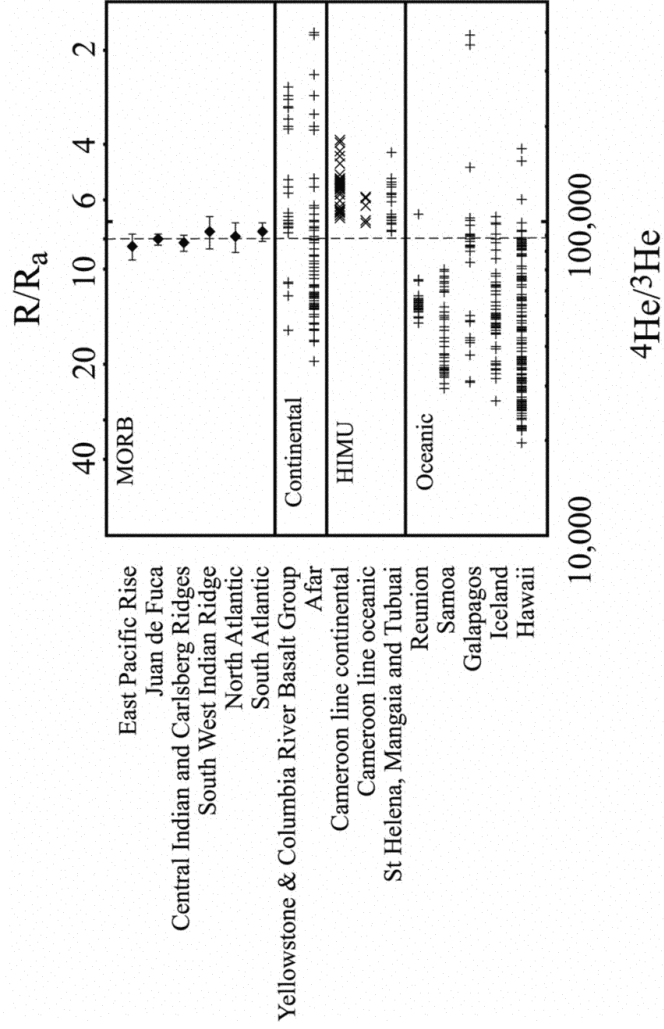
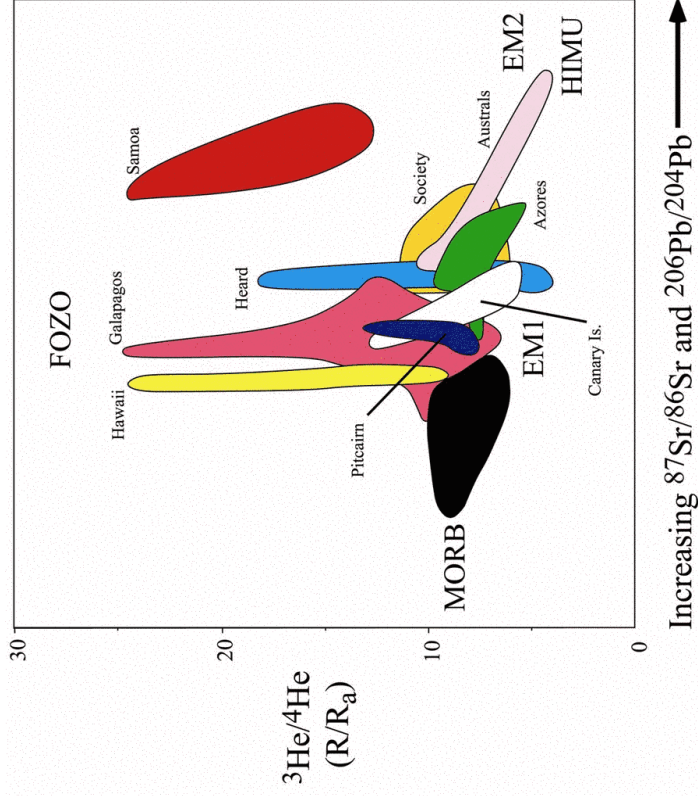


