

**Surface roughness & optical properties  
influence on the Casimir/vdW force**

**George Palasantzas**



# Acknowledgements



RUG: *P. Van Zwol (my student!)*

Collaborator → Vitali Svetovoy (MESA+, NL)



Steve Lamoreaux (Yale, USA)

Astrid Lambrect & Serge Reynaud (France)

R. Onofrio (USA)

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D. Iannuzzi (NL)

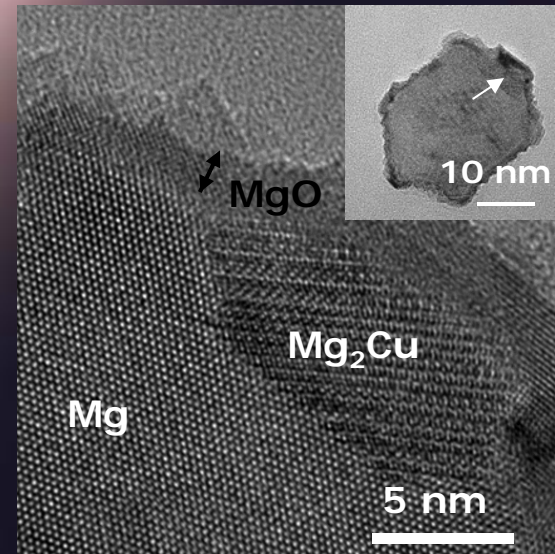
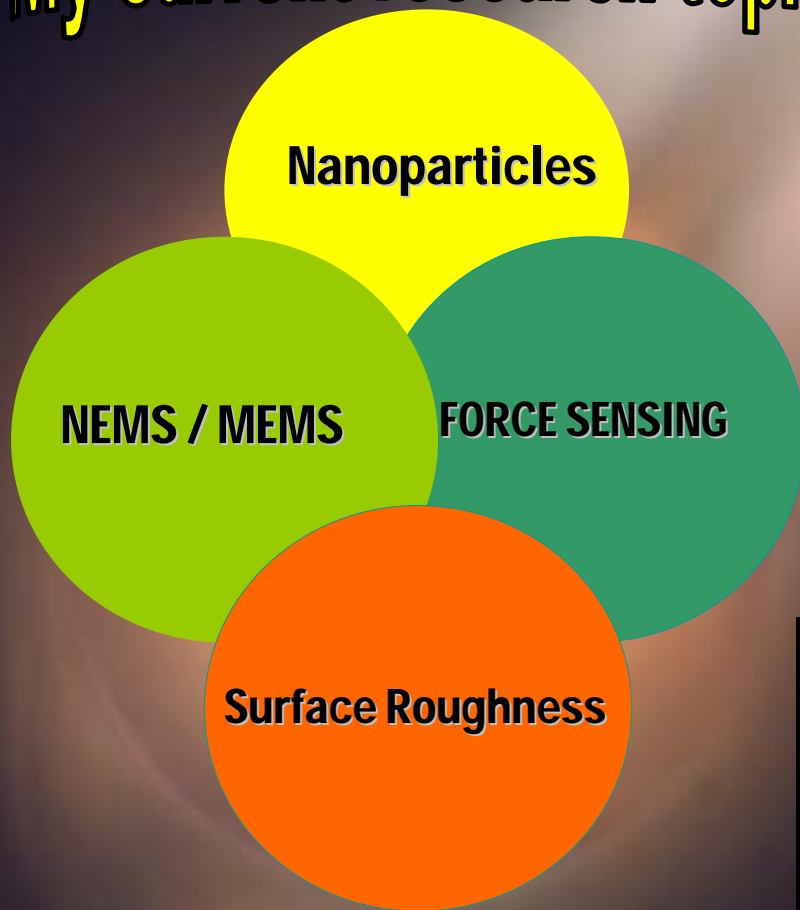
J. Munday, F. Capasso (USA)

Funding agency →

**M2i / TNO**



# My current research topics



## PhD projects

- *Casimir effect & NEMS*
- *Hydrogen storage-nanoparticles*
- *Piezoresponse Force Microscopy*
- *AFM optical data storage*
- *NEMS / MEMS (look for student!)*

<http://palasantzas.fmns.rug.nl>

*The main lines of this research involve:*

- **Casimir force in complex geometries and novel topologies such as pattered or corrugated surfaces, nanospheres or small spheroidal shaped bodies** (Barton, Binns, Bordag, Brevik-Hoye, Chevrier, Duplantier, Lambrecht-Reynaud)
- **"Vacuum torques" acting on irregular bodies** (Iannuzzi, Lambrecht- Reynaud, Svetovoy)
- **Studying the Casimir force using new materials such as superconductors, magnetic materials, meta-materials or Carbon nanotubes** (Binns, Bordag, Bimonte-Calloni, Lambrecht, Henkel)
- **Dispersion forces and NEMS Devices** (Andreucci, Iannuzzi, Palasantzas, Svetovoy)
- **Nanoscale surface/interface roughness influence on functional properties of MEMS, NEMS** (Andreucci, Emig, Lambrecht, Palasantzas)
- **Development of new instrumentation for surface force measurements in liquids, also at cryogenic temperatures** (Iannuzzi)
- **New forces**





## Recent Measurements of Casimir Force

Investigators	Year	Geometry	Method	Distance Scale (nm)	Materials	Pressure (mbar)	Temp (K)	Accuracy (%)
S. K. Lamoreaux	1997		Torsion pendulum	600 - 6000	Au(500nm)	$10^{-4}$	300	5
U. Mohideen & A. Roy	1998		AFM	100 - 900	Al (300nm) + AuPd (20nm)	$5 \times 10^{-2}$	300	
A. Roy and U. Mohideen	1999		AFM	100 - 900	Al (250nm)+ AuPd (8nm)	$5 \times 10^{-2}$	300	
G. L. Klimchitskaya, A. Roy, U. Mohideen and V. M. Mostepanenko	1999		AFM	100 - 900	Al (300nm) + AuPd (20nm)	$5 \times 10^{-2}$	300	
T. Ederth	2000		Piezo-tube manipulator	20 - 100	50 $\mu$ m Au wires coated in thiol SAM	1000	300	
H. B. Chan, V. A. Aksyuk, R. N. Kleiman, D. J. Bishop & F. Capasso	2001		MEMS torsion bar capacitance	90 - 1000	Au (200nm) + Cr underlayer	1000	300	1
G. Bressi, G. Carugno, R. Onofrio & G. Ruoso	2002		Interferometry	500 - 3000	Cr (50nm) on Si	$10^{-3}$	300	15
R. S. Decca, D. Lopez, E. Fischbach & D. E. Krause	2003		MEMS torsion bar capacitance	200 - 2000	Cu/Au	$10^{-4}$	300	
NANOCASE	2005-		AFM, MEMS	10 - 1000	Si, Au	$10^{-11}$	20 - 1000	

Today's topics

- *Real materials*
- *Roughness, Optical properties*
- *Critics .....Cosmology.....*
- *Conclusions.....*

# Real Materials → Corrections to the Casimir / vdW force

## Finite conductivity

$$\omega < \omega_p \quad (\lambda_p \sim 100 \text{ nm})$$

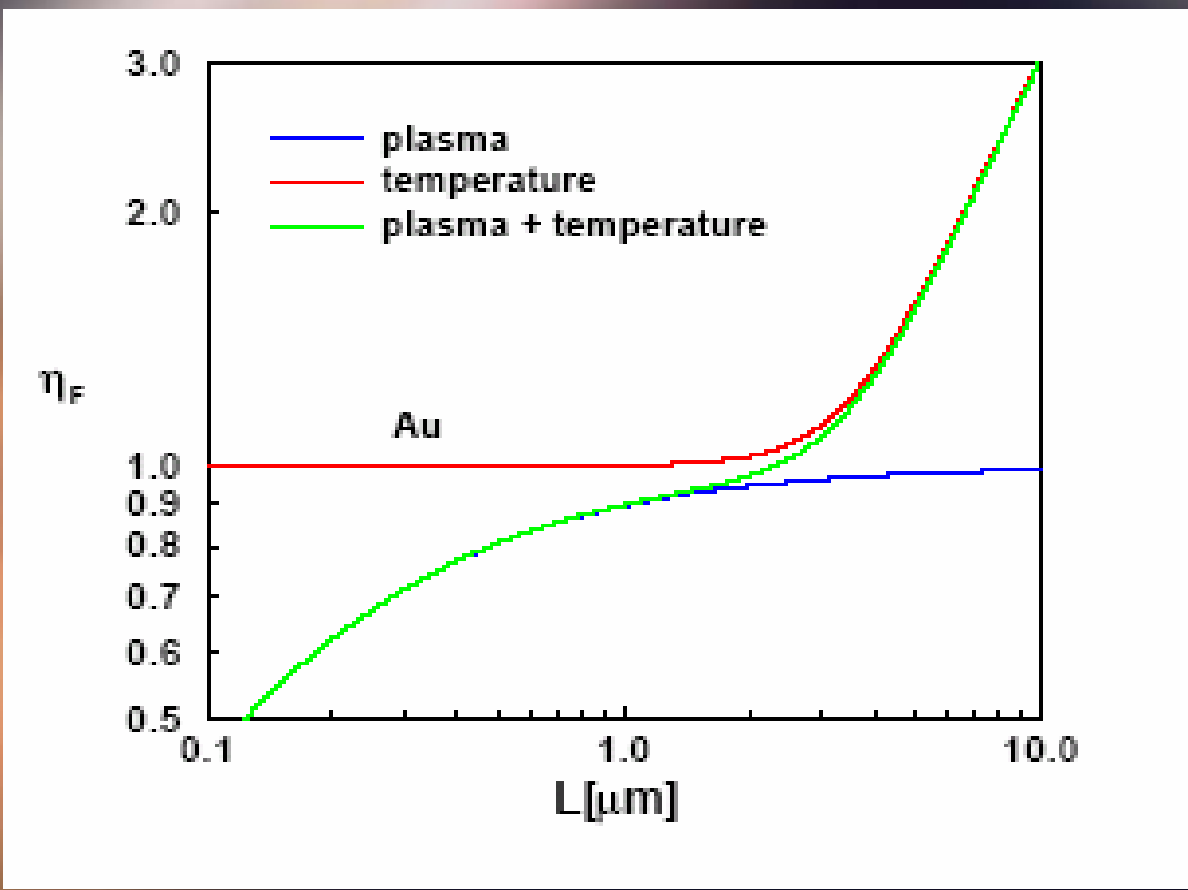
$$\lambda_p = 2\pi c / \omega_p$$

## Temperature correction

$$\lambda_T = \frac{\hbar c}{k_B T} \cong 7 \mu\text{m} \quad (T = 300 \text{ K})$$

## Roughness correction

$$\eta_F = \frac{F_{Cas}^{Real}}{F_{Cas}^{Perf}}$$



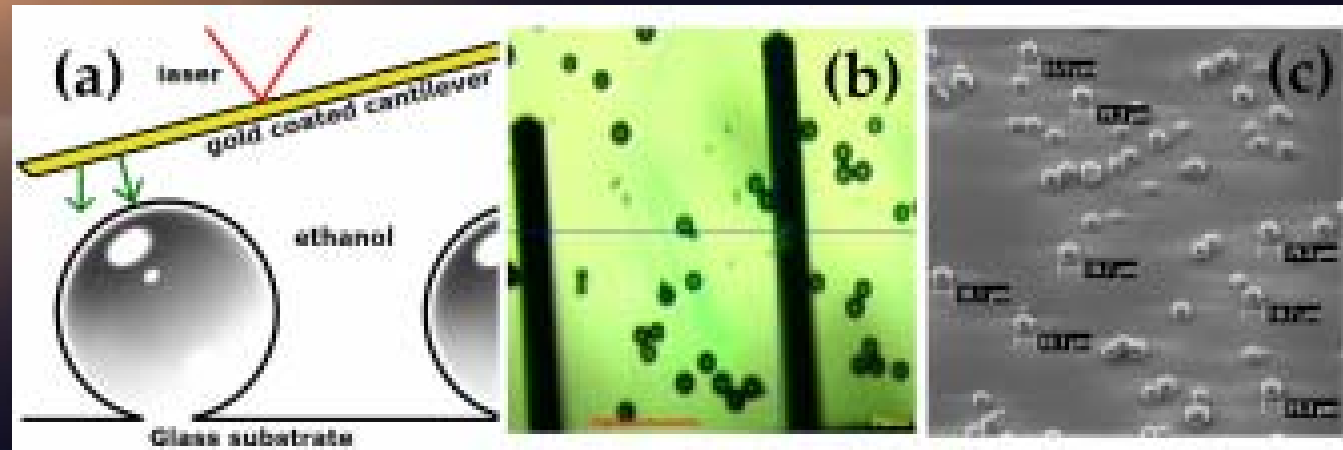
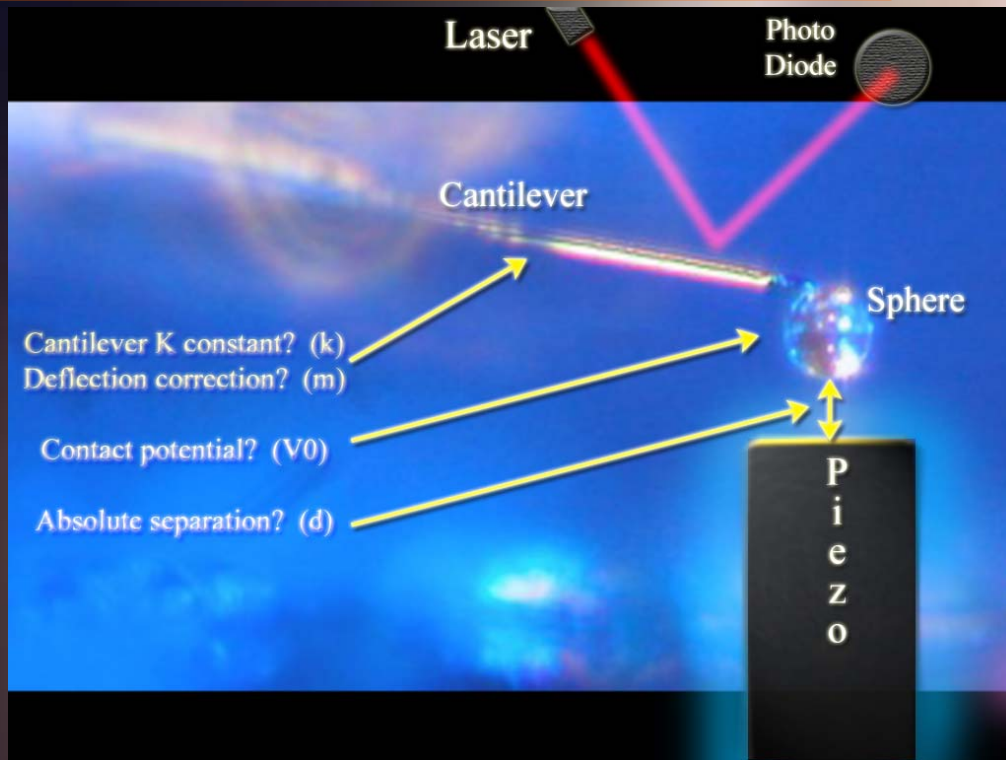
C. Genet, A. Lambrecht, and S. Reynaud, PRA (2000)

Temperature correction (accuracy > 1 %)

