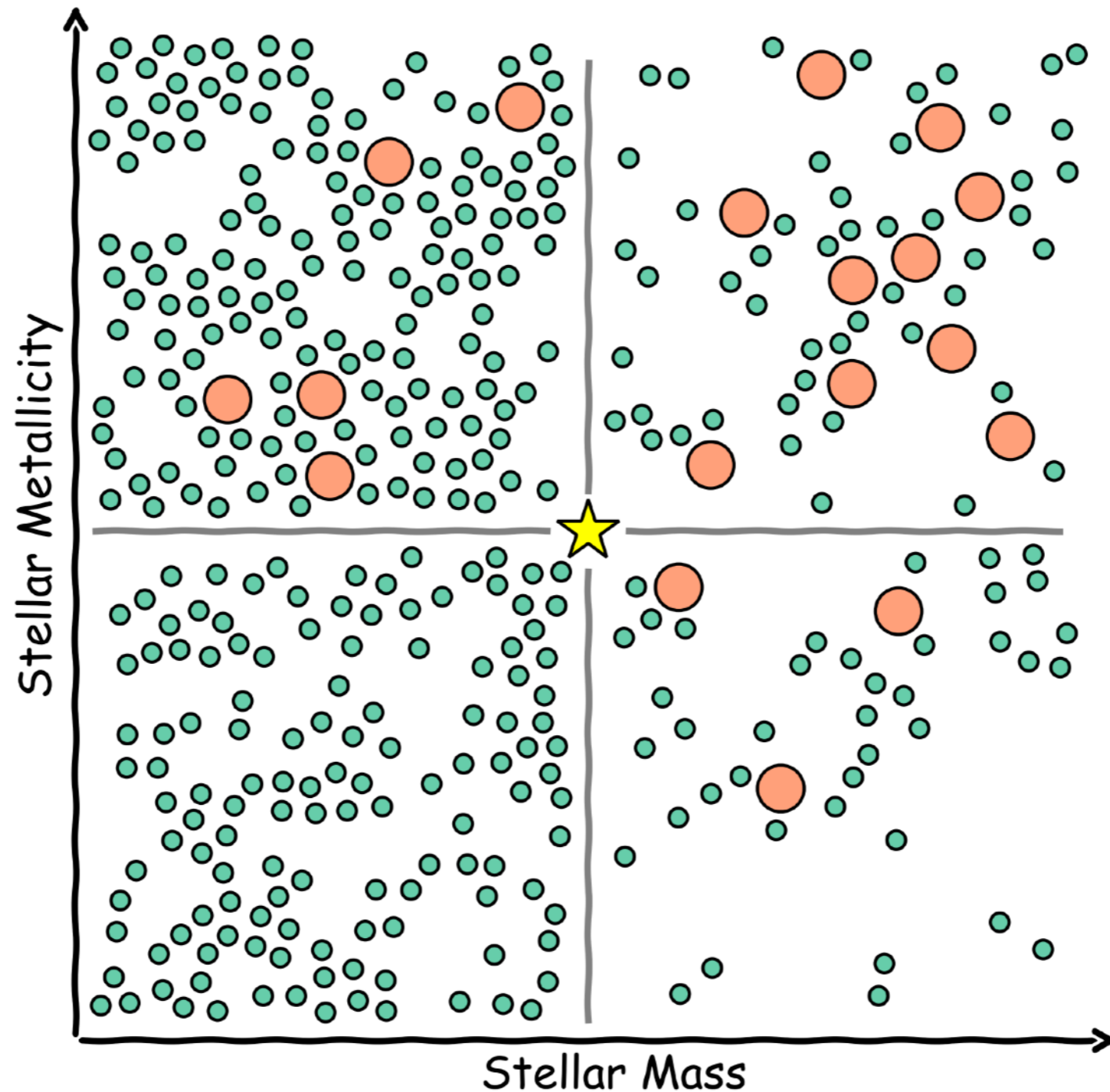


The exoplanet population around stars of different masses

Gijs Mulders
University of Chicago



The Kepler Exoplanet Population Around Different Types of Stars

Mulders 2018 (Handbook of Exoplanets)

Conclusion

**Planets occur more frequently
around low-mass stars...**

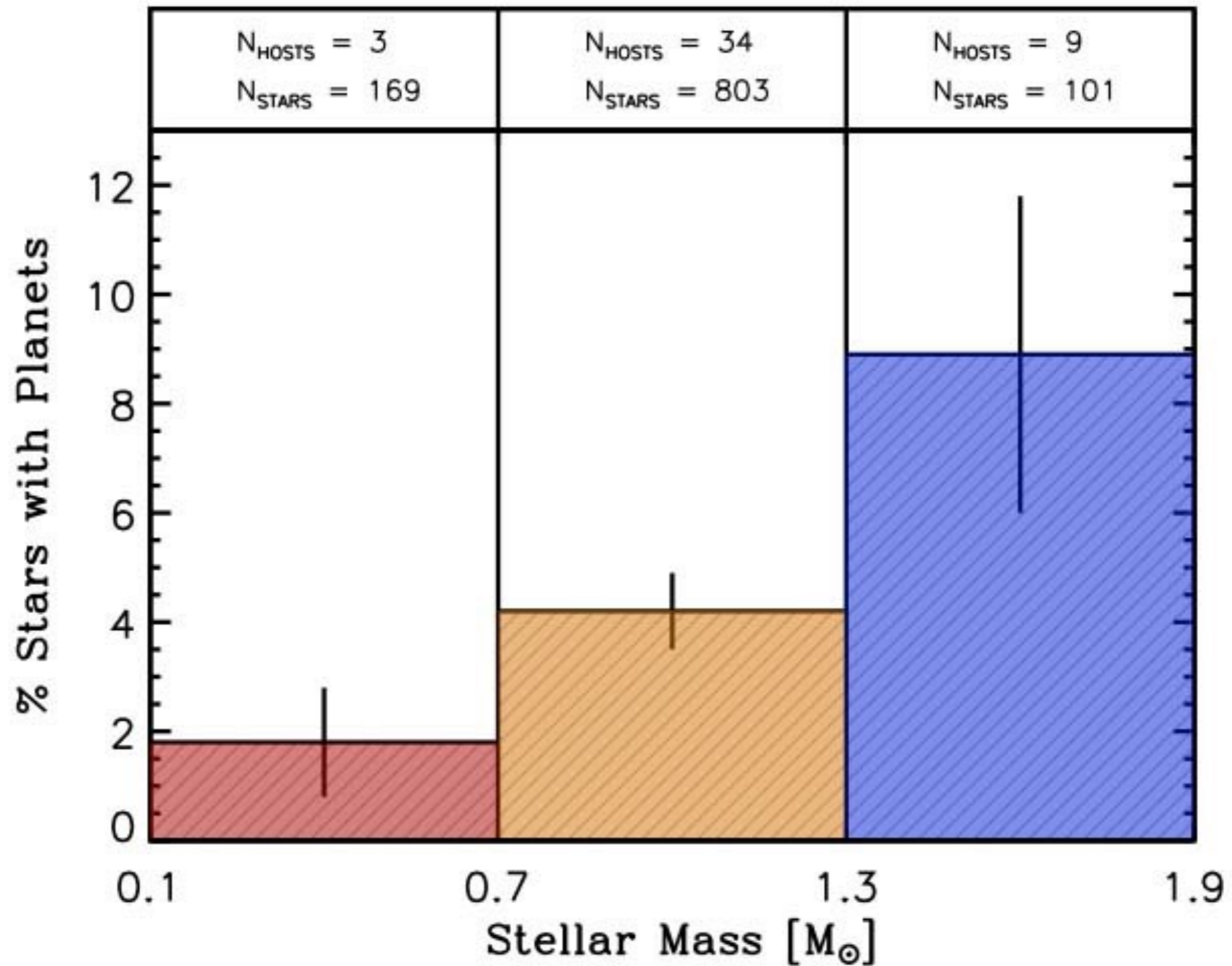
Conclusion

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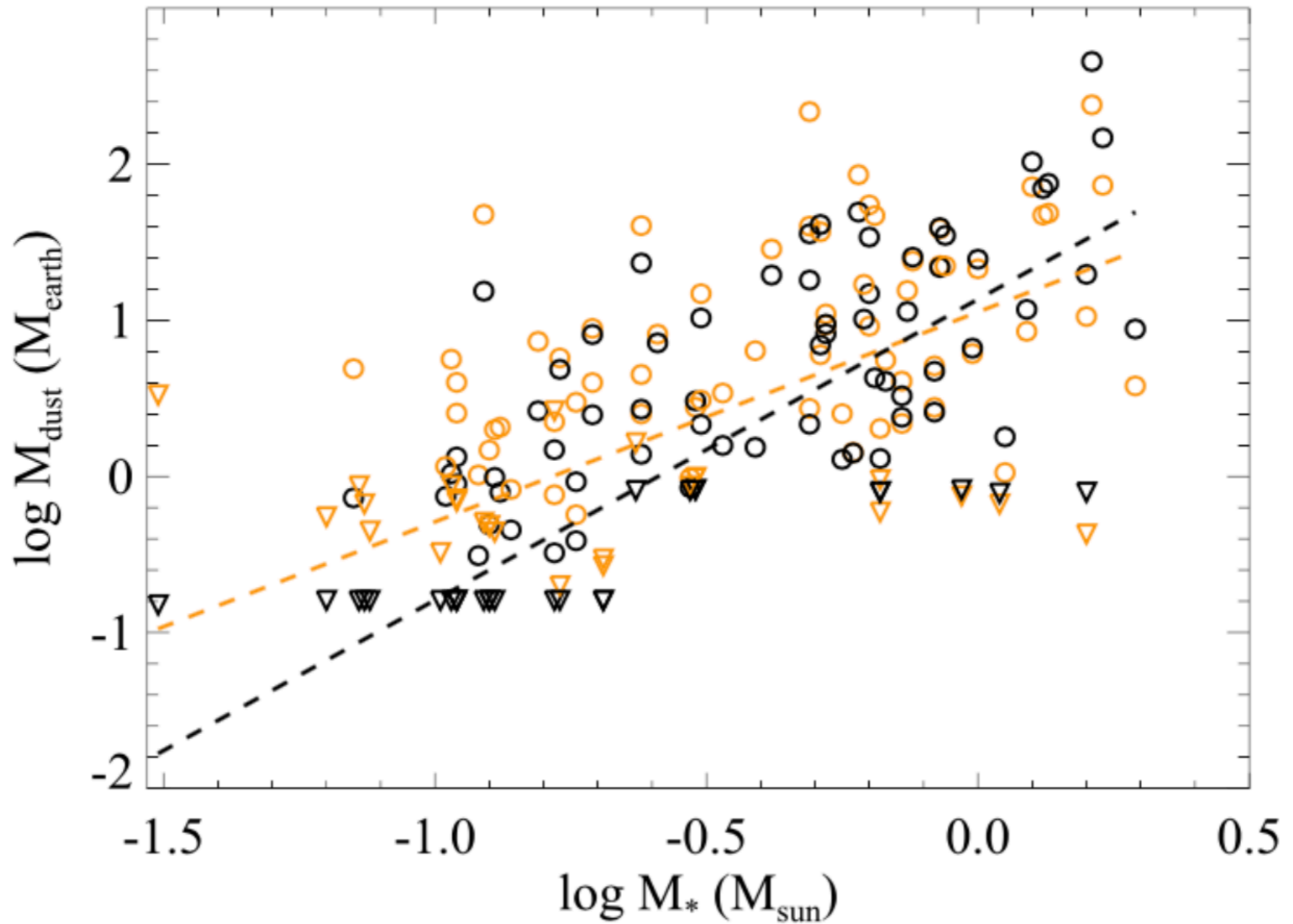
Discussion