## The C in Cider

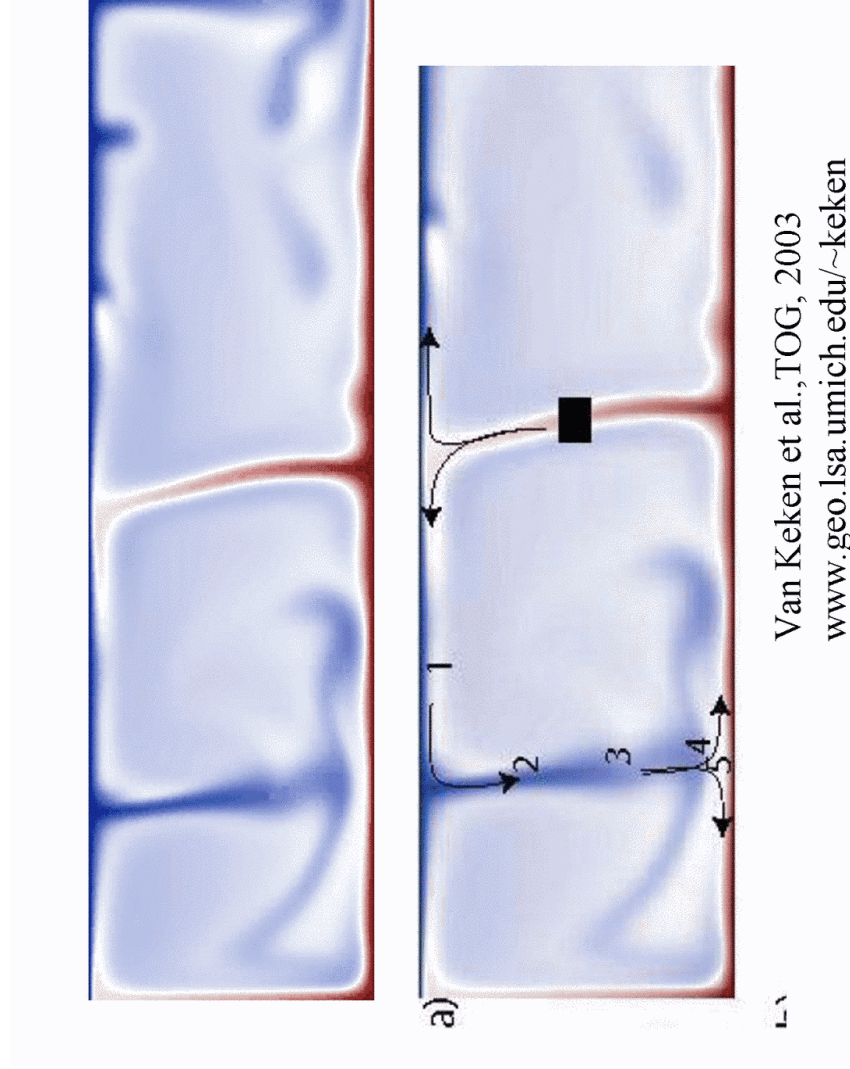


- Mantle mixing
  - Physics of ...
  - Tools for characterization
- Subduction zones
  - Thermal structure
  - EQ source mechanisms
  - Transport of H<sub>2</sub>O
- Plate tectonics & early Earth
  - Limits of uniformitarianism?

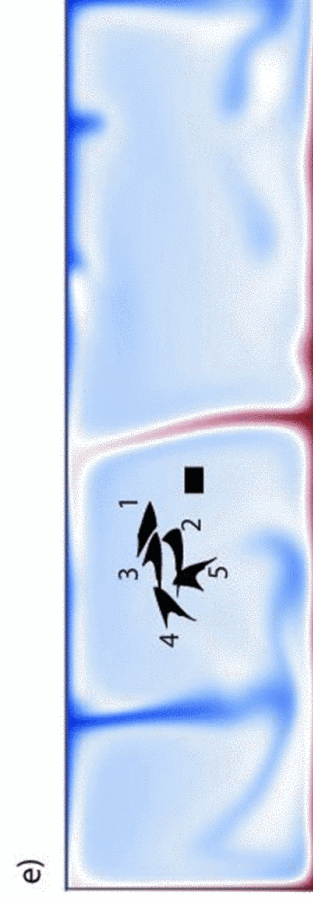
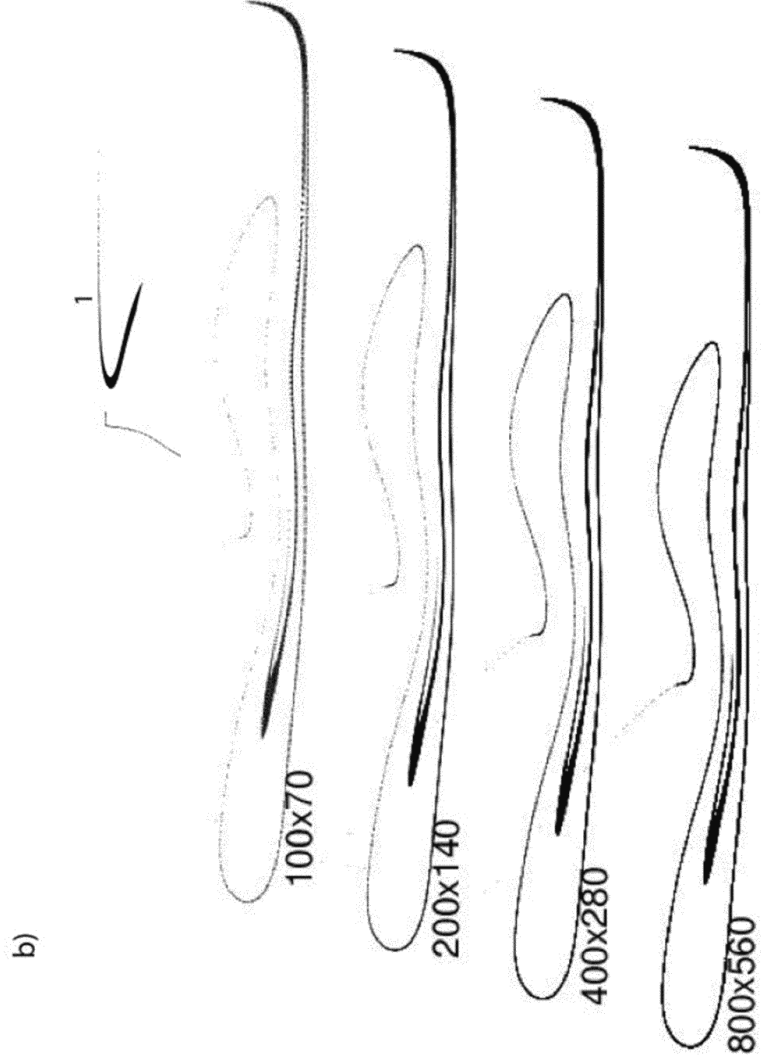