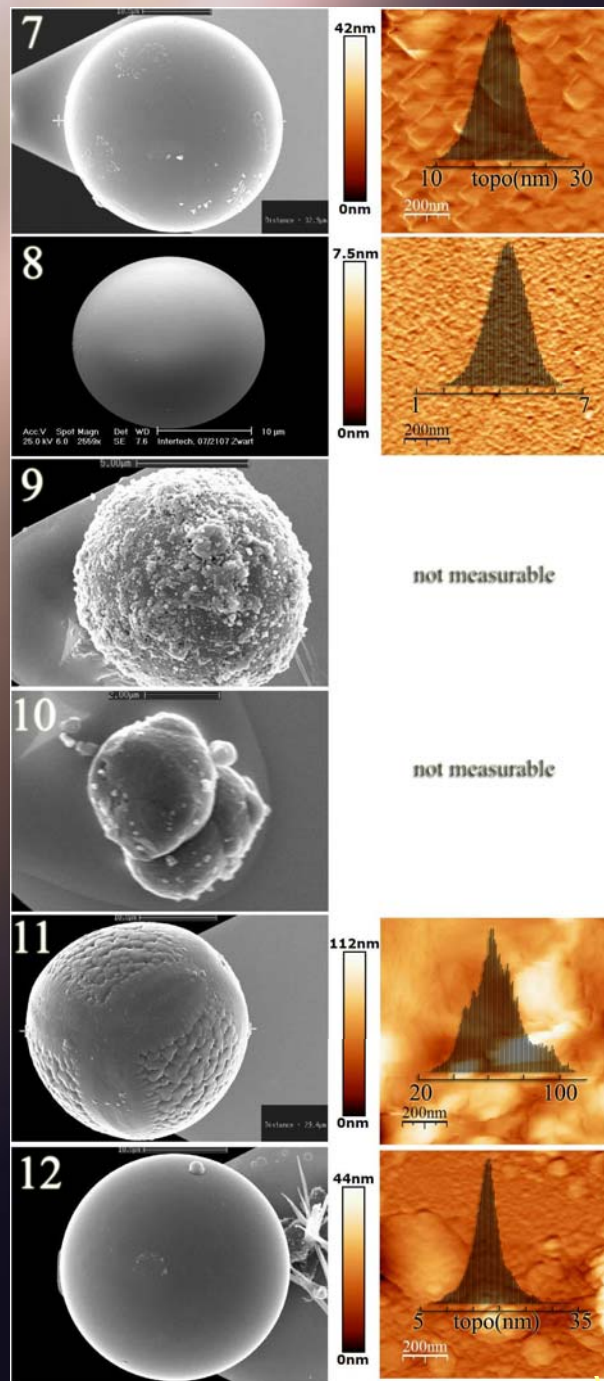
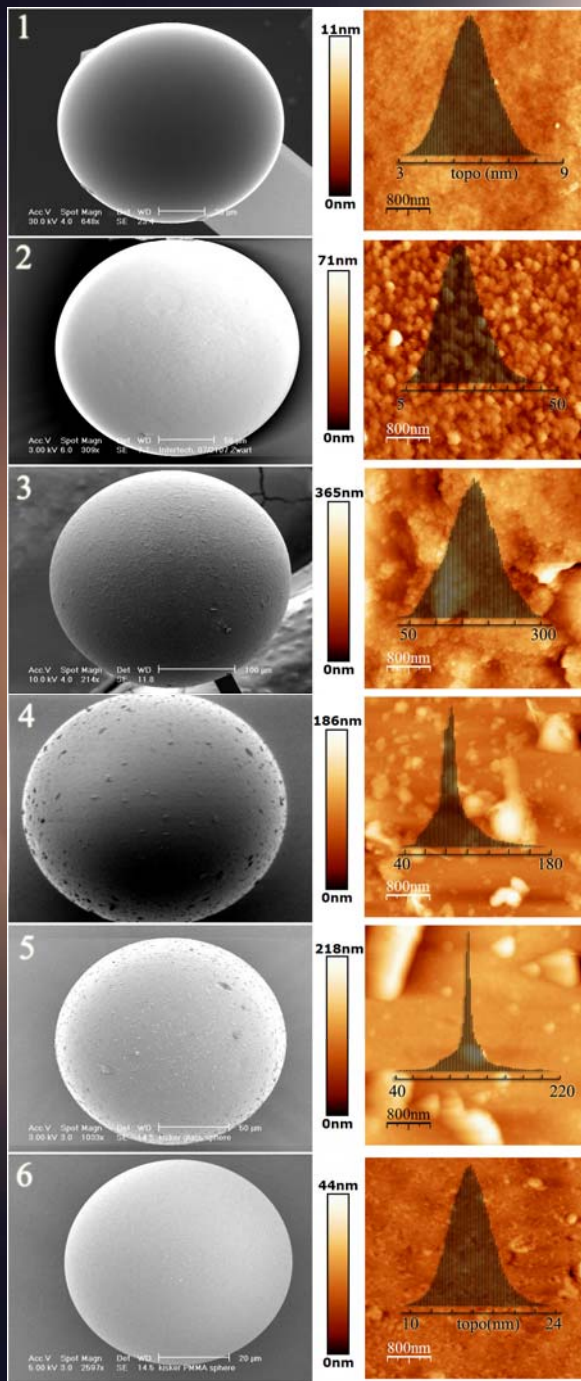
$$F^T(T, d) = \begin{cases} 1 + \frac{720}{\pi^2} \left\{ \left( \frac{K_B T d}{\hbar c} \right)^3 \frac{\zeta(3)}{2\pi} - \frac{45}{\pi^2} \left( \frac{K_B T d}{\hbar c} \right)^4 \right\} & \text{if } K_B T d / \hbar c < 1/2 \\ \left( \frac{K_B T d}{\hbar c} \right) \frac{\zeta(3)}{8\pi} - \frac{\pi^2}{720} & \text{if } K_B T d / \hbar c > 1/2 \end{cases}$$



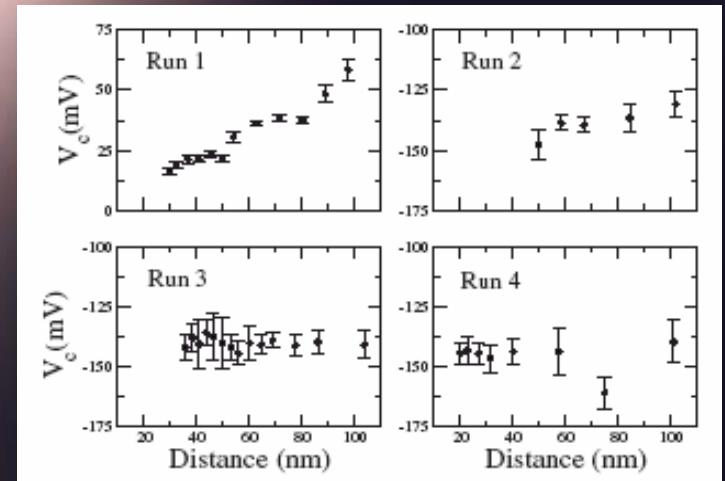
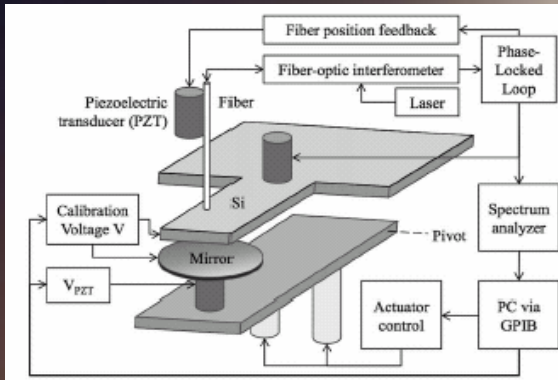
# Force measurement by AFM



After R. Onofrio, PRA 2008

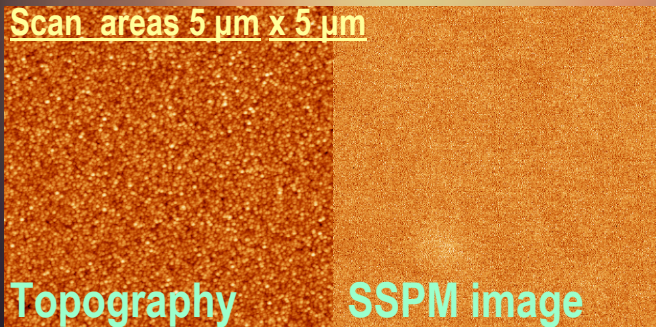






~~$$F'_{el} = -\pi\epsilon_0 R V^2 / d^2$$~~

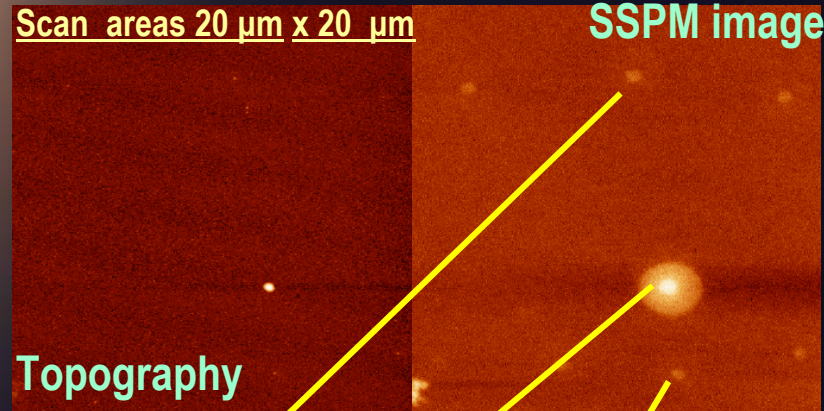
Scan areas 5  $\mu\text{m} \times 5 \mu\text{m}$



Topography

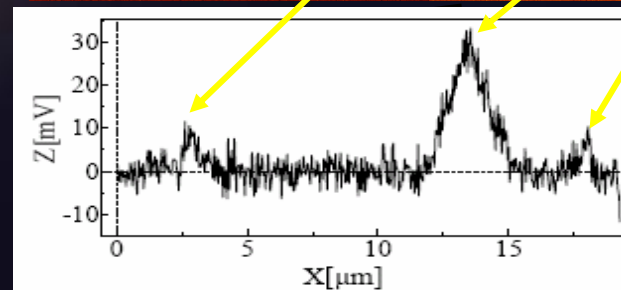
SSPM image

Scan areas 20  $\mu\text{m} \times 20 \mu\text{m}$



Topography

SSPM image



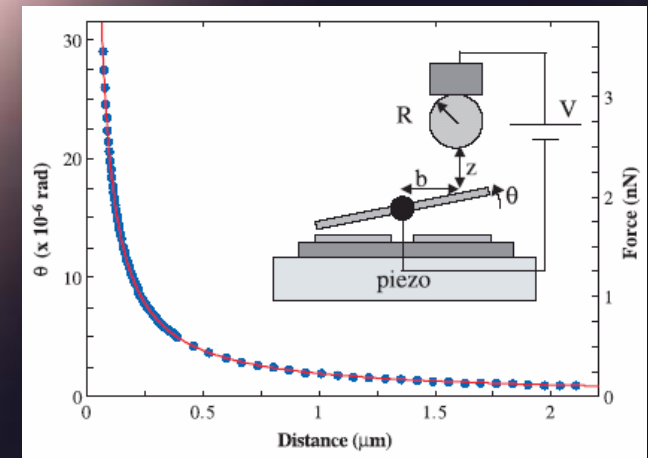
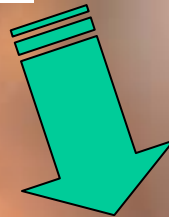
Iannuzzi et al, arxiv 2008

$$F'_{el} = -\pi\epsilon_0 R V^2 / d^2$$

# Quantum Mechanical Actuation of Microelectromechanical Systems by the Casimir Force

H. B. Chan, V. A. Aksyuk, R. N. Kleiman, D. J. Bishop, Federico Capasso\*

SCIENCE VOL 291 9 MARCH 2001



values to better than 1%. As pointed out by Lamoreaux (26, 27), a fundamental test of the theory would require a direct measurement of the optical properties of the metal under study.

26. S. L. Lamoreaux, *Phys. Rev. A* **59**, 3149 (1999).

27. \_\_\_\_\_, *Phys. Rev. Lett.* **83**, 3340 (1999).

# Sample dependence of the Casimir force

I Pirozhenko<sup>1</sup>, A Lambrecht<sup>1</sup> and V B Svetovoy<sup>2,3</sup> *New Journal of Physics* 8 (2006) 238

**Table 1.** The Drude parameters found by fitting the available infrared data for  $\epsilon'(\omega)$  and  $\epsilon''(\omega)$  with equation (10). The error is statistical.

N	$\omega_p$ (eV)	$\omega_r \cdot 10^2$ (eV)	$\mathcal{P}$	
1	$7.50 \pm 0.02$	$6.1 \pm 0.07$	$-27.67 \pm 5.79$	Palik, 66 points, ·
2	$8.41 \pm 0.002$	$2.0 \pm 0.005$	$7.15 \pm 0.035$	Weaver, 20 points, ■, □
3	$8.84 \pm 0.03$	$4.2 \pm 0.06$	$12.94 \pm 16.81$	Motulevich, 11 points, ●, ○
4	$6.85 \pm 0.02$	$3.6 \pm 0.05$	$-12.33 \pm 9.13$	Padalka 11 points, ▼, ▽

$$\eta_F = \frac{120L^4}{c\pi^4} \int_0^\infty dk k^2 \int_0^k d\xi \sum_\mu \frac{r_\mu^2}{e^{2k\xi} - r_\mu^2}$$

$\omega_p, \omega_r$ (eV) \ $L(\mu\text{m})$	0.1	0.3	0.5	1.0	3.0
1. $\omega_p = 7.50, \omega_r = 0.061$	0.43	0.66	0.75	0.85	0.93
2. $\omega_p = 8.41, \omega_r = 0.02$	0.45	0.69	0.79	0.88	0.95
3. $\omega_p = 8.84, \omega_r = 0.0422$	0.46	0.69	0.78	0.87	0.94
4. $\omega_p = 6.85, \omega_r = 0.0357$	0.42	0.65	0.75	0.84	0.93
5. $\omega_p = 9.00, \omega_r = 0.035$	0.47	0.71	0.79	0.88	0.95
6. $\omega_p = 7.50 \pm 15\%$ $\omega_r = 0.061$	0.45	0.68	0.77	0.86	0.94
7. $\omega_p = 7.50$ $\omega_r = 0.061 \pm 30\%$	0.42	0.65	0.74	0.84	0.92
	0.44	0.67	0.76	0.86	0.93

**Variation of optical data & associated Drude parameters → variation in the Casimir force : 5.5% at 100 nm → 1.5% at 3μm**



$$E_{ppflat} = -\hbar A \sum_P \int [d^2k / 4\pi^2] \int_0^\infty [d\Phi / 2\pi] \ln[1 - r^P(k, \Phi)^2 e^{-2\kappa d}]$$

$$r_s = -\frac{\sqrt{k^2 + \varepsilon(i\zeta) \frac{\zeta^2}{c^2} - c\kappa}}{\sqrt{k^2 + \varepsilon(i\zeta) \frac{\zeta^2}{c^2} + c\kappa}}$$

$$r_p = \frac{\sqrt{k^2 + \varepsilon(i\zeta) \frac{\zeta^2}{c^2} - c\kappa\varepsilon(i\zeta)}}{\sqrt{k^2 + \varepsilon(i\zeta) \frac{\zeta^2}{c^2} + c\kappa\varepsilon(i\zeta)}}$$

$$\varepsilon(i\zeta) = 1 + \frac{2}{\pi} \int_0^\infty d\omega \frac{\omega E''(\omega)}{\omega^2 + \zeta^2}$$

**Ellipsometry measurement of the dielectric function**

**Ellipsometry:**  $\rightarrow$  ratio of *p*-polarized and *s*-polarized complex Fresnel reflection coefficients is obtained

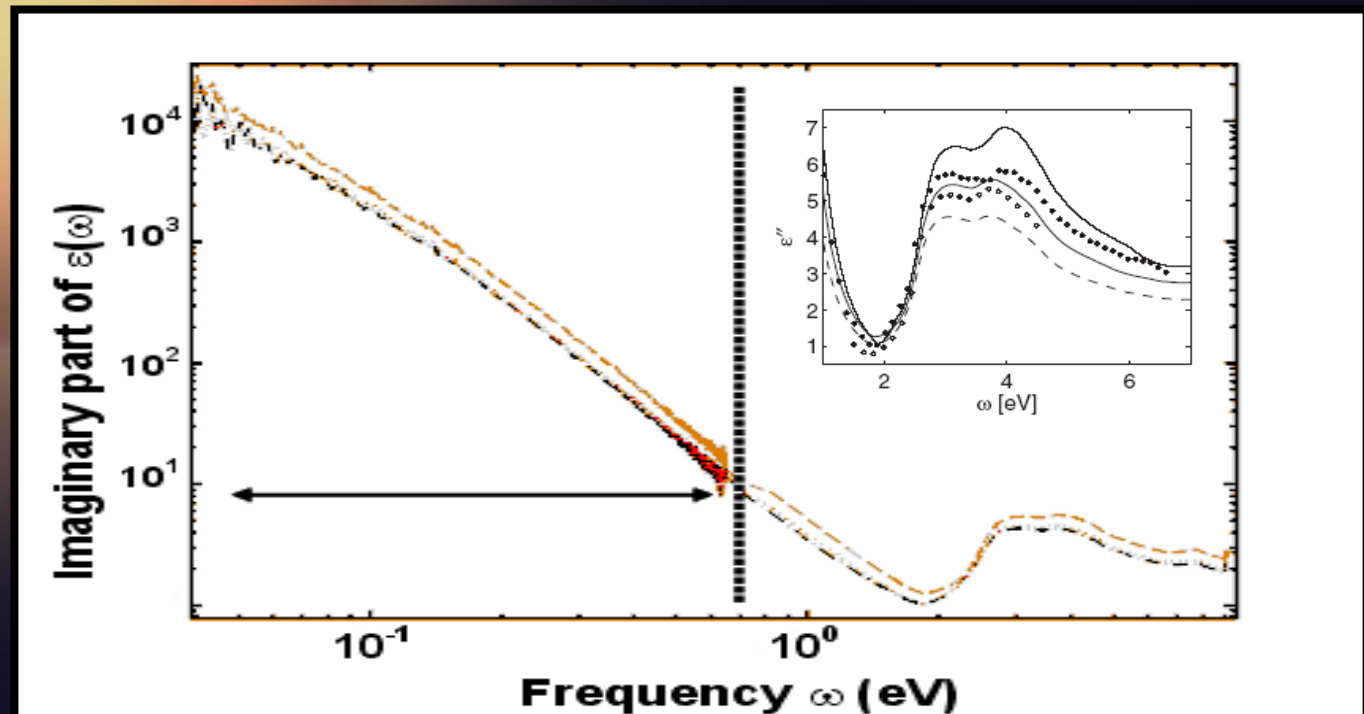
$$\rho = \frac{r_p}{r_s} = \tan \Psi e^{i\Delta}$$