**...but we do not understand the
implications for planet formation**

pre-Kepler



**Giant Planets Occurrence increases
with stellar mass**



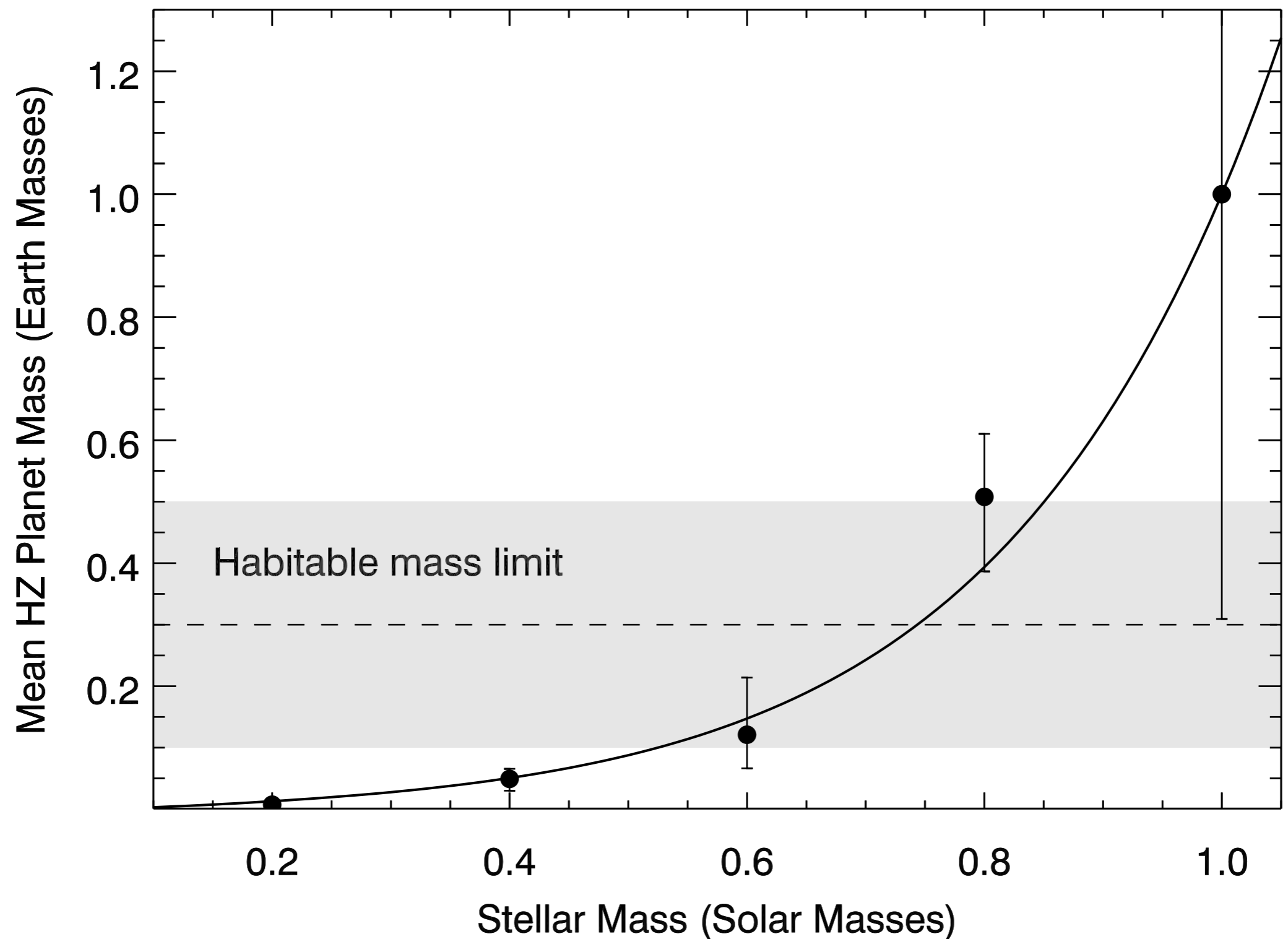
Disk Mass correlated with stellar mass

Star-Planet Connection

Star-Disk-Planet Connection



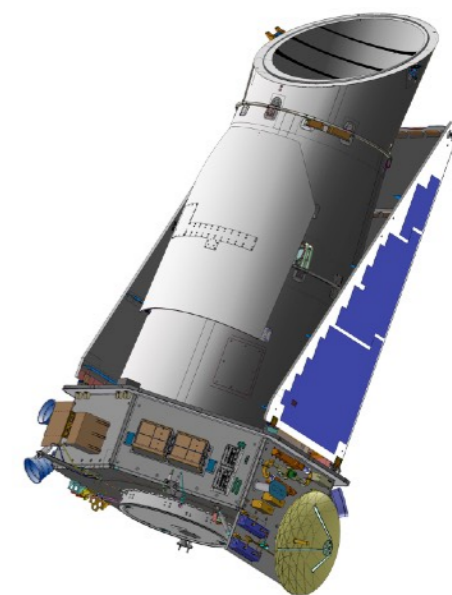
Star-Disk-Planet Connection

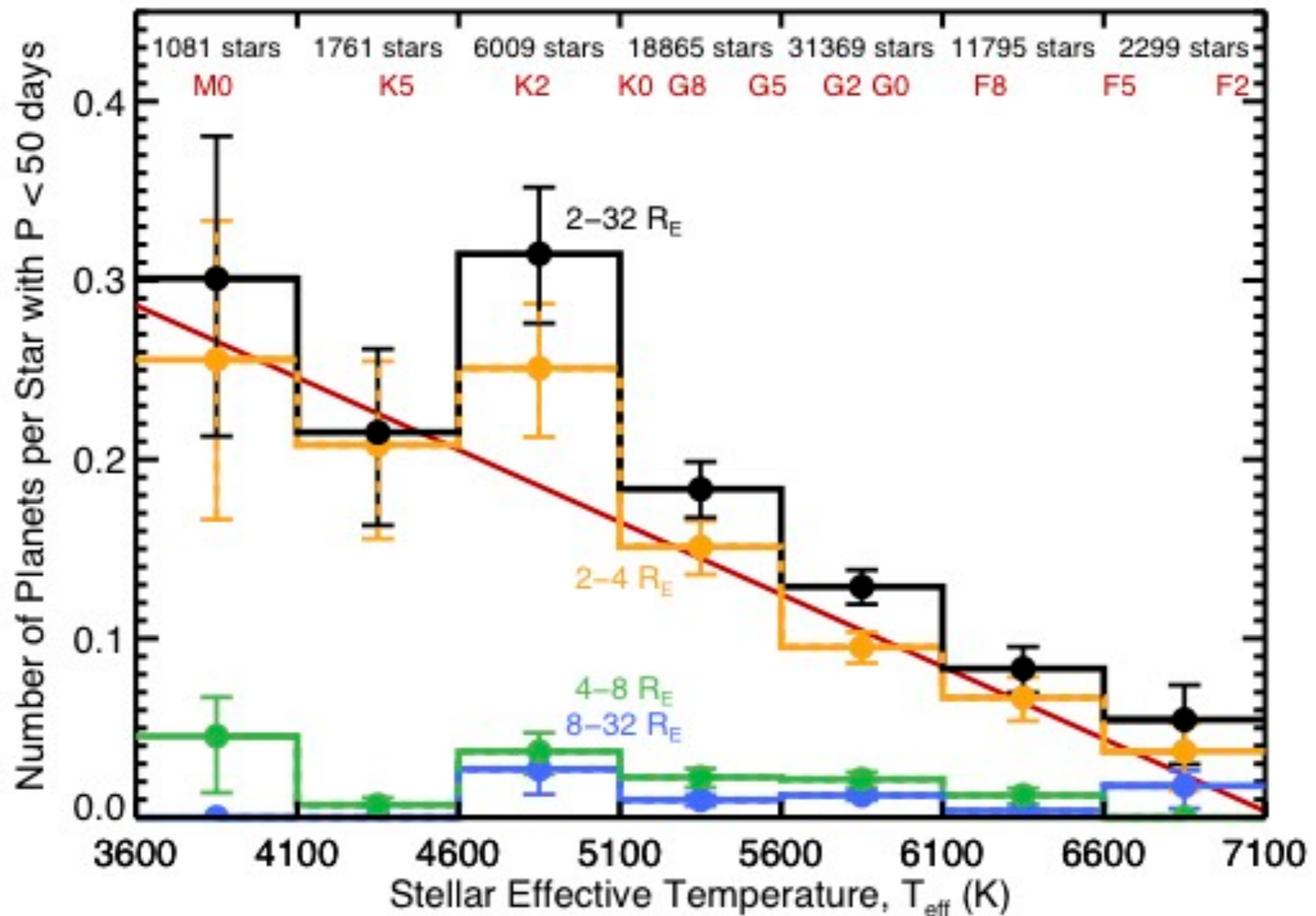


**Expectation: Habitable zone planets
around M stars are small and dry**

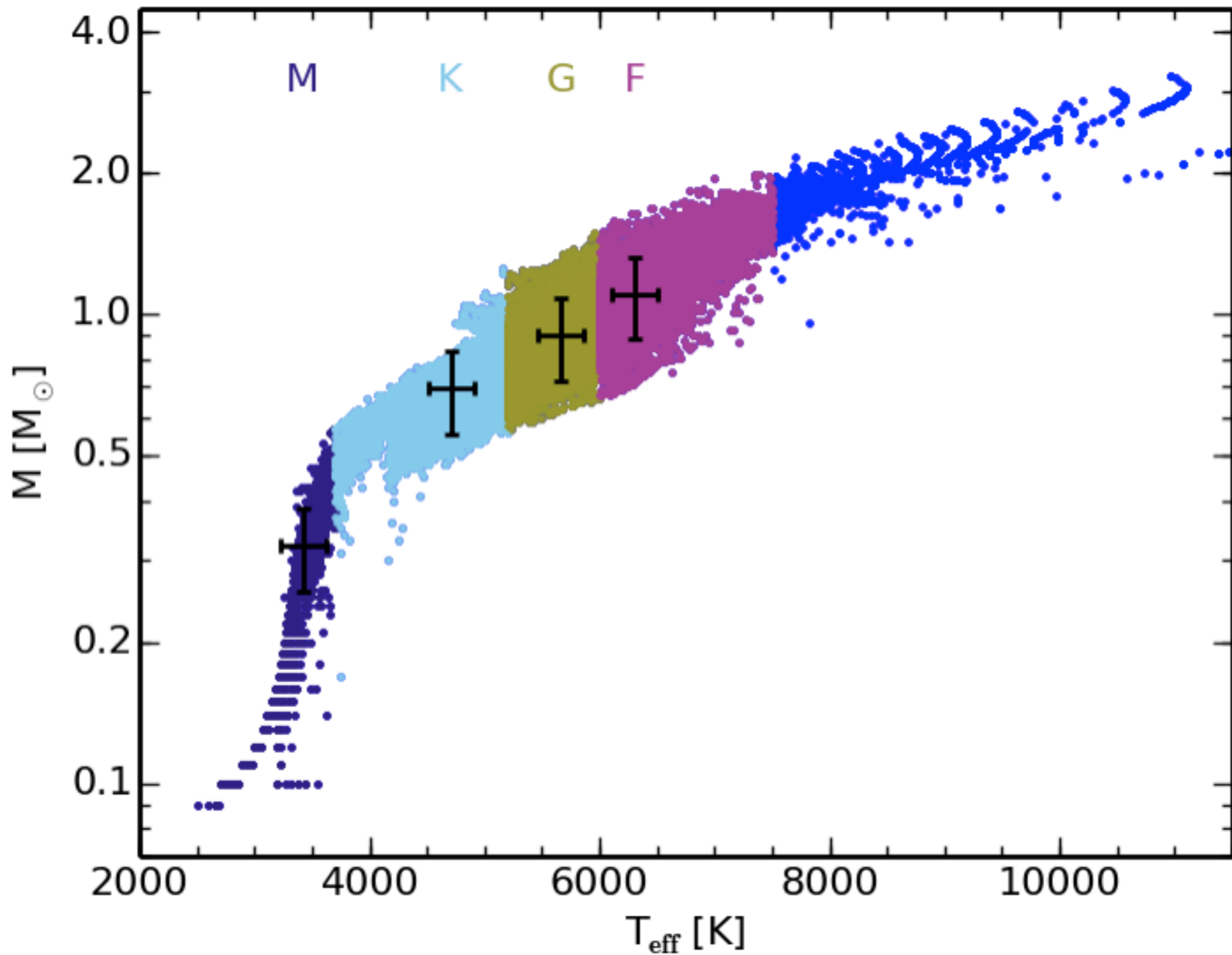
Raymond+ 2007, Lissauer 2007

Kepler