Van Keken, Ballentine and Hauri, Ann Rev., 2002



Van Keken et al., TOG, 2003  
[www.geo.lsa.umich.edu/~keken](http://www.geo.lsa.umich.edu/~keken)





Mixing = stretching + folding



## Physics of mixing

- Mixing processes
  - **Stretching**
  - **Folding**
  - Break up
  - Diffusion
- Efficient mixing
  - Strong deformation (velocity gradients)
  - Reorientation
  - Lagrangian turbulence



Physics of mixing

$$dX \cdot \longrightarrow dx$$

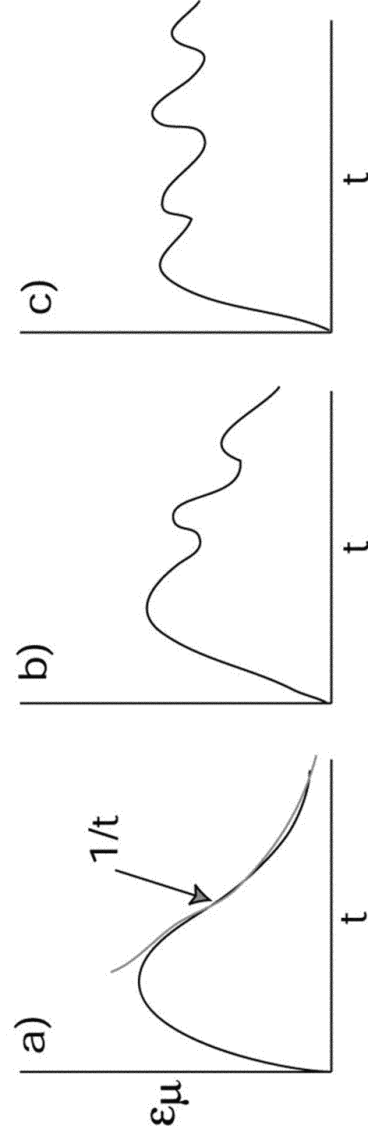
Stretching length  $\mu = dx/dX \quad dX \rightarrow 0$

Efficient stretching  $d(\ln \mu)/dt > 0$ .

Stretching tensor  $\mathbf{D} = \frac{1}{2}(\nabla \mathbf{v} + (\nabla \mathbf{v})^T)$

$$d(\ln \mu)/dt \leq (\mathbf{D}:\mathbf{D})^{1/2}$$

Stretching efficiency  $\varepsilon_\mu = \frac{d(\ln \mu)/dt}{(\mathbf{D}:\mathbf{D})^{1/2}} \leq \{(n-1)/n\}^{1/2}$