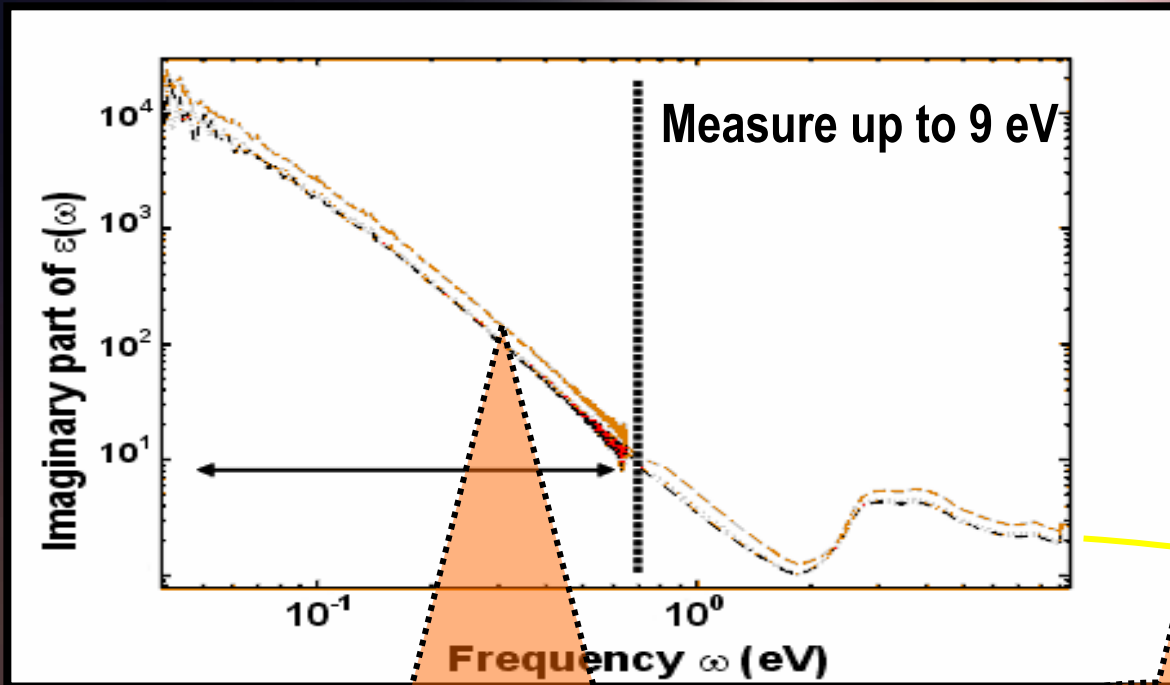
$$r_p = \frac{\langle \epsilon \rangle \cos \vartheta - \sqrt{\langle \epsilon \rangle - \sin^2 \vartheta}}{\langle \epsilon \rangle \cos \vartheta + \sqrt{\langle \epsilon \rangle - \sin^2 \vartheta}}, \quad r_s = \frac{\cos \vartheta - \sqrt{\langle \epsilon \rangle - \sin^2 \vartheta}}{\cos \vartheta + \sqrt{\langle \epsilon \rangle - \sin^2 \vartheta}}$$

$\theta$ : is the angle of incidence

$137 \text{ nm} < \lambda < 32 \mu\text{m}$

$\langle \epsilon \rangle = \langle \epsilon(\lambda) \rangle$  "pseudodielectric" function





above 100 eV  
 $\epsilon(\omega) = 1 - \omega_p^2 / \omega^2 + iA / \omega^3$

$$\epsilon'(\omega) = 1 - \frac{\omega_p^2}{\omega^2 + \omega_\tau^2}, \quad \epsilon''(\omega) = \frac{\omega_p^2 \omega_\tau}{\omega(\omega^2 + \omega_\tau^2)}$$

$9 < \omega < 100$  eV  
 handbook data

Kramers-Kronig (KK) consistency.

$0.01 < \omega < 100$  eV

$$\epsilon'(\omega) - 1 = \frac{2}{\pi} P \int_0^\infty dx \frac{x \epsilon''(x)}{x^2 - \omega^2}$$

$$\varepsilon(i\zeta) = 1 + \varepsilon_{cut}(i\zeta) + \varepsilon_{expt}(i\zeta)$$

$$\omega_{cut} = 0.038 \text{ eV}$$

$$\varepsilon_{cut}(i\zeta) = \frac{2}{\pi} \int_0^{\omega_{cut}} d\omega \frac{\omega \varepsilon''(\omega)}{\zeta^2 + \omega^2}, \quad \varepsilon_{expt}(i\zeta) = \frac{2}{\pi} \int_{\omega_{cut}}^{\infty} d\omega \frac{\omega \varepsilon''(\omega)}{\zeta^2 + \omega^2}$$

• **Our films** → **contribution from the extrapolated region,  $\omega_{cut}$  dominates  $\zeta < 0.2$  eV**

• **handbook data** it dominates up to  $\zeta = 4$  eV

• **Result of reduced  $\omega_{cut}$  (hanbook: 0.125 eV) → calculations based on our data are more reliable**  
 ⇔ *smaller part of  $\varepsilon(i\zeta)$  depends on the extrapolation.*

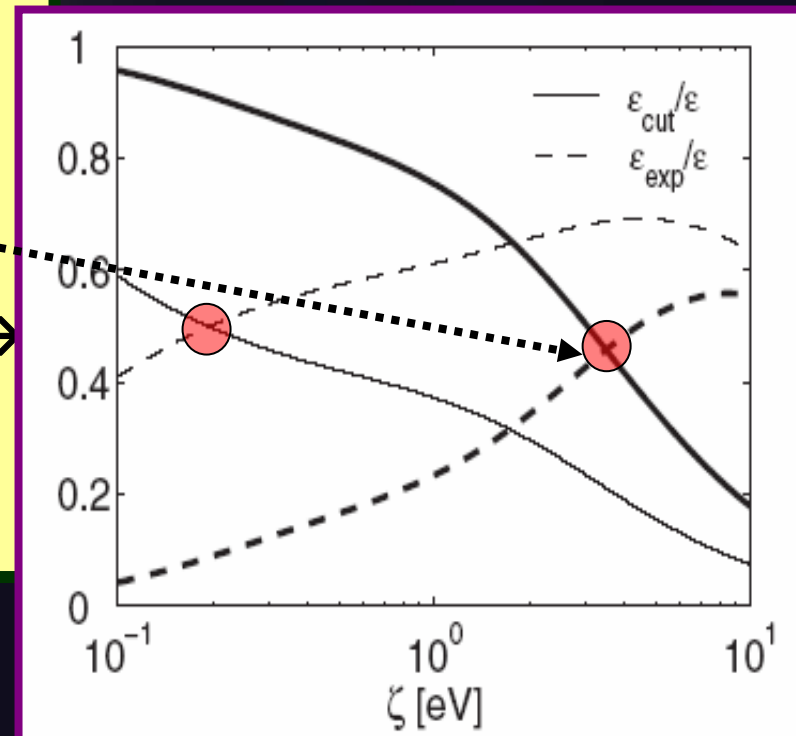


TABLE II. The Drude parameters determined by different methods described in the text. In all cases the statistical errors in the parameters are on the same level: 0.01–0.03 meV for  $\omega_p$  and 0.2–0.5 meV for  $\omega_\tau$ . The last column shows the values of the parameters averaged on different methods and the corresponding rms errors.

Sample	Parameter	Joint $\epsilon', \epsilon''$	Joint $n, k$	KK $\epsilon'$	KK $n$	Average
1	$\omega_p$ [eV]	6.70	6.87	6.88	6.83	$6.82 \pm 0.08$
400 nm/Si	$\omega_\tau$ [meV]	38.4	43.3	40.2	39.9	$40.5 \pm 2.1$
2	$\omega_p$	6.78	7.04	6.69	6.80	$6.83 \pm 0.15$
200 nm/Si	$\omega_\tau$	40.7	45.3	36.1	36.0	$39.5 \pm 4.4$
3	$\omega_p$	7.79	7.94	7.80	7.84	$7.84 \pm 0.07$
100 nm/Si	$\omega_\tau$	48.8	52.0	47.9	47.4	$49.0 \pm 2.1$
4	$\omega_p$	7.90	8.24	7.95	7.90	$8.00 \pm 0.16$
120 nm/Si	$\omega_\tau$	37.1	41.4	35.2	29.2	$35.7 \pm 5.1$
5	$\omega_p$	8.37	8.41	8.27	8.46	$8.38 \pm 0.08$
120 nm/mica	$\omega_\tau$	37.1	37.7	34.5	39.1	$37.1 \pm 1.9$

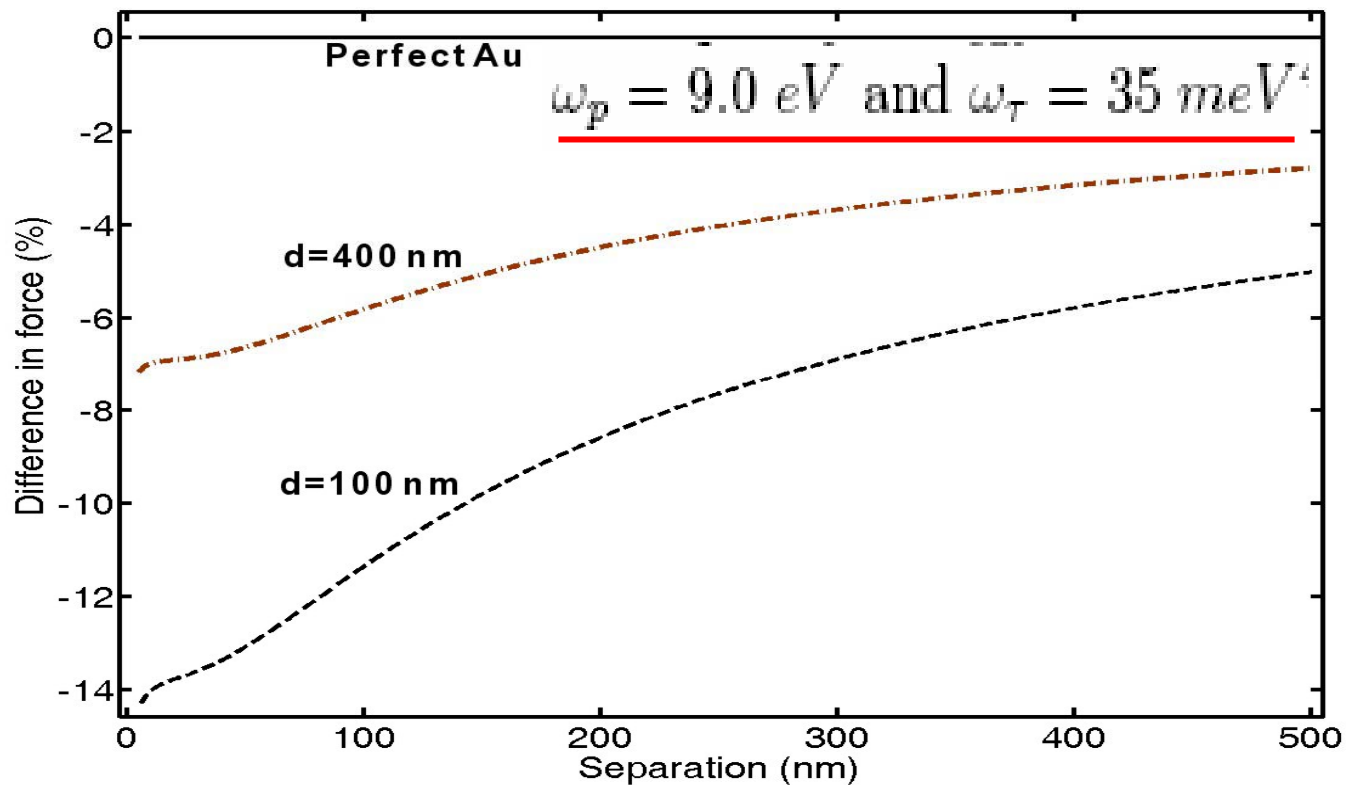


Important contribution to Casimir force comes from imaginary frequencies around

→

$$\zeta_{ch} = c/2a.$$

$$10 \text{ nm} \lesssim a \lesssim 1 \text{ } \mu\text{m} \longleftrightarrow 0.1 \lesssim \zeta_{ch} \lesssim 10 \text{ eV}$$

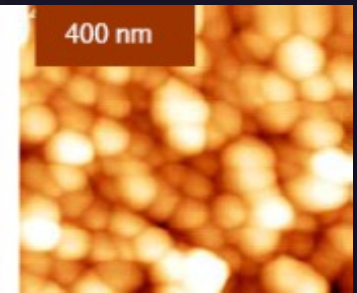
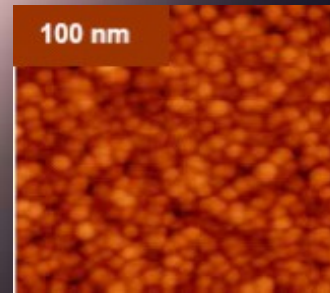
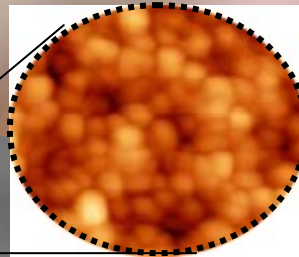
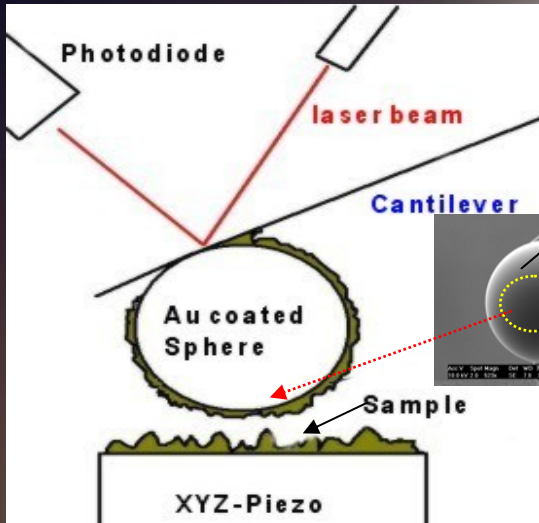


Noise in the optical data is responsible for the force variation within 1%

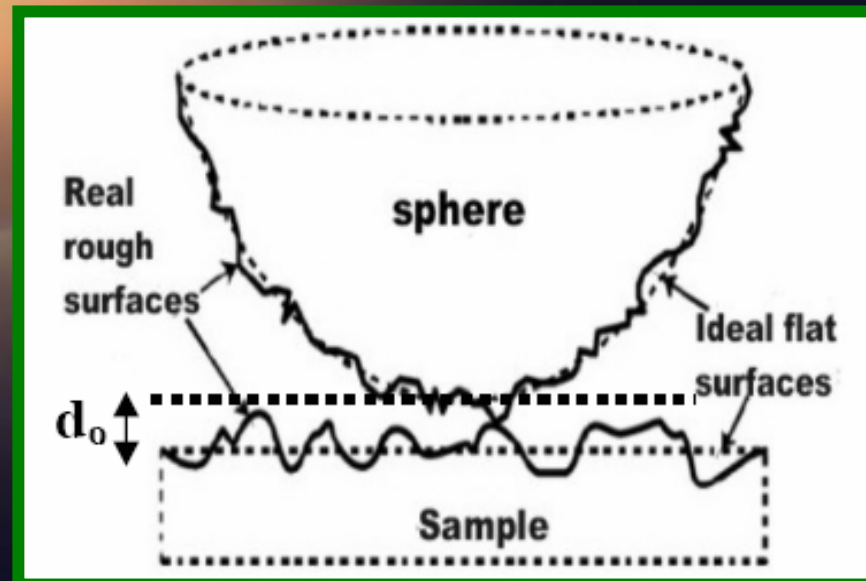
**Optical properties of gold films and the Casimir force**

Precise optical properties of metals are very important for accurate prediction of the Casimir force acting between two metallic plates. Therefore we measured ellipsometrically the optical responses of Au films in a wide range of wavelengths from 0.14 to 33  $\mu\text{m}$ . The films at various thicknesses were deposited at different conditions on silicon or mica substrates. Considerable variation of the frequency dependent dielectric function from sample to sample was found. Detailed analysis of the dielectric functions was performed to check the Kramers-Kronig consistency, and extract the Drude parameters of the films. It was found that the plasma frequency varies in the range from 6.8 to 8.4 eV. It is suggested that this variation is related with the film density. X-ray reflectivity measurements support qualitatively this conclusion. The Casimir force is evaluated for the dielectric functions corresponding to our samples, and for that typically used in the precise prediction of the force. The force for our films was found to be 5%–14% smaller at a distance of 100 nm between the plates. Noise in the optical data is responsible for the force variation within 1%. It is concluded that prediction of the Casimir force between metals with a precision better than 10% must be based on the material optical response measured from visible to mid-infrared range.

