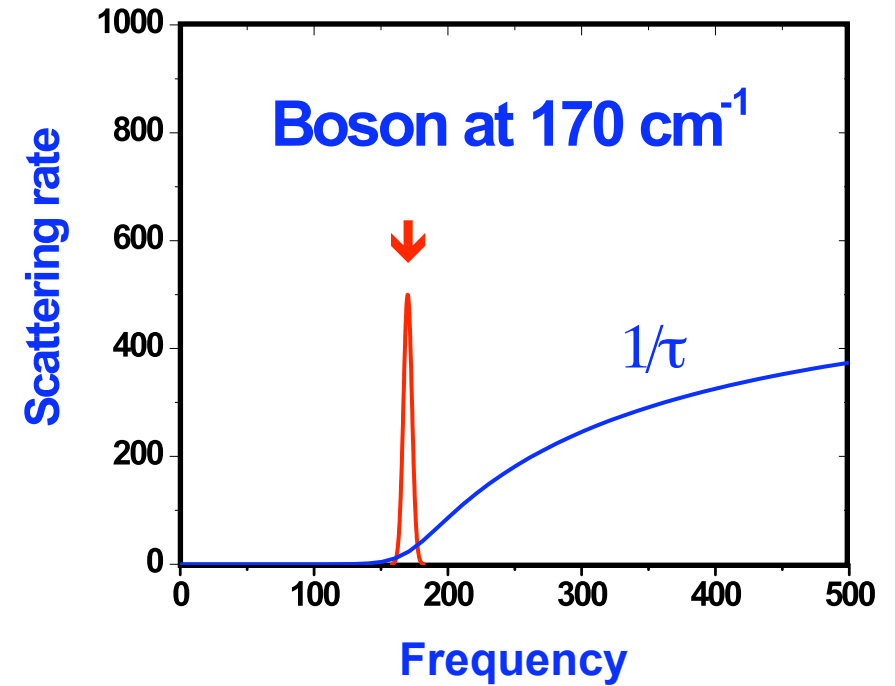
Standard optical formulas:

$$r = Re^{i\theta} = (\varepsilon - 1)/(\varepsilon + 1)$$

$$\omega (\varepsilon - \varepsilon_H) = 4\pi\sigma i$$

# Drude conductivity, extended

$$4\pi\sigma(\omega) = \frac{\omega_p^2}{1/\tau(\omega) - i\omega(1 + \lambda(\omega))}$$



# Extended Drude Model

$$4\pi\sigma(\omega) = i\omega_p^2 / (\omega - 2\Sigma^{\text{op}}(\omega, T))$$

$\Sigma^{\text{op}} = \Sigma_1^{\text{op}} + i\Sigma_2^{\text{op}}$  Optical self energy

$1/\tau(\omega, T)$  Frequency dependent scattering rate

$\lambda$  Mass enhancement  $m^* = m(1 + \lambda)$

$\omega_p = 4\pi ne^2/m$  Plasma frequency

$\Sigma_1(\omega, T) = -\lambda\omega/2$  Real part of self energy

$\Sigma_2(\omega, T) = -1/(2\tau)$  Imaginary part of self energy

The parameters  $1/\tau(\omega, T)$  and  $\lambda(\omega, T)$  can be obtained from the experimental  $\sigma(\omega) = \sigma_1 + i\sigma_2$

$$1/\tau(\omega, T) = ne^2/m \operatorname{Re}(1/\sigma(\omega))$$

$$\omega(1 + \lambda(\omega)) = ne^2/m \operatorname{Im}(1/\sigma(\omega))$$

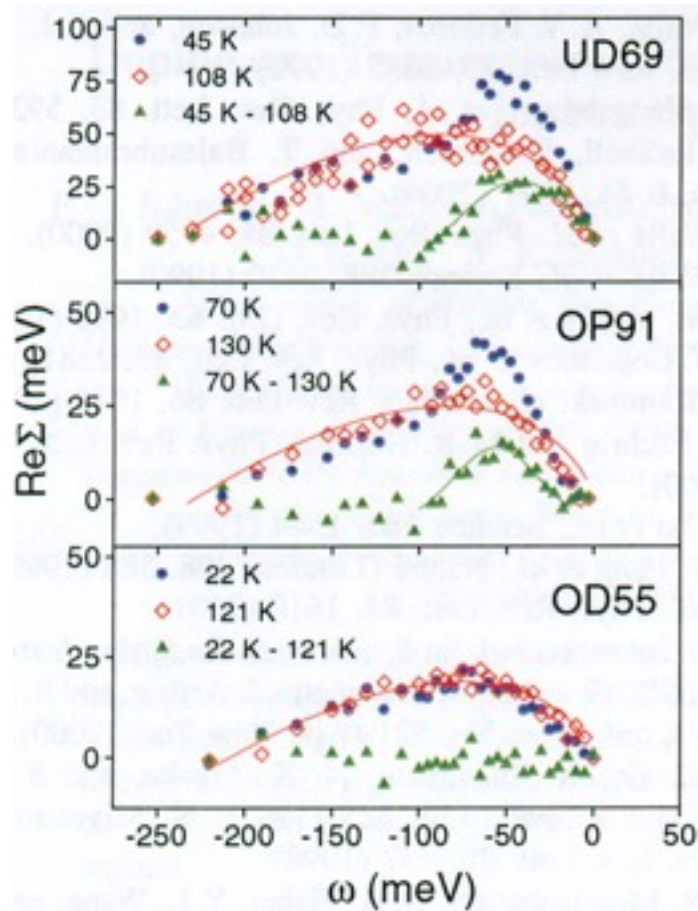
# Caveats

- $\Sigma^{\text{op}}$  is not exactly the self energy  $\Sigma$  as measured by ARPES.
  - The optical conductivity is an average over the Fermi surface
  - There is an additional factor of  $(1-\cos \theta)$  where  $\theta$  is the scattering angle
- The plasma frequency  $\omega_p$  must be known
- For the extended Drude formalism there must only be channel of conductivity.
- $\alpha^2_{\text{tr}}F(\Omega)$  is not the same as  $\alpha^2F(\Omega)$

# Comparison of ARPES and infrared $\Sigma(\omega)$

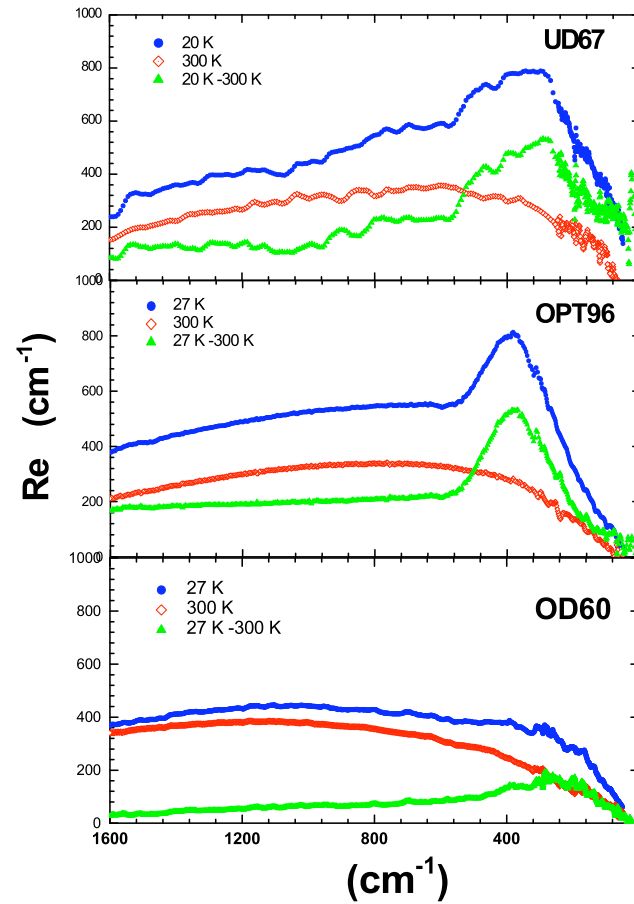
ARPES

P.D. Johnson et al.



Infrared

J. Hwang et al.

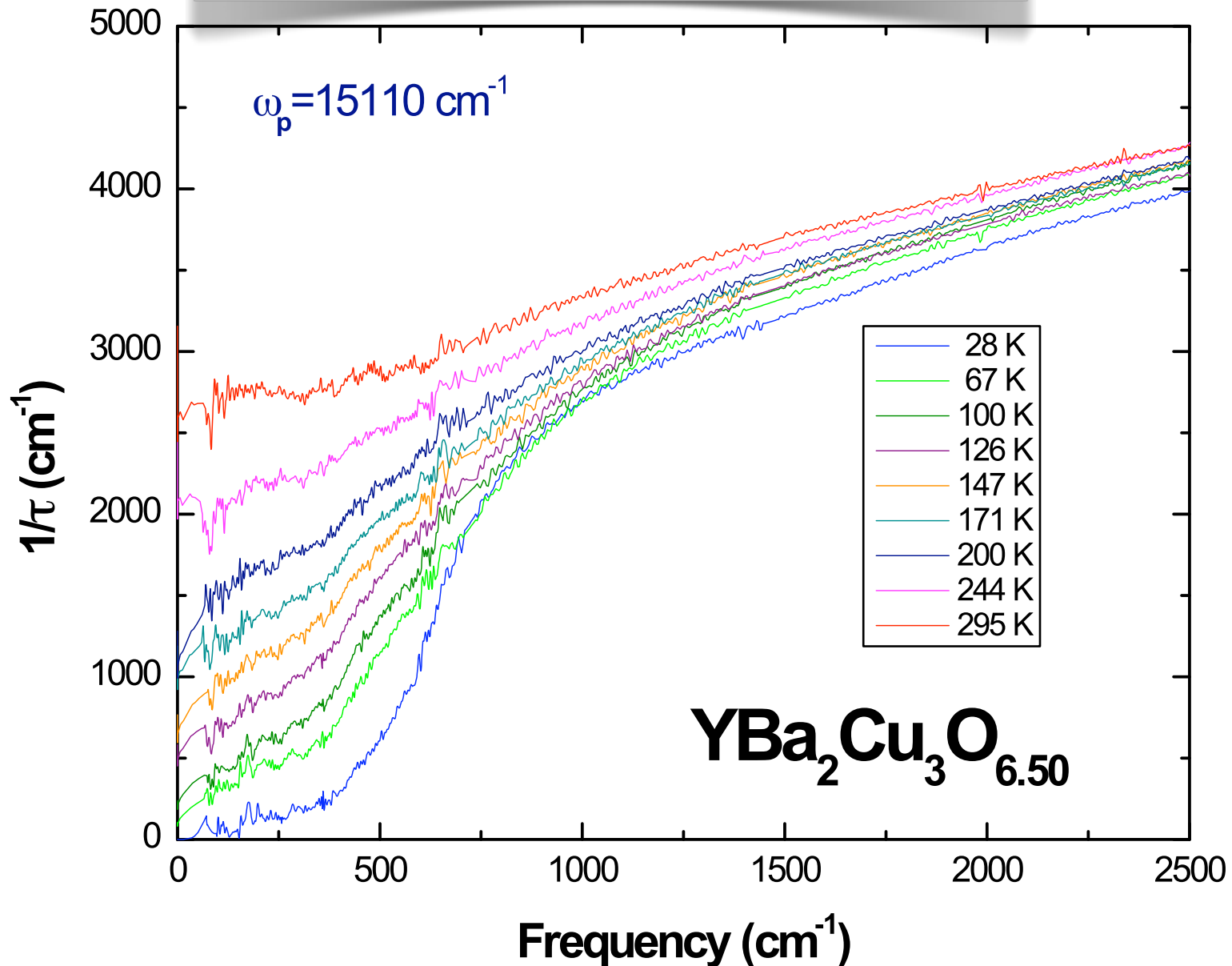


Bosonic modes in YBCO,  
evidence for magnetic  
scattering

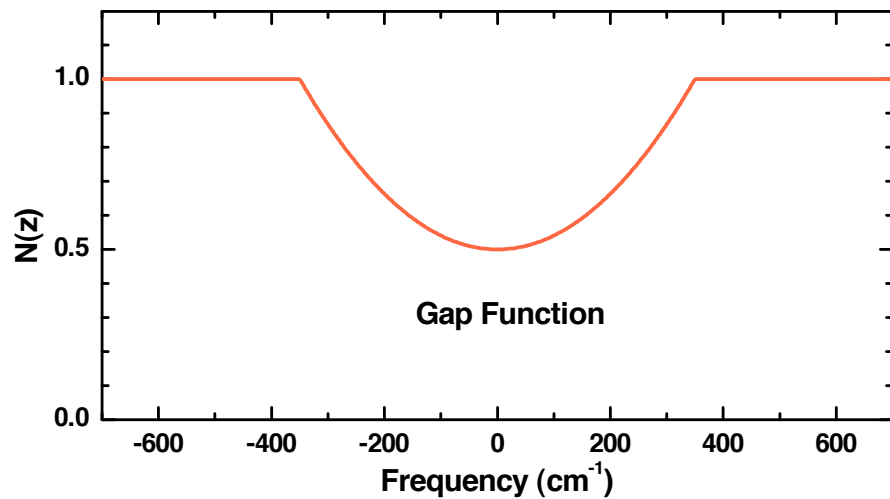
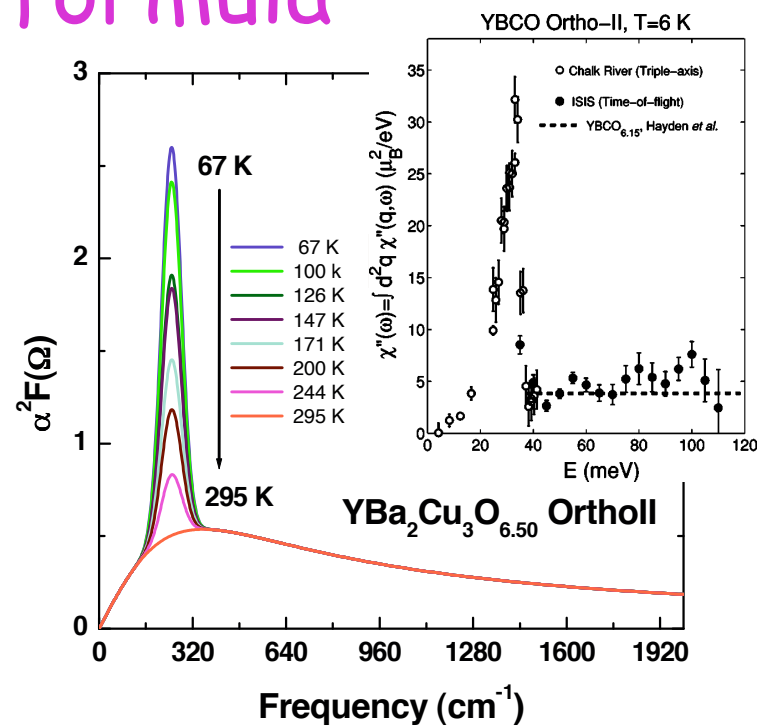
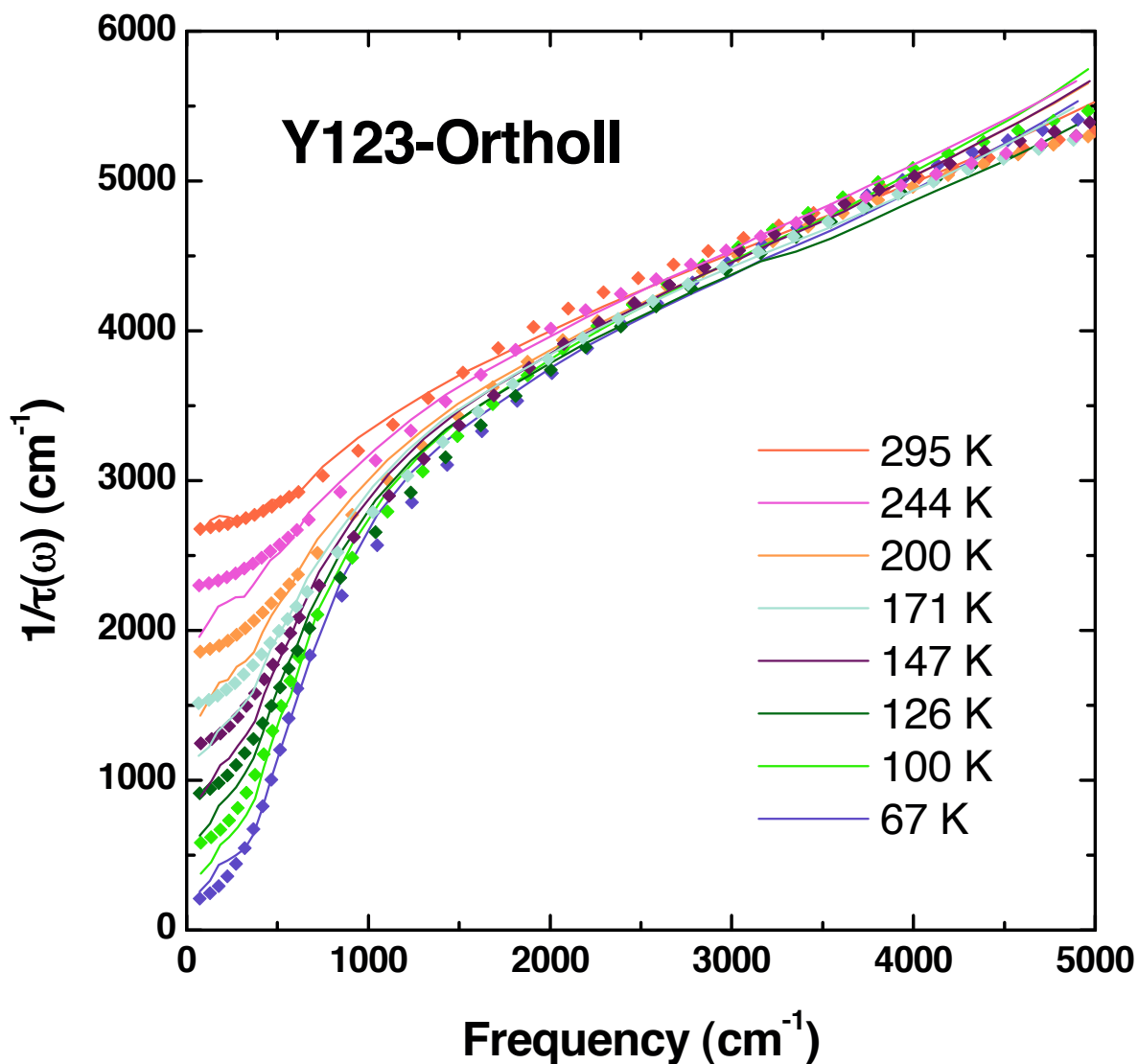