Simple stretching

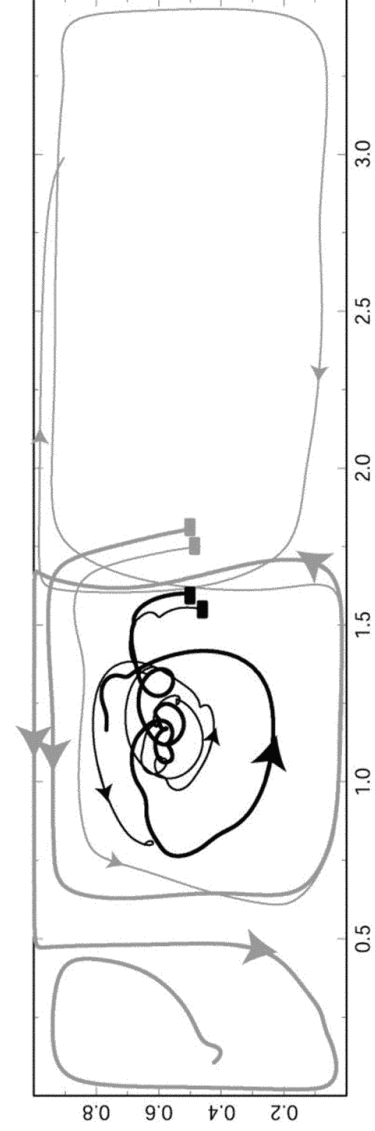
Stretching + occasional ('weak') reorientation

Stretching + frequent ('strong') reorientation

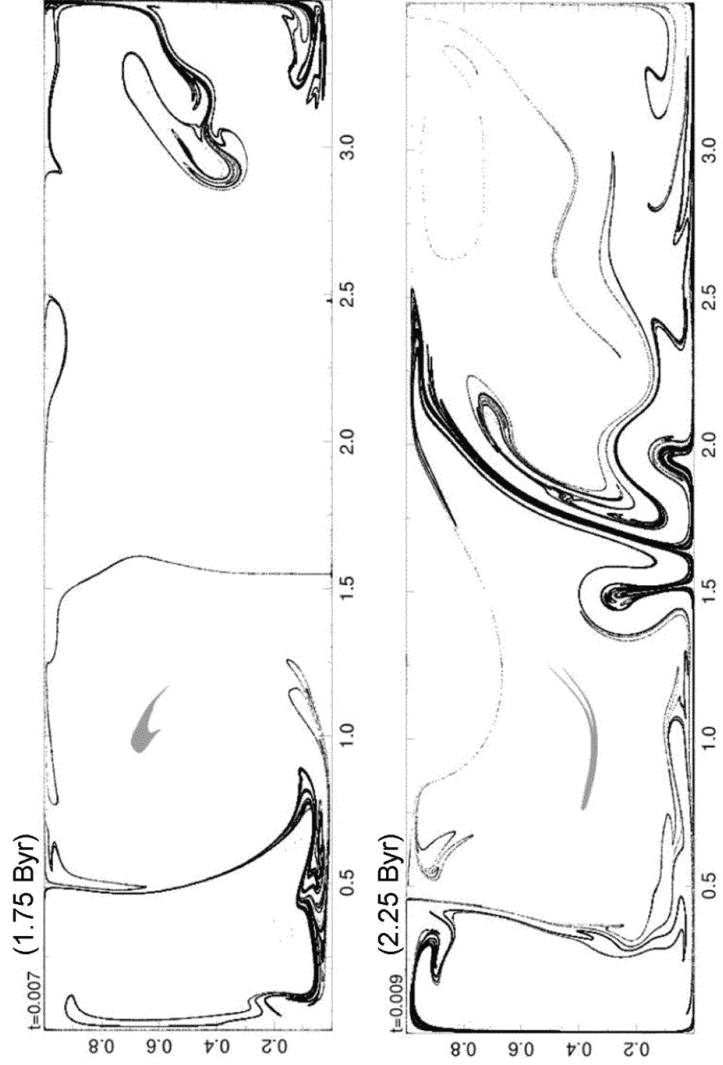
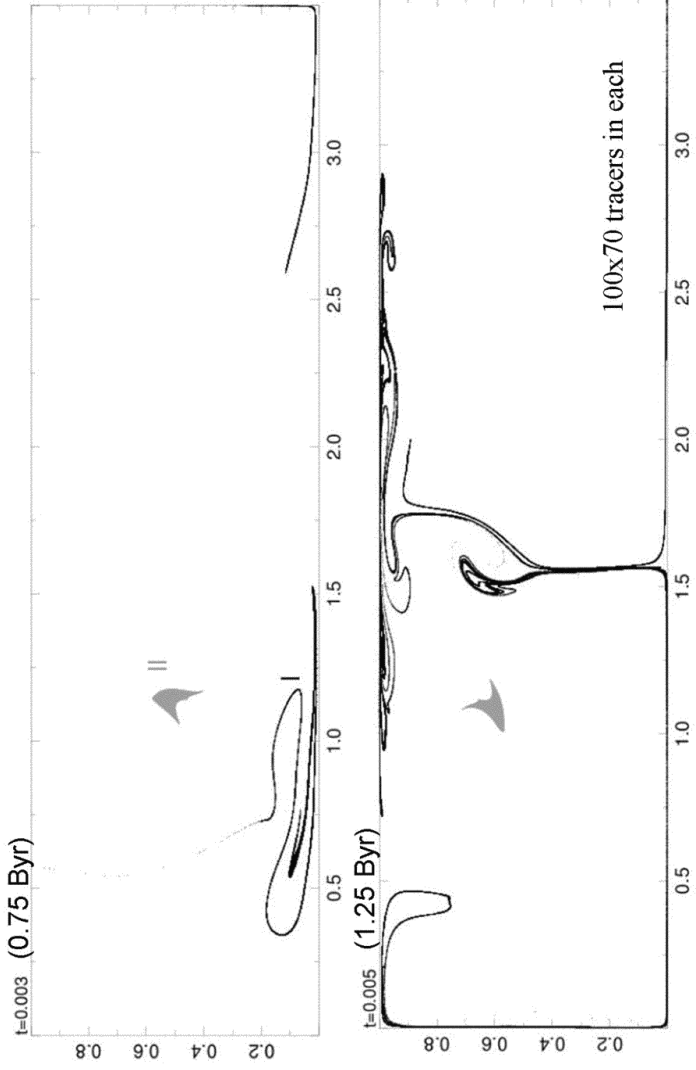
### Visualizing mixing: streamlines