# Surface roughness influence



Sphere is usually coated by Au (~100 nm thick) → optically bulk



# Accuracy down to < 1 % ...is it possible ...yes but!....

$K$ : spring constant (~ 1-3%: Electrostatic)

$V_c$ : contact potential (Au/Polystyrene sphere: 5-20 mV)

Measure  $\varepsilon(\omega)$ :  $100 \text{ nm} \leq \lambda \leq 30 \mu\text{m}$  (Ellipsometry)

$d = Z_{piezo} + d_o - Z_{defl}$  (surface separation)

$Z_{piezo}$ : piezo movement,  $Z_{defl}$ : cantilever deflection correction

$d_o$ : contact distance due to roughness  $\approx 3.7 (w_{plate} + w_{sphere})$



Accuracy (%)
5
2
2
1
1
1
15
1
<1

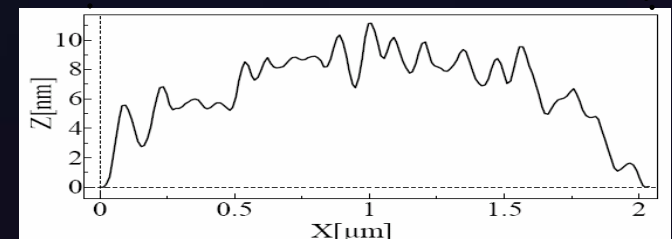
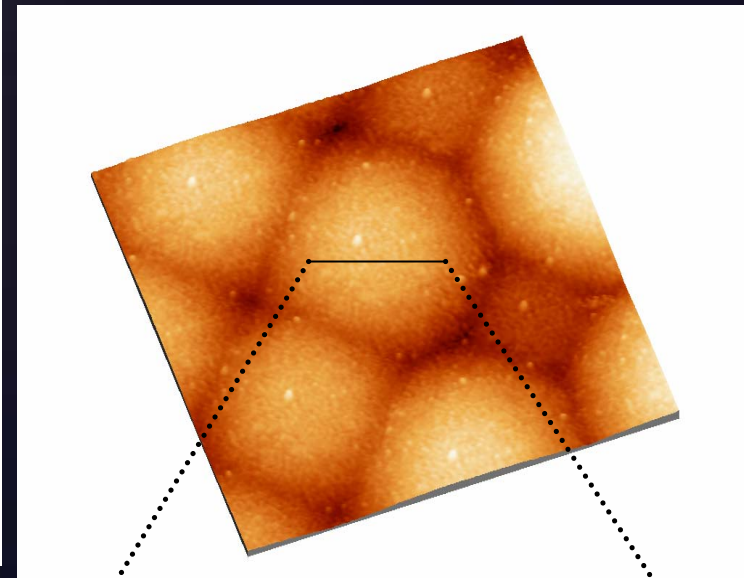
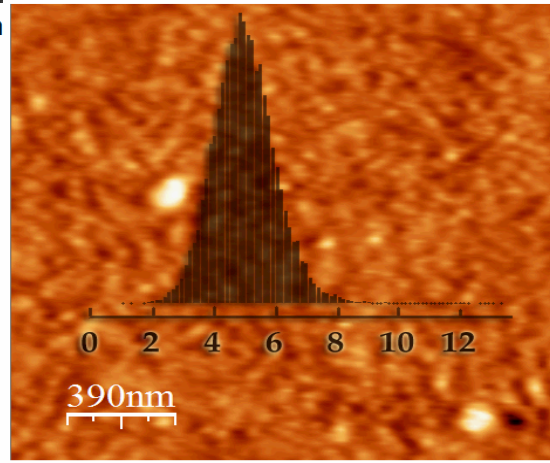
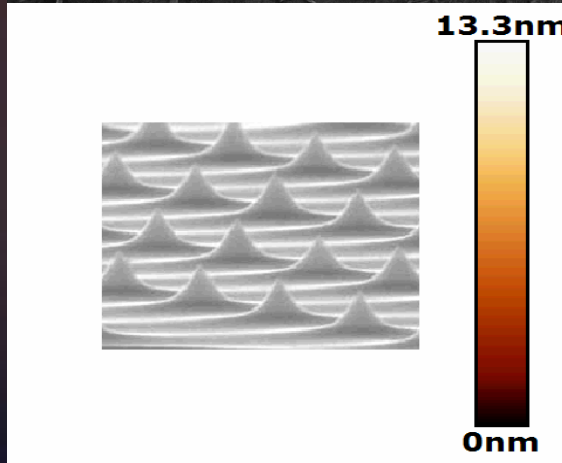
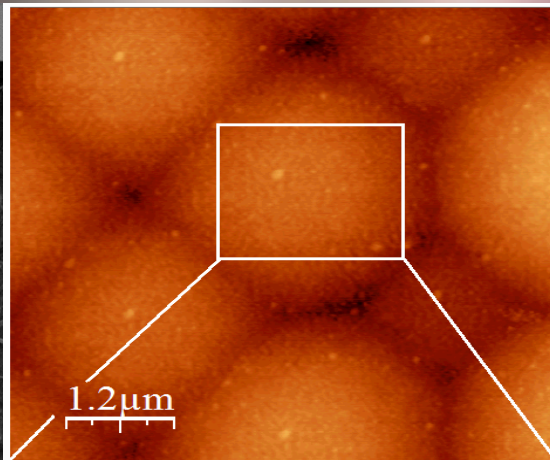
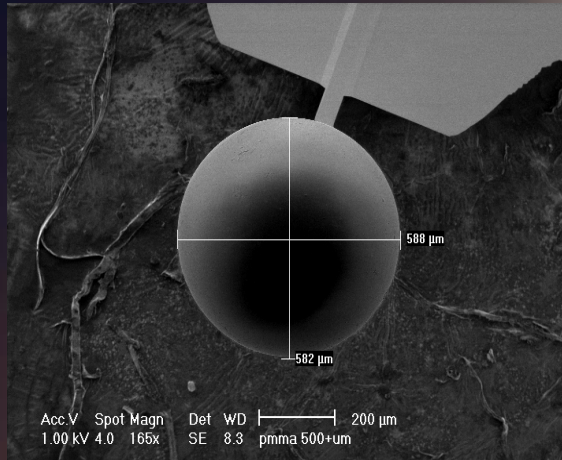
Error:  $\Delta F \approx (m\Delta z / d)F$  ( $m \approx 2.5 - 3$ )

$\Delta z = 1 \text{ nm} \Leftrightarrow d_o = 30 \pm 1 \text{ nm} \rightarrow \Delta F / F \approx 10 \%$

$\Delta z = 0.5 \text{ nm} \Leftrightarrow d_o = 30 \pm 0.5 \text{ nm} \rightarrow \Delta F / F \approx 5 \%$



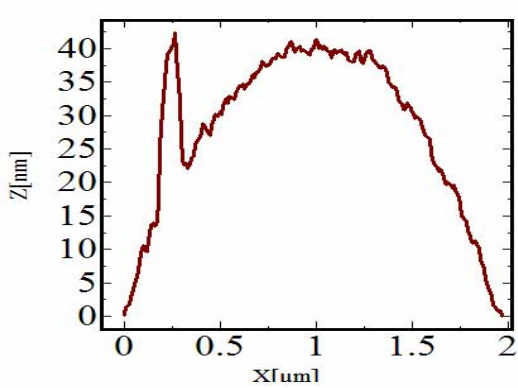
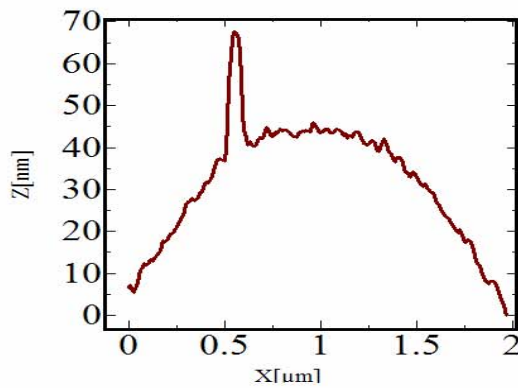
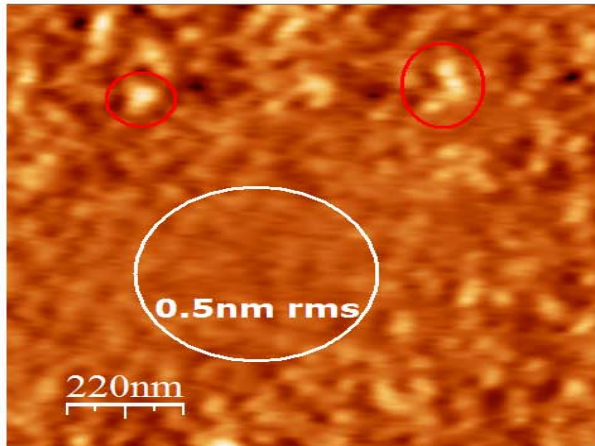
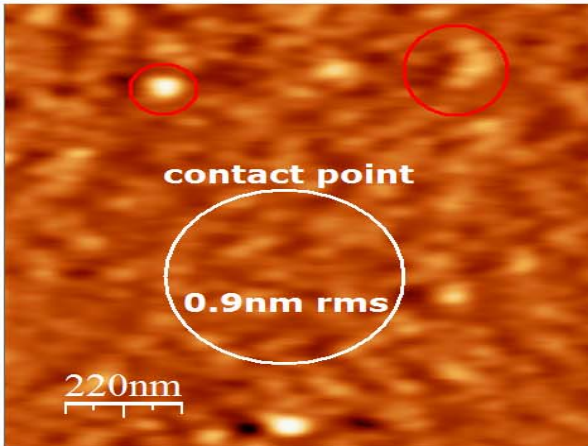
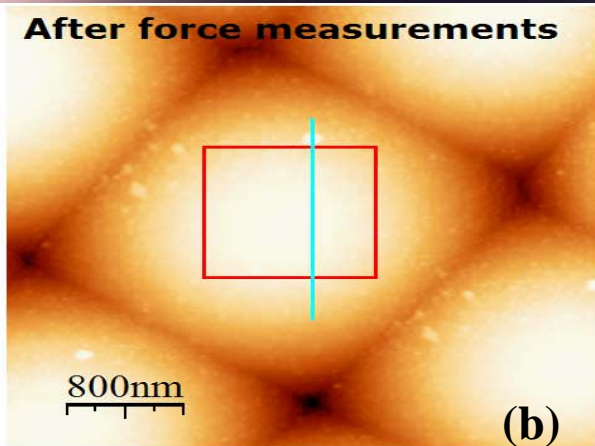
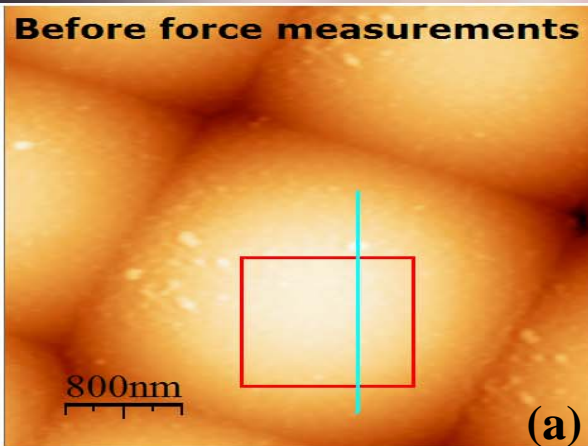
# Reverse imaging, so we have knowledge of the roughness at the contact point of sphere and plate.



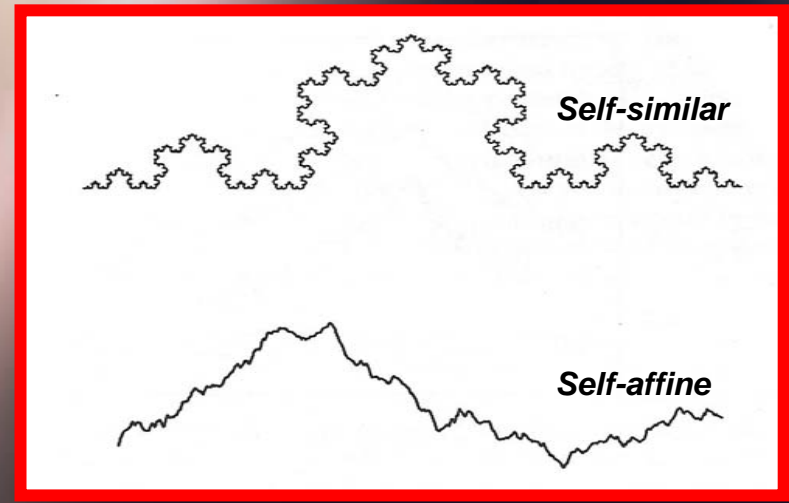
Van Zwol et al, Langmuir 2008

P.J. van Zwol, et al., APPL 2008





# Definition & characterization & growth of roughness



◆ Topography  $\Rightarrow G(r) = \langle [h(r) - h(0)]^2 \rangle$

◆ For self-affine roughness:

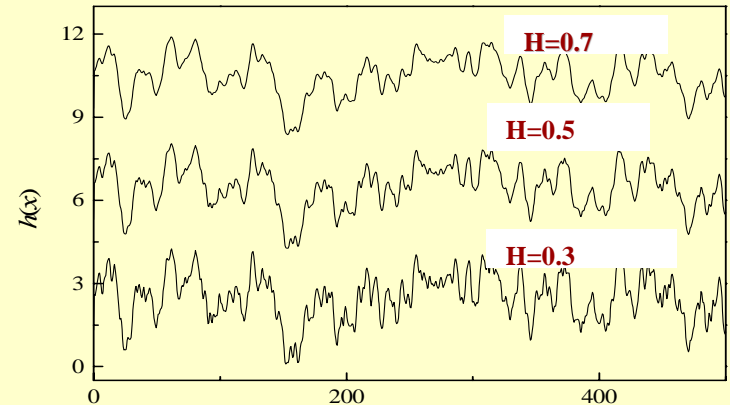
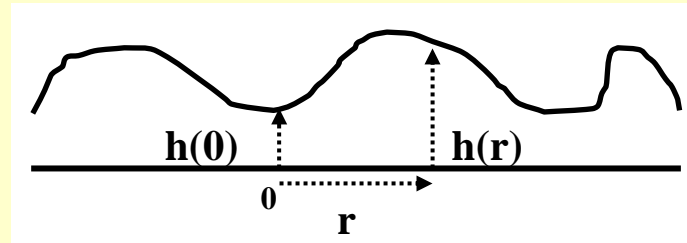
$$G(r) \propto r^{2H} \text{ if } r \ll \xi$$

$$G(r) = 2w^2 \text{ if } r \gg \xi$$

**w**: Rms roughness amplitude

**$\xi$** : In-plane correlation length

**H**: Roughness exponent ( $0 < H < 1$ )



## Normal self-affine growth

If  $\beta = H/z$  ( $H < 1$ )

$$w \propto t^\beta \quad \xi \propto t^{1/z}$$

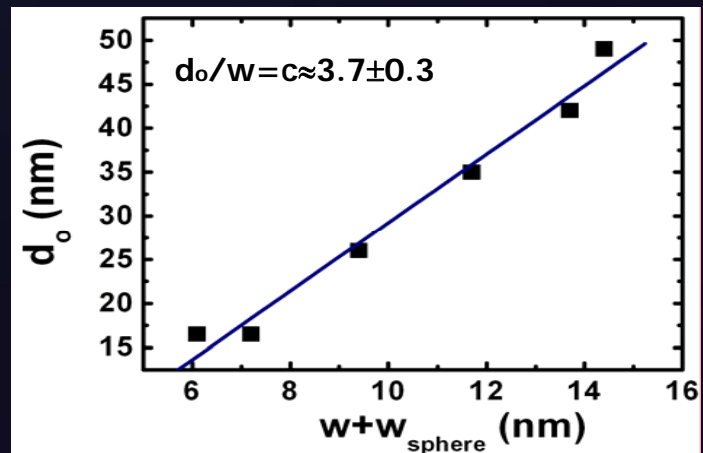
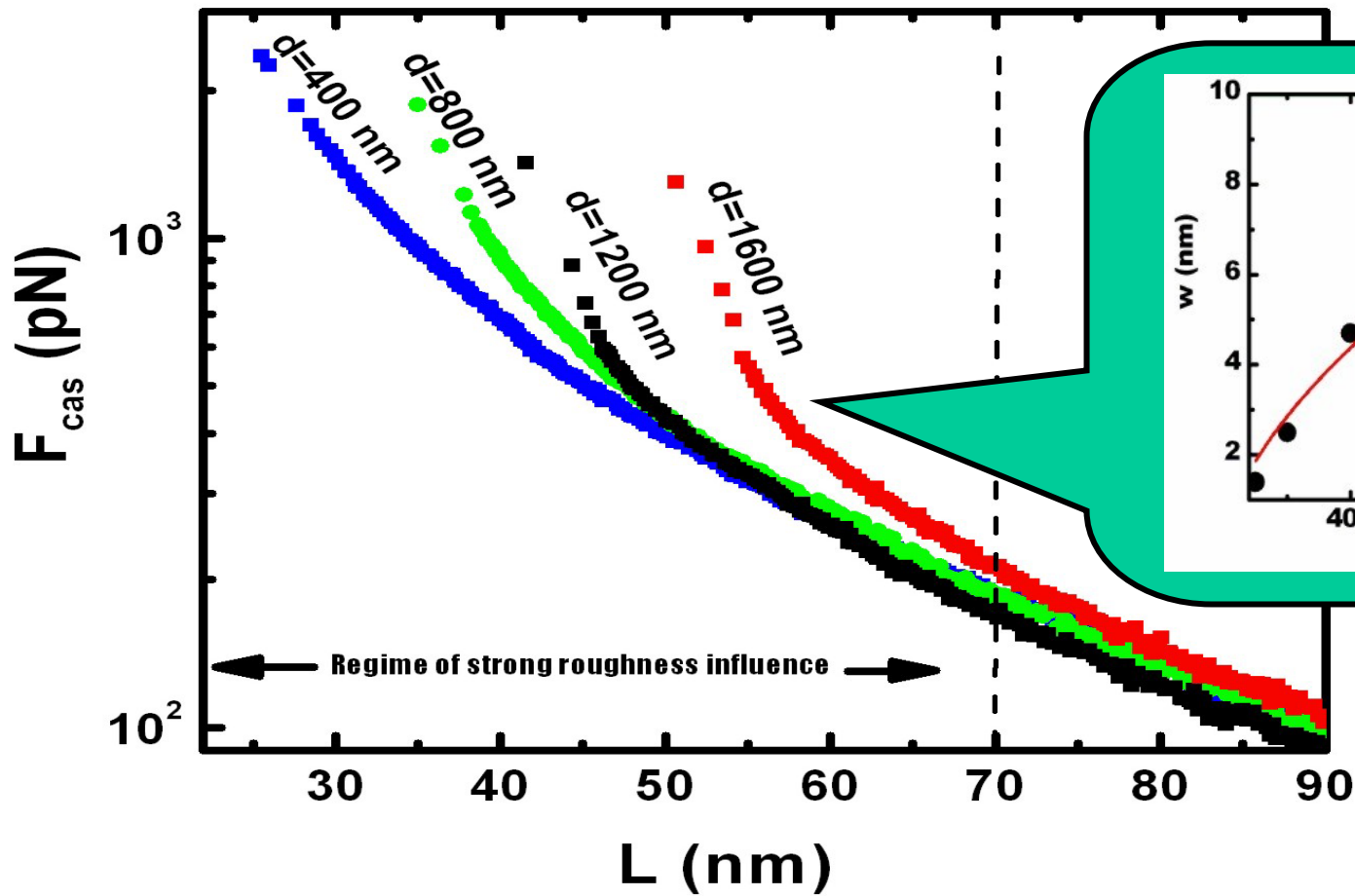
# Summary basic growth models

[Krim & Palasantzas, Int. J. Mod. Phys. B (1995)]

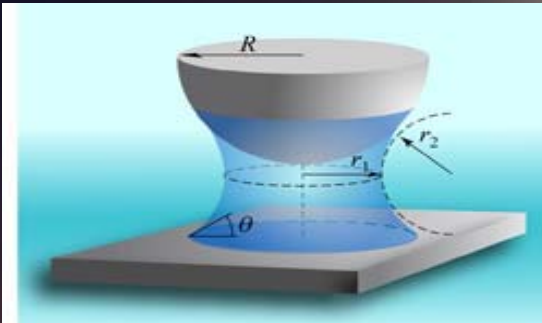
	<b>H</b>	<b><math>\beta</math></b>	<b>[HD]</b>
$\partial h / \partial t = \eta(r, t)$	-	0.5	Gaussian
$\partial h / \partial t = \nu \nabla^2 h + \eta(r, t)$	0	0	Gaussian
$\partial h / \partial t = -\kappa \nabla^4 h + \eta(r, t)$	1	0.25	Gaussian
$\partial h / \partial t = \nu \nabla^2 h + \lambda (\nabla h)^2 / 2 + \eta(r, t)$	0.38	0.24	Non-Gaussian
$\partial h / \partial t = -k \nabla^4 h + \nu \nabla^2 (\nabla h)^2 + \eta(r, t)$	2/3	0.20	Non-Gaussian

**Gaussian:  $h \leftrightarrow -h$  symmetry**

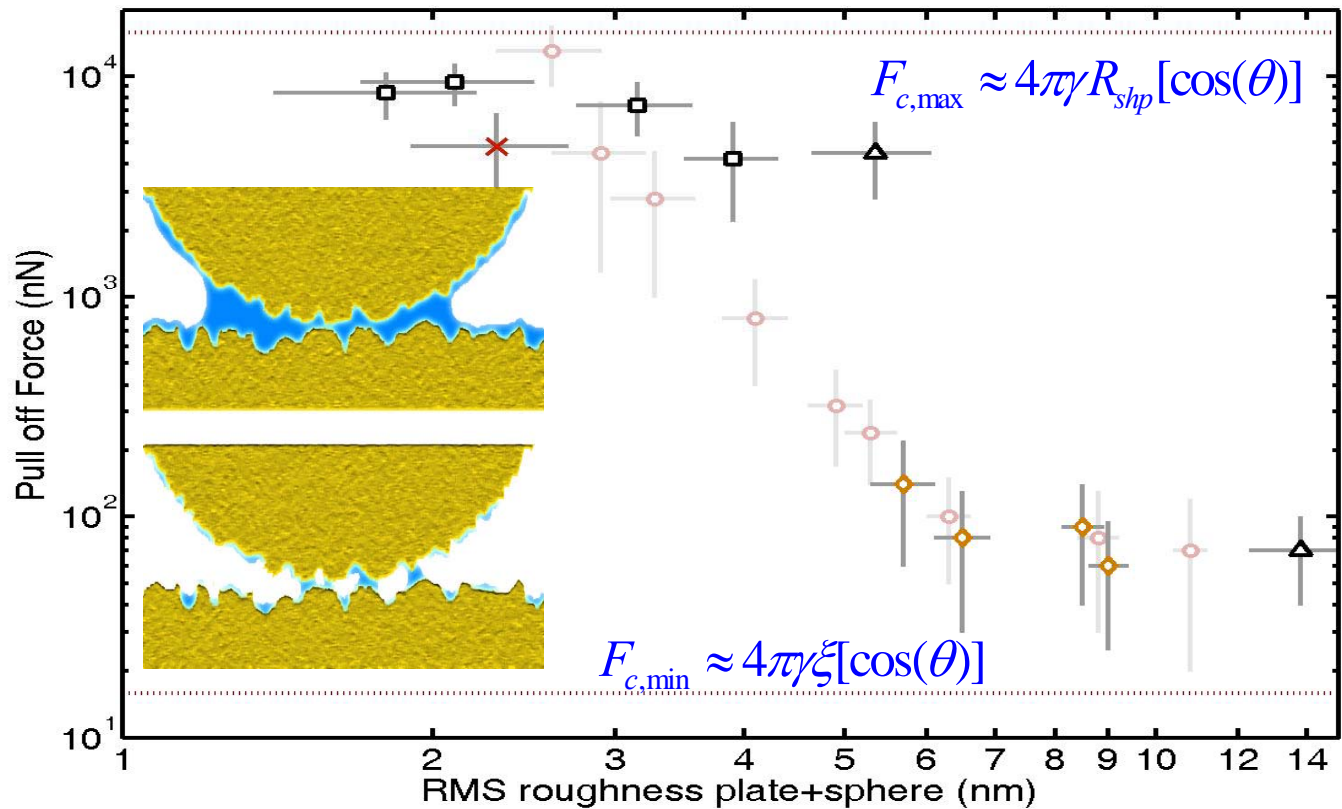
$$\begin{cases} \langle \eta(r, t) \rangle = 0 & \text{white-noise} \\ \langle \eta(r, t) \eta(r', t') \rangle = 2D \delta(r - r') \delta(t - t'). \end{cases}$$



# Operate in air so upon contact we have capillary condensation



$$F_{ad} = \pi\gamma R_{sph} \left\{ -\sin\varphi + [\cos(\theta_1 + \varphi) + \cos(\theta_2)] \sin^2\varphi (1 - \cos\varphi + D/R_{sph})^{-1} \right\} + 2\pi\gamma R_{sph} \sin\varphi \sin(\theta_1 + \varphi)$$





$$F_{theory} = (2\pi R/A) E_{pp,rough}$$

$$E_{pp,rough} = E_{ppflat} + \delta E_{pp,rough}$$

$$E_{ppflat} = \hbar A \sum_P \int [d^2k / 4\pi^2] \int_0^\infty [d\Phi / 2\pi] \ln[1 - r^P(k, \Phi)^2 e^{-2\kappa d}]$$

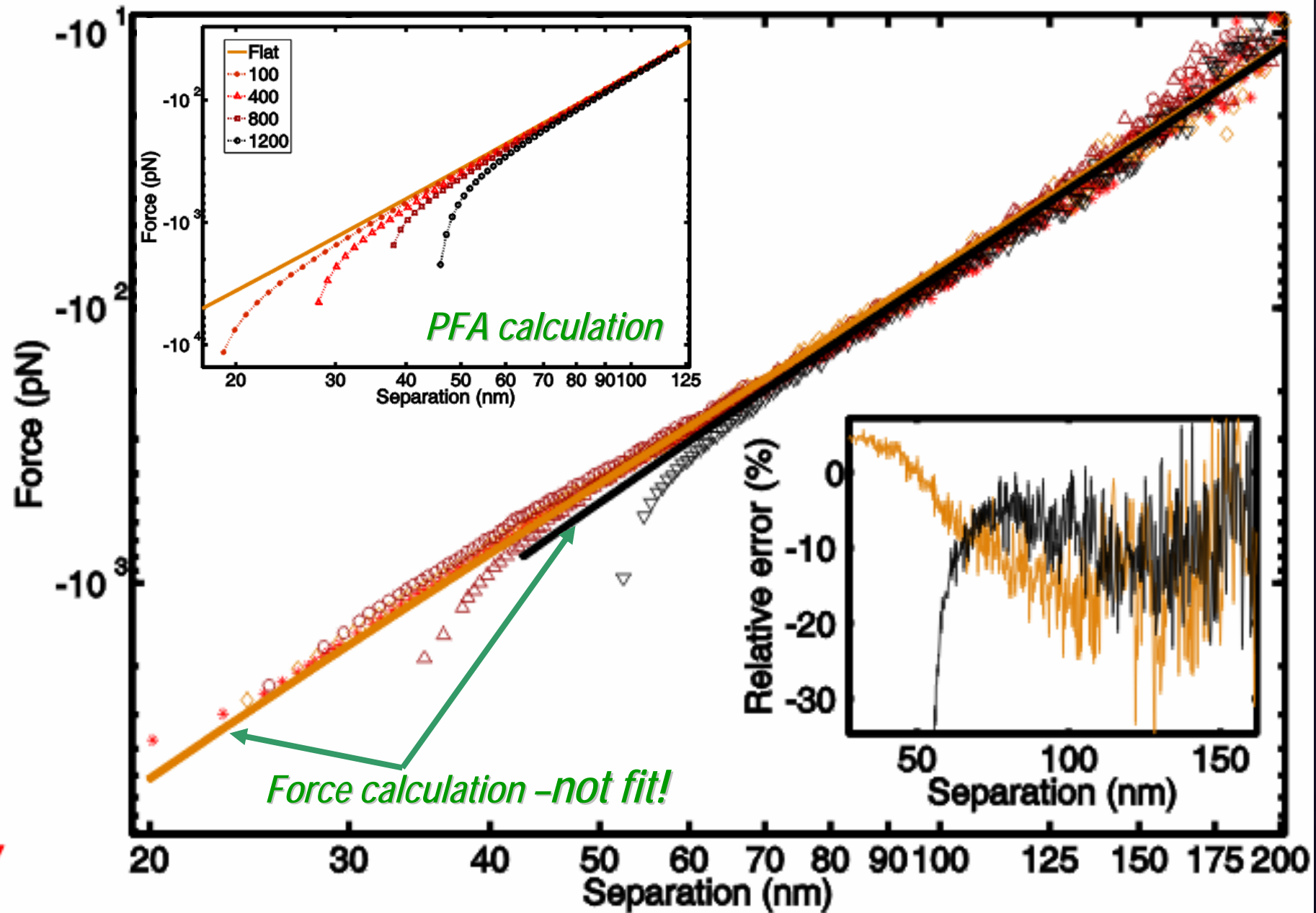
$$\delta E_{pp,rough} = \int [d^2k / 4\pi^2] G(k) \sigma(k)$$

Neto, Lambrecht, Reynaud, PRA 2005

$$\sigma(k) = (AHw^2 \xi^2) / (1 + k^2 \xi^2)^{1+H}$$

Palasantzas, PRB 1993

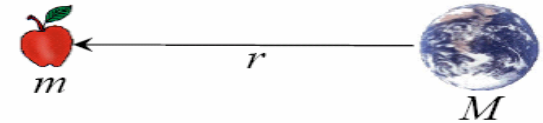
# Roughness influence: Experiment & Perturbation theory



# A Short-Range Test of Newton's Gravitational Inverse-Square Law

Isaac Newton's Gravitational Inverse-Square Law (ISL)

$$F = \frac{GMm}{r^2}$$



Gravity is an *attractive* force that acts between all mass/energy in the Universe.

Expect the expansion of the Universe to slow.