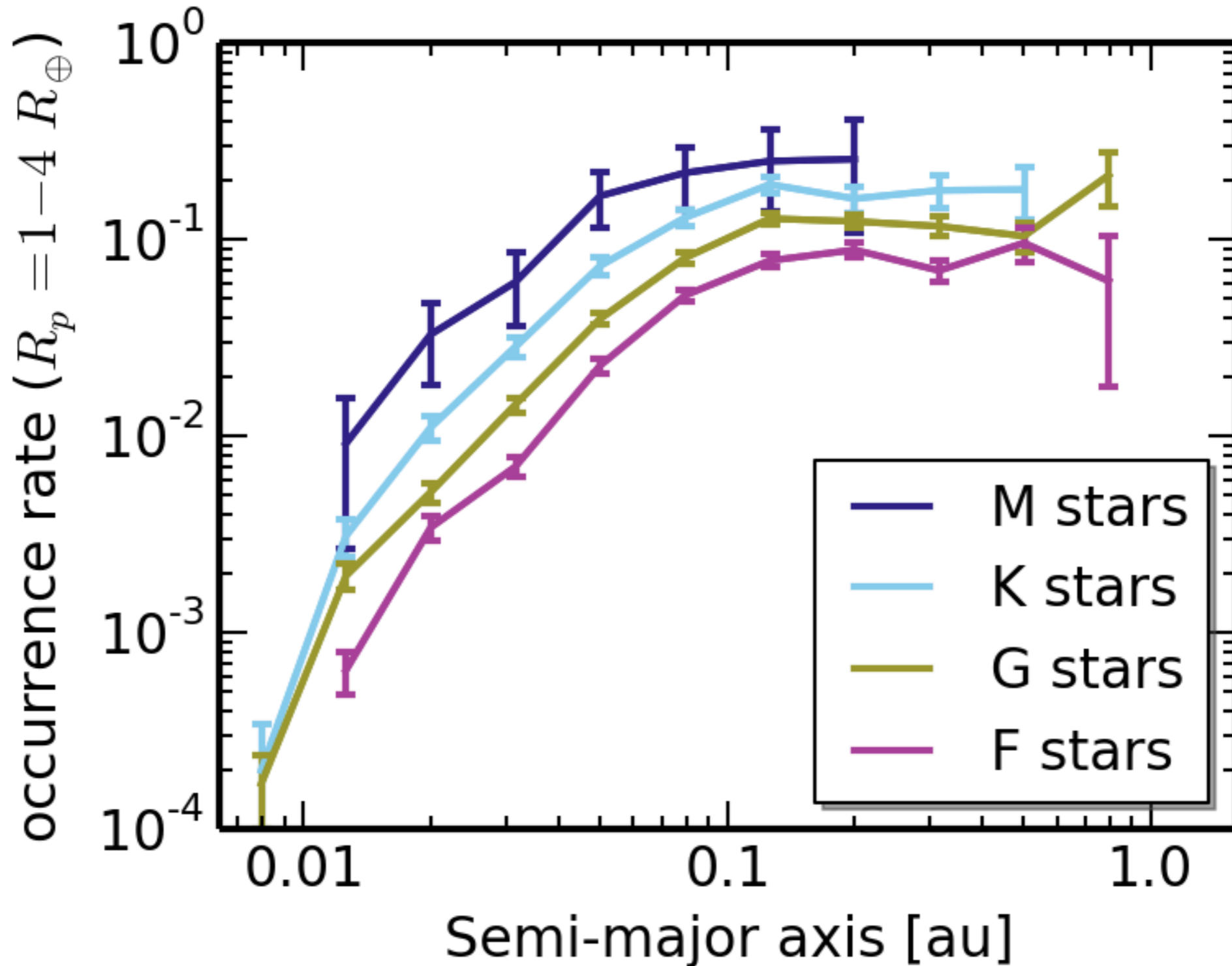


Planet occurrence decreases with effective temperature

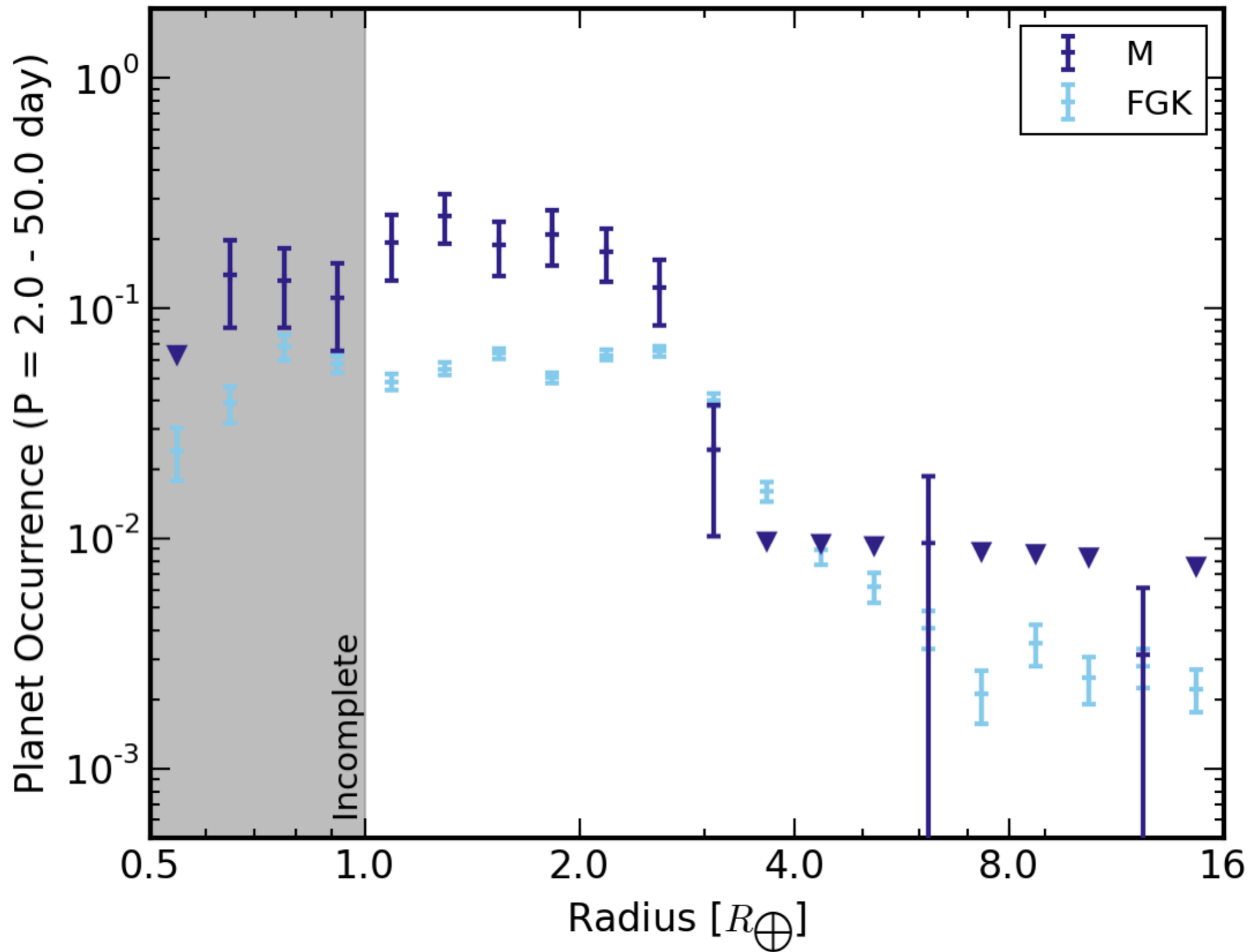


Spectral type as proxy for Stellar Mass

Huber+ 2014, Dressing & Charbonneau 2013



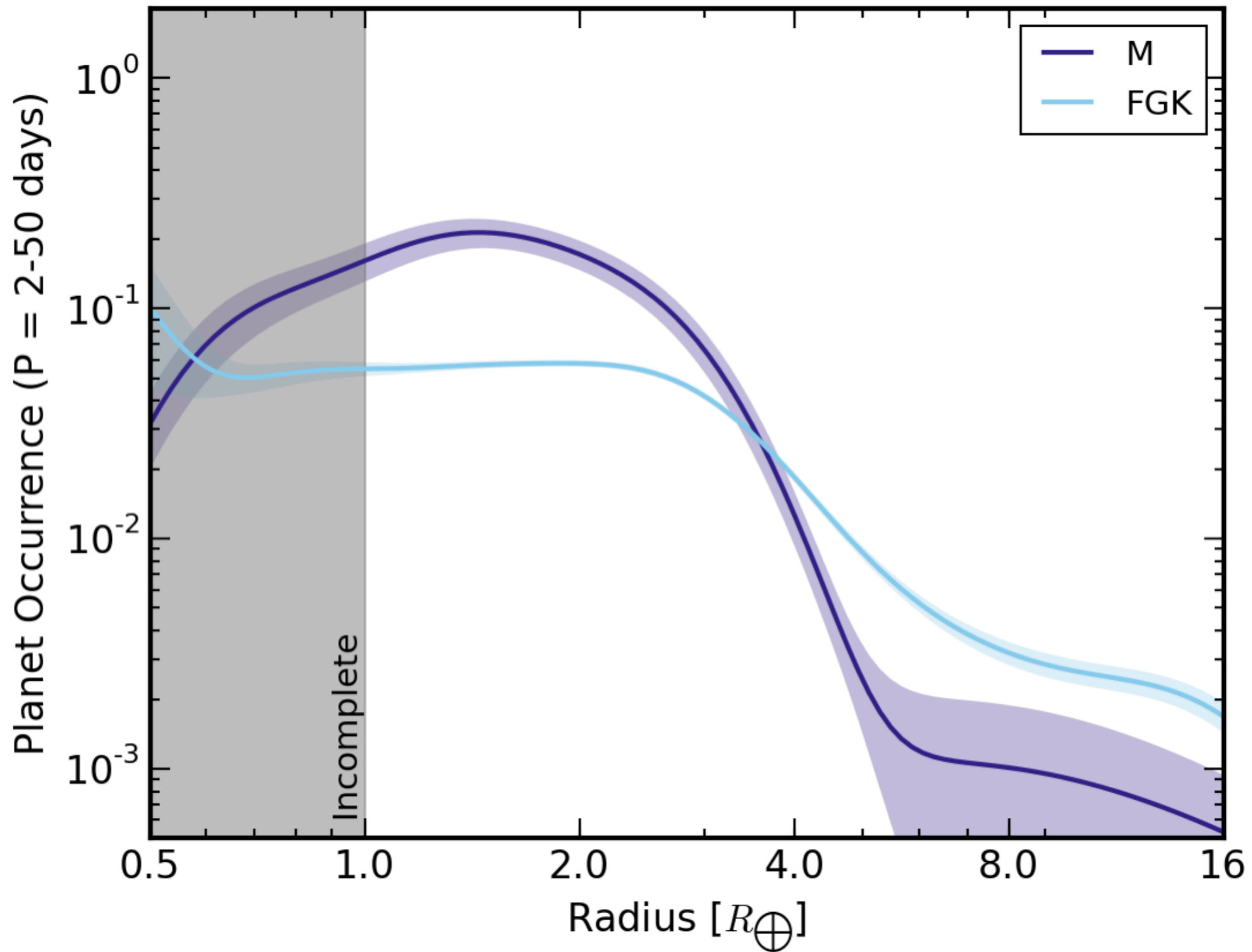
**Same spatial distribution:
Low-mass stars have more planets!**



Planet Radius Distribution

Low-mass stars have more sub-Neptunes!

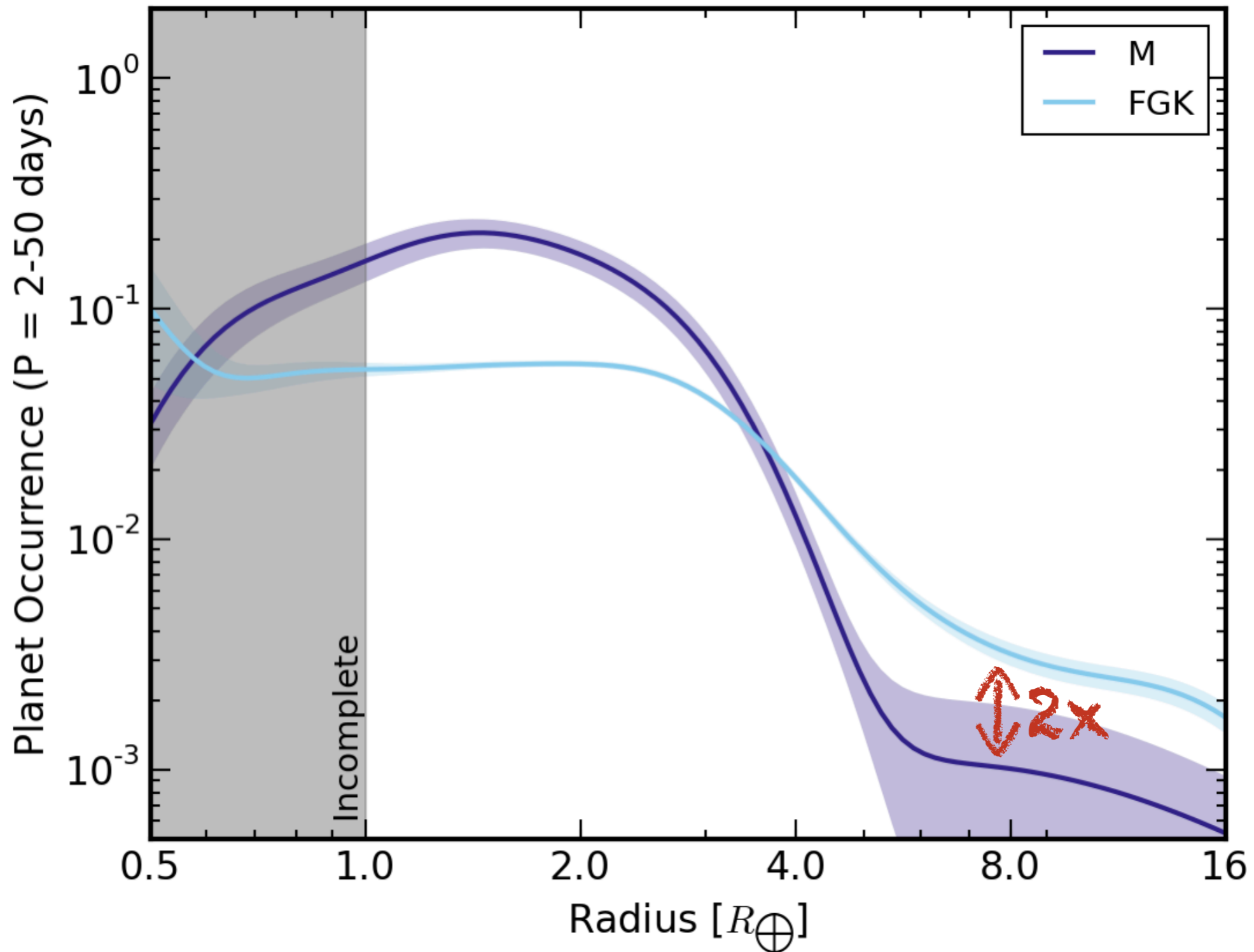
Mulders et al. 2015b



Planet Radius Distribution

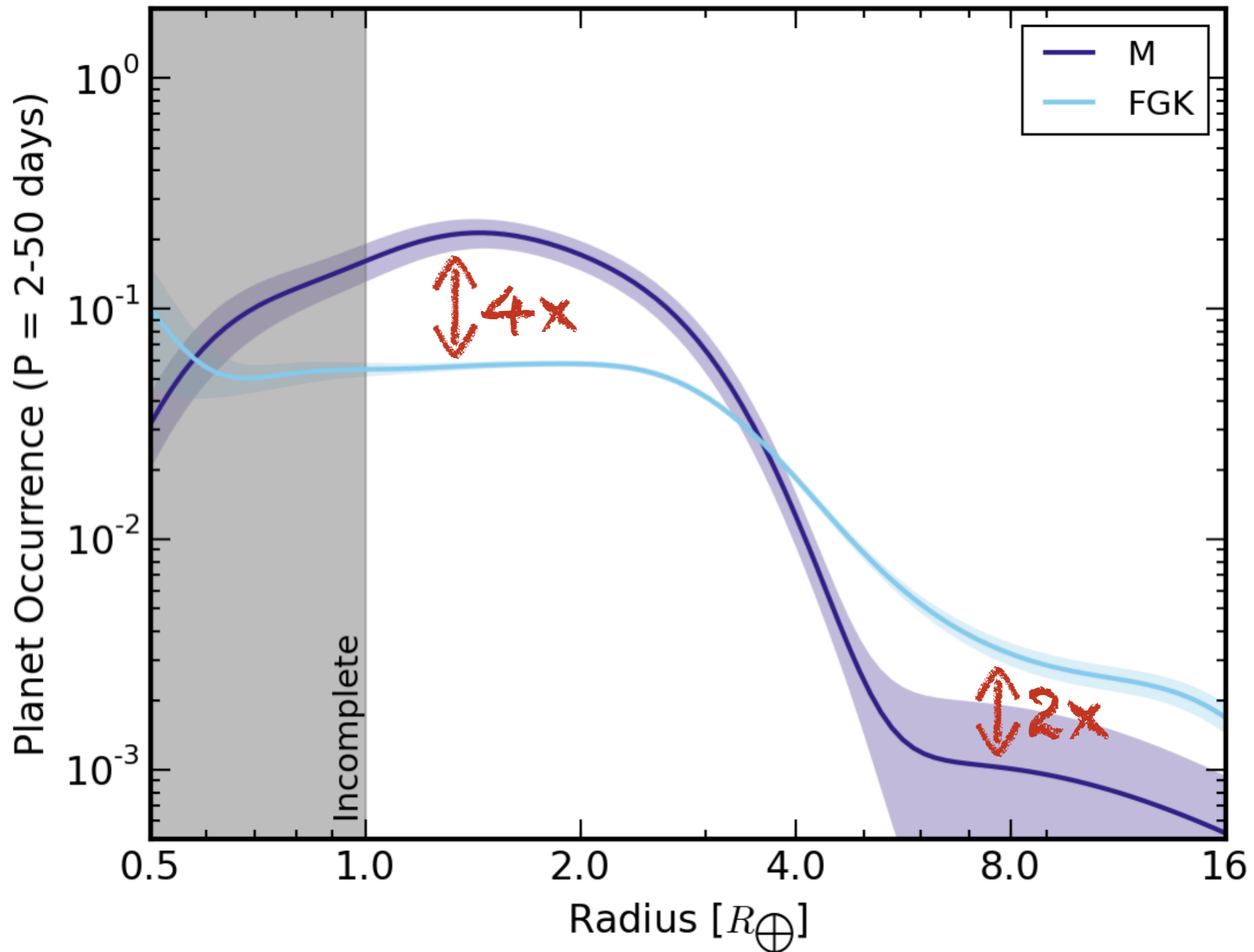
Low-mass stars have more sub-Neptunes!