### Visualizing mixing: particle paths



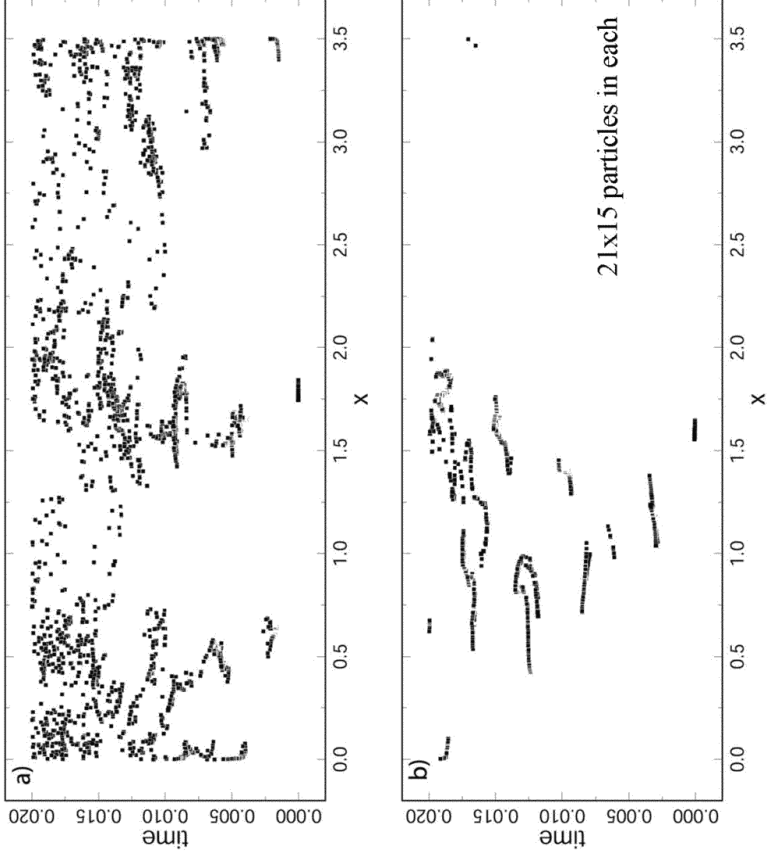
Visualizing mixing: dispersal of heterogeneity