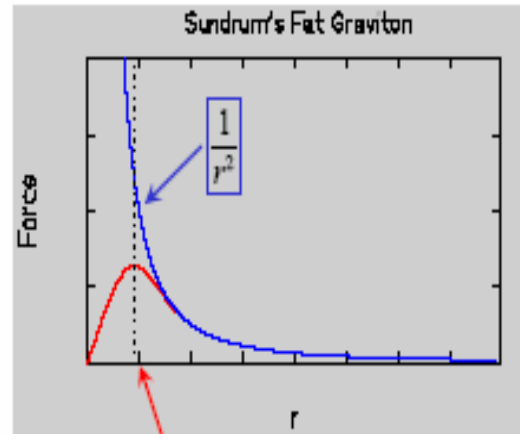
The expansion is accelerating, which implies a dark energy density :

$$\rho_d = 3.8 \text{ keV/cm}^3$$

• And a natural length scale :

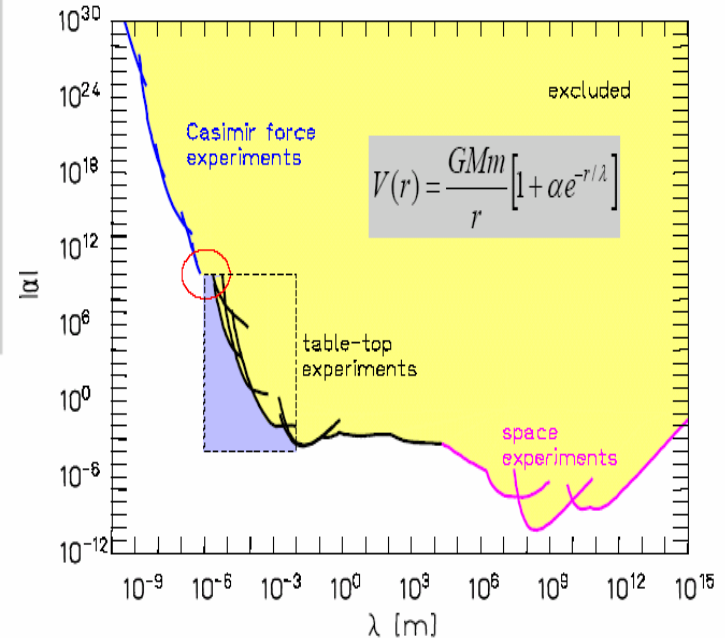
$$\lambda_d = \sqrt[4]{hc/\rho_d} \approx 85\mu\text{m}$$

A modified ISL could potentially link the dark energy density with vacuum fluctuations, a much higher energy density. Gravity might be blind to small-scale (high-energy) physics.



This length scale is unknown, but should be  $> 20\mu\text{m}$ .

Measured over a huge range of scales in search of new physics :



•Fifth force & Higher dimensions → short range force with coupling of order of gravity

$$V_N(r) = -G \frac{M}{r}$$

$$V_{\text{tot}} = V_N + V_Y$$

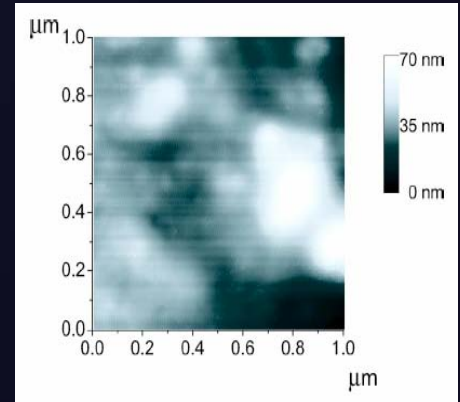
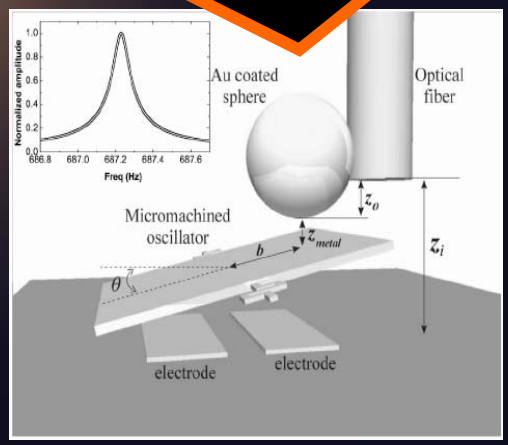
$$V_Y(r) = -\alpha G \frac{M}{r} e^{-r/\lambda}$$

$N = 4 + n$  dimensions

$\lambda \sim R_n$   $R_n$  is the size of the compact dimensions,  
 $r \gg R_n \sim \frac{1}{M_{Pl}^{(N)}} \left( \frac{M_{Pl}}{M_{Pl}^{(N)}} \right)^{2/n} \sim 10^{32/n - 17}$  cm.  ~~$R_1 \sim 10^{15}$  cm~~  
 $R_2 \sim 1$  mm or  $R_3 \sim 5$  nm.

Casimir force:  
 background force to be removed → search for hypothetical forces

**MEMS set-up for hypo-forces**



Decca et al, PRD 2003

•Fifth force & Higher dimensions → short range force with coupling of order of gravity

**Novel constraints on light elementary particles and extra-dimensional physics from the Casimir effect**

R.S. Decca<sup>1</sup>, D. López<sup>2</sup>, E. Fischbach<sup>3</sup>, G.L. Klimchitskaya<sup>4,\*</sup>, D.E. Krauss<sup>5,3</sup>, V.M. Mostepanenko<sup>4,b,c</sup>

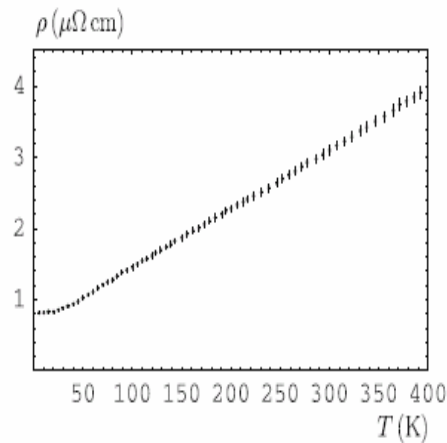


Fig. 2. Resistivity of the Au films (measured with an error of about 2%) as a function of temperature

$$\rho(T) = \frac{4\pi}{\omega_p^2 \tau(T)} = \frac{4\pi v_F}{\omega_p^2 l(T)} = \frac{CT}{\omega_p^{3/2}}. \quad (9)$$

Here  $\tau(T) = l(T)/v_F$  is the relaxation time,  $l(T) \sim T$  is the mean free path of an electron,  $v_F \sim \omega_p^{1/2}$  is the Fermi velocity, and  $C = \text{const}$ . The fit results in  $C/\omega_p^{3/2} = (8.14 \pm 0.16) \text{ n}\Omega \text{ cm K}^{-1}$ . On the other hand, using the resistivity data for pure Au as a function of temperature [67] and the previously used value of the plasma frequency  $\tilde{\omega}_p = 9.0 \text{ eV}$  [49, 68] we obtain  $C/\tilde{\omega}_p^{3/2} = 8.00$ . As a result we find for the Au film used in our experiment  $\omega_p = (8.9 \pm 0.1) \text{ eV}$ . Here, the absolute error of 0.1 eV arises from

quire knowledge of the relaxation parameter. The smooth Drude extrapolation of the imaginary part of the Au dielectric permittivity, given by the tabulated optical data [49], yields the relaxation parameter at room temperature  $\gamma = 0.0357 \text{ eV}$  (which compares with  $\tilde{\gamma} = 0.035 \text{ eV}$  used in previous work [43, 44, 68]).

$$V(r) = -\frac{Gm_1m_2}{r} \left(1 + \alpha e^{-r/\lambda}\right)$$

$$P^{\text{hyp}}(z) = -2\pi G\alpha\lambda^2 e^{-z/\lambda} \quad (24)$$

$$\times \left[ \rho_g - (\rho_g - \rho_c) e^{-\Delta_g^{(s)}/\lambda} - (\rho_c - \rho_s) e^{-(\Delta_g^{(s)} + \Delta_c)/\lambda} \right]$$

$$\times \left[ \rho_g - (\rho_g - \rho_c) e^{-\Delta_g^{(p)}/\lambda} - (\rho_c - \rho_{si}) e^{-(\Delta_g^{(p)} + \Delta_c)/\lambda} \right].$$

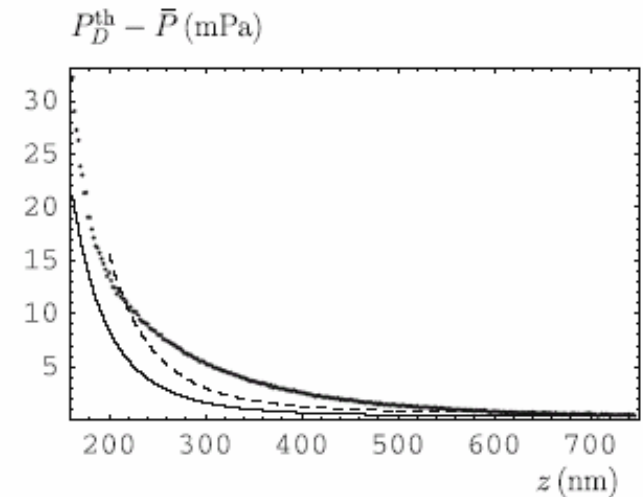
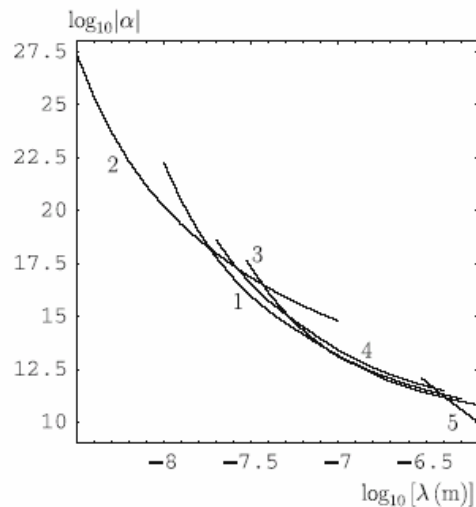
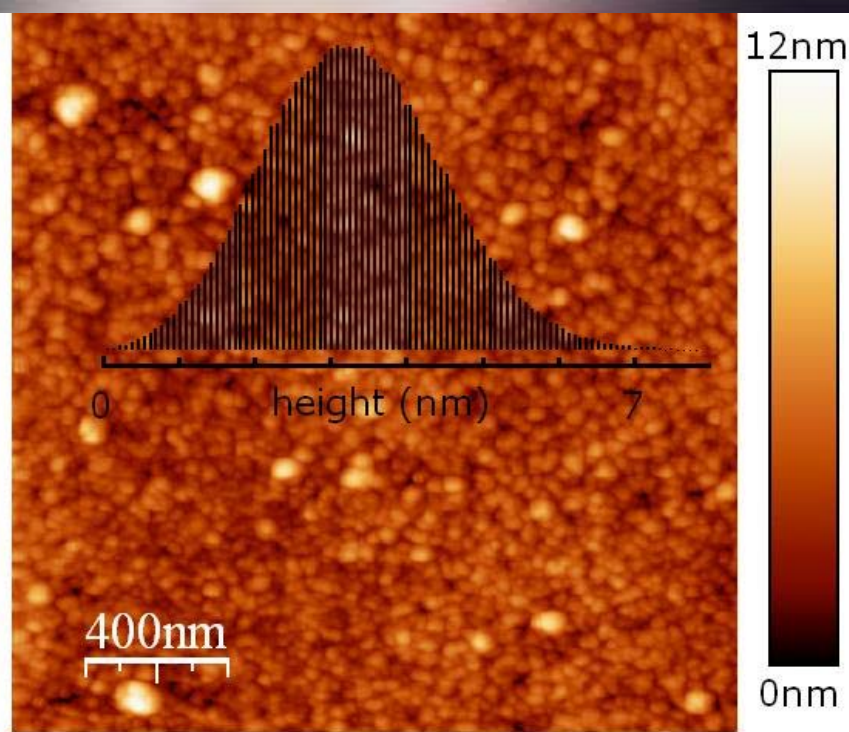
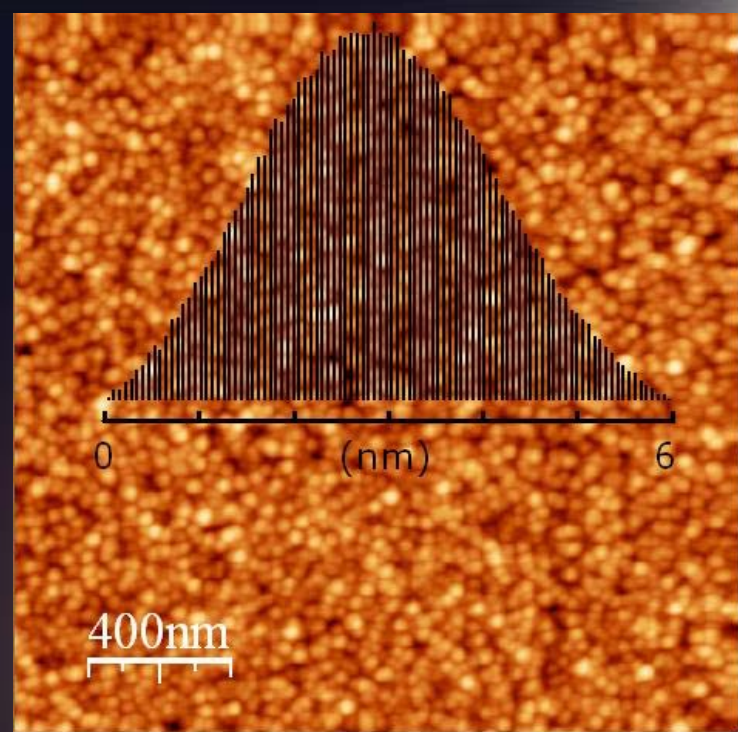


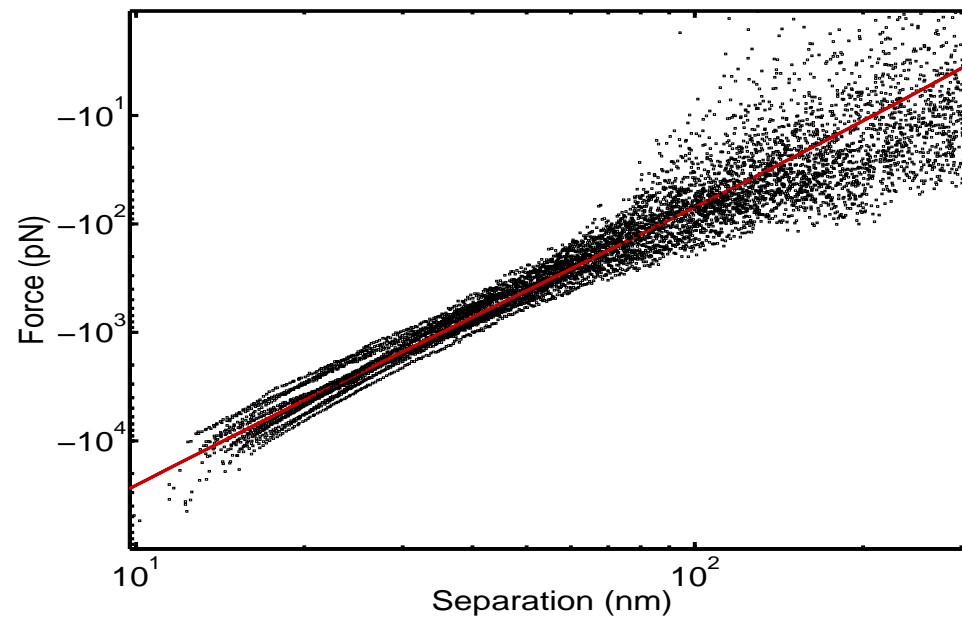
Fig. 5. Differences between theoretical Casimir pressures computed using the Drude model approach and mean experimental Casimir pressures (*dots*) versus separation. The *solid line* indicates the limits of the 95% confidence intervals, while the *dashed line* indicates the limits of the 99.9% confidence intervals

relevant theory. The strongest constraints within the interaction region  $10 \text{ nm} \leq \lambda \leq 56 \text{ nm}$  are obtained from the comparison of measurements with theory at a separation  $z = 180 \text{ nm}$ . With the increase of  $\lambda$ , the strongest con-