Mulders et al. 2015b



Planet Radius Distribution

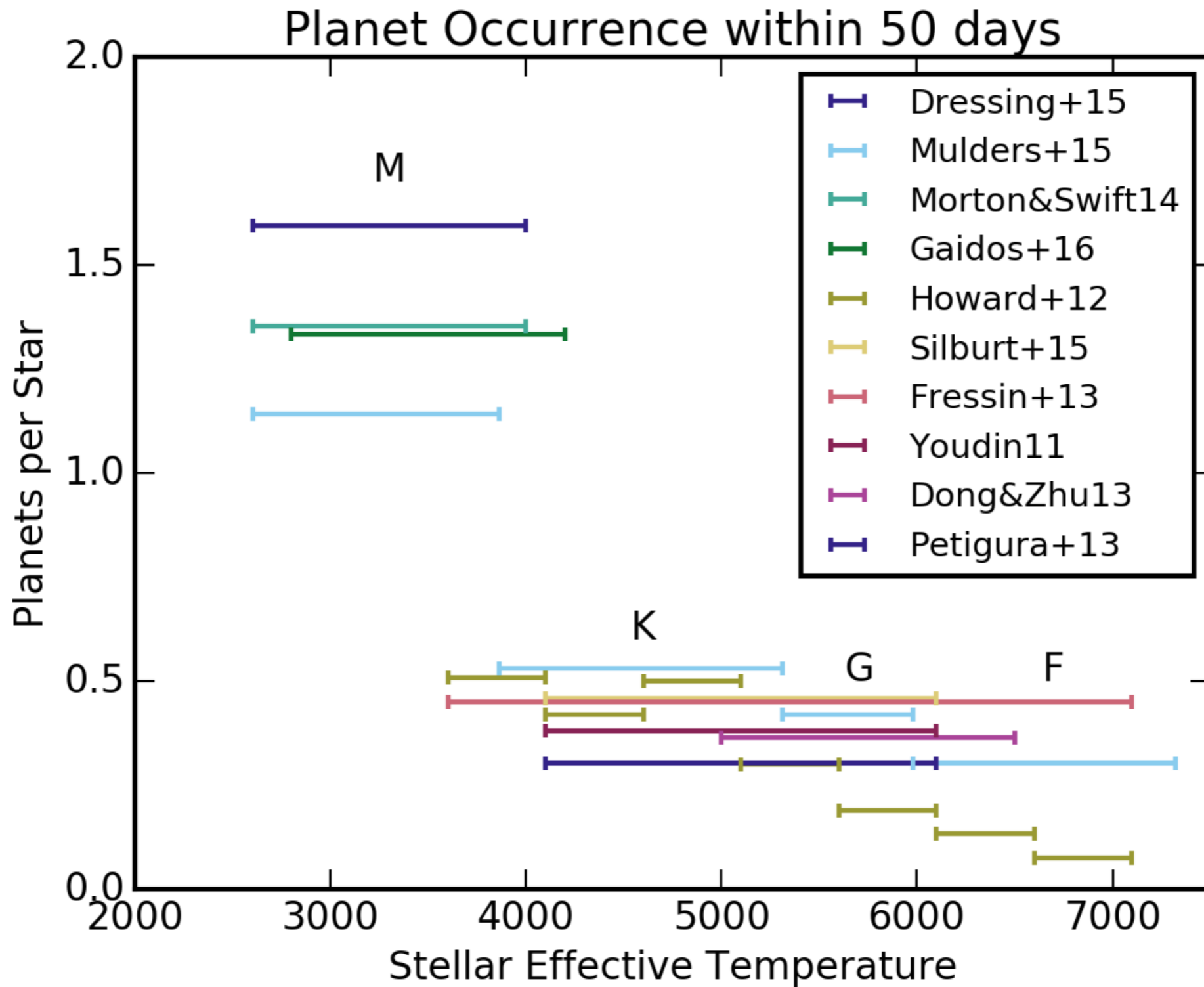
Low-mass stars have more sub-Neptunes!



Planet Radius Distribution

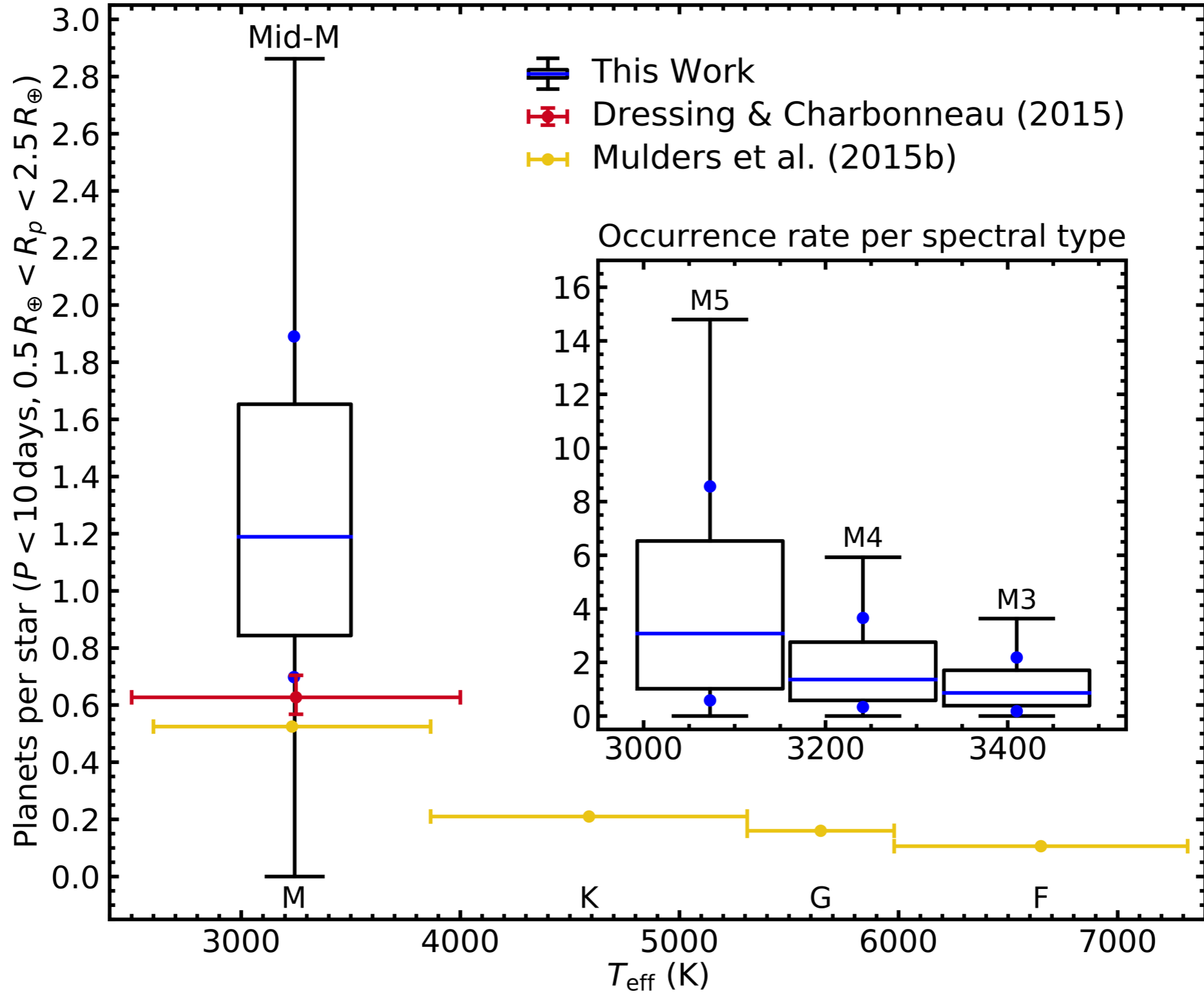
Low-mass stars have more sub-Neptunes!

Mulders et al. 2015b



Low-mass stars have more sub-Neptunes

Mulders 2018, Handbook of Exoplanets

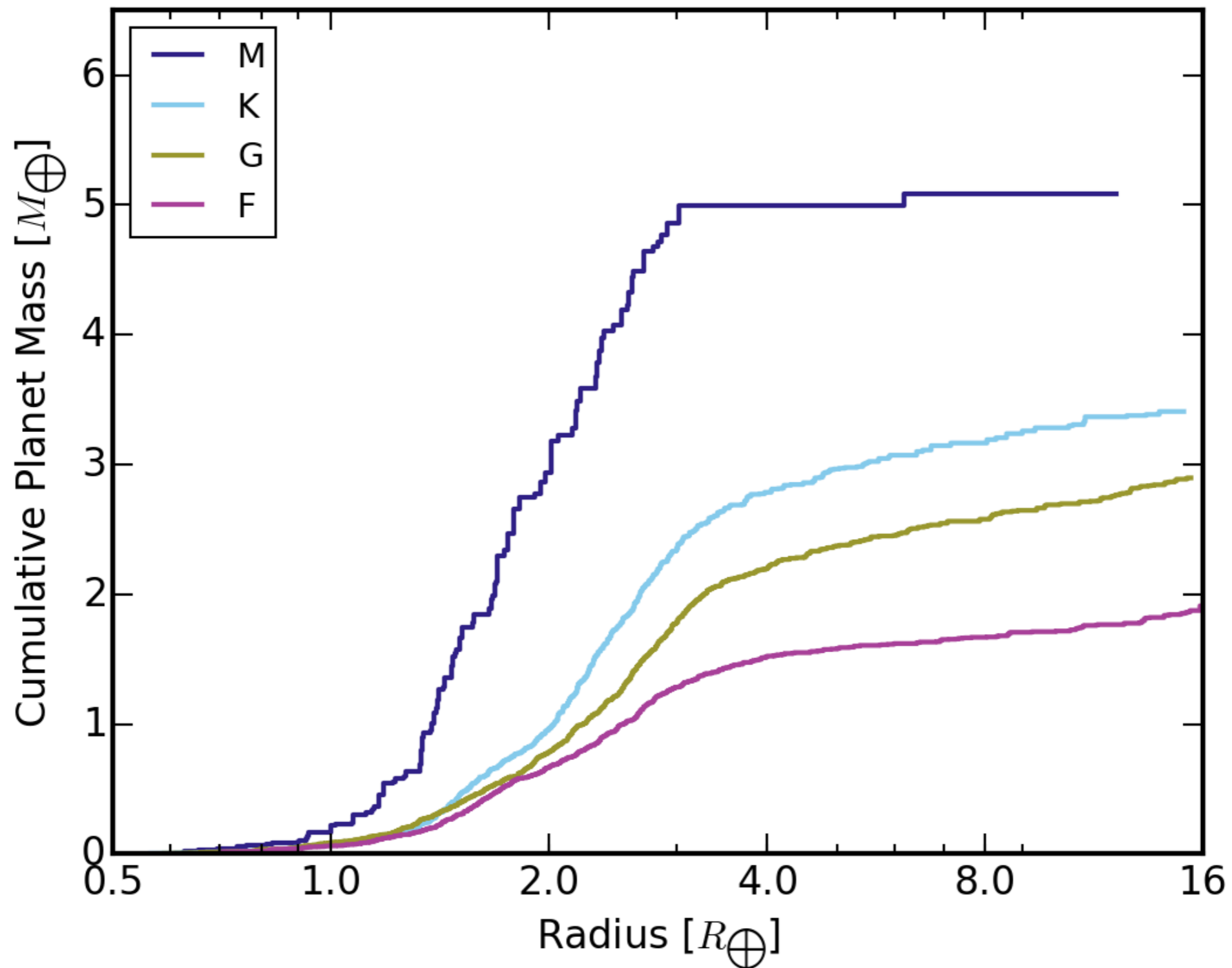


Trend continues for mid-M dwarfs

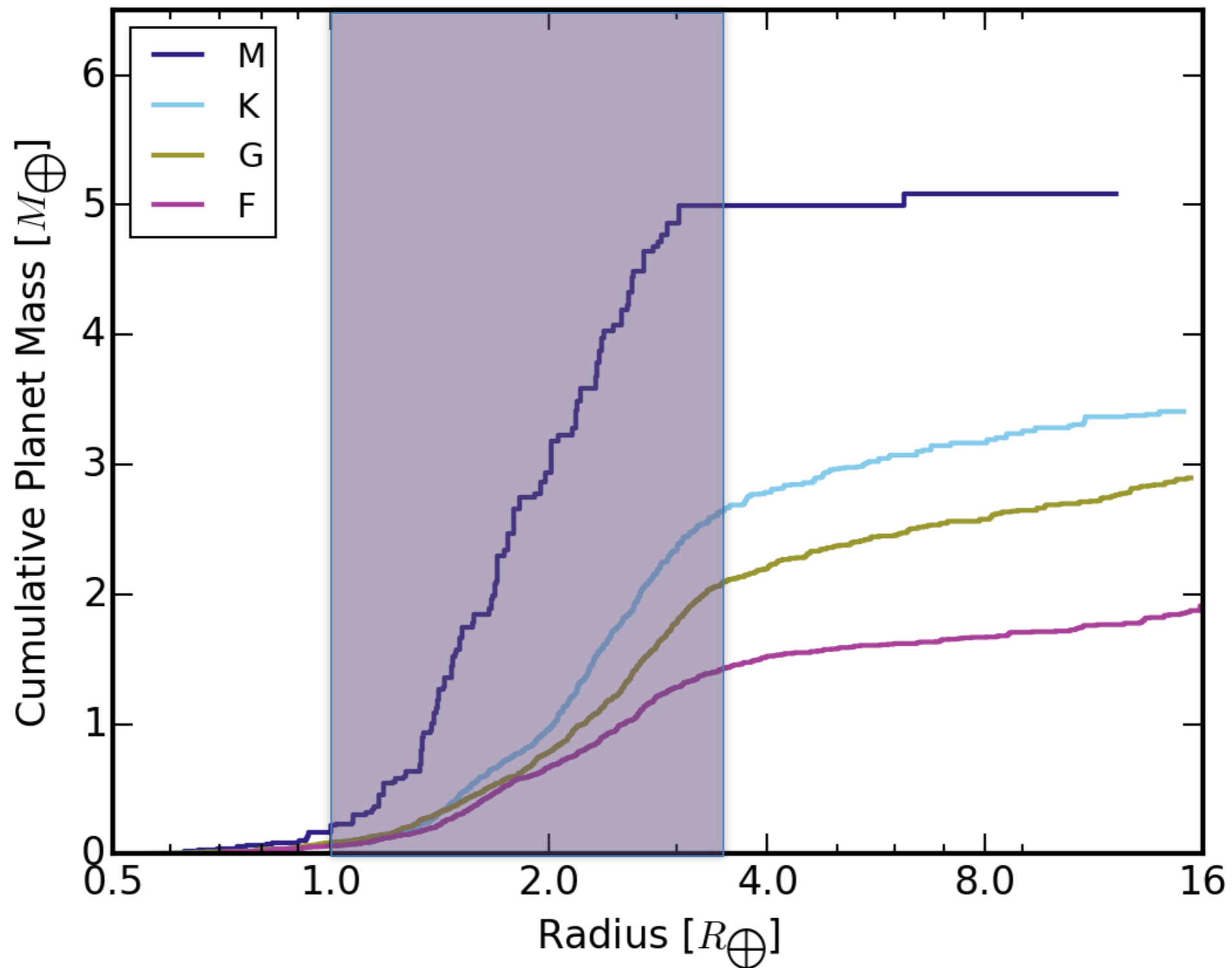
Why?

How to explain elevated occurrence rate for M dwarfs planets

- Detection biases **X**
- Spatial or size distribution **X**
- Trade-off with giant planets?
- Binaries?
- Planet Formation?