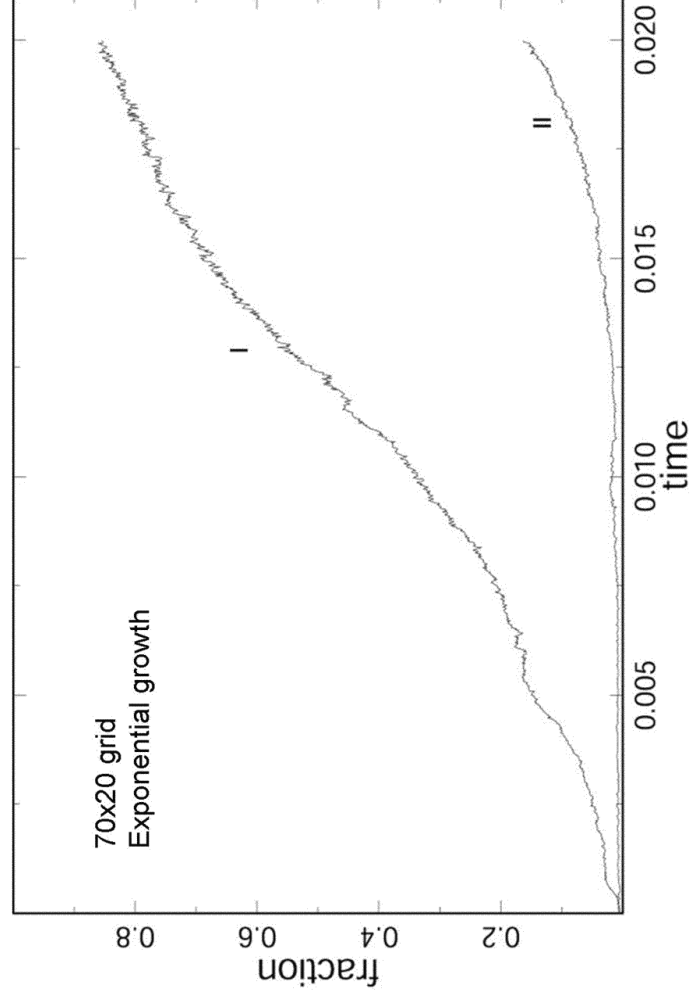
Van Keken et al., TOG, 2003



### Visualizing mixing: Poincaré maps

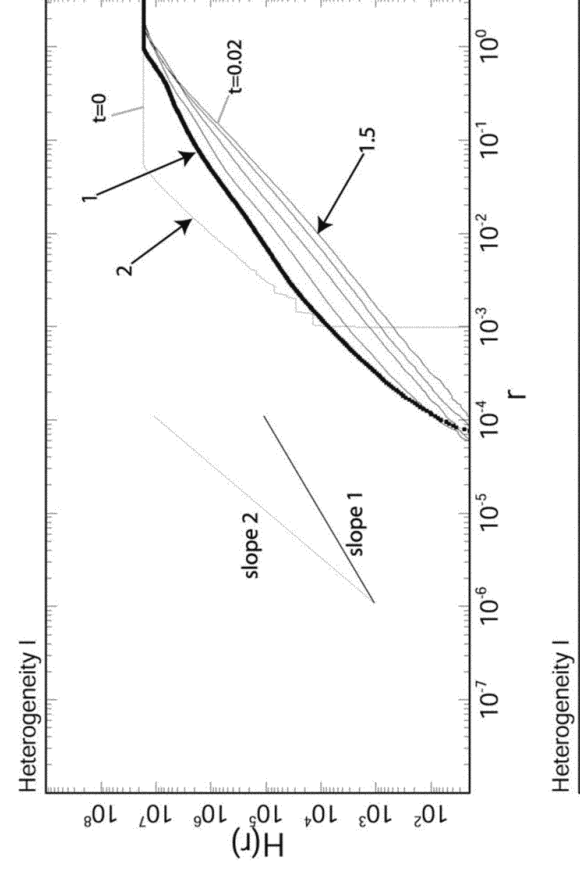
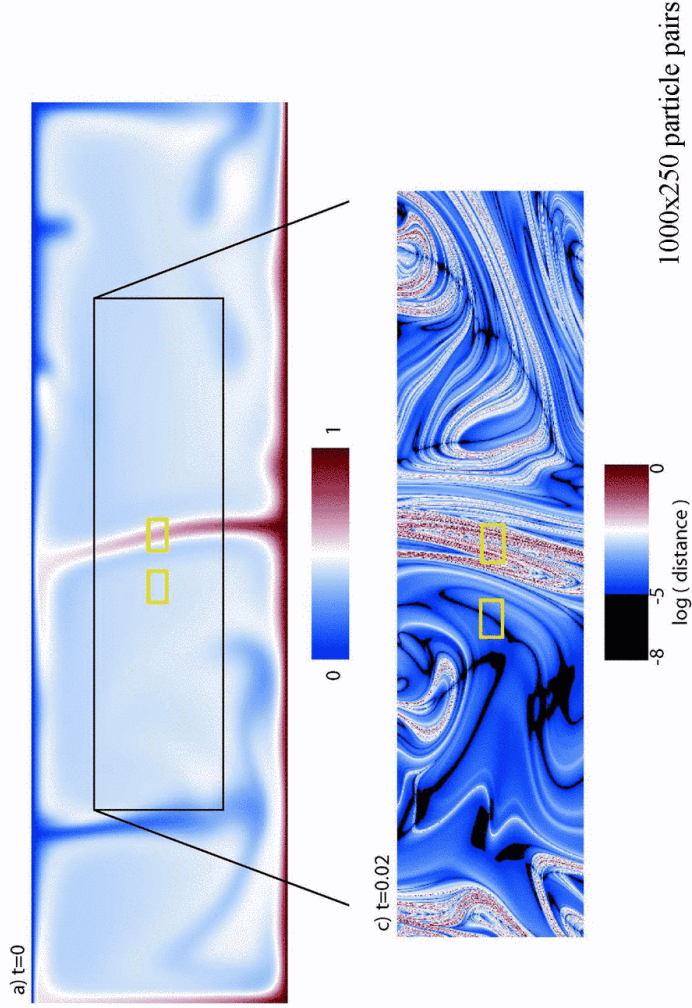