# Ortho II YBCO scattering rate

$$1/\tau(\omega, T) = ne^2/m \operatorname{Re}(1/\sigma(\omega))$$



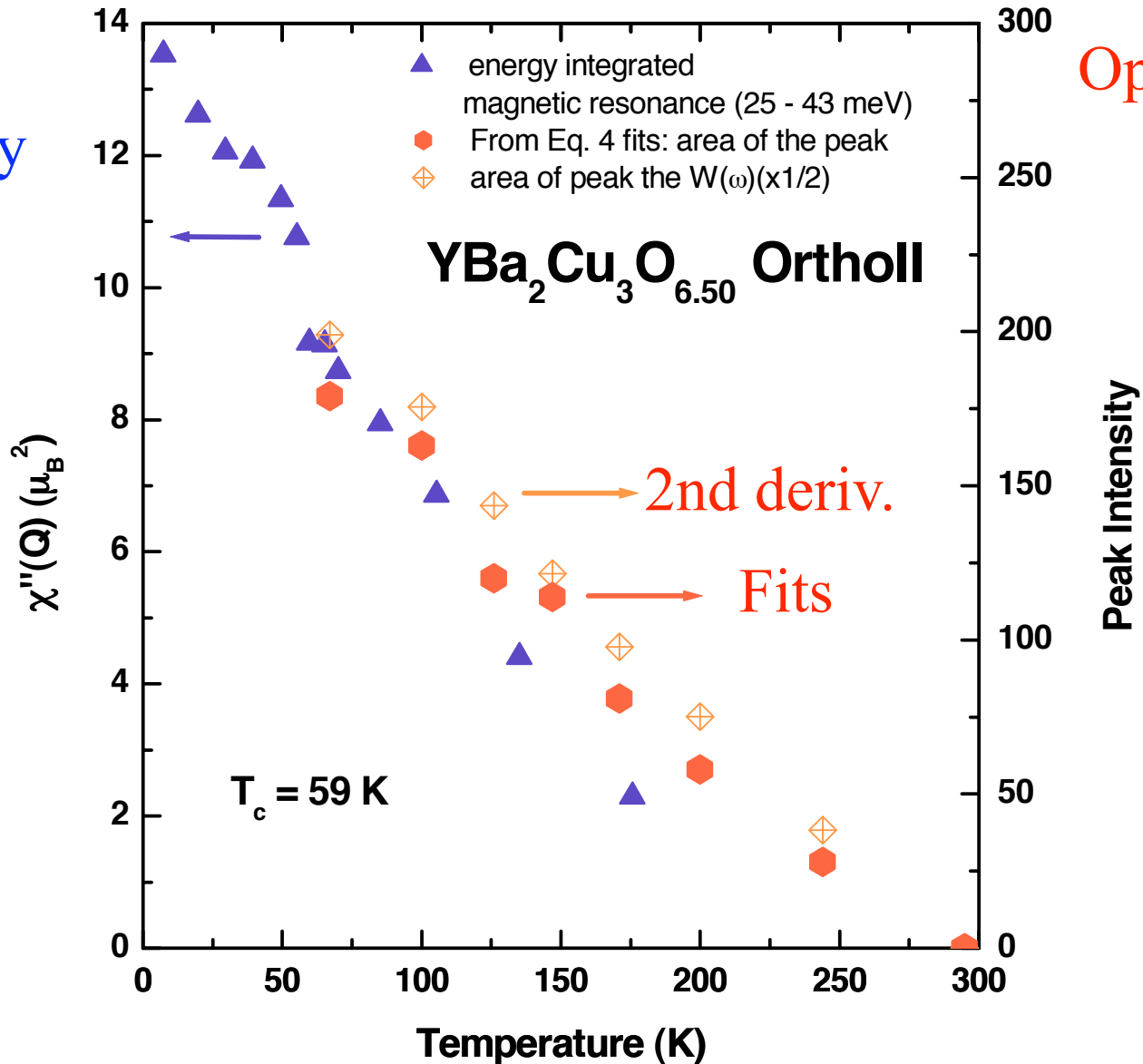
# Fits to Shulga's formula



# Temperature dependence of the mode spectral weight

Neutron susceptibility

Optical amplitude

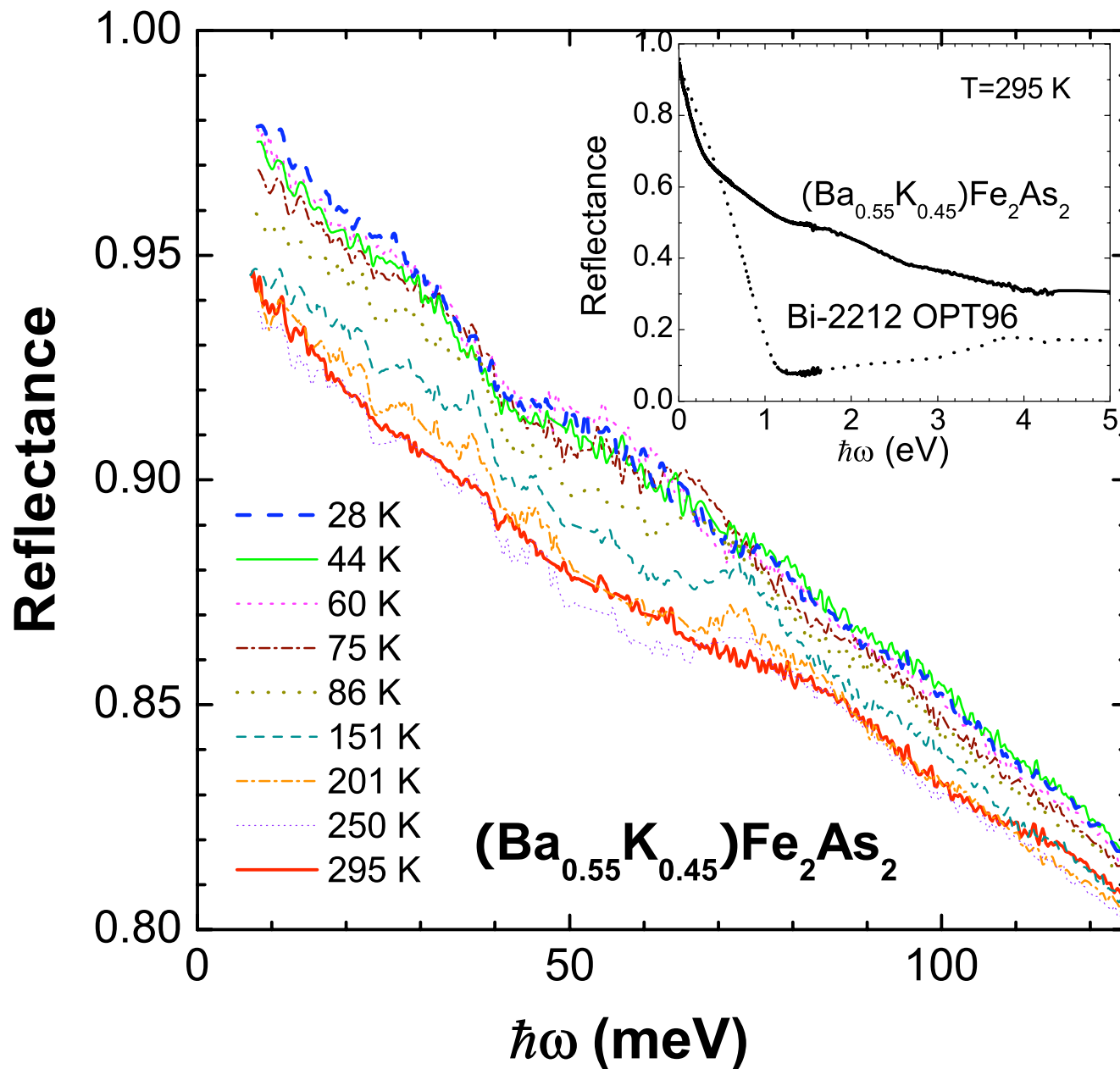


## Inversion of the $1/\tau$ spectra

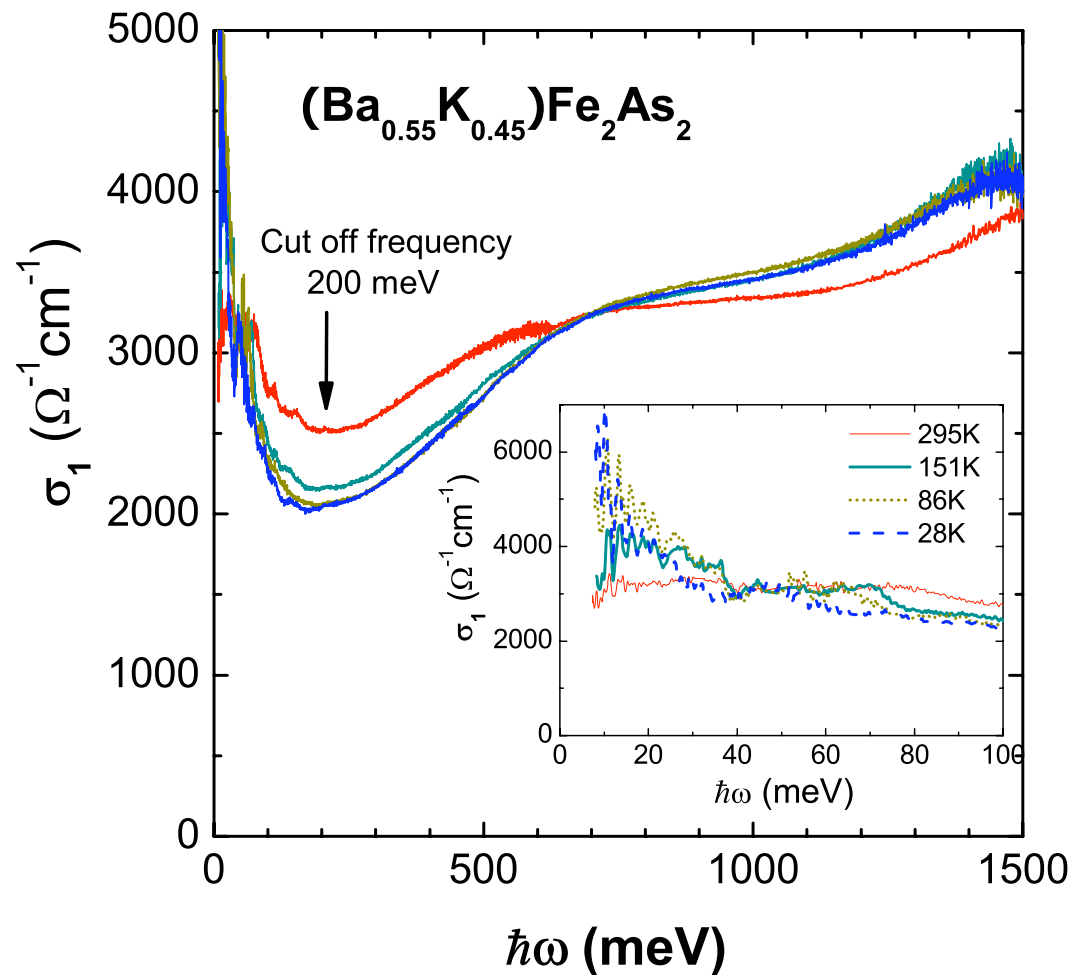
- Allen approximation for  $1/\tau$  solved by maximum entropy inversion to give  $I^2\chi(\omega)$
- Use Kubo's formula with  $I^2\chi(\omega)$  to fit  $1/\tau$
- d-wave full Eliashberg inversion gives  $T_c$

# Bosonic modes in $\text{Ba}_{0.35}\text{K}_{0.45}\text{Fe}_2\text{As}_2$

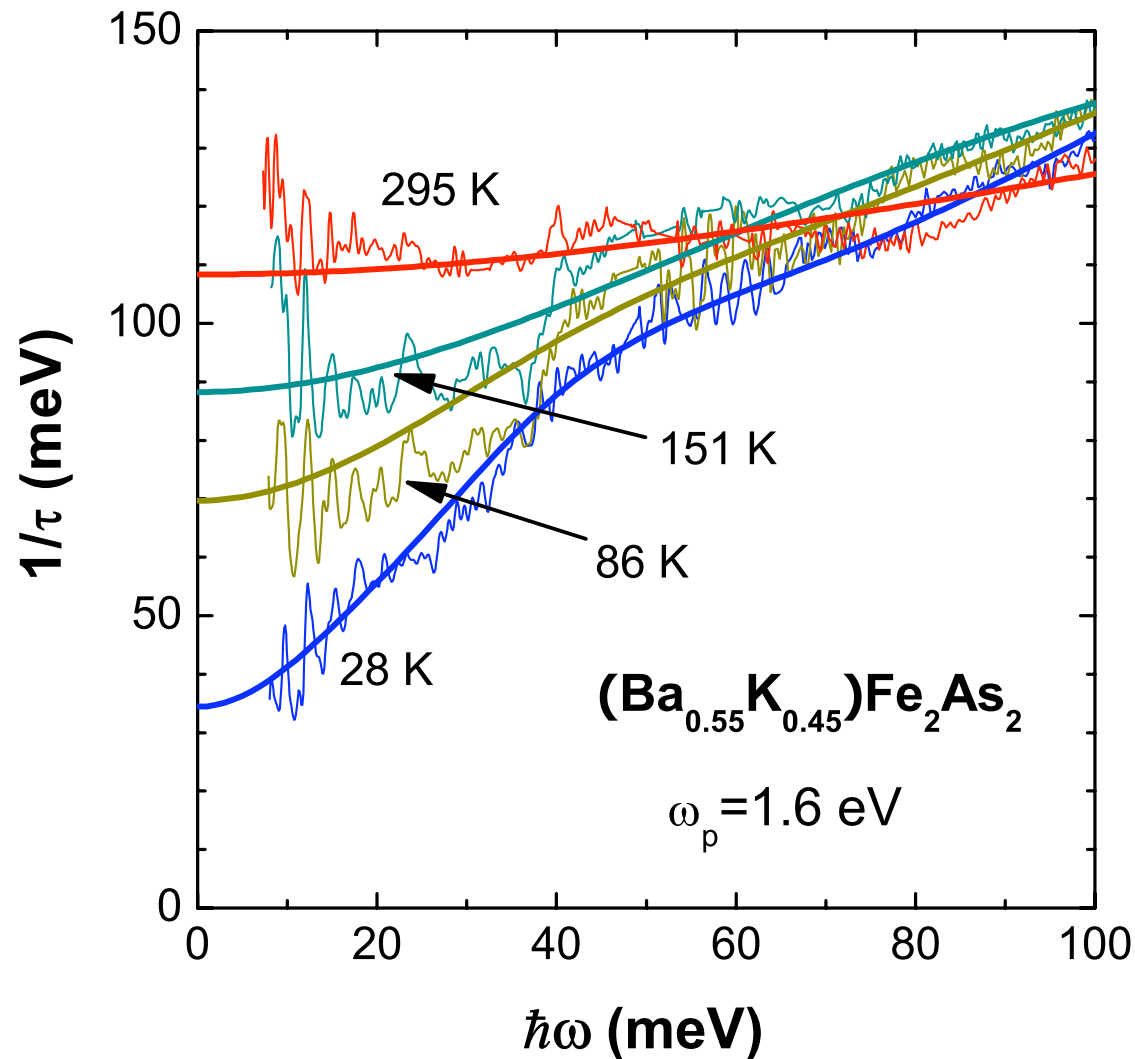
# Reflectance of $\text{Ba}_{0.35}\text{K}_{0.45}\text{Fe}_2\text{As}_2$