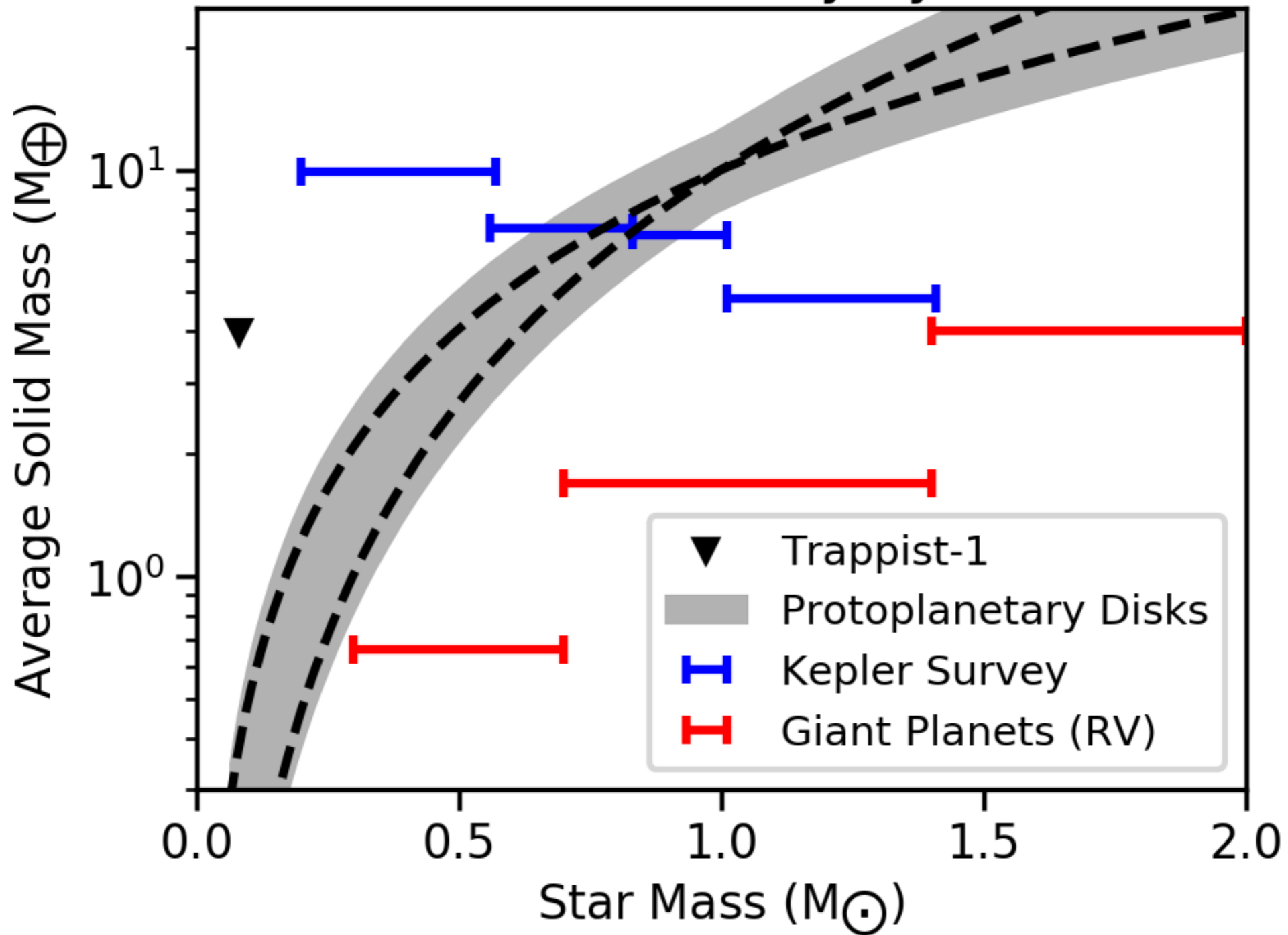


**Heavy Elements Mass, $P < 50$ days
(mass-radius relation)**



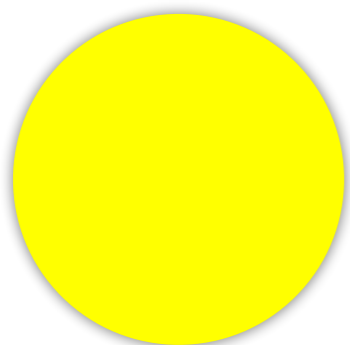
Heavy Elements Mass, $P < 50$ days (mass-radius relation)

Solids in Planetary Systems



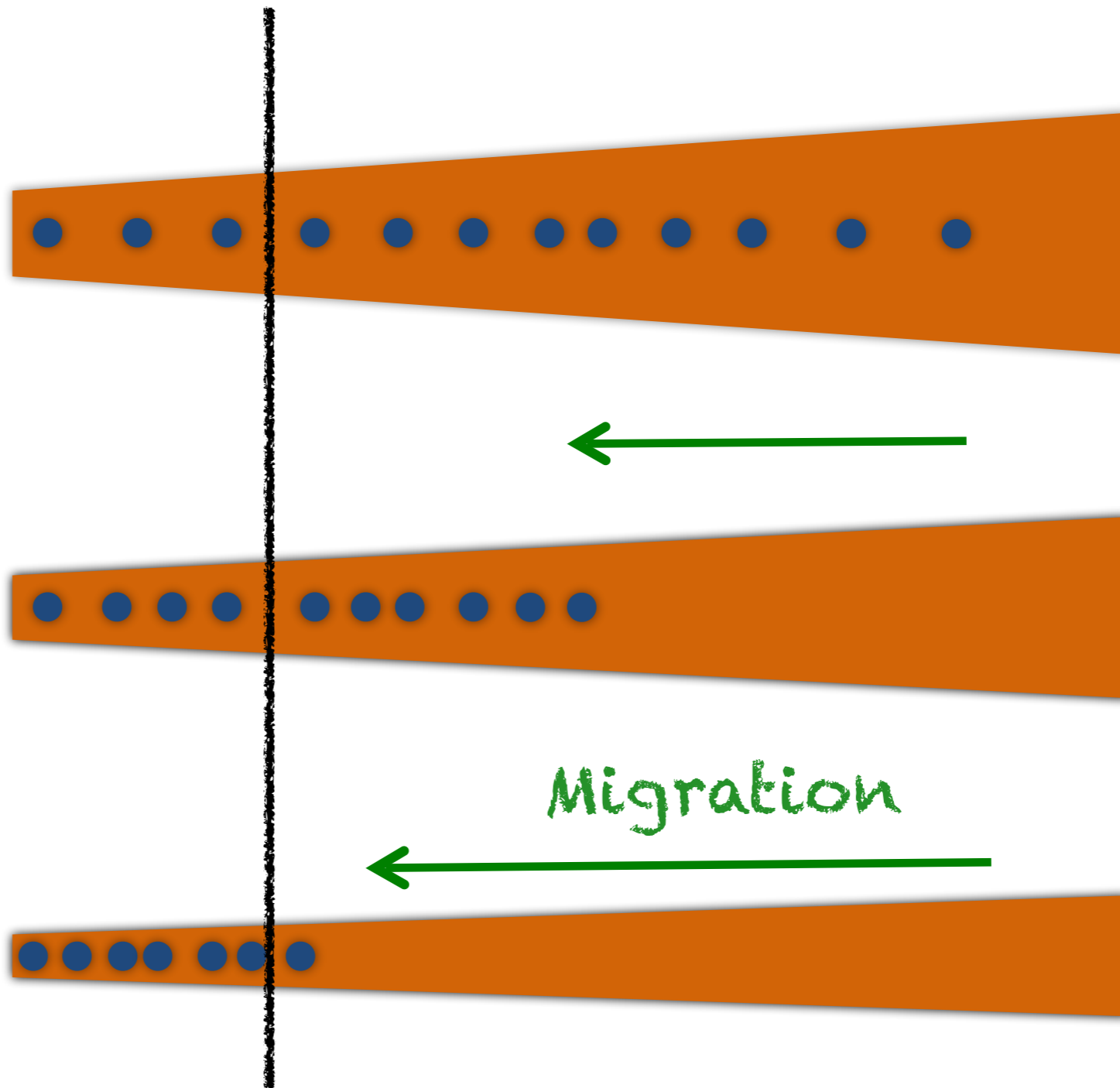
Comparison with disk solids

Sunlike
Star

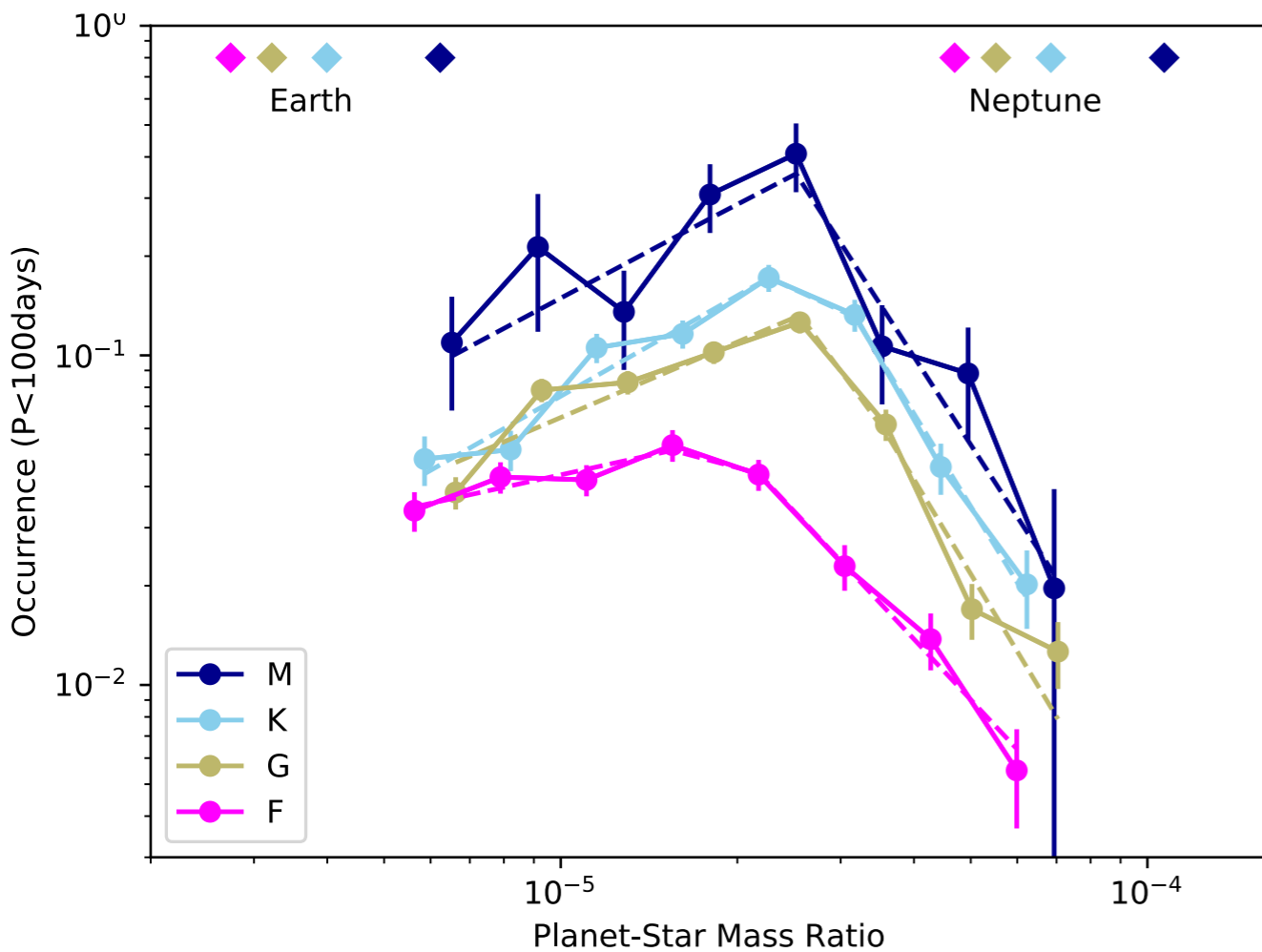


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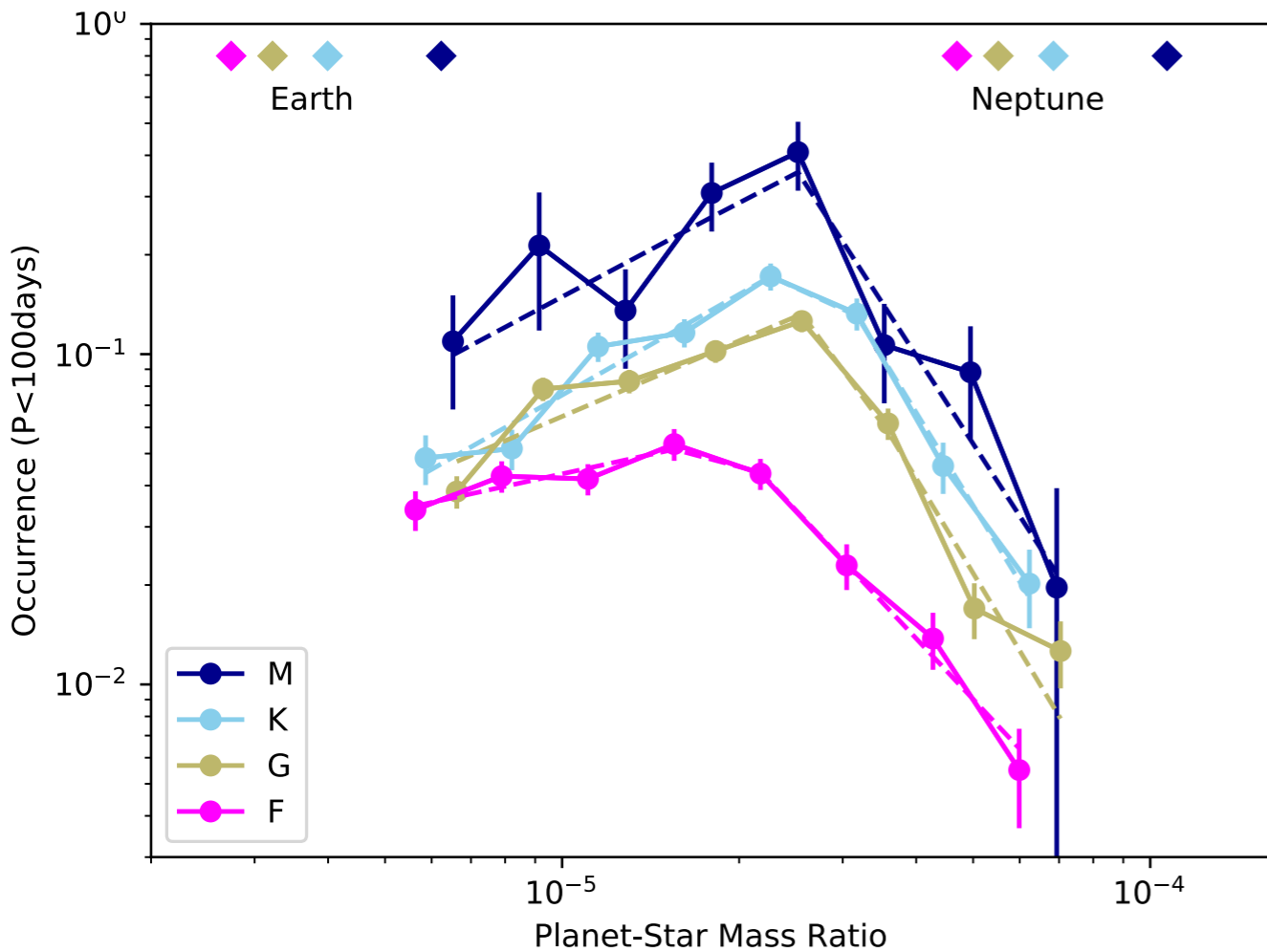
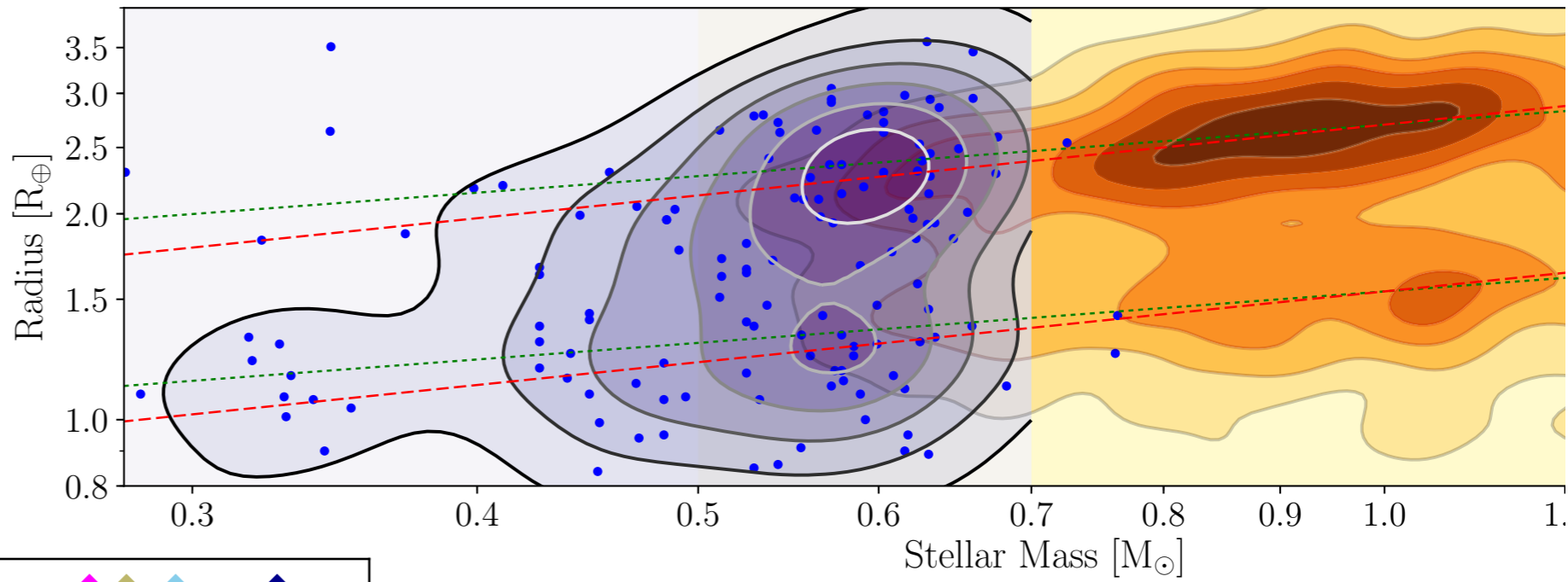
Protoplanetary Disk