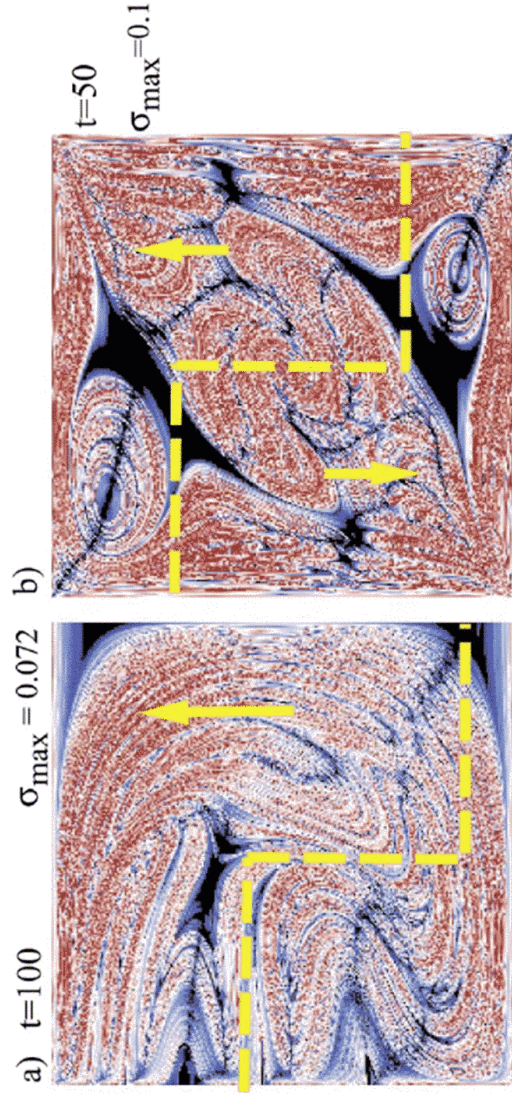
### Quantifying mixing: box counting



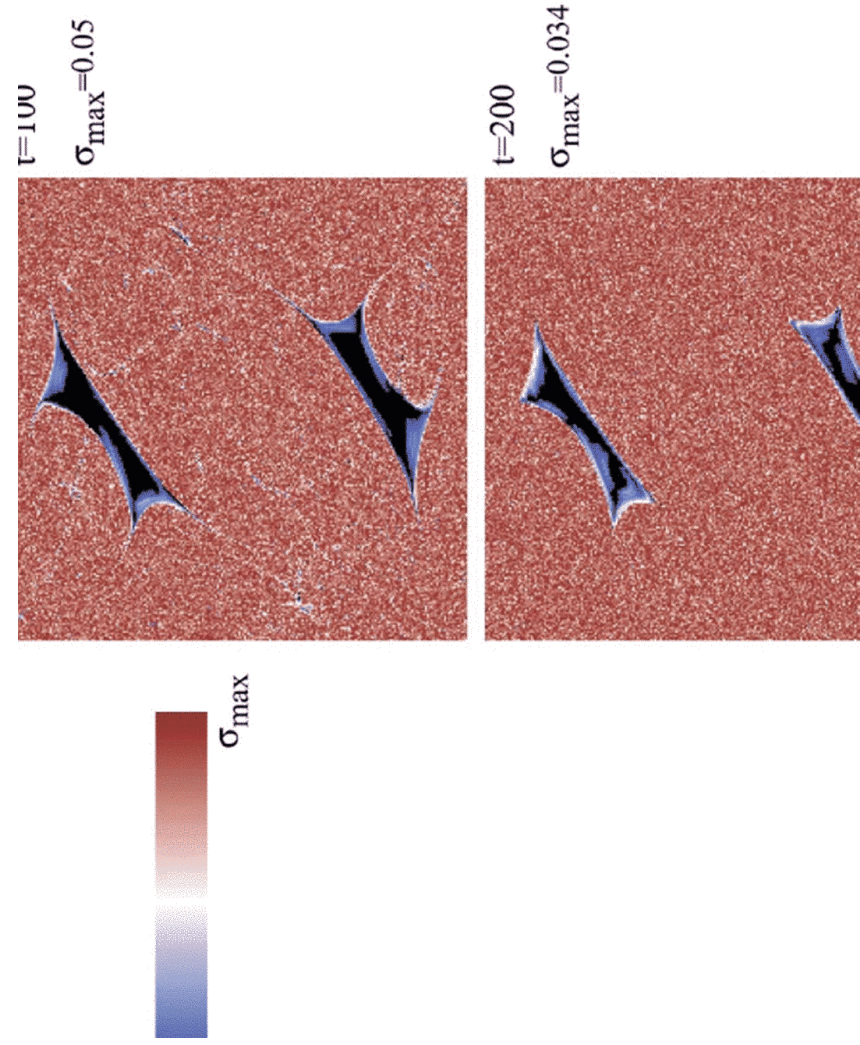
### Quantifying mixing: Lyapunov exponent