# Conductivity

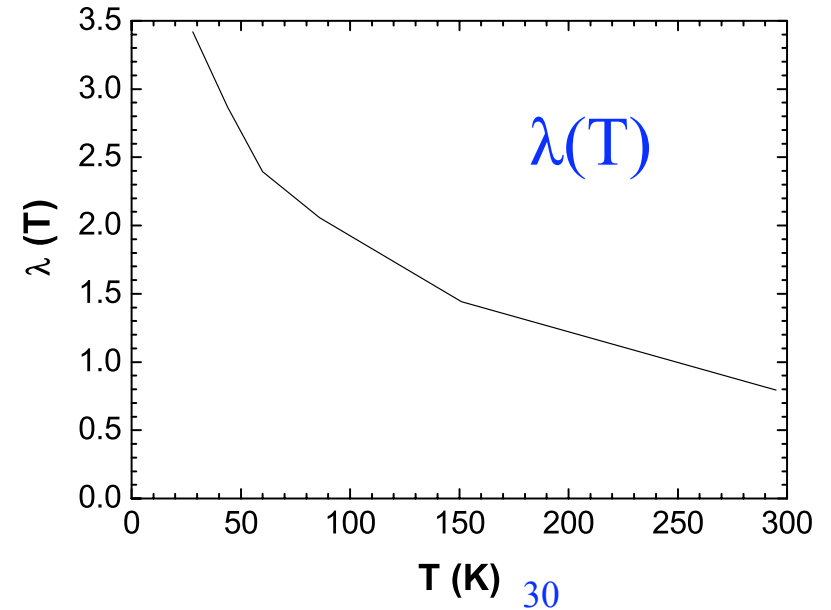
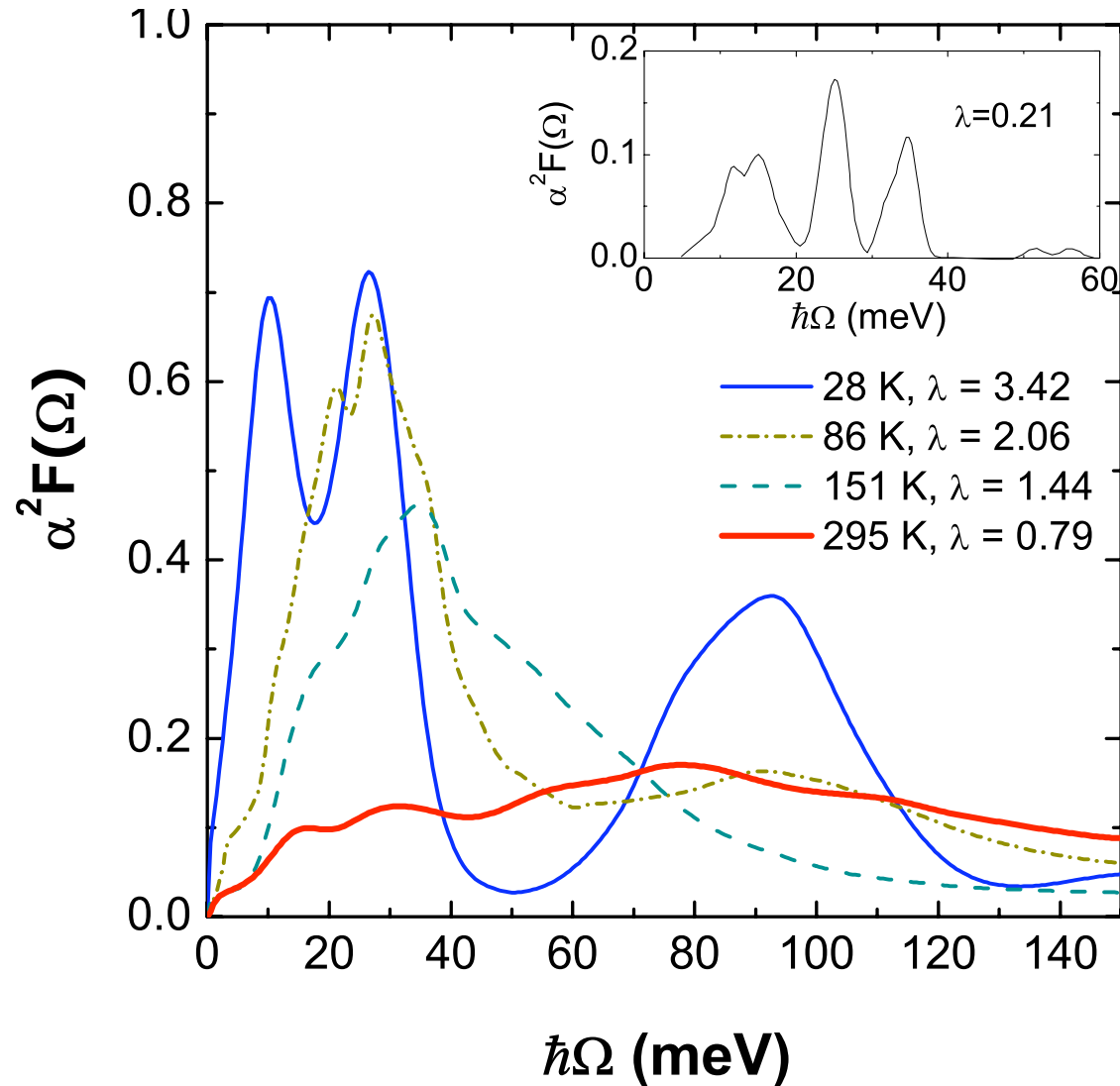


# The scattering rate

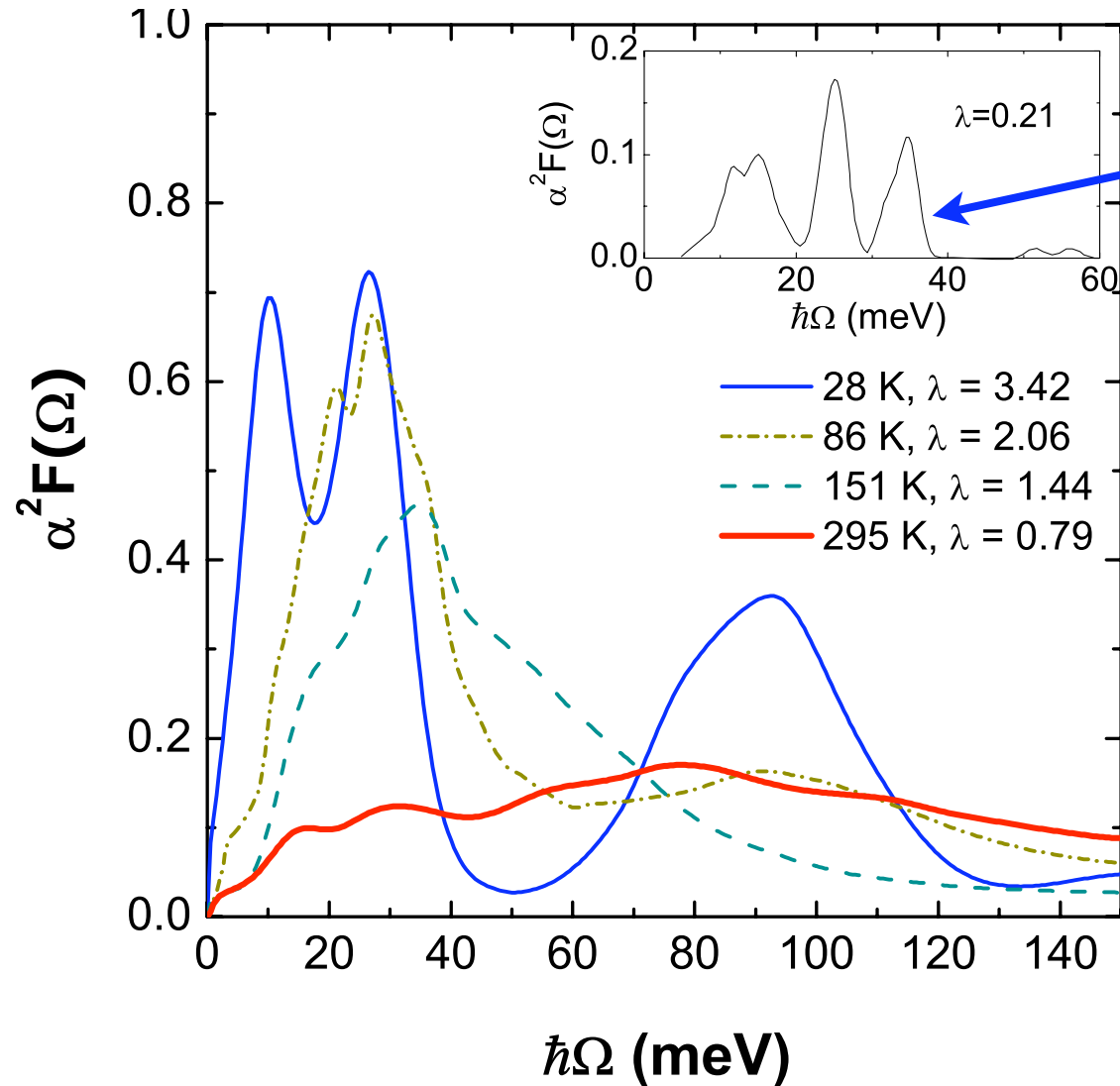




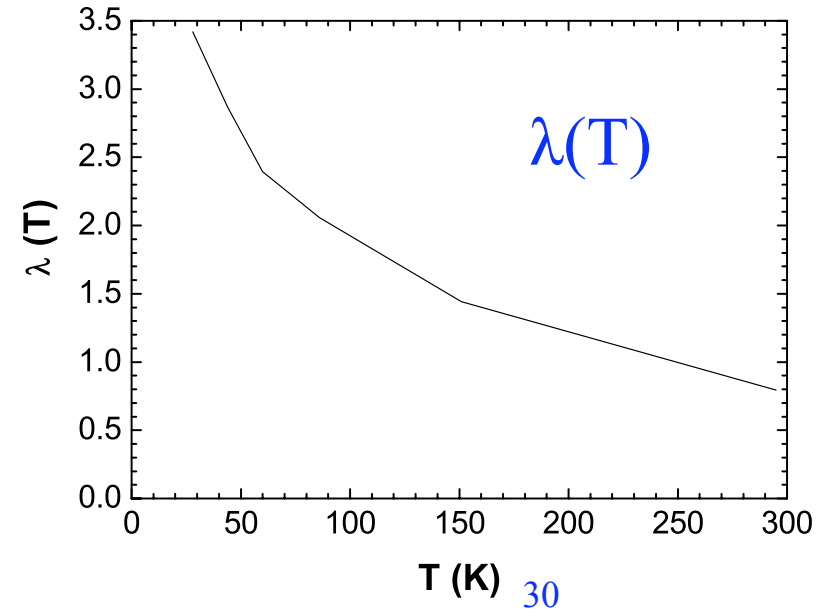
# The bosonic spectral function



# The bosonic spectral function



Phonons, theory

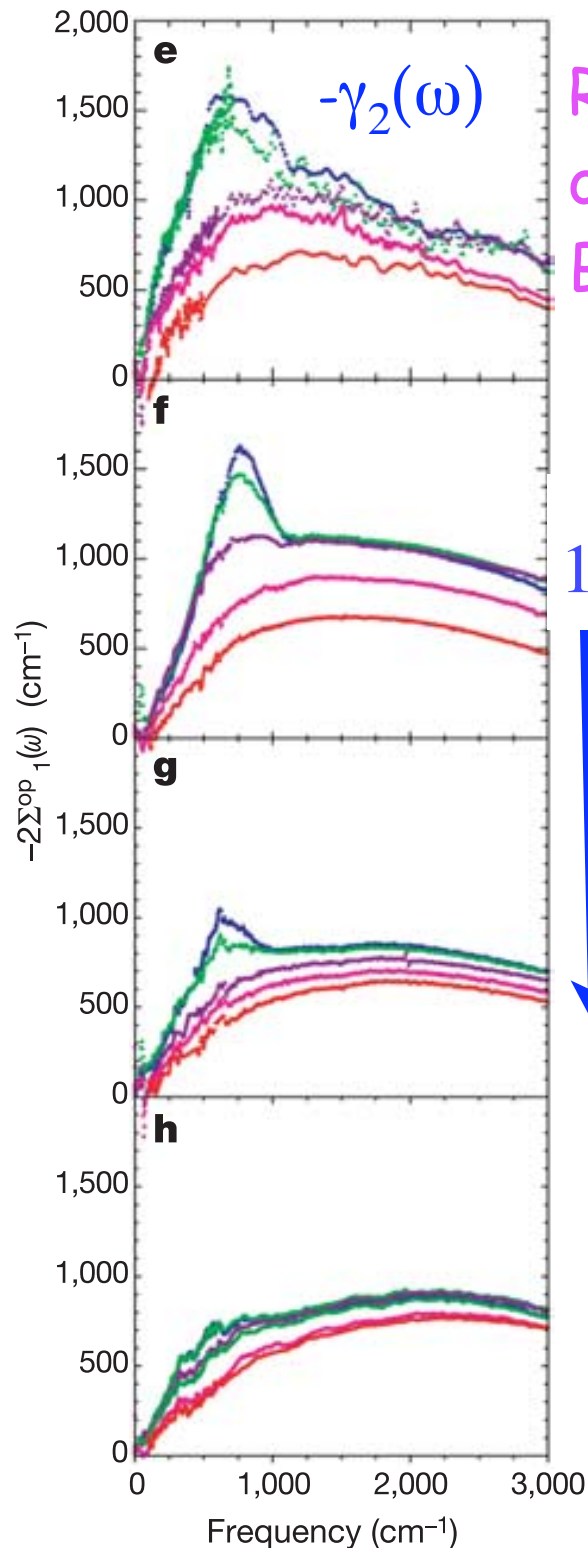
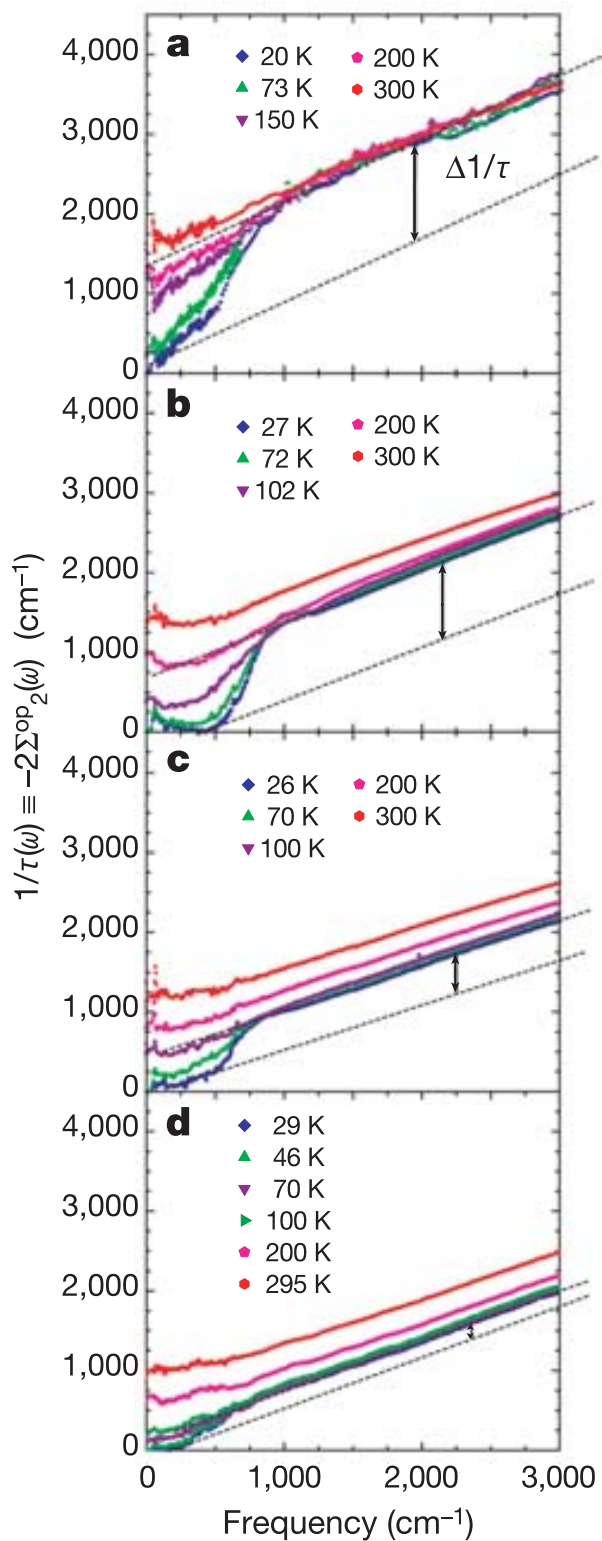


$\lambda(T)$

T (K) 30

# Bi-2212

- Doping dependence of the self energy
- Comparison with ARPES
- New results on the bosonic spectral function



Real and imaginary parts  
of the scattering rate in  
Bi-2212    Hwang et al. 2003

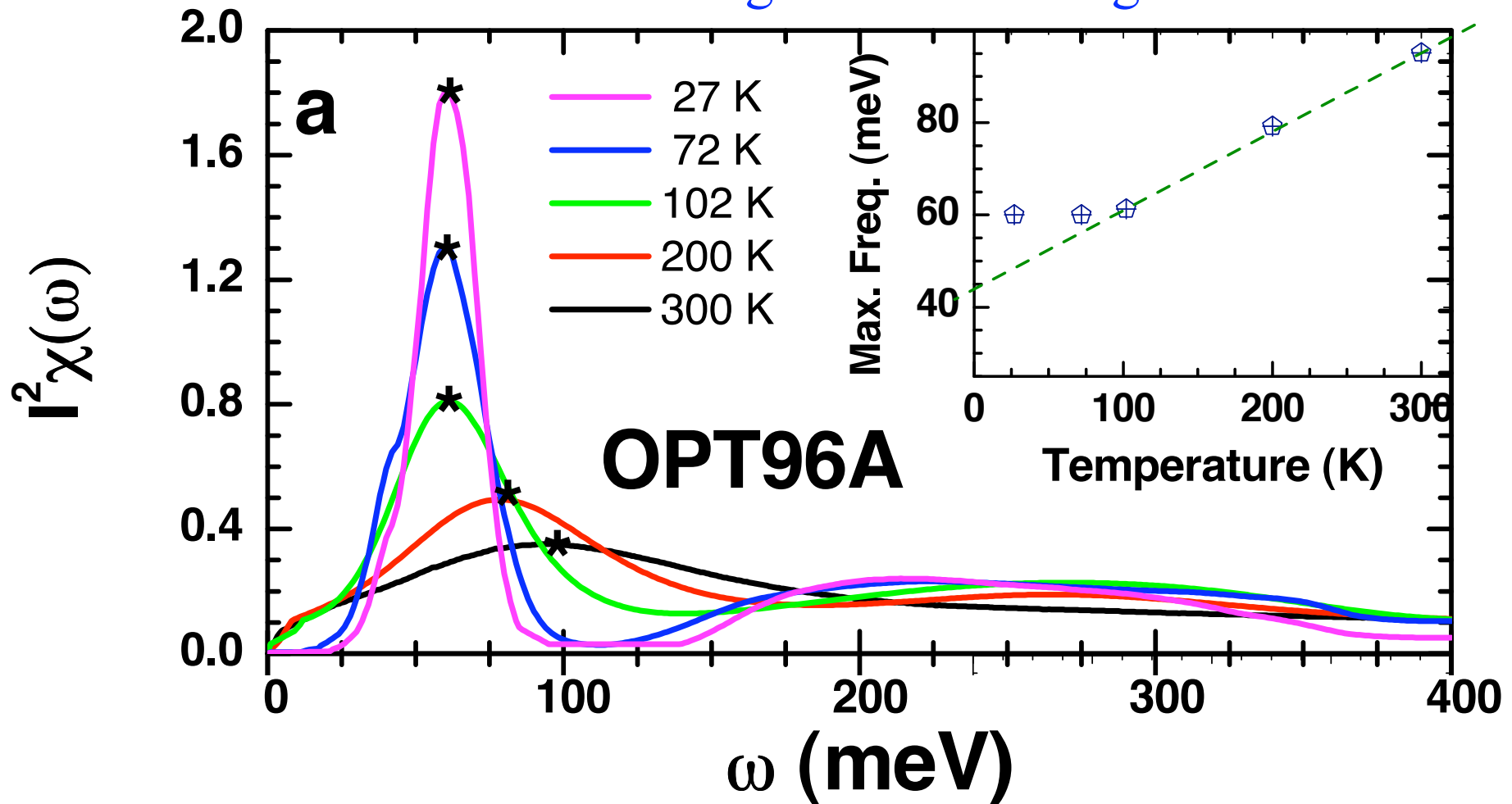
Optical self energy:

$$1/\tau = -2(\Sigma_2^{op}(\omega) + i\Sigma_1^{op}(\omega))$$

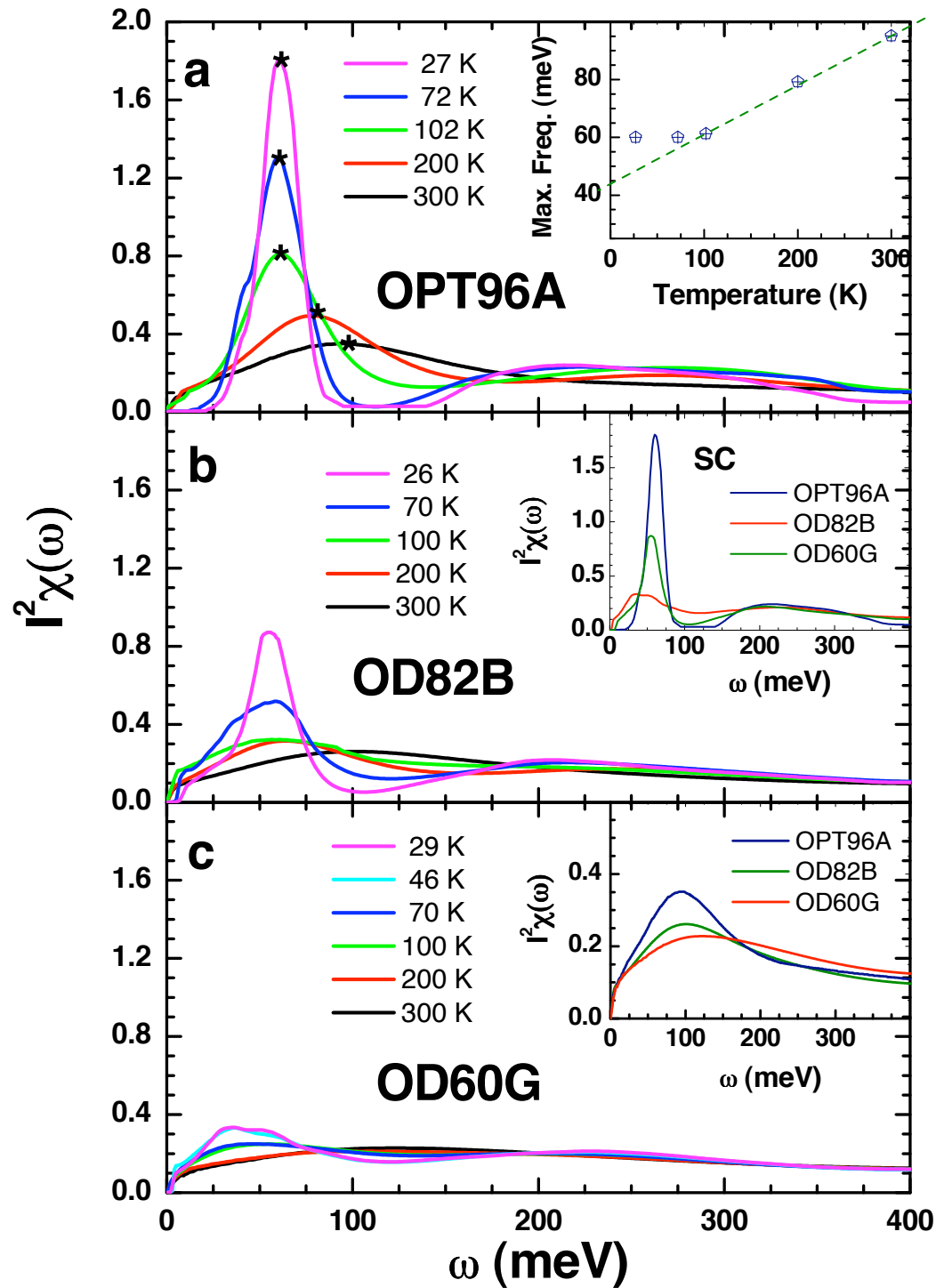
Doping

# Bi-2212 Eliashberg inversion

Hwang with Schachinger and Carbotte

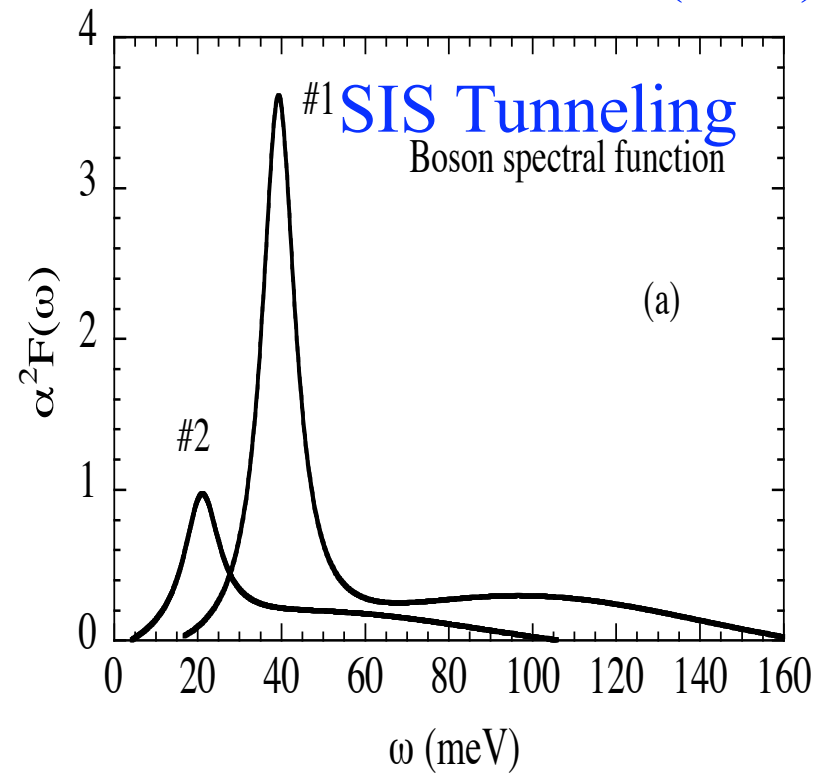
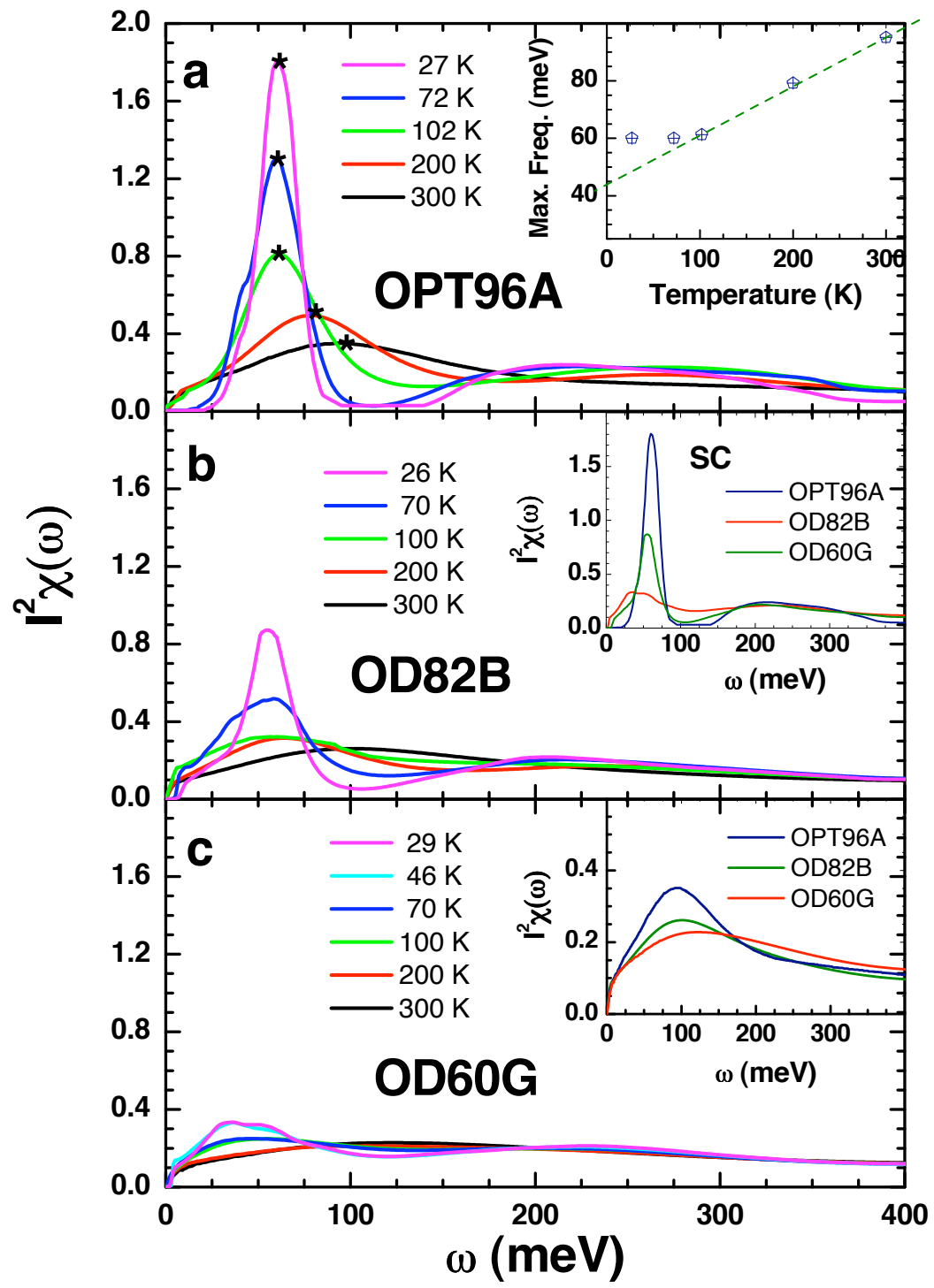


# Eliashberg inversion, doping dependence



# Comparison with tunneling data

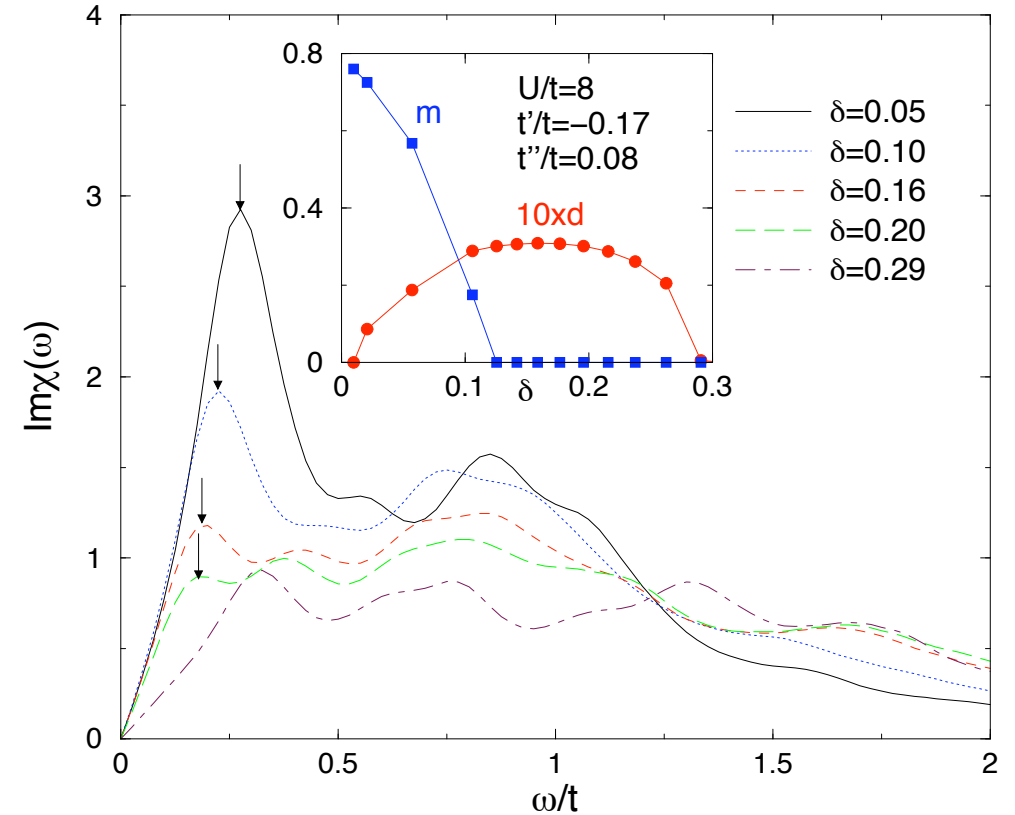
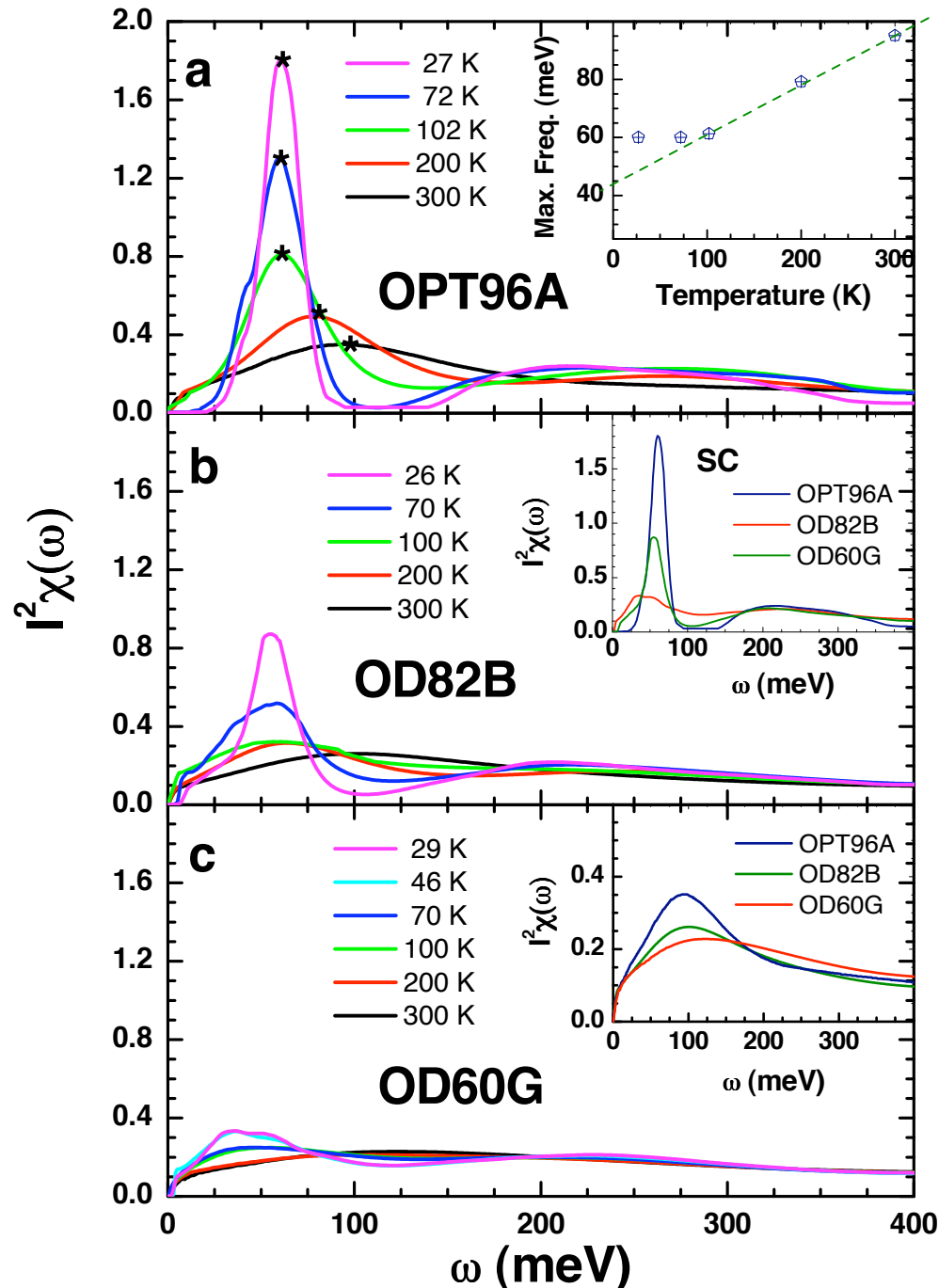
Zasadzinski  
*et al.* (2005)



