## GRB Jet Propagation Inside A Star

Weiqun Zhang (KIPAC, Stanford)

S.E. Woosley (UCSC) A.I. MacFadyen (IAS) A. Heger (LANL)

### We Know That

- at least some GRBs are made from the death of massive stars (Fact)
- GRB outflows are relativistic jets (Theory)
- the jets need to escape from massive stars

- Progenitor: Collapsar, Magnetar, He-BH, He-He, .....
- How to make jets?
- Poynting Flux?
- Supernova
- Let's not worry about the above issues first. Let's ask a simple question.
   Can jets get out of the stars?

# Yes, they can!

- A common misconcept: All stellar material on its way will be swept by the jet.
- The magic is the bow shock, which pushs stellar material sideways, and the cocoon, which can protect the jet beam.
- Relativity helps too. Instabilities do not have enough time to grow.

Relativistic Jets From Collapsars S.E. Woosley's Group Inital Model: he15 480 radial zones, 200 angular zones Energy Deposition Rate: 10<sup>51</sup> ergs/s Half Opening Angle: 20  $f_{e}(E_{th}/E_{tot}): 0.67$ Lorentz Factor; 50







- Ultrarelativistic core (~200) Mildly relativistic wings (~5) Viewing Angle ==> ???
- It takes ~10 s to break out.
  Star: 3 light-second
  "Wasted" energy: 10^50 erg/s \* 7s~10^51
  Even if the jet is dominated by B, .....
- Lorentz factor Initial injection: ~10 Final: 200 How? Initial jet is hot.

(Model dependent)

#### 3D





3e50 erg/s t = 6.6s

Ε

0.08

0.4

2

### Bigger Cocoon

### Kelvin-Helmholtz

- Relativity makes it hard to develope
- It can grow to the nonlinear regime when the jet is thin.
- Too much is bad, but a little bit is good.



#### Resolution is the key!



High-resolution

Low-resolution

Grow exponentially, then saturate

### **Blue** Supergiant



If the engine can last a few hundred seconds, ..... Instabilities?

### Jet outside



#### AMR: 14 levels of refinement