Red
Dwarf



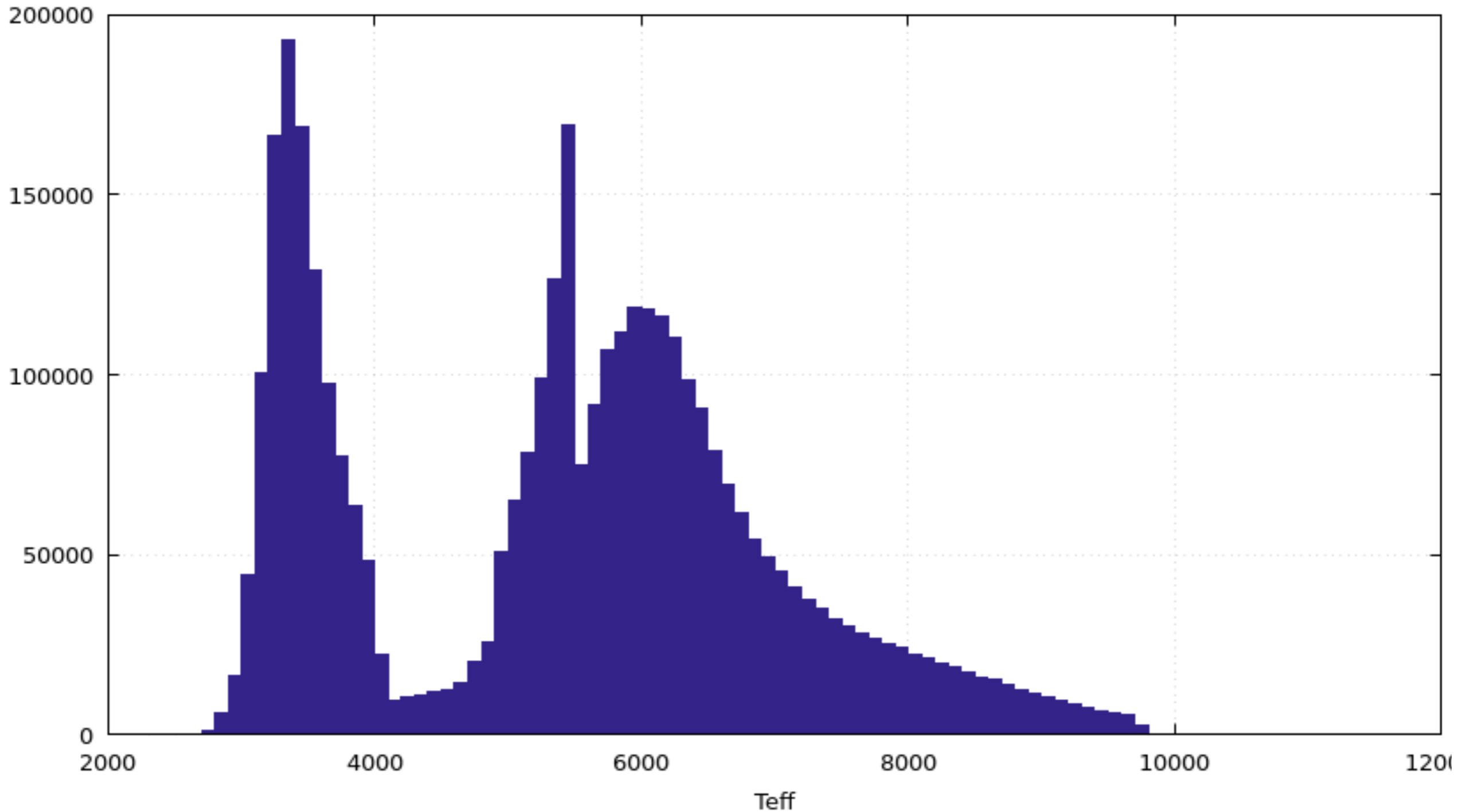
Planet Mass Dependence?



Planet Mass Dependence?

post-Kepler

TESS stars



K2 / TESS

Homogenous planet search & detection efficiency

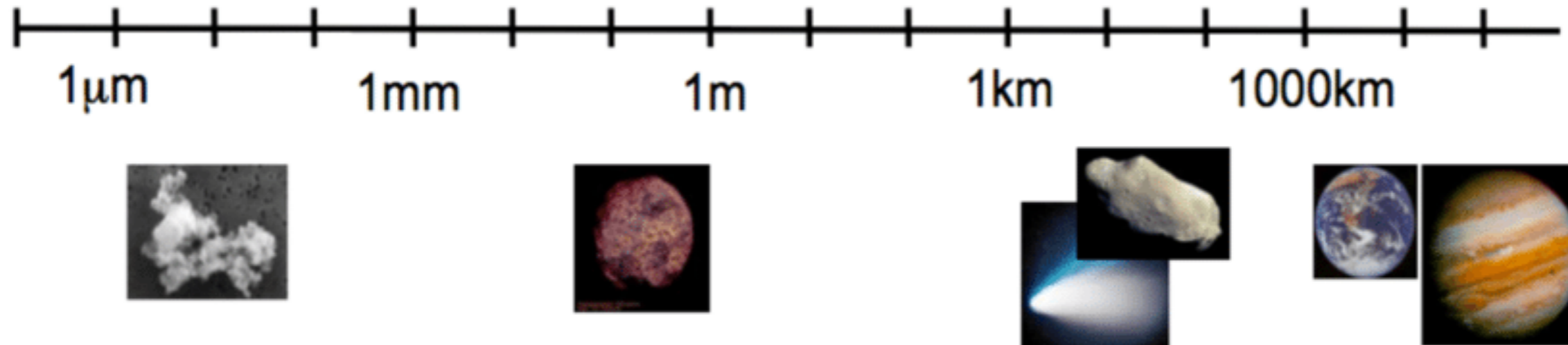
Protoplanetary Disk

Exoplanet

**Which planet formation processes has
right stellar mass dependence?**

image credit: Kees Dullemond

Protoplanetary Disk



Exoplanet

Which planet formation processes has right stellar mass dependence?

image credit: Kees Dullemond

Protoplanetary Disk

Coagulation

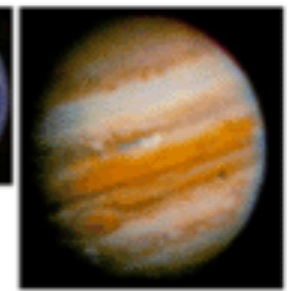
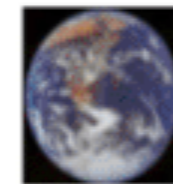
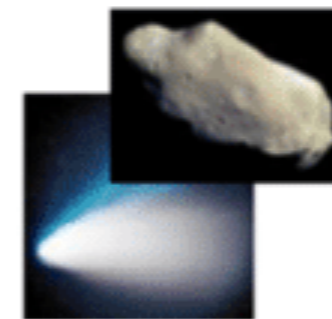
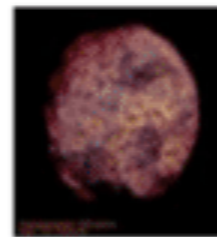
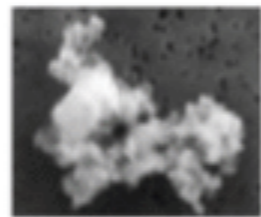
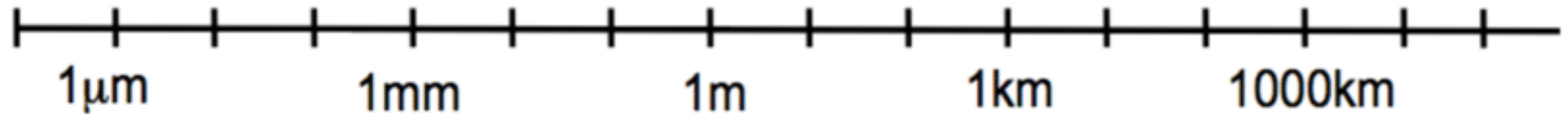
Streaming Instability