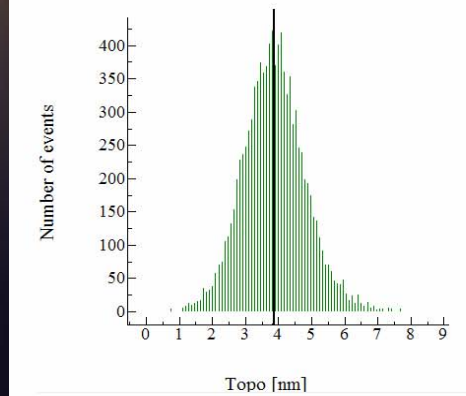
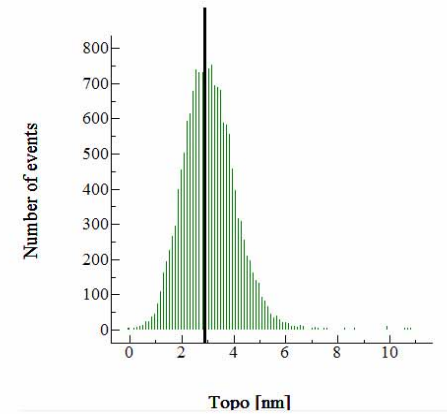
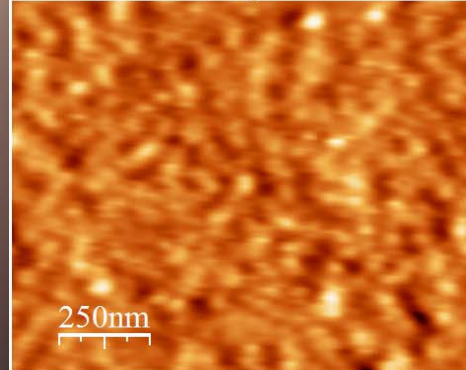
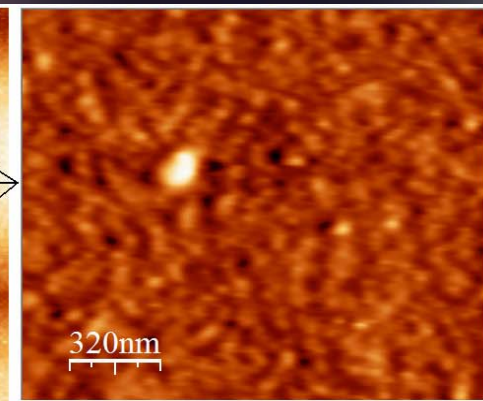
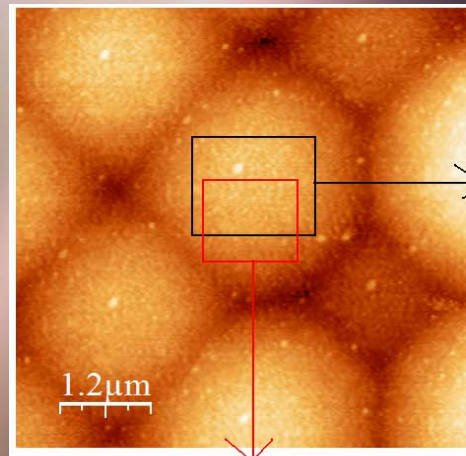
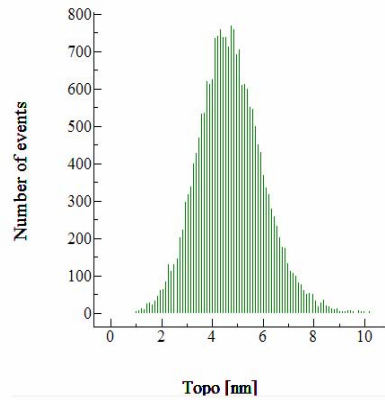
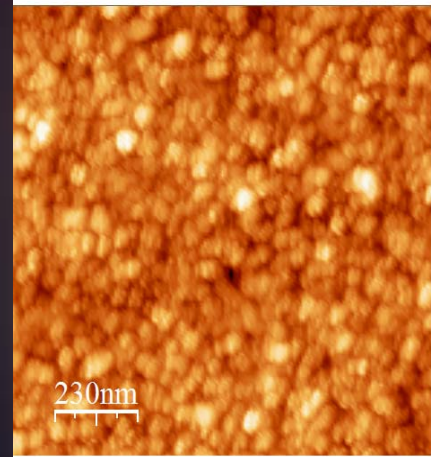
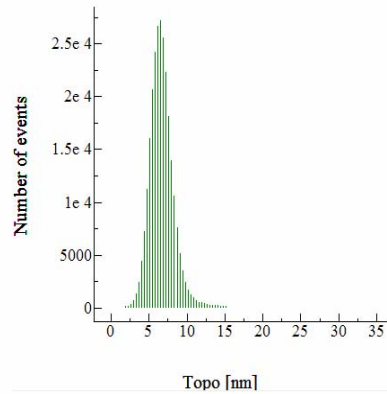
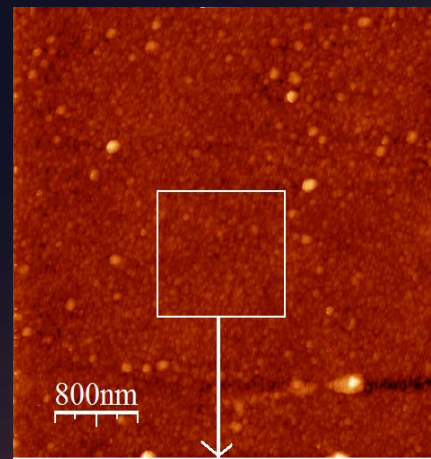




- Stiff-cantilever  $\rightarrow K=4 \text{ N/m}$
- Reduce jump-to-contact



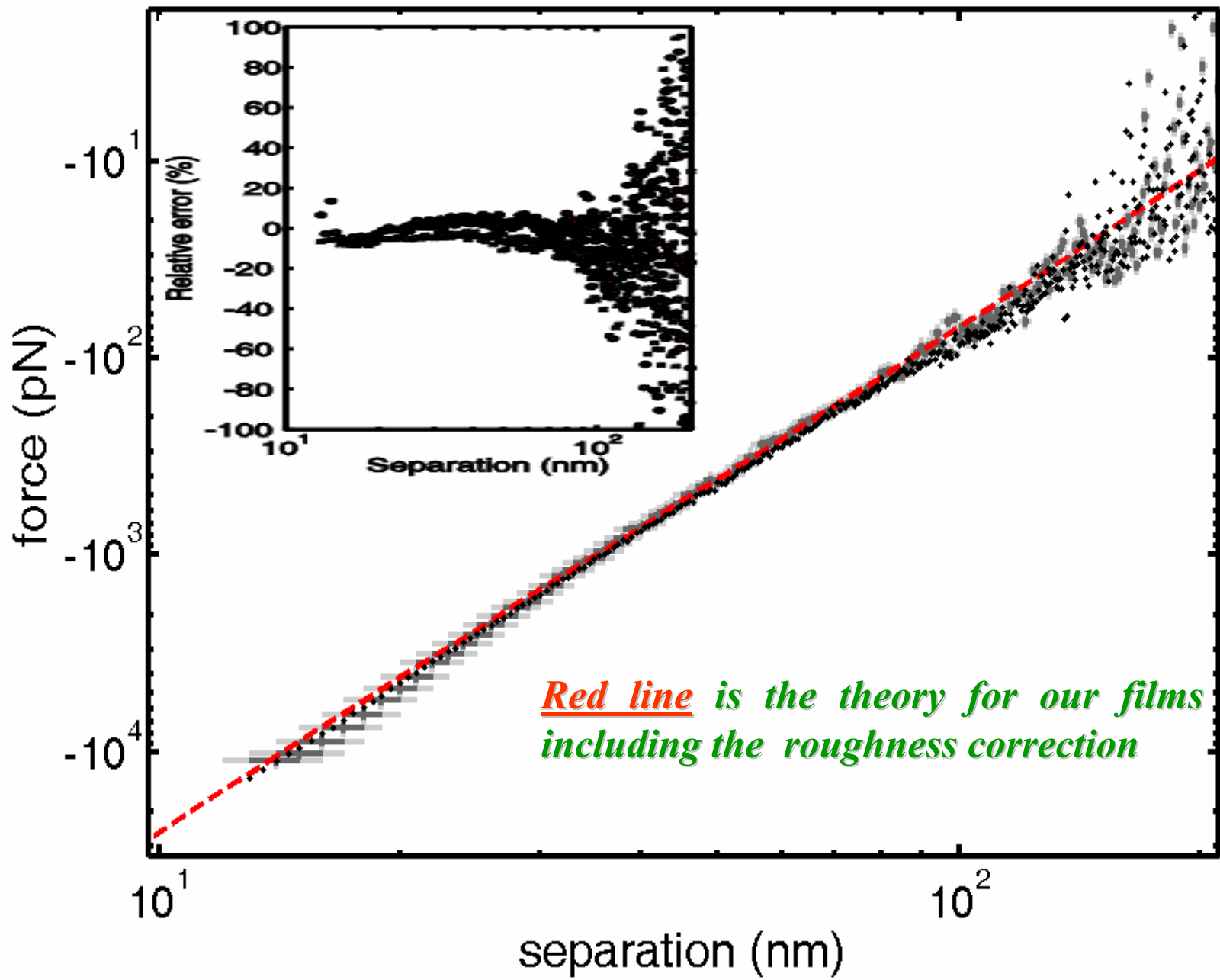




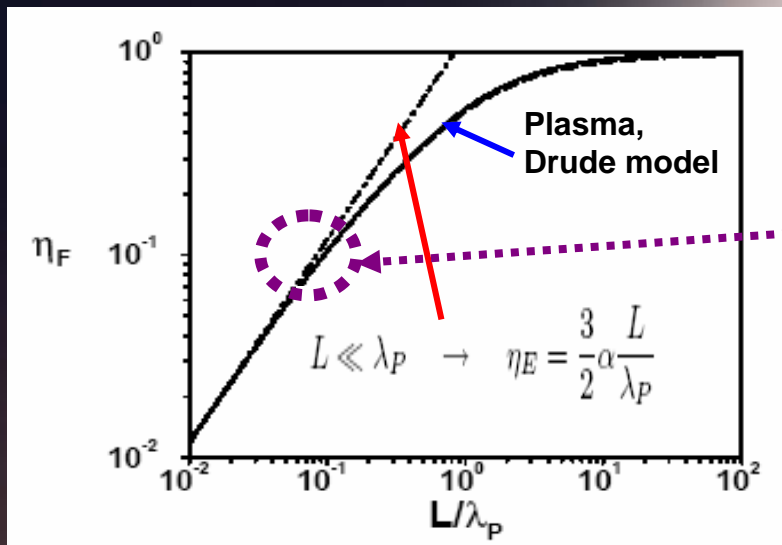
*Rms roughness of 2.2 nm without spots → decreases to 1.3 nm.*

*we get  $d_0 = 7.5 \text{ nm} \pm 1 \text{ nm}$*

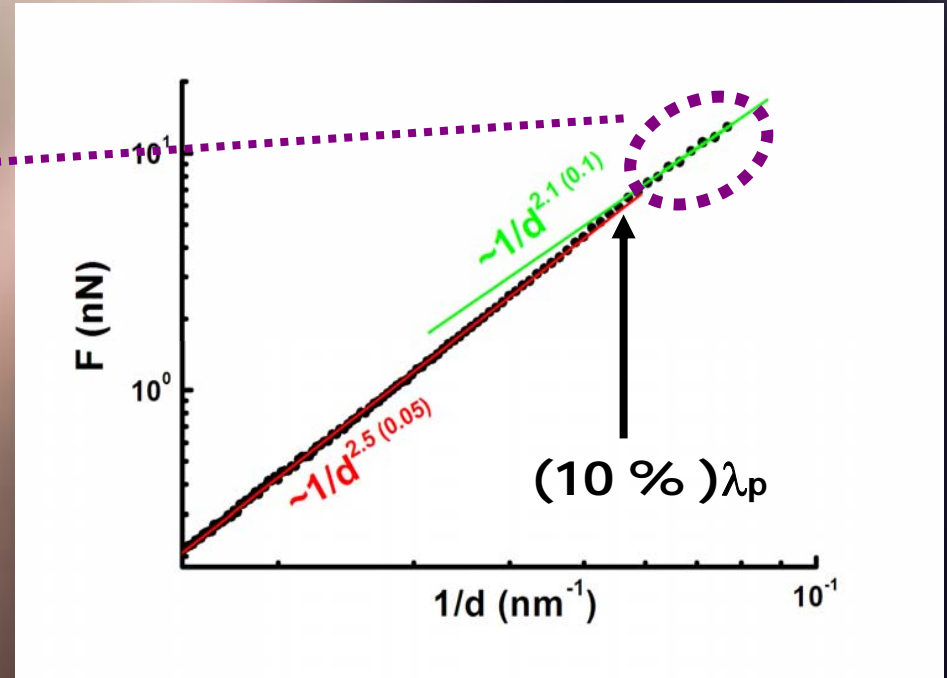
*• Top to bottom roughness of the sphere (without the spots) is about 6 nm*



**Water layer  $\cong$  1-2 nm thick influence?**



Lambrecht, Reynaund, E. J. Phs. D (2000)



Palasantzas et al., APL 2008

$$F_{vdW} \cong - \frac{A_H R}{6d^2} \rightarrow A_H \approx 29 \times 10^{-20} \text{ J}$$

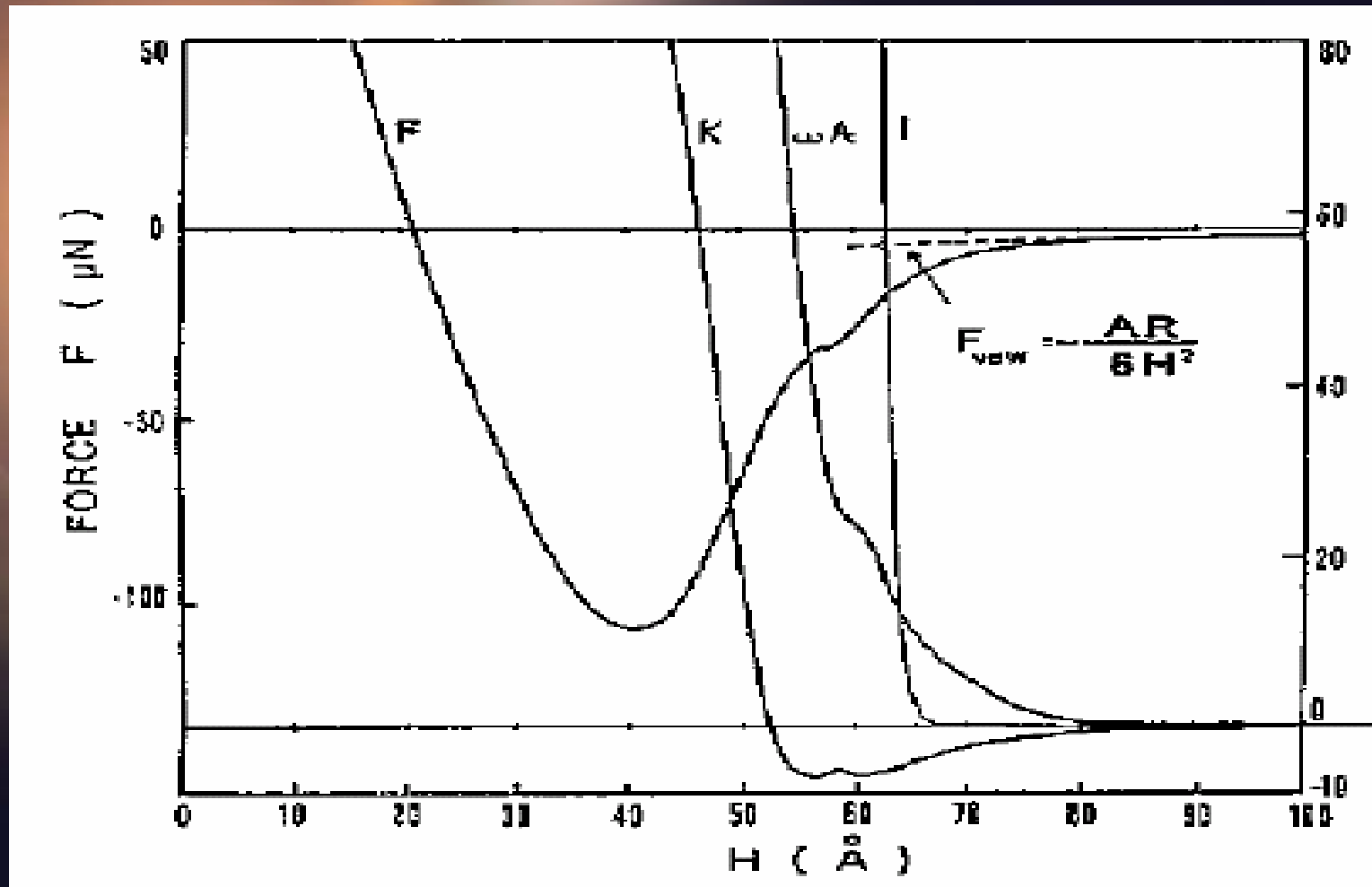
surface force apparatus (SFA)  $\rightarrow$   $A_H \approx 28 \times 10^{-20} \text{ J}$

For  $d > 8.5 \text{ nm}$  for Au - roughness of 5–6 nm peak to peak (Tonck et al., j. Phys. Con. Mat 1991)

surface force apparatus (SFA)  $\rightarrow A_H \approx 28 \times 10^{-20} \text{ J}^2$