# Eliashberg inversion, doping dependence

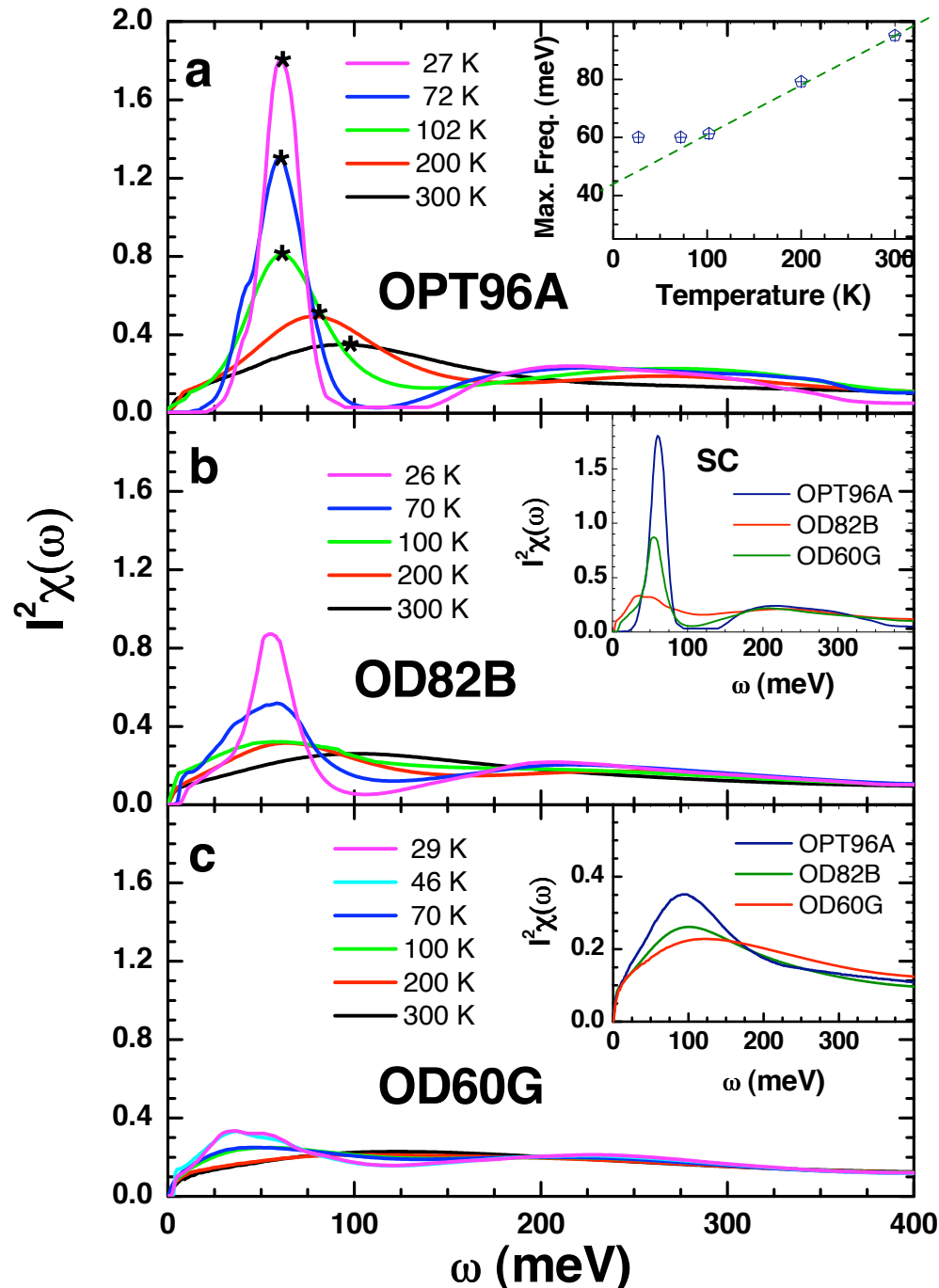
# Comparison with DMF theory

Kyung *et al.* arXiv:  
812.1228

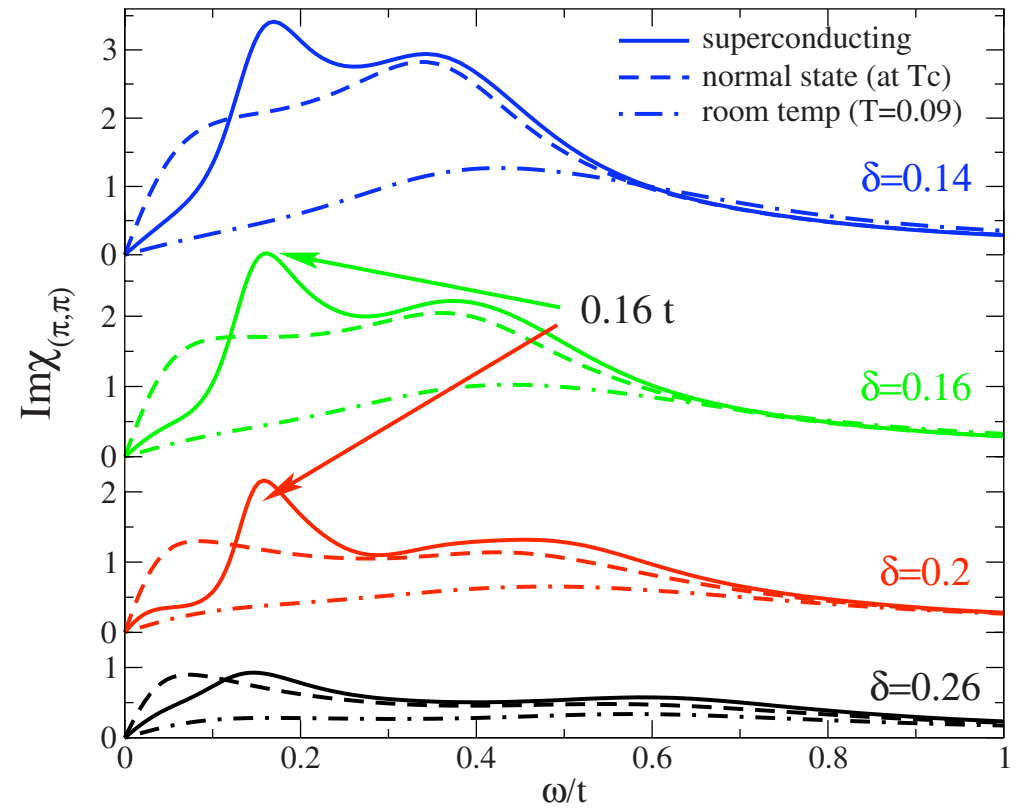




# Eliashberg inversion, doping dependence



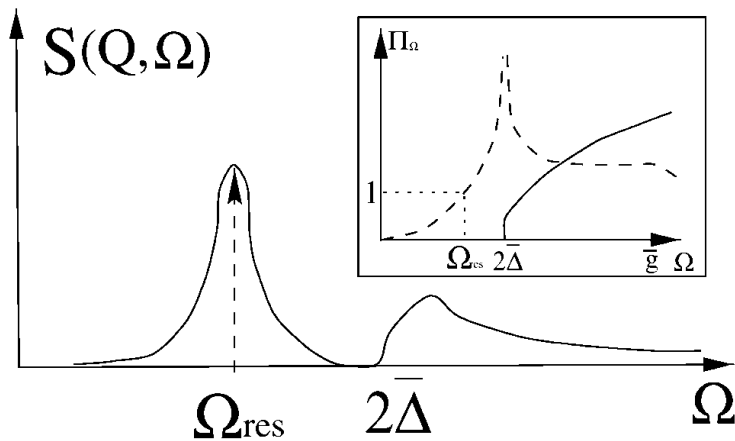
Haule & Kotliar PRB 76, 104509 (2007)



# Other theoretical ideas for the magnetic susceptibility

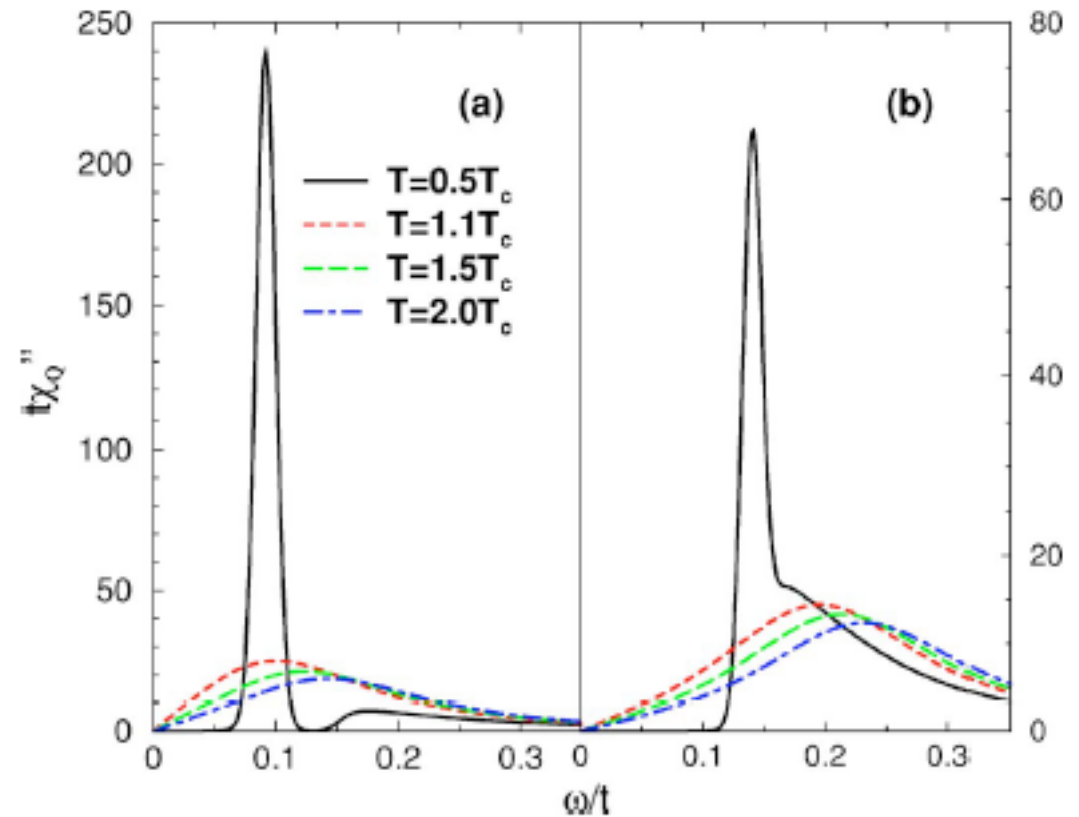
Exciton model

Eschrig & Norman  
Abanov & Chubukov



Memory function

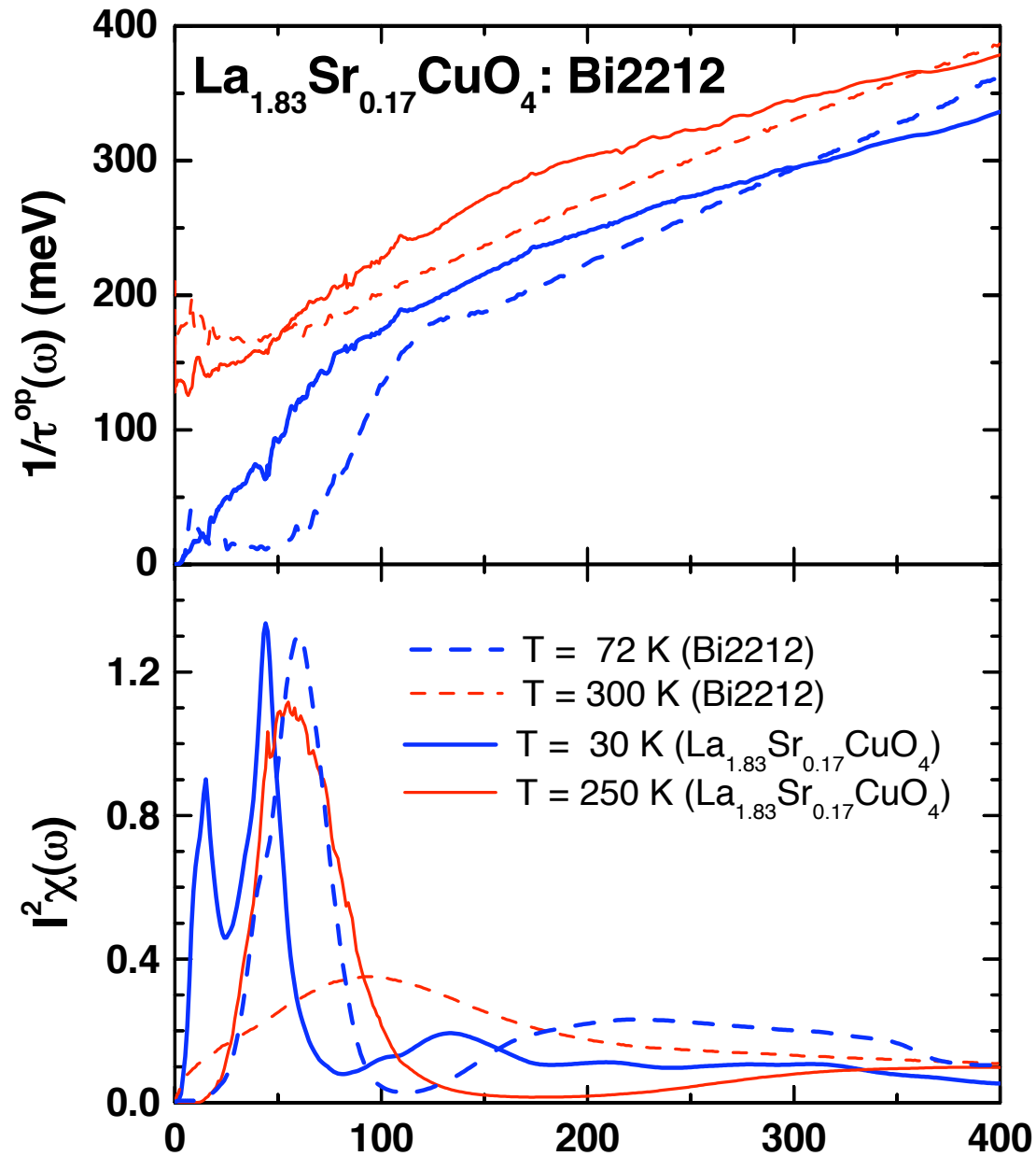
Prelovsek & Sega



# LSCO

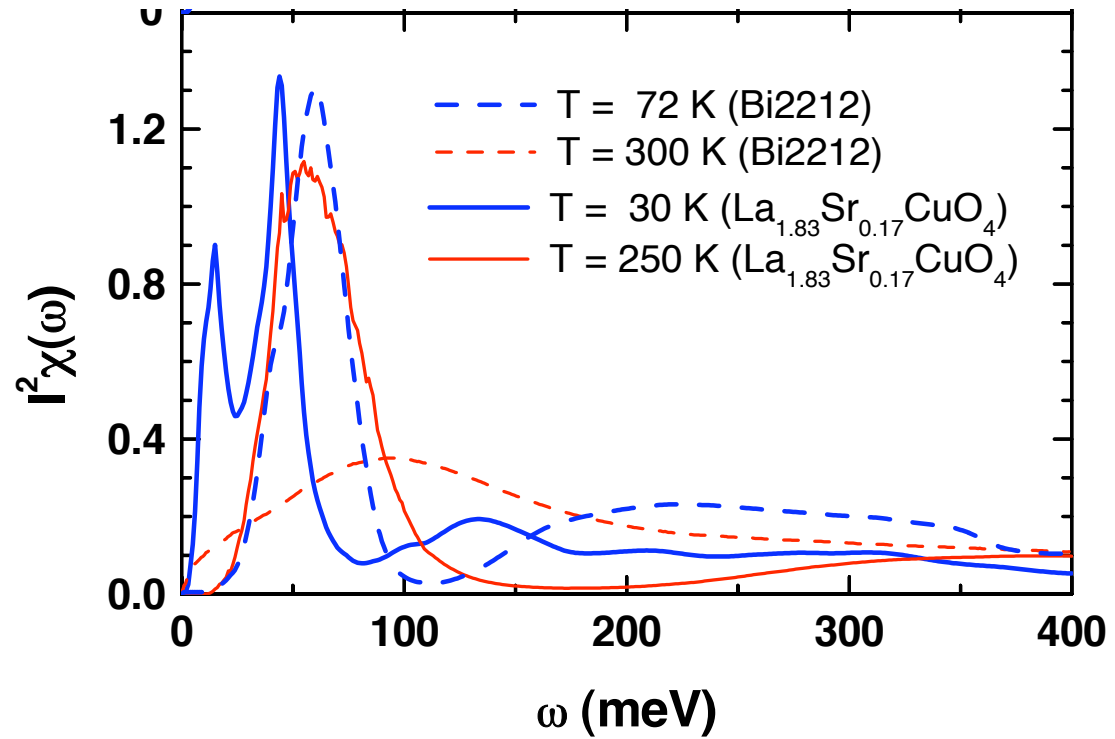
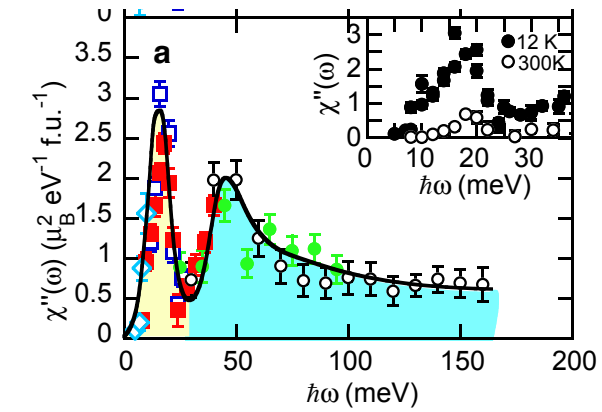
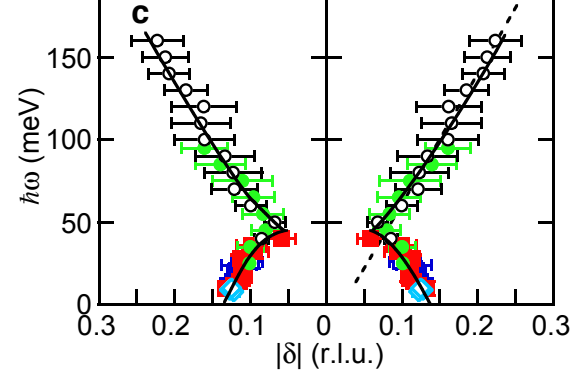
- Odd material,  $T_c = 30$  K
- Comparison with neutron scattering
- "Prediction of  $T_c$ "