Gravitational Collisions

Radial Drift

Pebble Accretion

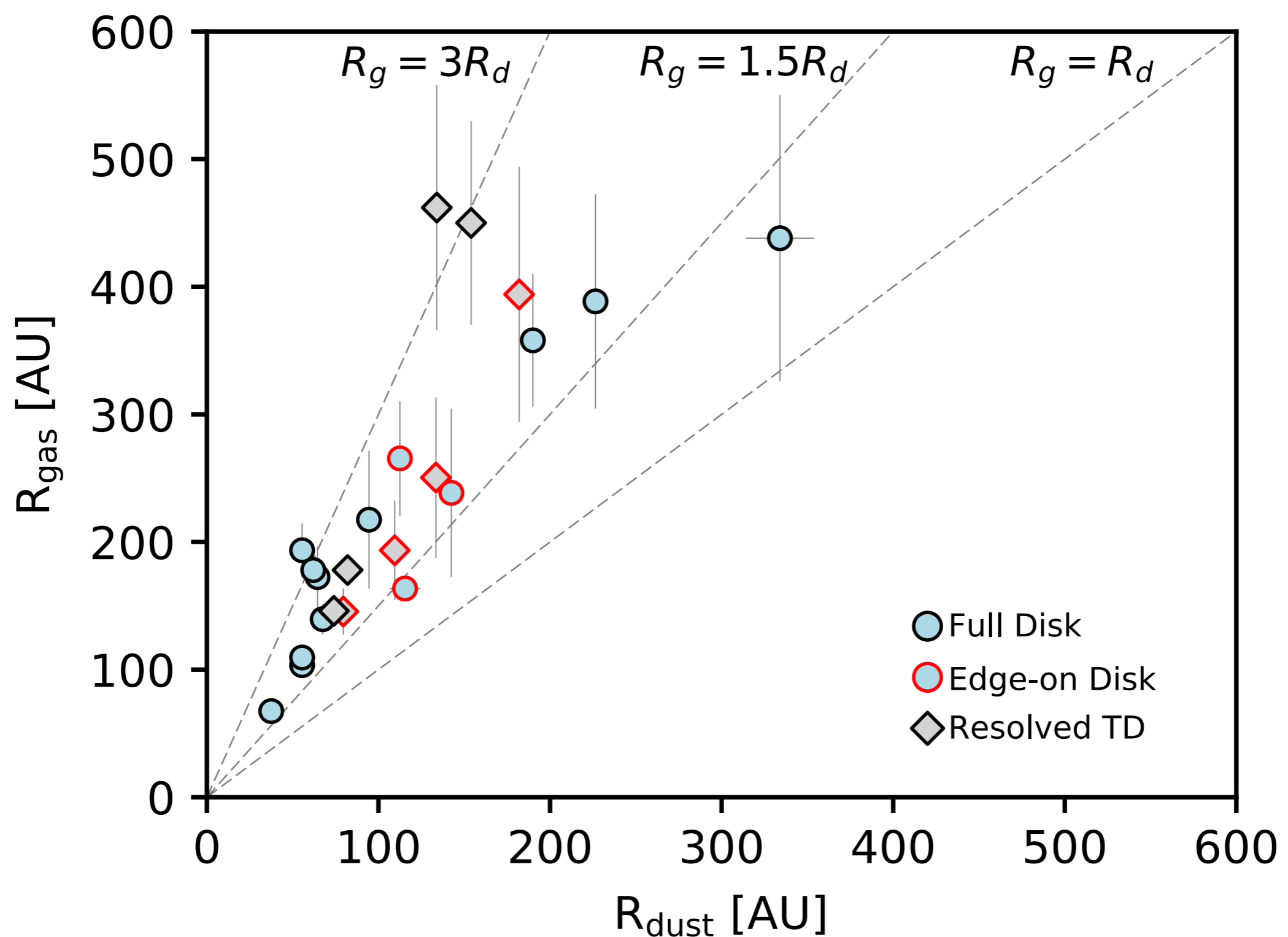
Planet Migration



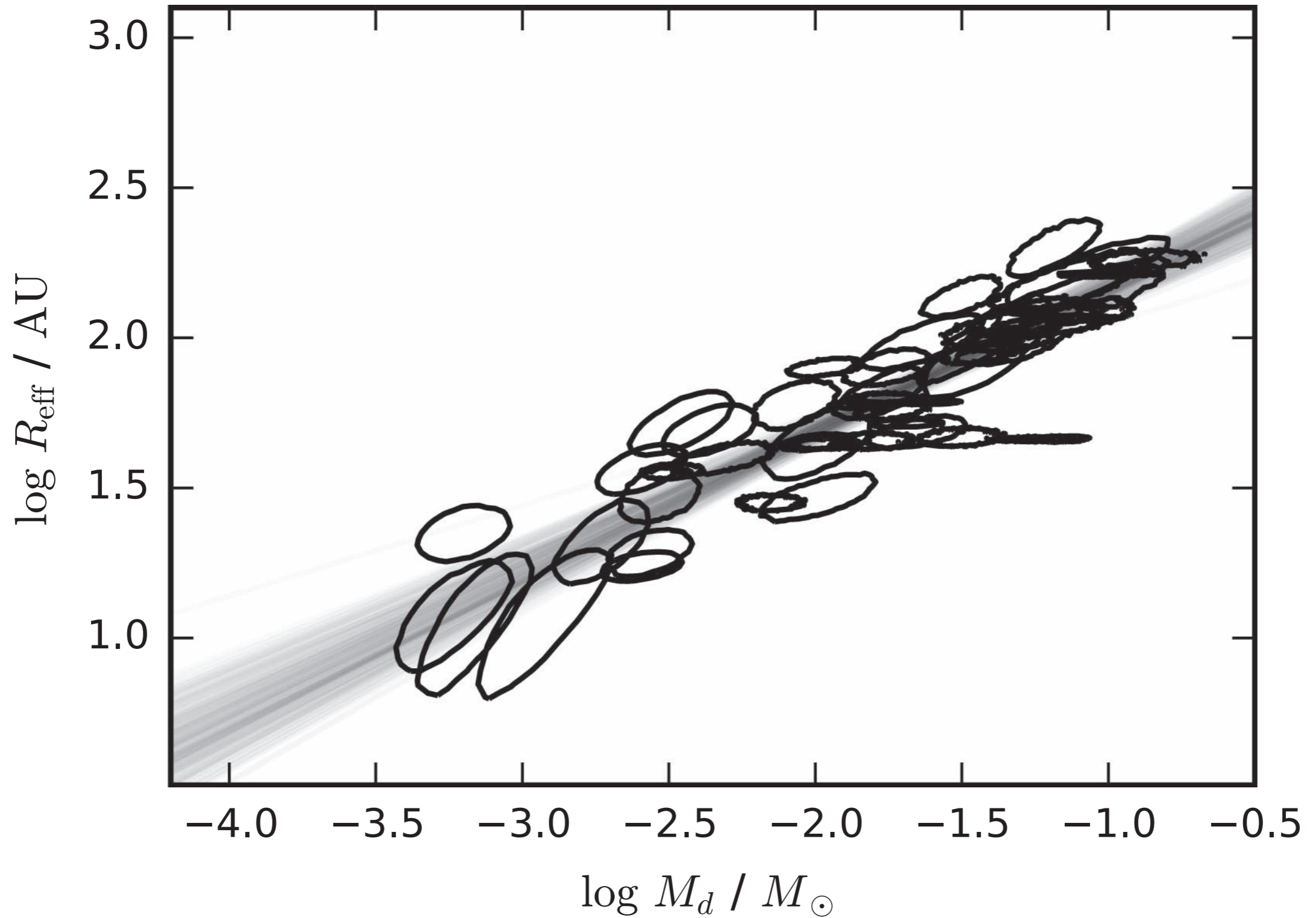
Exoplanet

Which planet formation processes has right stellar mass dependence?