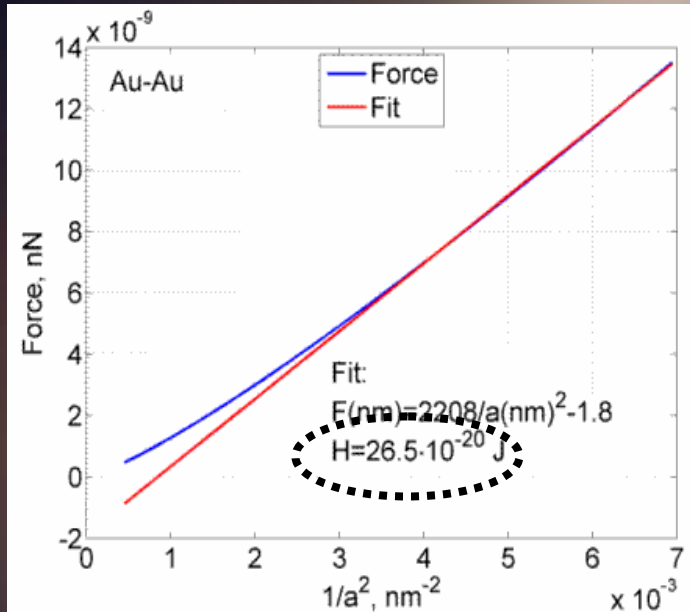
For  $d > 8.5 \text{ nm}$  for Au - roughness of 5–6 nm peak to peak (Tonck et al., *J. Phys. Con. Mat* 1991)

600 Å gold terminal coating. A STM examination of such surfaces shows irregular connected clusters producing a gently bumpy corrugation similar to the 'blackberry' roughness already observed on comparable Au films [12], with peak-to-valley heights of 50–60 Å and a mean distance between adjacent bumps of about 500 Å [13]. After

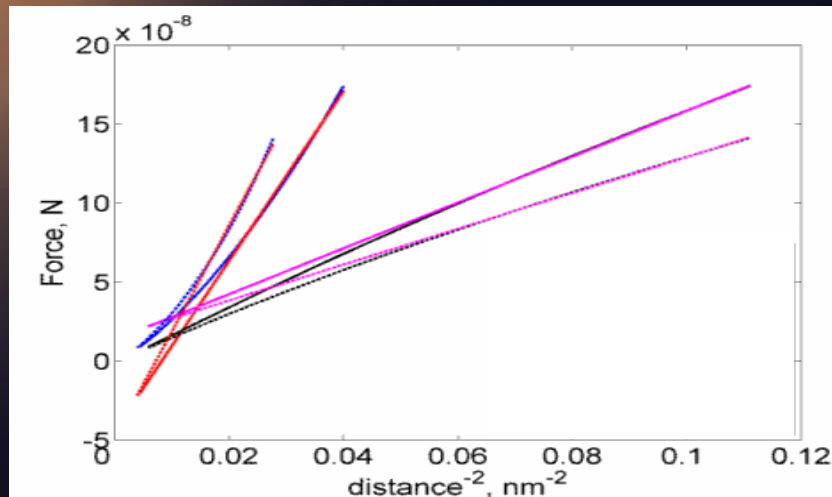
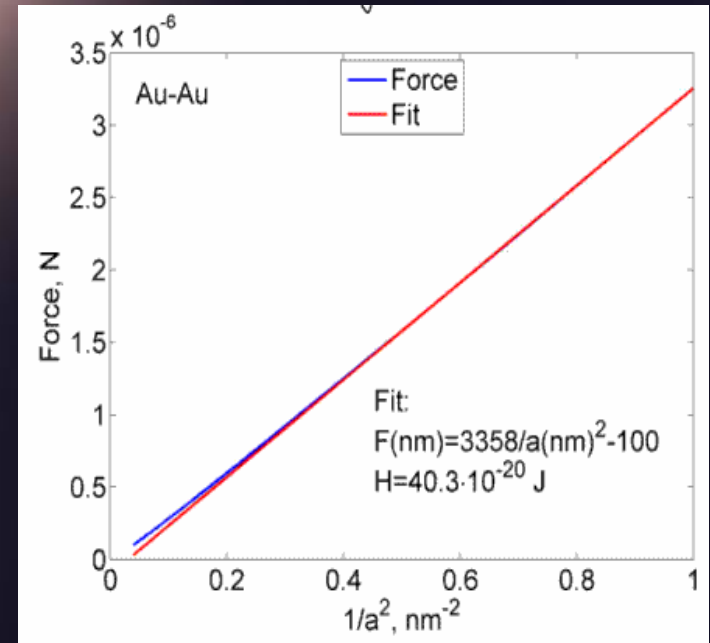
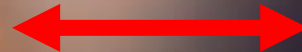




# Water layer $\cong$ 1-2 nm thick influence?



No water



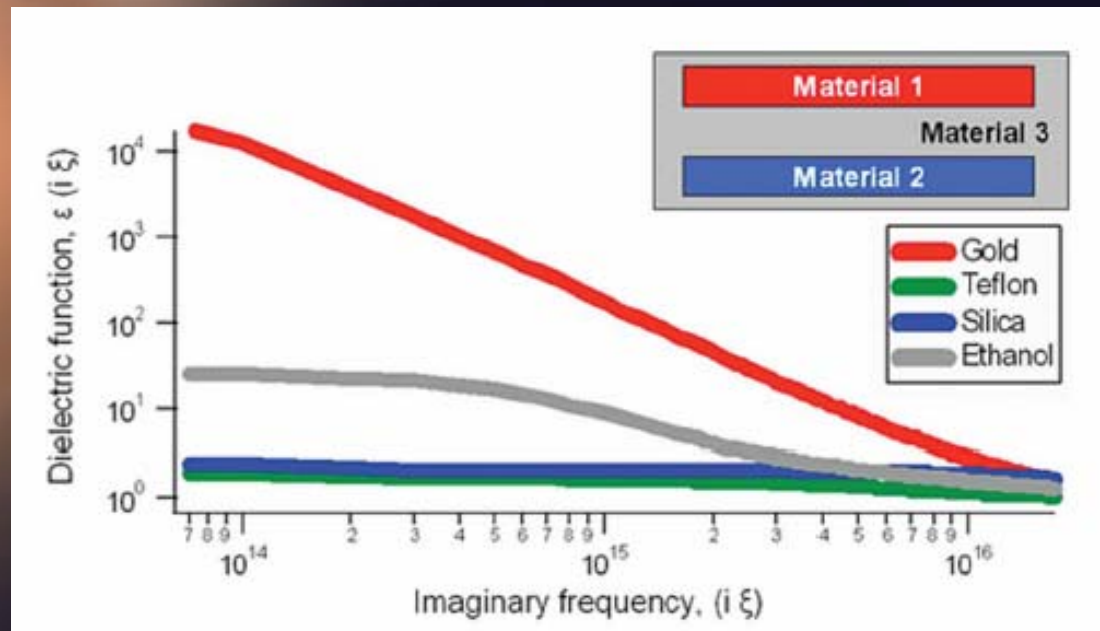
Svetovoy, Palasantzas et al, in preparation (2008)

# Measurement in liquids-Repulsive – attractive forces

- The Casimir–Lifshitz force is always attractive between two identical materials
- The force can become repulsive when two different materials are submerged in a third medium if→

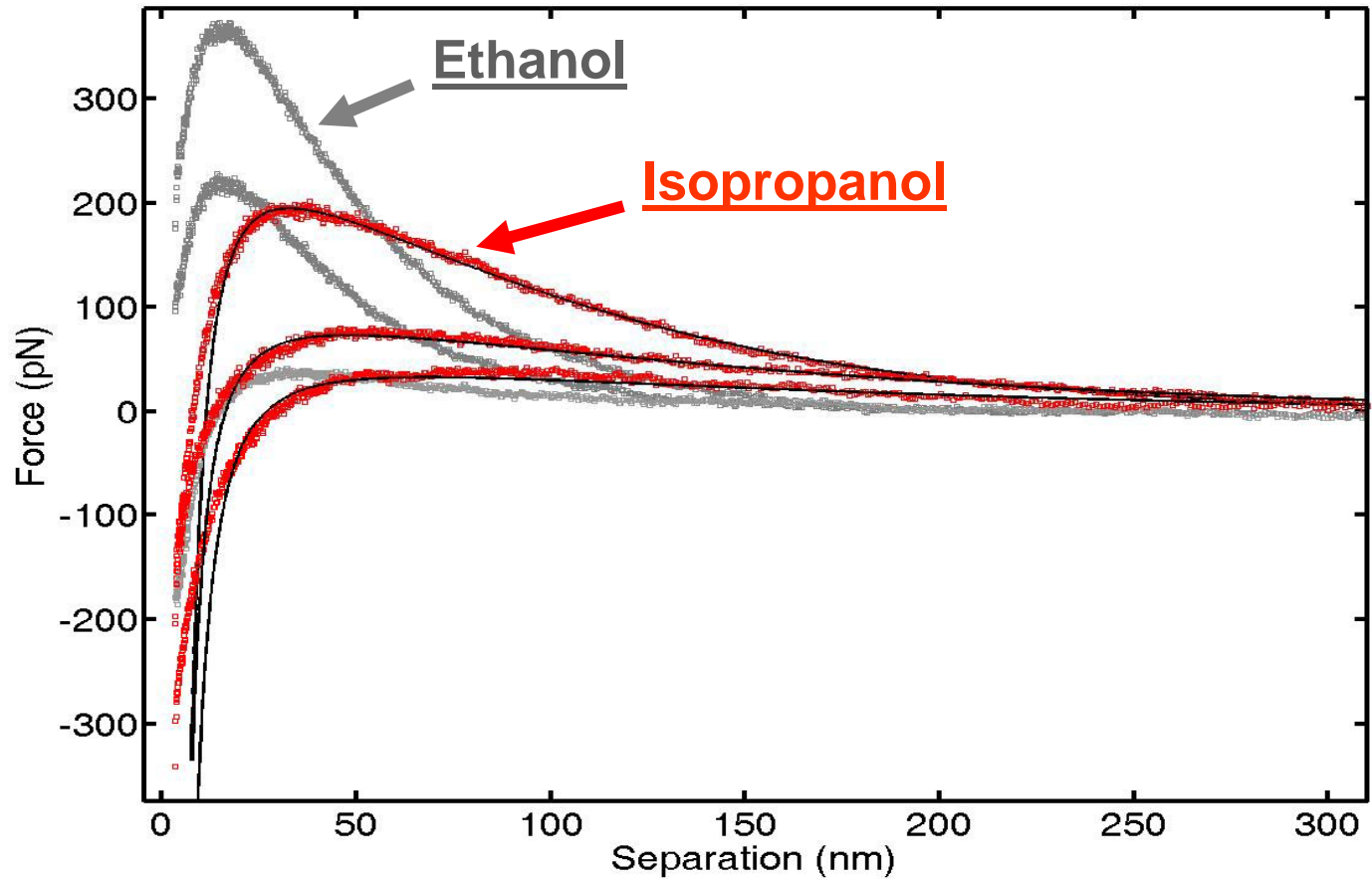
$$\varepsilon_1(i\xi) < \varepsilon_3(i\xi) < \varepsilon_2(i\xi)$$

$$\varepsilon_2(i\xi) < \varepsilon_3(i\xi) < \varepsilon_1(i\xi)$$



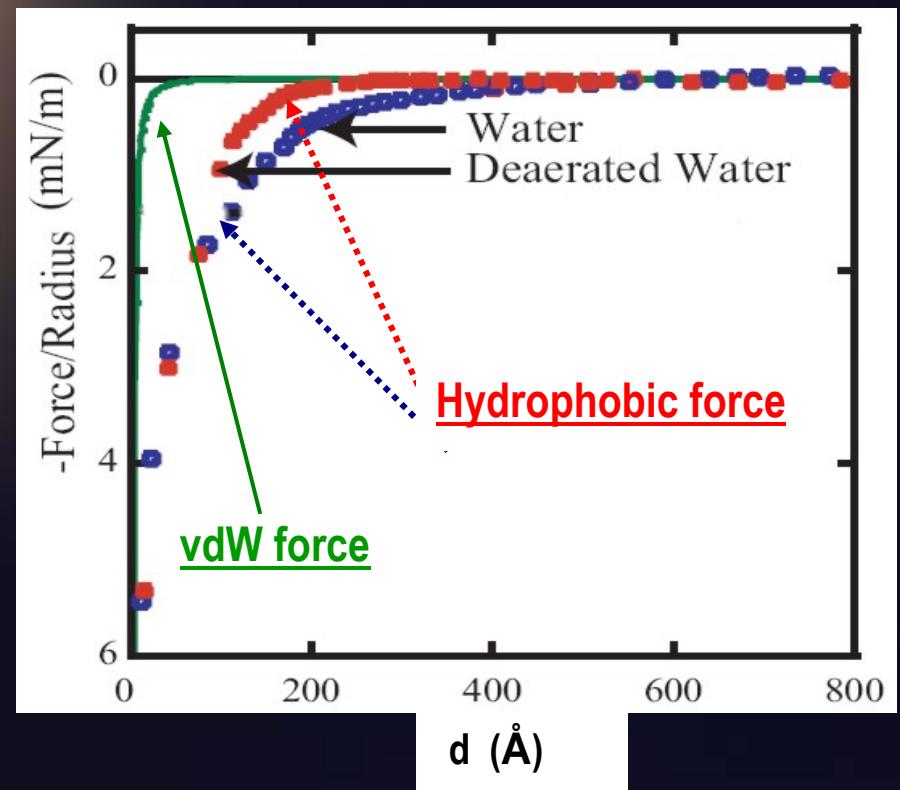
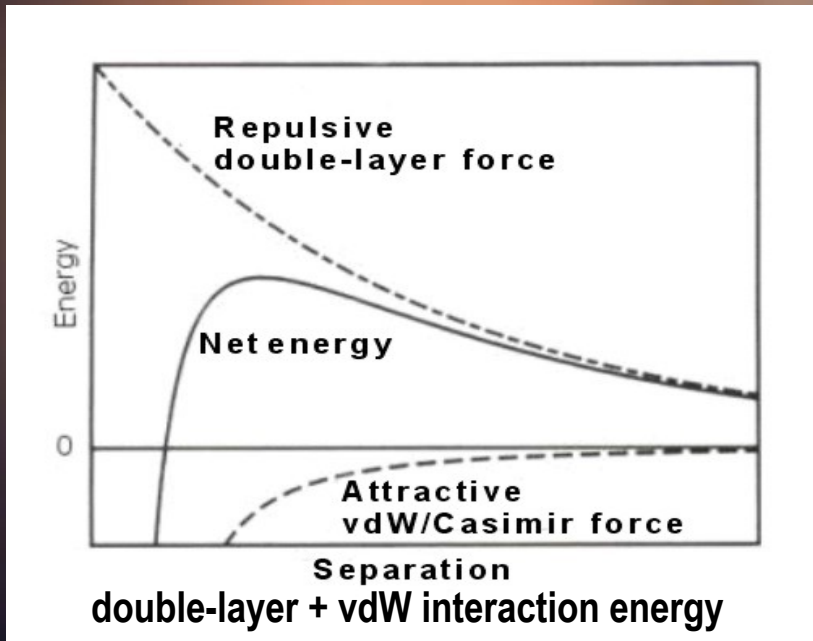
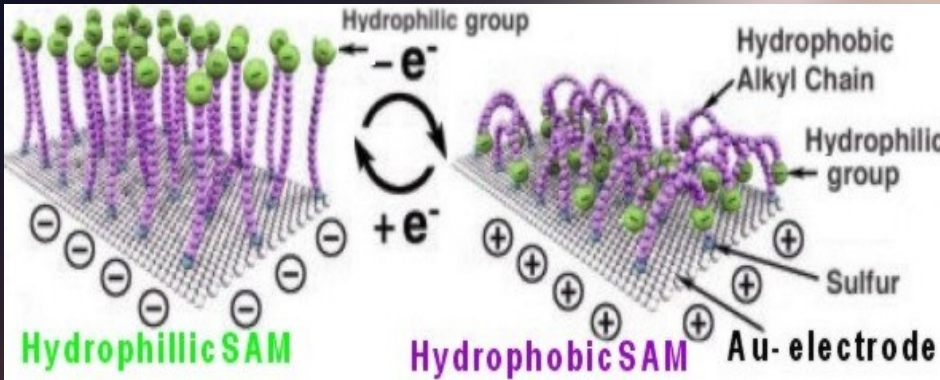
*F. Capasso et al., IEEE (2007)*

# Measurement in liquids



# •Control surface chemistry

- Electrostatic forces & double layers
- Hydrophobic interactions



# CONCLUSIONS

- → *Surface roughness & optical properties affect forces*
- → *Short range forces → do better estimation – Inverse AFM*
- → *Change optical properties → modulate forces*
- → *Understanding water layer influence & force measurement below 10 nm*
- → *Mesurements in liquids .....separations < 10 nm*
- → *For the future .....stll a lot to be understood for vacuum engines, dark energy etc... !!!*