# LSCO self energy



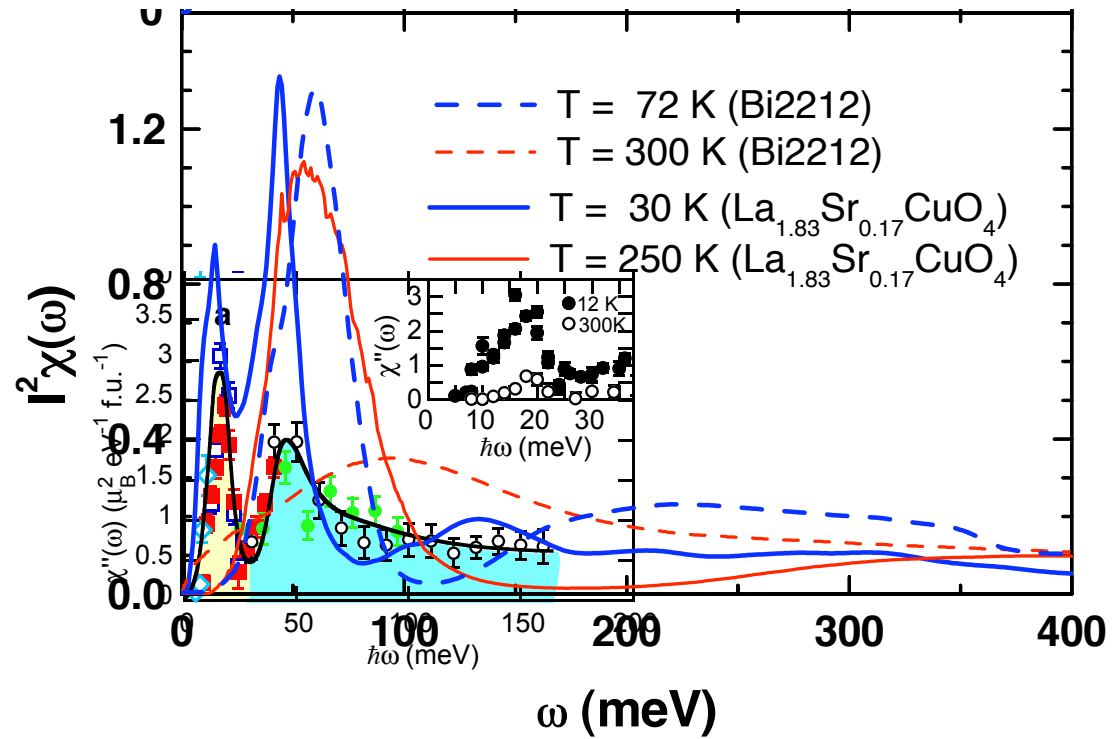
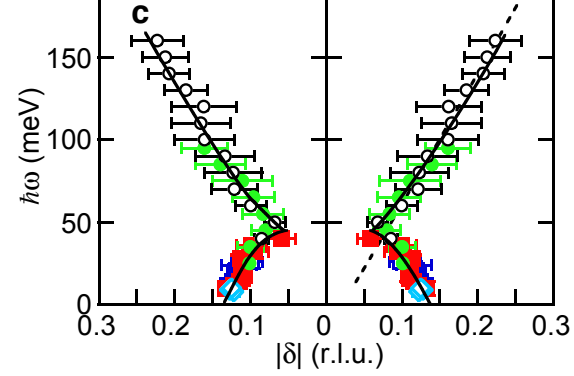
# LSCO neutron scattering

Vignolle *et al.*



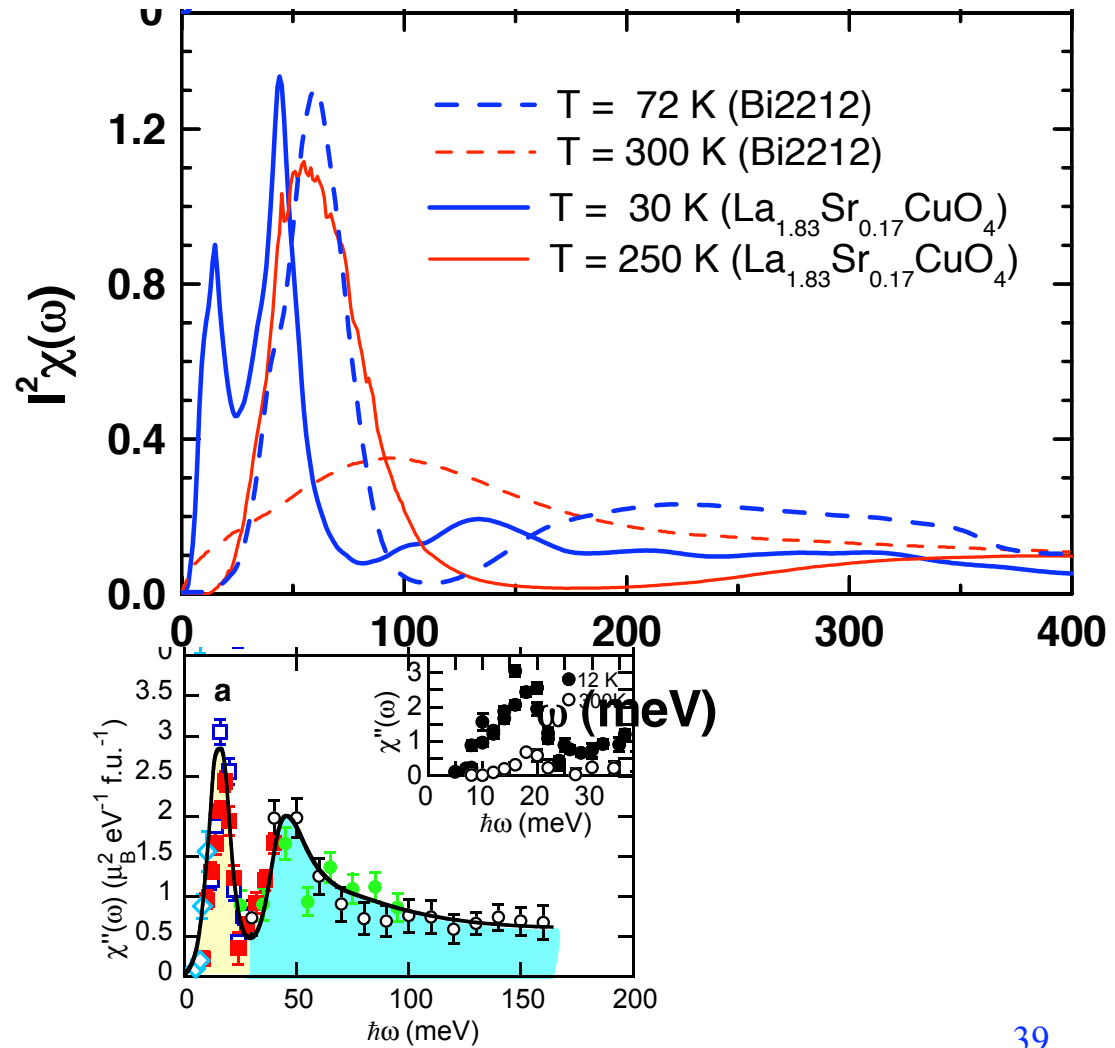
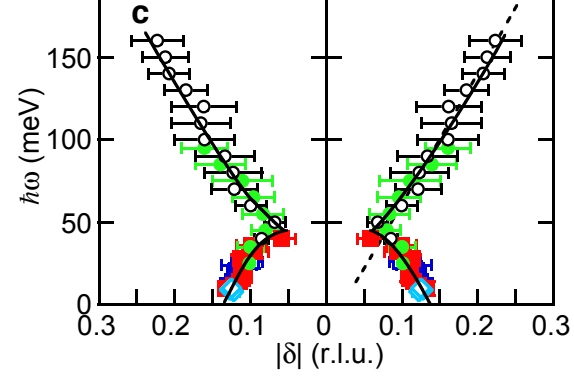
# LSCO neutron scattering

Vignolle *et al.*



# LSCO neutron scattering

Vignolle *et al.*



# Why is LSCO $T_c$ so low?

McMillan equation for  $T_c$ :

$$k_B T_c \cong \hbar \omega_{ln} \exp \left[ - \frac{1 + \lambda^s}{\lambda^d} \right]$$

	LSCO	Bi-2212
$\lambda_d$	1.90	1.85
$\lambda_s$	3.40	2.50
$\omega_{ln}$	25	50 meV
$T_c$	29	89 K

where

$$\omega_{ln} \equiv \exp \left[ \frac{2}{\lambda} \int_0^\infty \ln \omega \frac{I^2 \chi(\omega)}{\omega} d\omega \right]$$

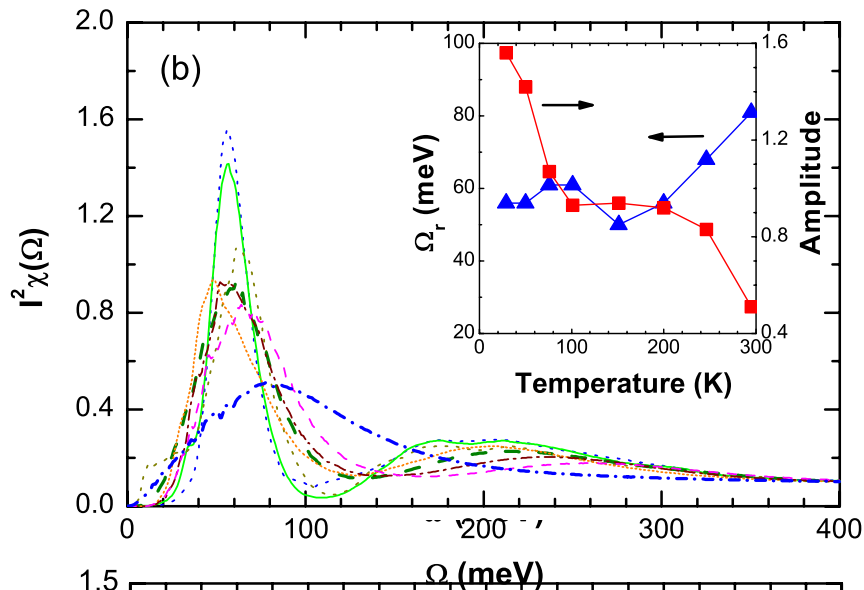
$$\lambda^{(s,d)} = 2 \int_0^{\omega_c} \bar{I}_{(s,d)}^2 \chi(\omega) / \omega d\omega$$



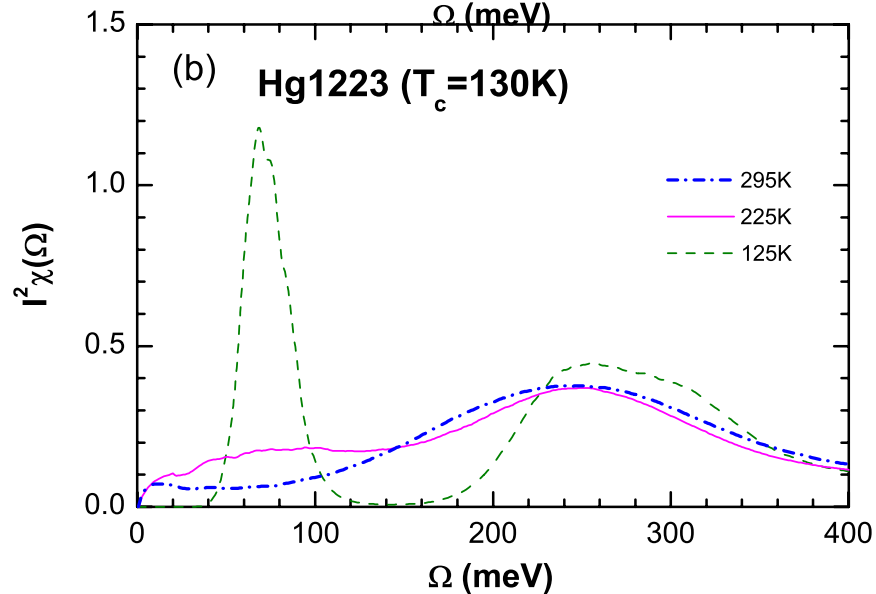
# Mercury based superconductors

# The mercury based superconductors

One layer  $T_c = 90$  K

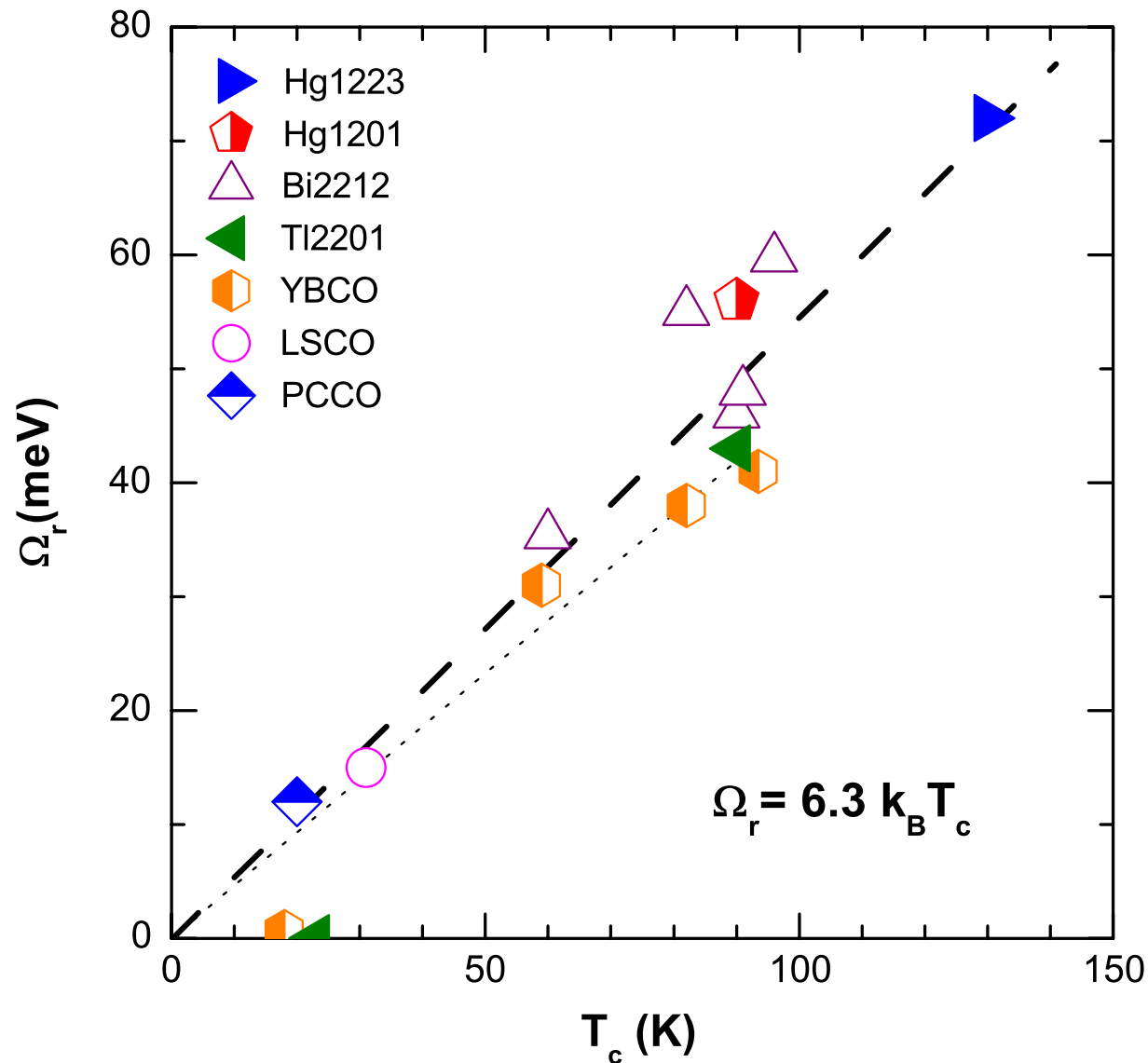


Three layer  $T_c = 130$  K



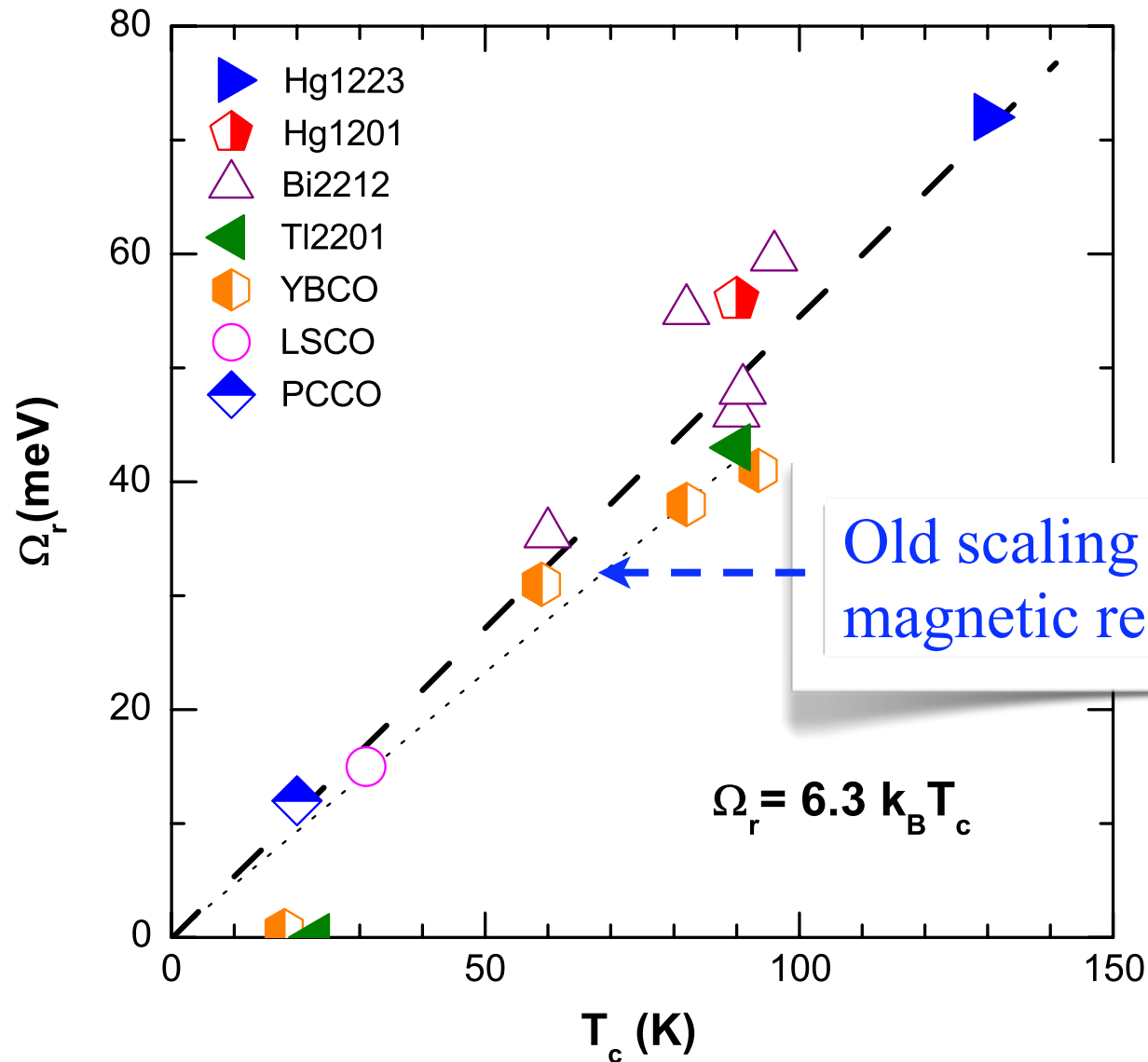
# A "new" scaling law

The peak of the bosonic spectral function scales with  $6.6 k_B T_c$ .