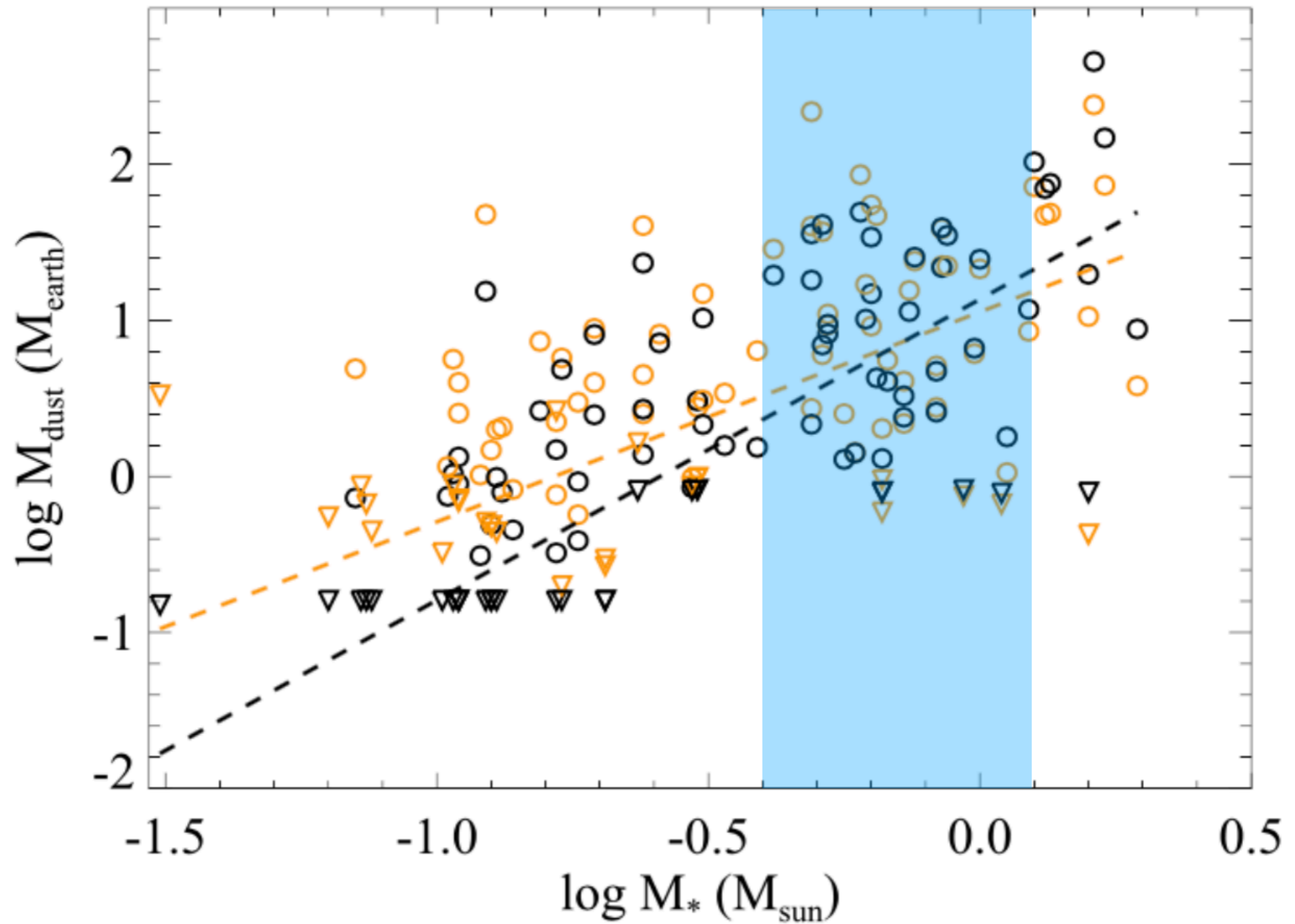
image credit: Kees Dullemond



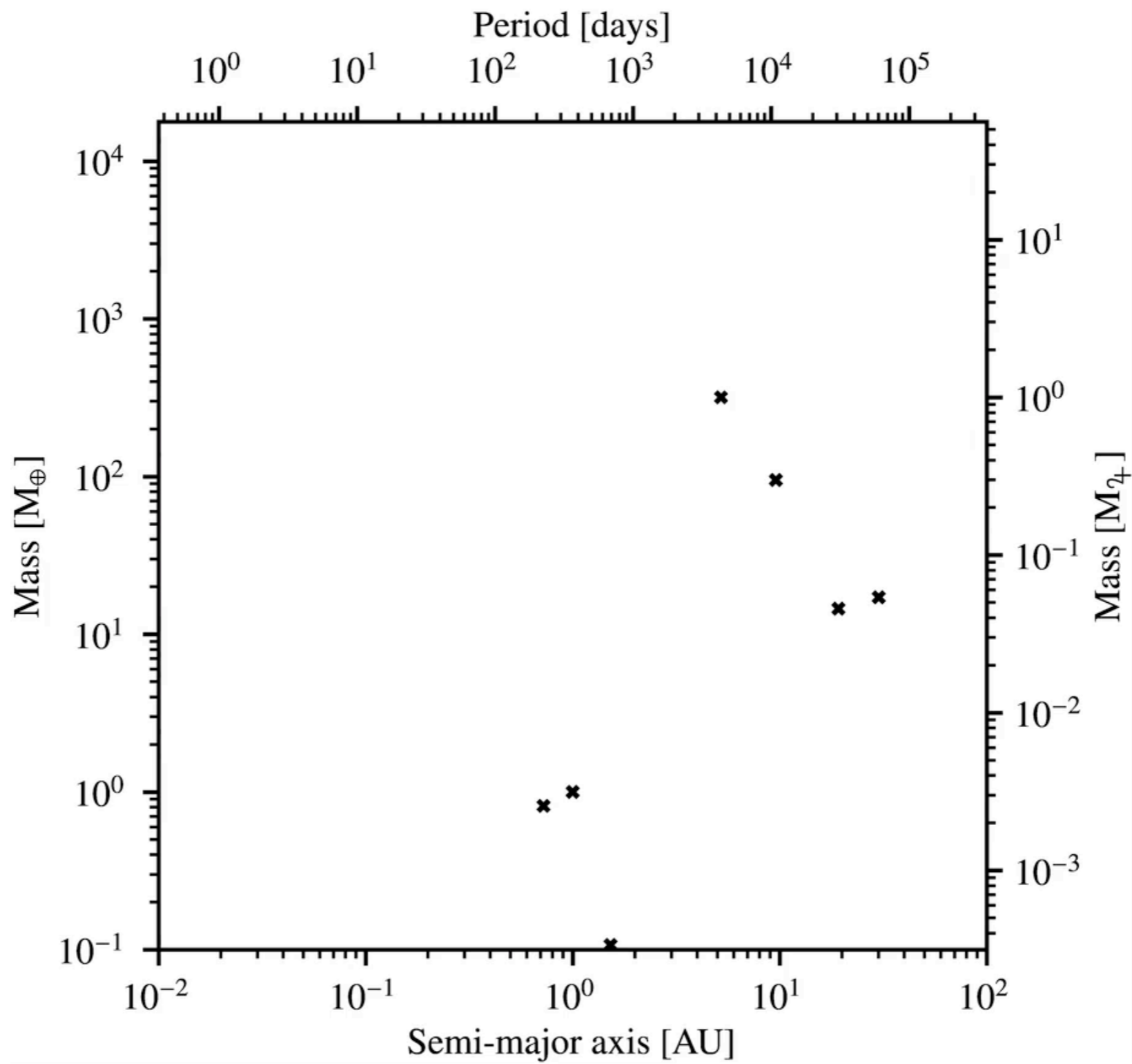
ALMA: Constraints on radial drift



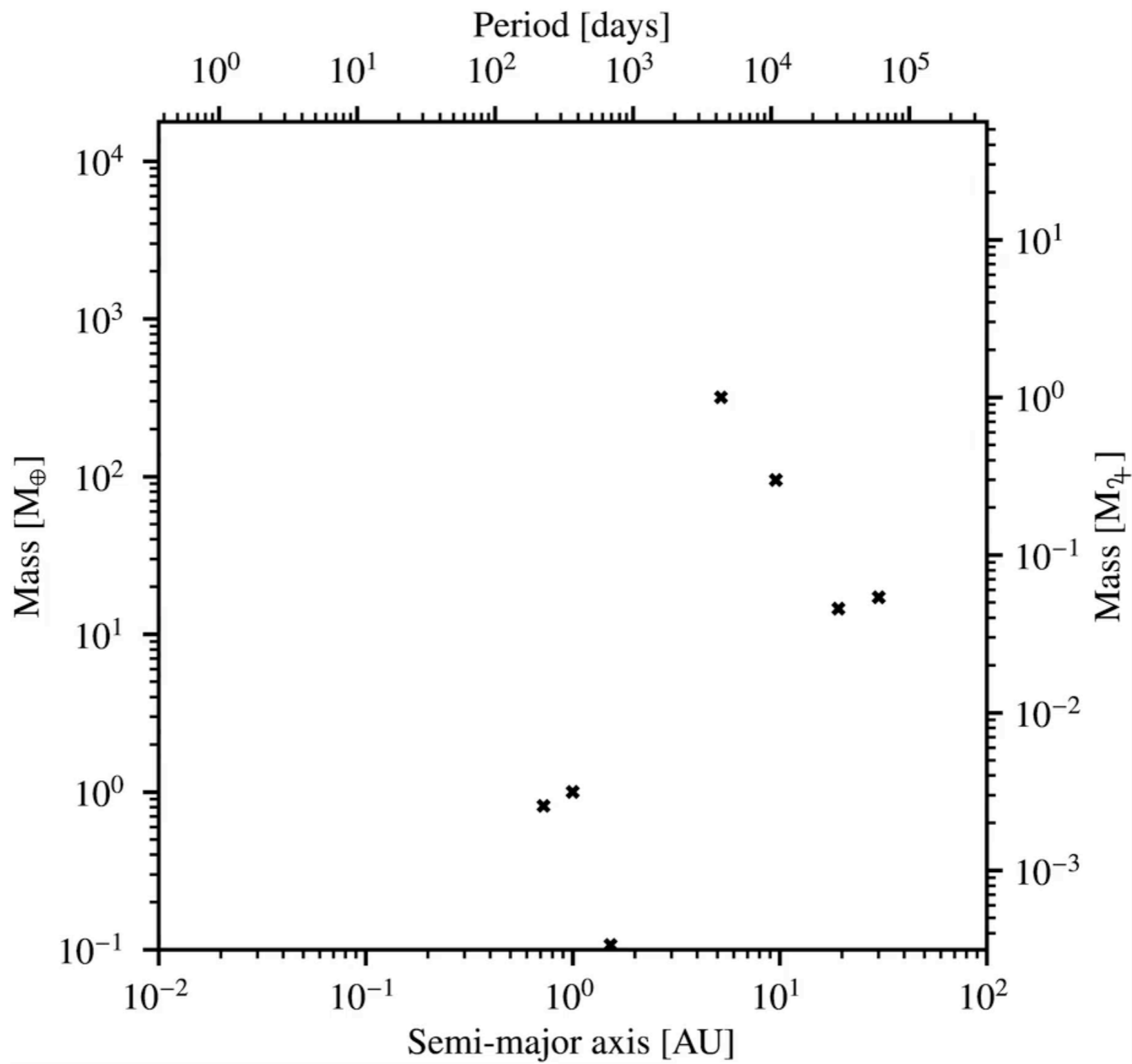
Low-mass disks are more compact



Dispersion in disk mass



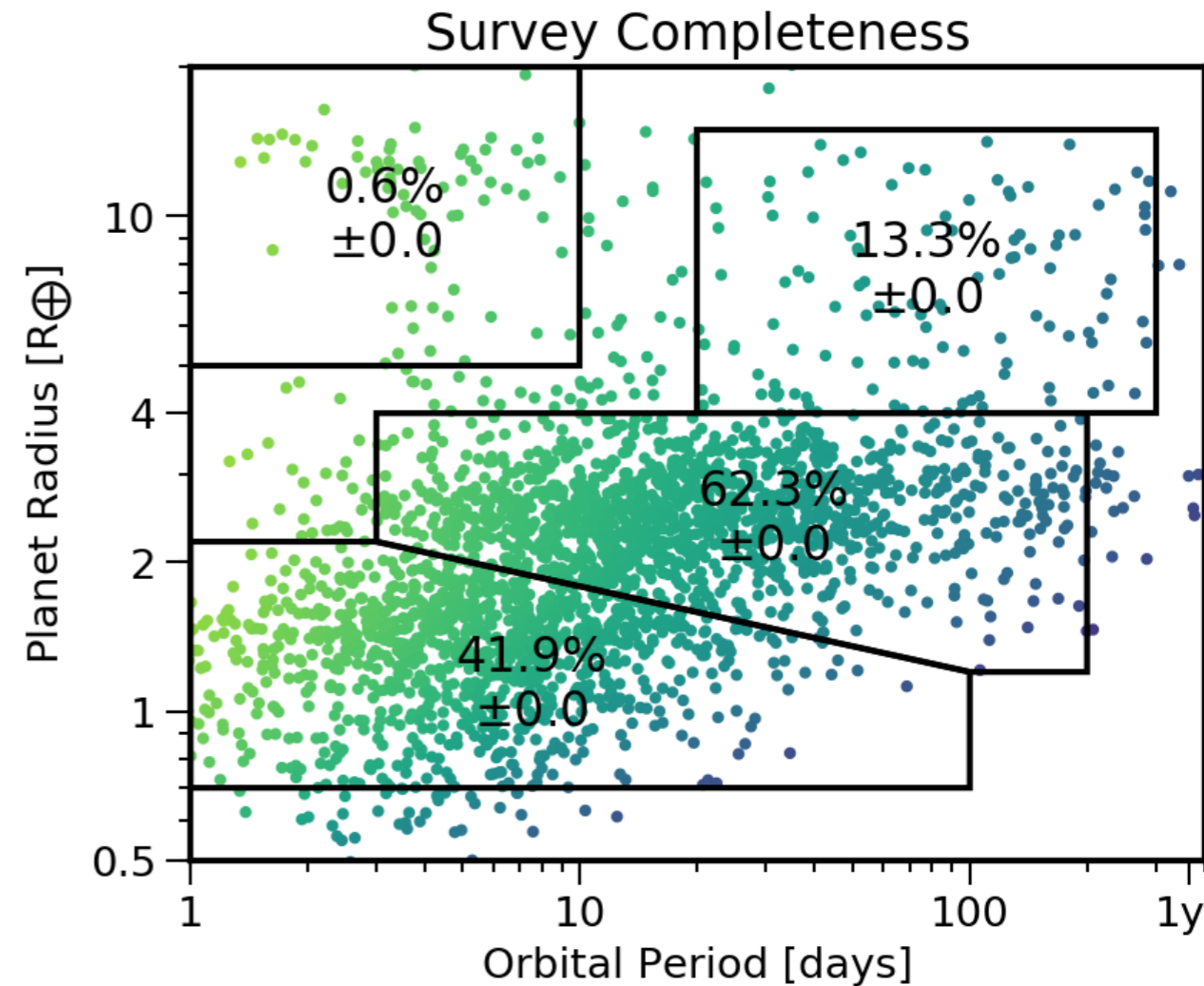
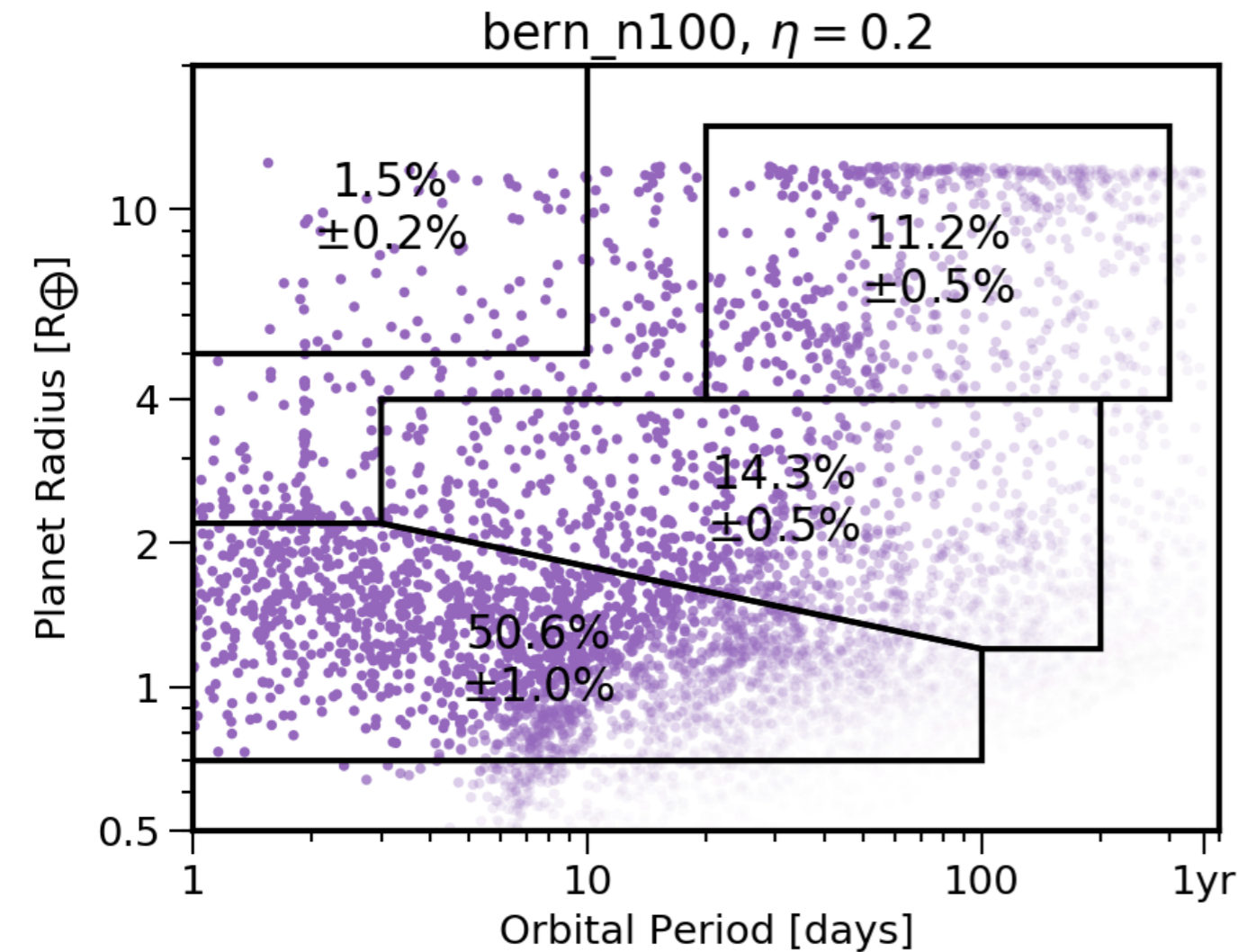
Population Synthesis (Bern models)



Population Synthesis (Bern models)

Bern Model

Observed Planets

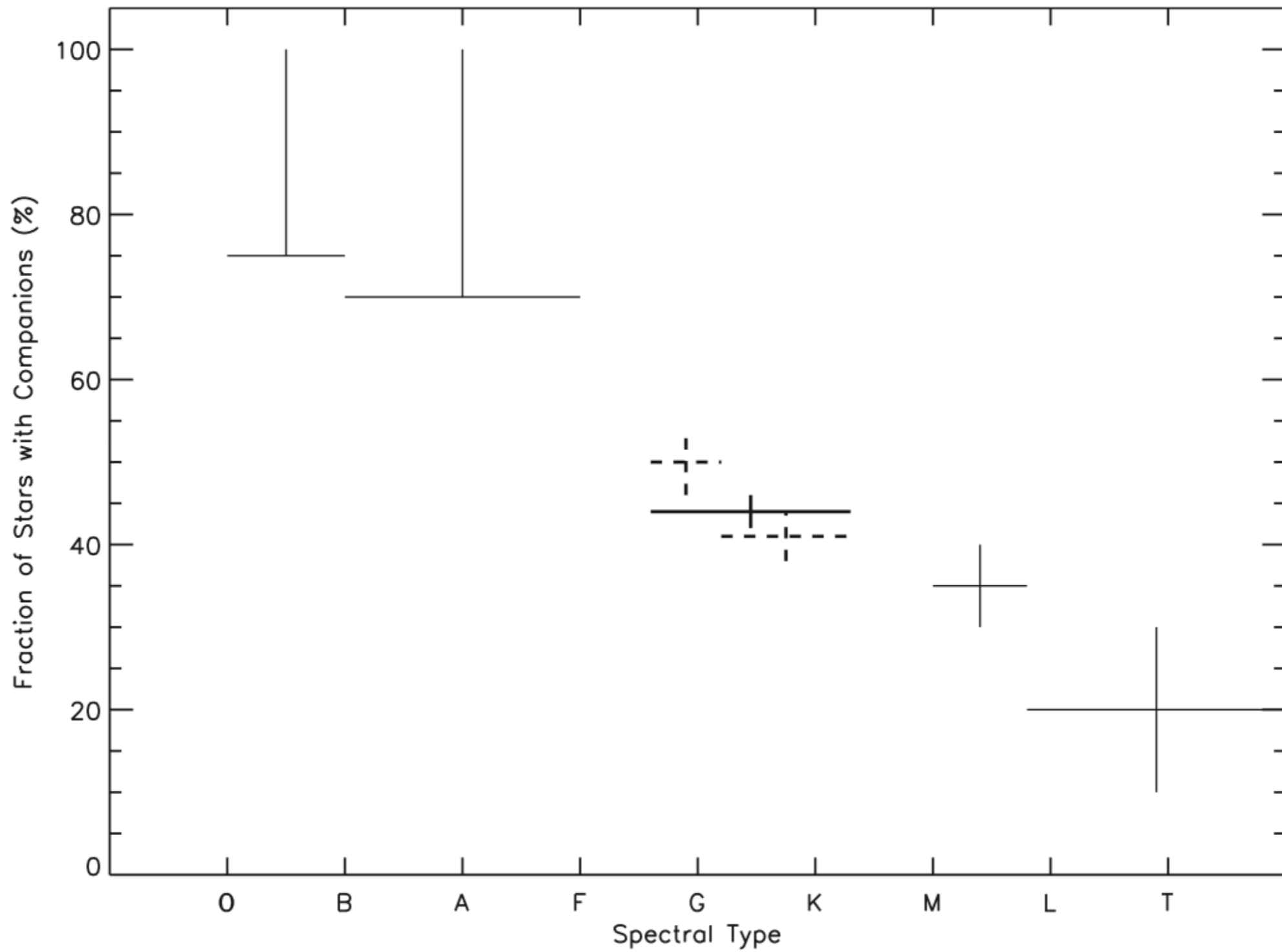


Evaluating Population Synthesis Models
arXiv:1905.08804

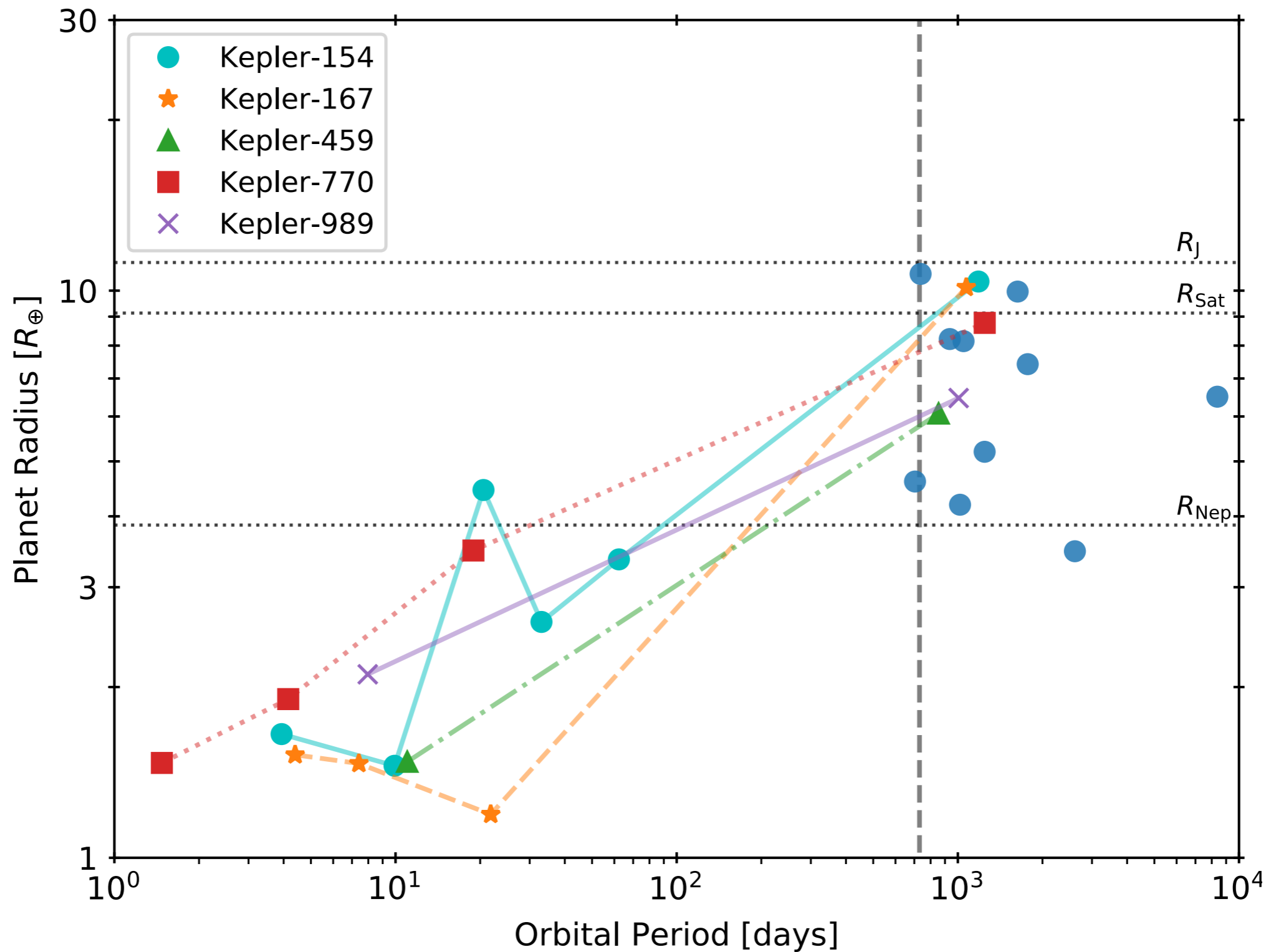
The exoplanet population around stars of different masses

- Giant planets more common around more massive stars
- Sub-Neptunes are more common around low-mass stars
- Inconsistent with protoplanetary disk mass scaling laws
- Efficient radial drift or migration?

Backup Slides



Binary Fraction



Cold Giants and Close-in Planets correlated

Herman+ 2019, Bryan 2016 & 2019