# A "new" scaling law

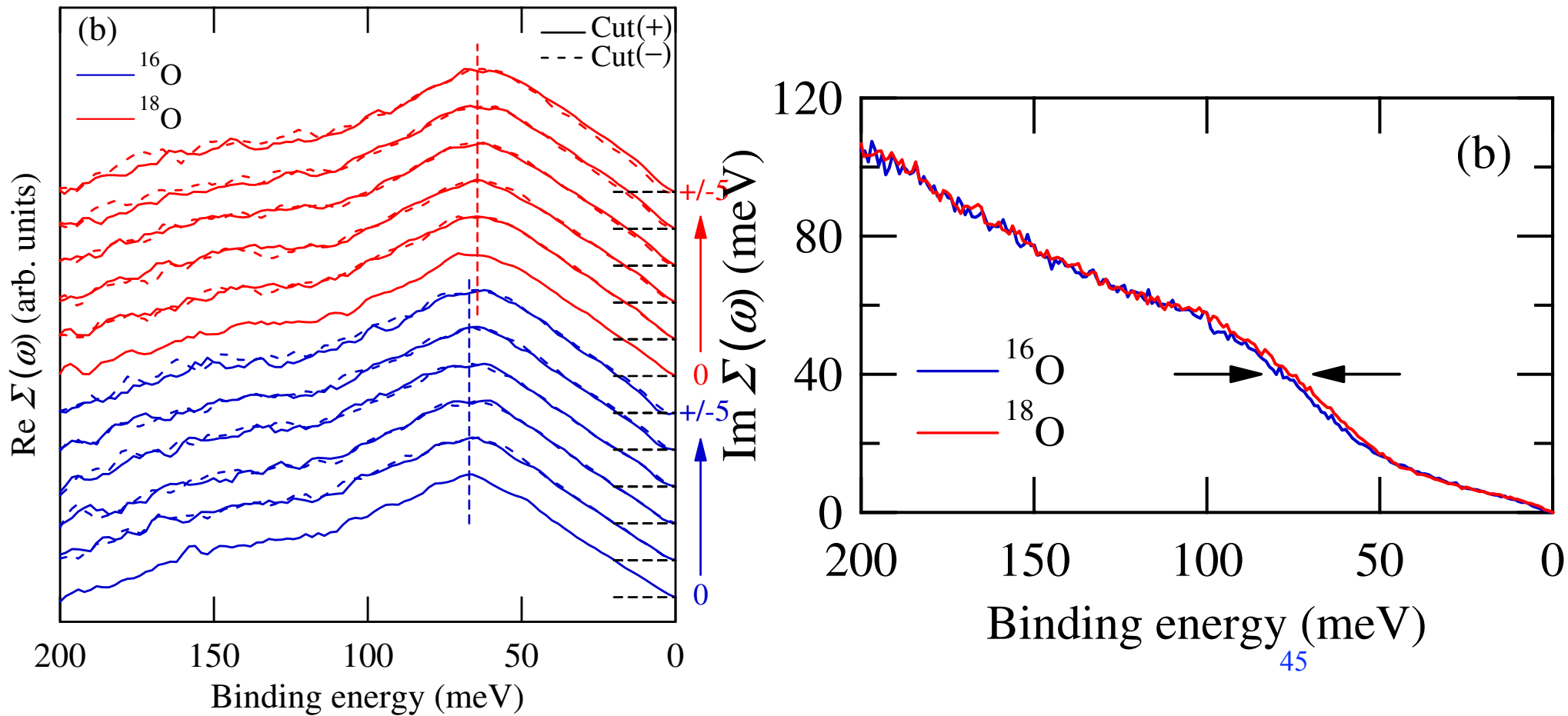
The peak of the bosonic spectral function scales with  $6.6 k_B T_c$ .



# The isotope effect

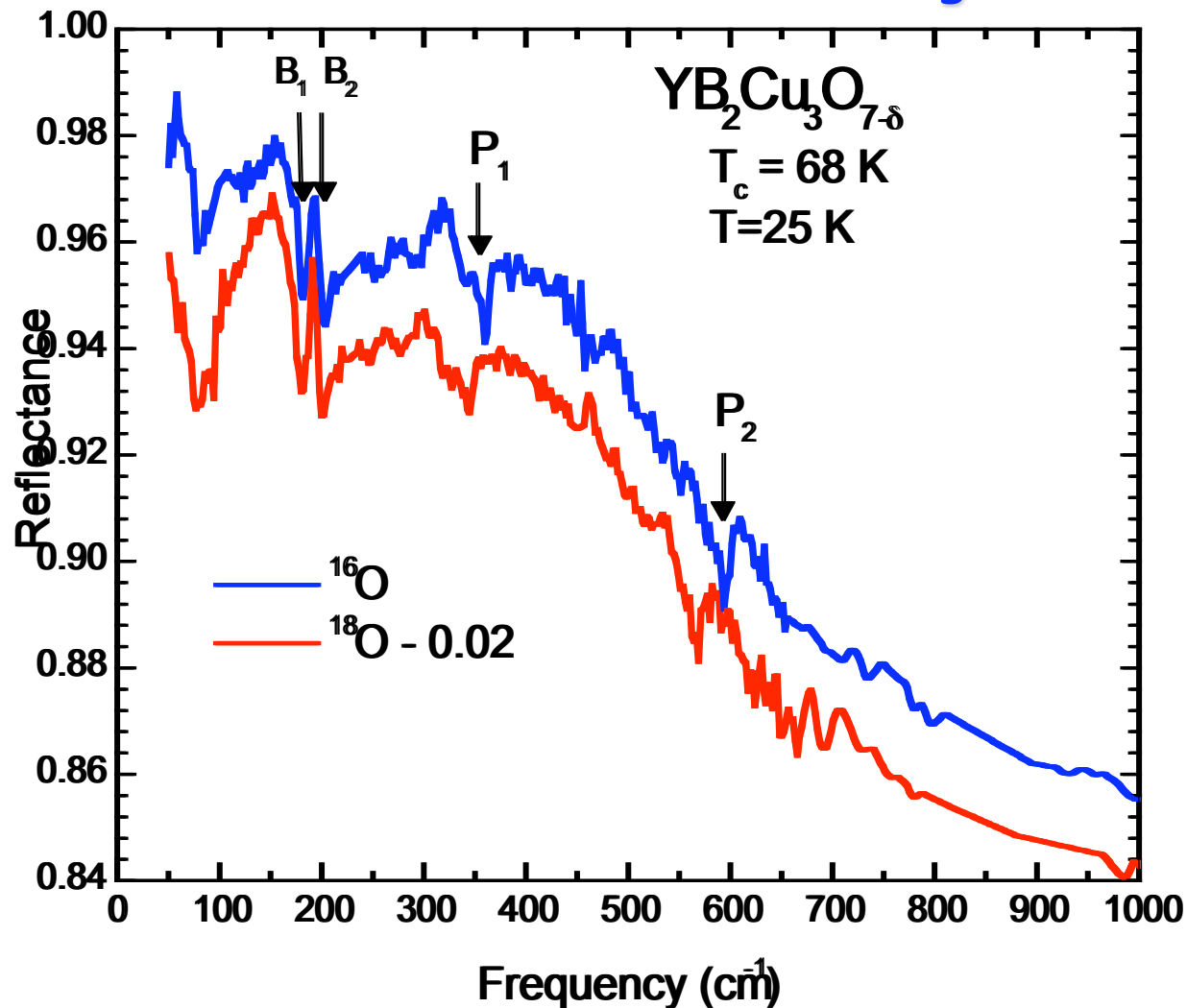
# The isotope effect: new ARPES data

Iwasawa *et al.* PRL 101, 157005 (2008).

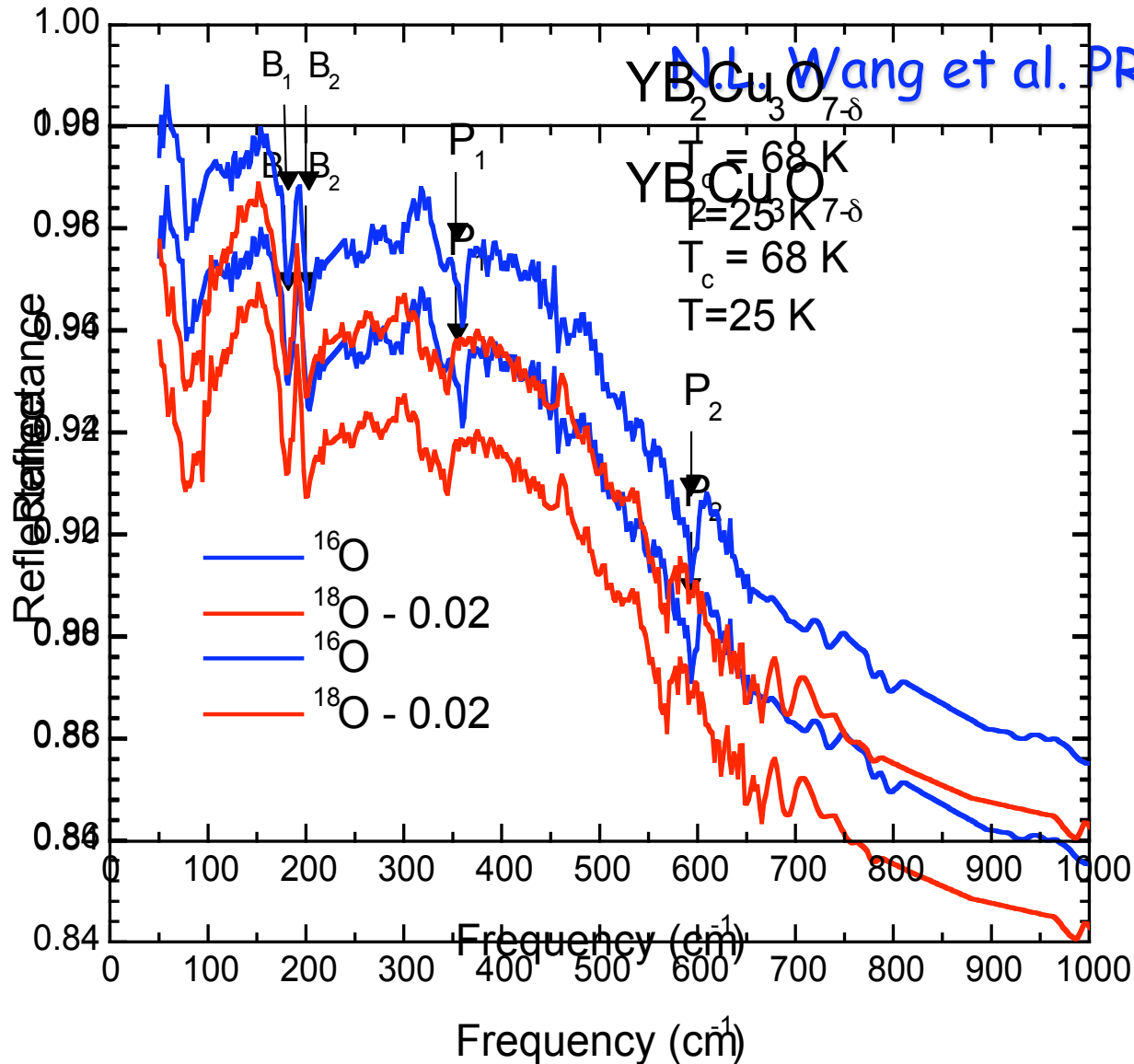


# Isotope effect in YBCO from optics

N.L. Wang et al. PRL 2002



# Isotope effect in YBCO from optics





# Isotope effect in YBCO

N.L. Wang et al. PRL (2002).

- Underdoped samples with  $^{16}\text{O}$  and  $^{18}\text{O}$
- Isotope effect:  $\Delta \omega / \omega = -\alpha \Delta m / m$
- The phonon lines show isotope effect with  $\alpha = -0.5$
- The normal state shoulder shows  $\alpha = -0.1 \pm 0.1$
- The superconducting state shoulder  $\alpha = -0.23 \pm 0.1$

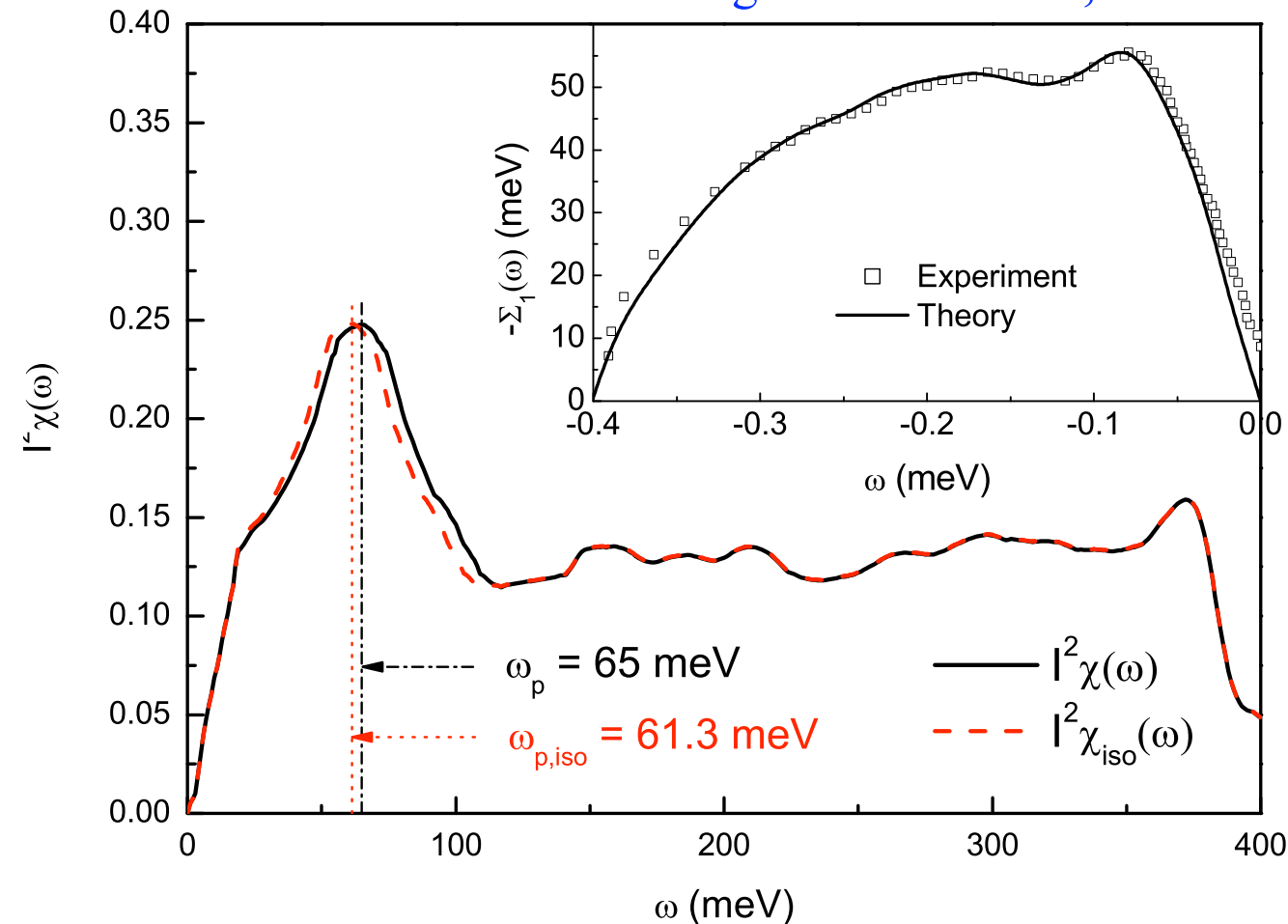
The mode giving rise to the shoulder  
is unlikely to be of phonon origin.

# Contradiction !

- ARPES finds  $\alpha = -0.5$  for nodal kink
- Optics finds  $\alpha = -0.1$  Fermi surface average kink

# Maximum entropy inversion of nodal ARPES

Wentao Zhang *et al.* PRL **101**, 107002 (2008).



Bi-2212 nodal  $\Sigma_1$

—  $O^{16}$  spectral function

- - -  $O^{18}$  spectral function

Model: Shift peak of bosonic function by 6 % and calculate  $\Sigma_1$  change

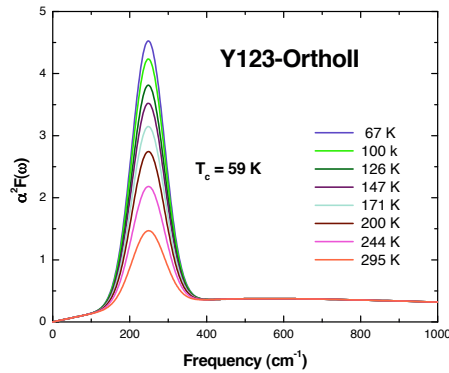
$\lambda_{ep} = 0.2$  Predicted shift 4.5%

$\lambda_{mag} = 0.8$  Observed shift of  $5 \pm 0.8$  %

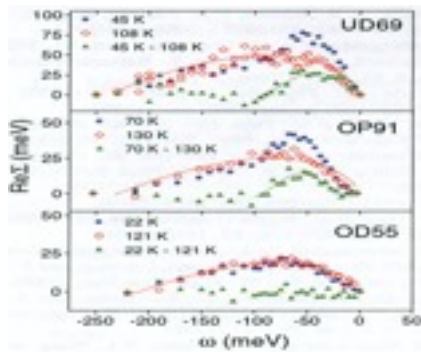
## Controversy resolved

Both optics and ARPES are consistent with a 10 to 20 % self energy contributions from phonons.

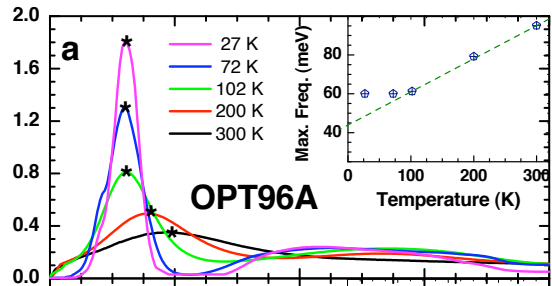
# Summary of self energy spectroscopy



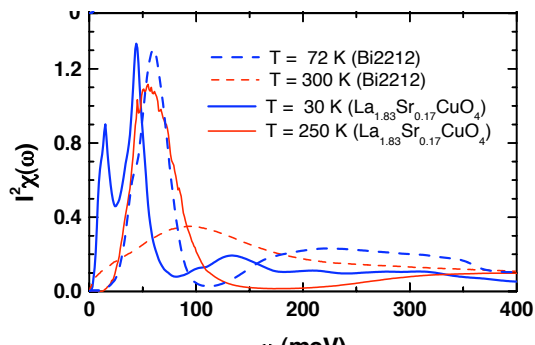
A bosonic peak and a background peak depends on temperature



Bi-2212 cf. ARPES

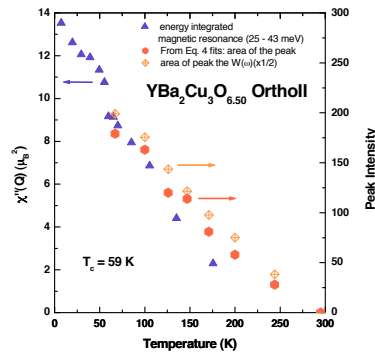
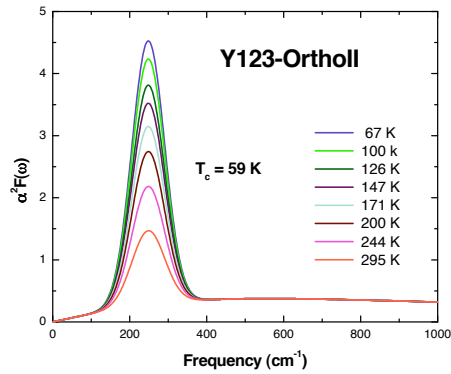


Bi-2212 Shift of spectral weight cf. tunneling

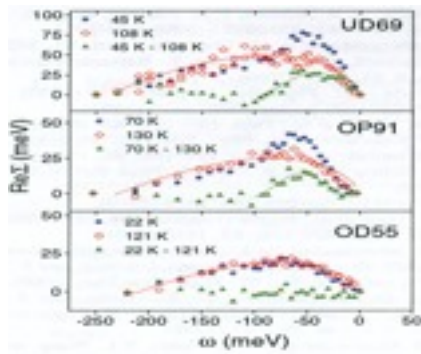


LSCO cf. magnetic neutron scattering

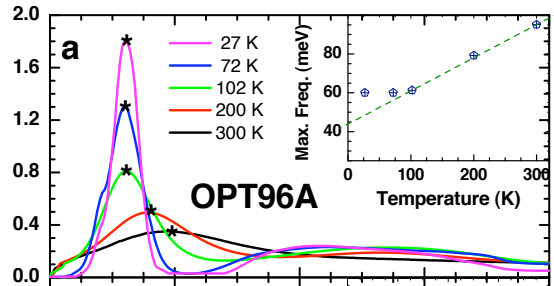
# Summary of self energy spectroscopy



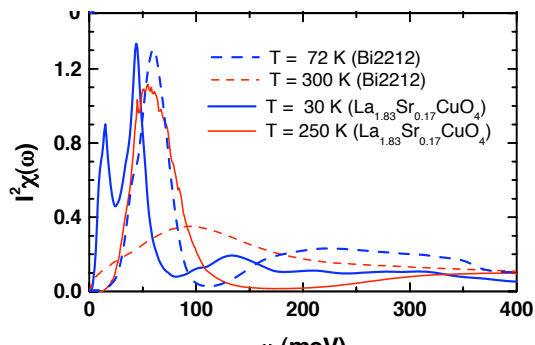
A bosonic peak and a background peak depends on temperature



Bi-2212 cf. ARPES

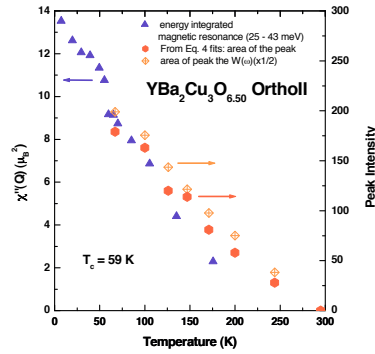
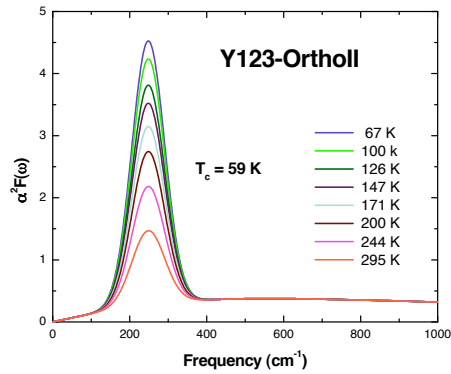


Bi-2212 Shift of spectral weight cf. tunneling

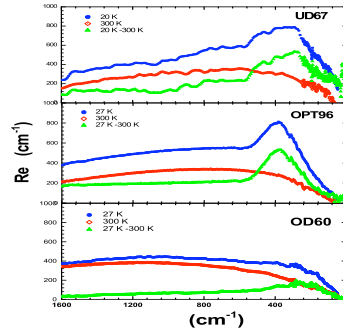
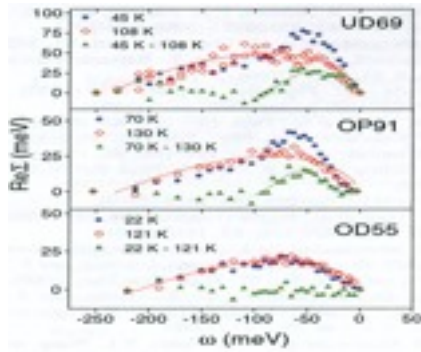


LSCO cf. magnetic neutron scattering

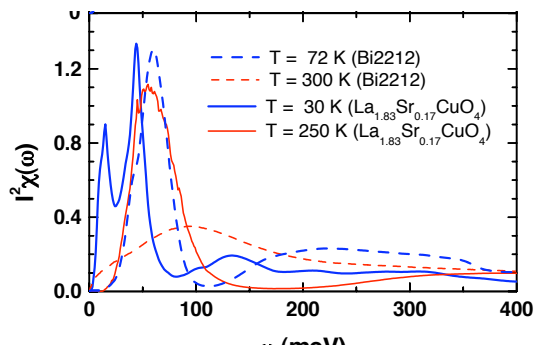
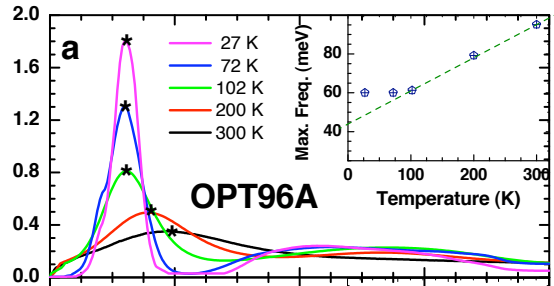
# Summary of self energy spectroscopy



A bosonic peak and a background peak depends on temperature



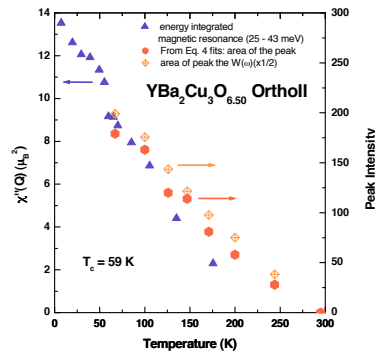
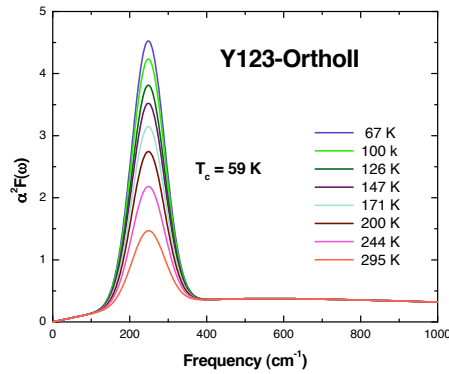
Bi-2212 cf. ARPES



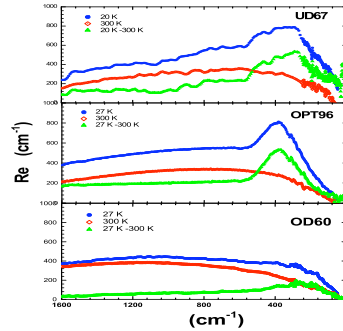
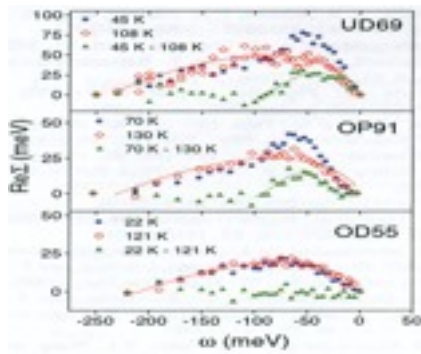
Bi-2212 Shift of spectral weight cf. tunneling

LSCO cf. magnetic neutron scattering

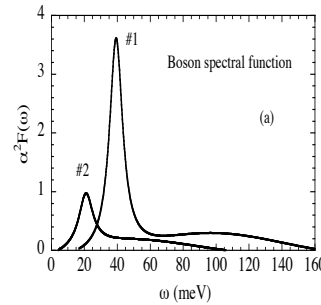
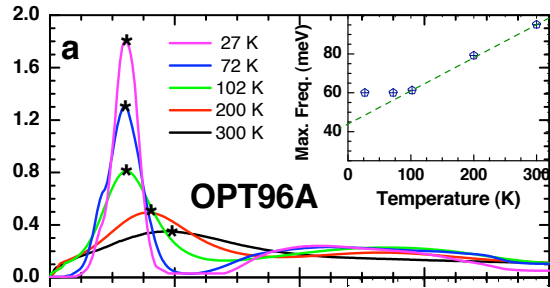
# Summary of self energy spectroscopy



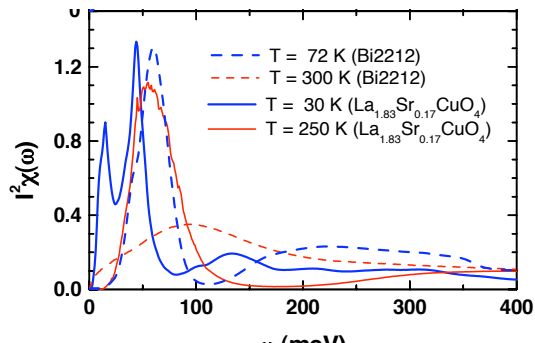
A bosonic peak and a background peak depends on temperature



Bi-2212 cf. ARPES



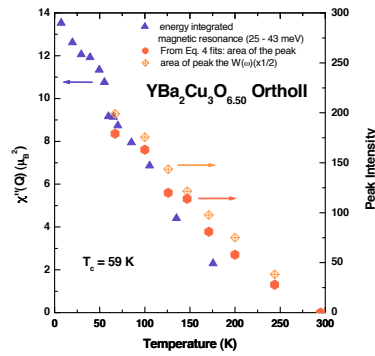
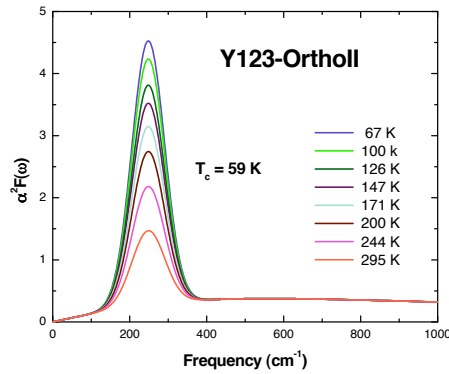
Bi-2212 Shift of spectral weight cf. tunneling



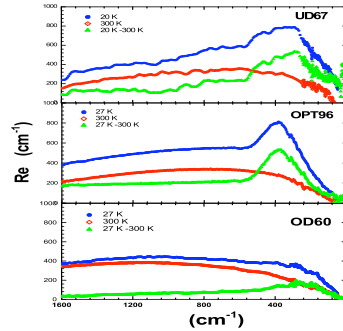
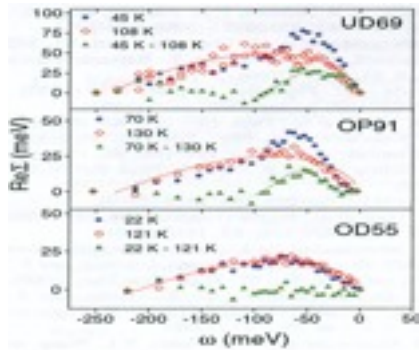
LSCO cf. magnetic neutron scattering



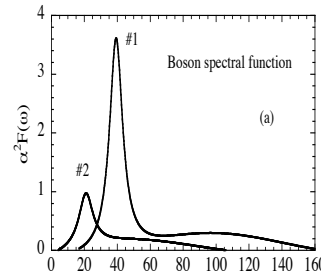
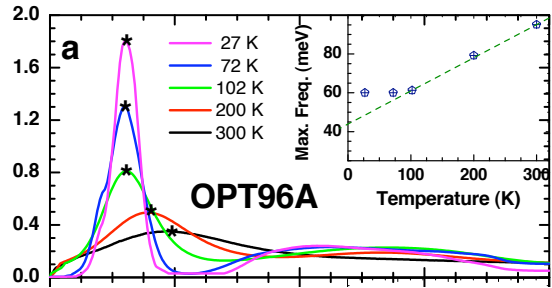
# Summary of self energy spectroscopy



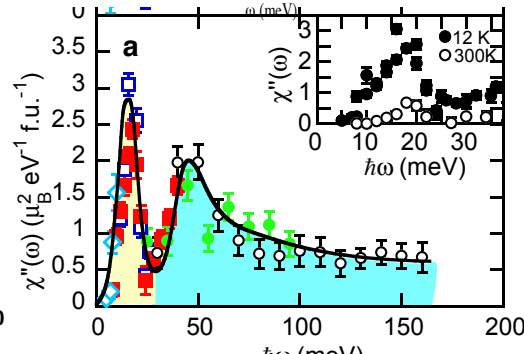
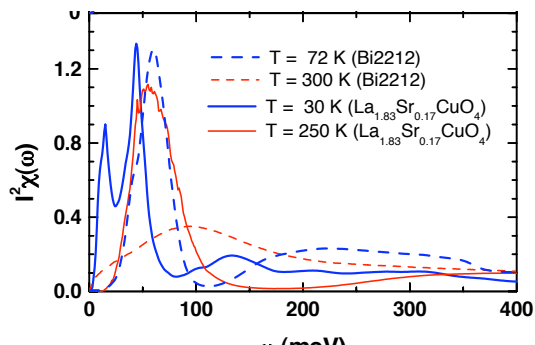
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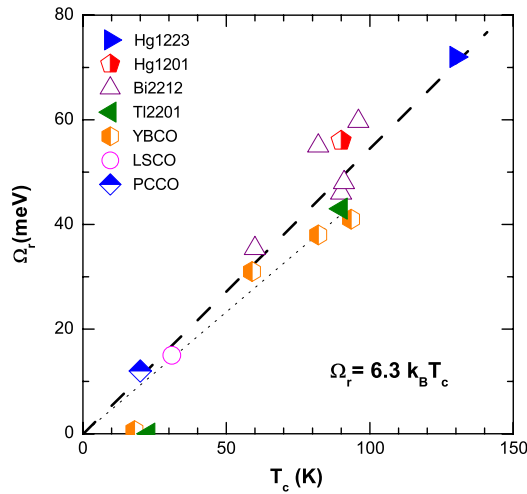
Bi-2212 cf. ARPES



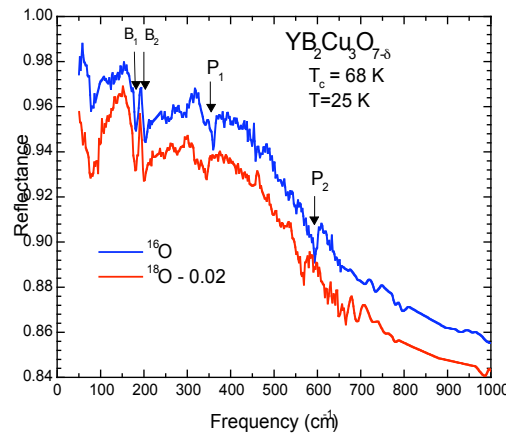
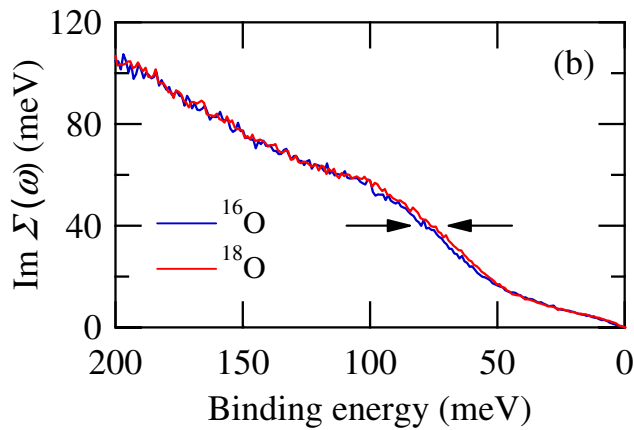
Bi-2212 Shift of spectral weight cf. tunneling



LSCO cf. magnetic neutron scattering



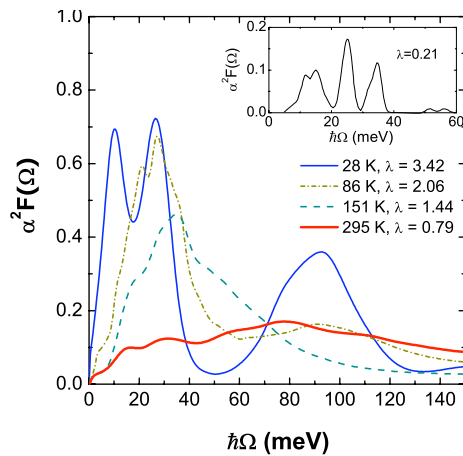
A scaling law  $\Omega_{\text{res}} \propto 6.3 k_B T_c$



Phonons vs. magnetism

$$\lambda_{\text{ep}} = 0.2$$

$$\lambda_{\text{mag}} = 0.8$$



## Remaining tests:

No magnetic mode in:

- Highly underdoped YBCO
- $\text{LSCO}_{0.22}$

## Remaining problems:

- Origin of magnetic susceptibility
- Materials differences
- Pseudogap, friend or foe?

Thank you