The  $\lambda$  in  $dx = dx \exp(\lambda t)$



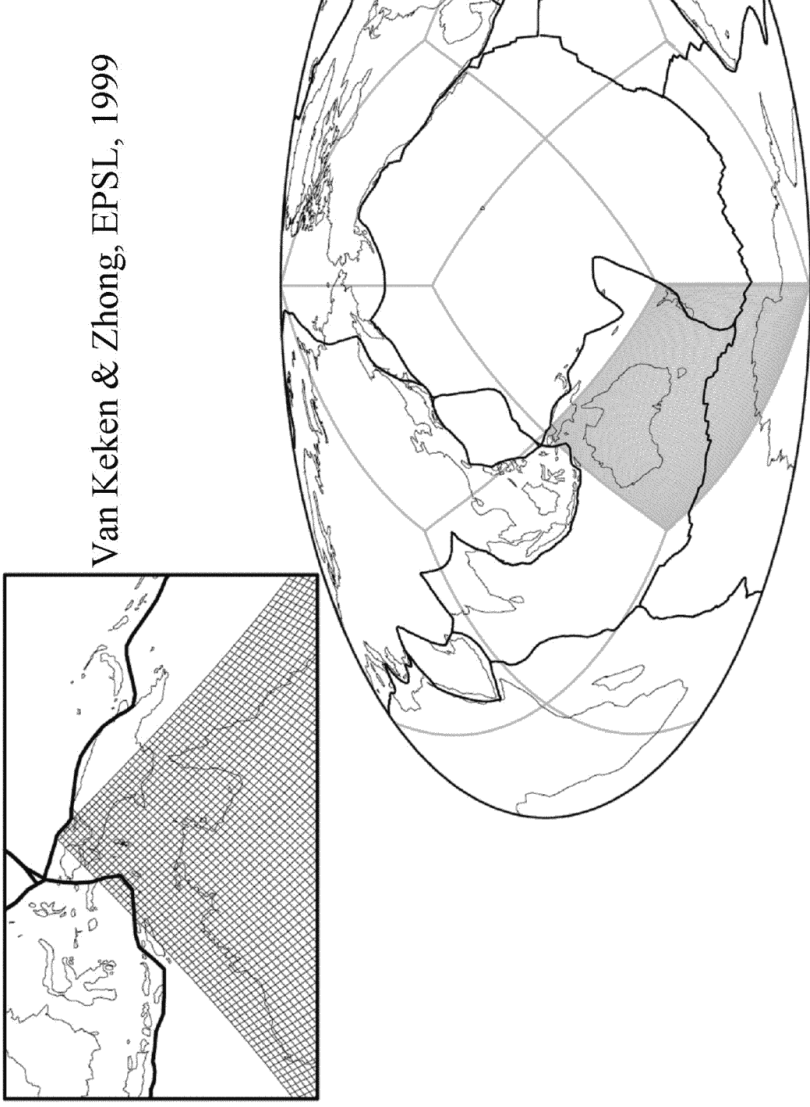


Reproduction of Ferrachat & Ricard, EPSL, 1999

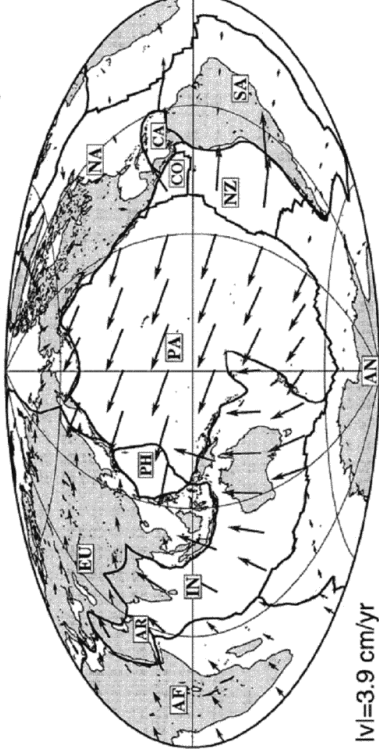




Van Keken & Zhong, EPSL, 1999

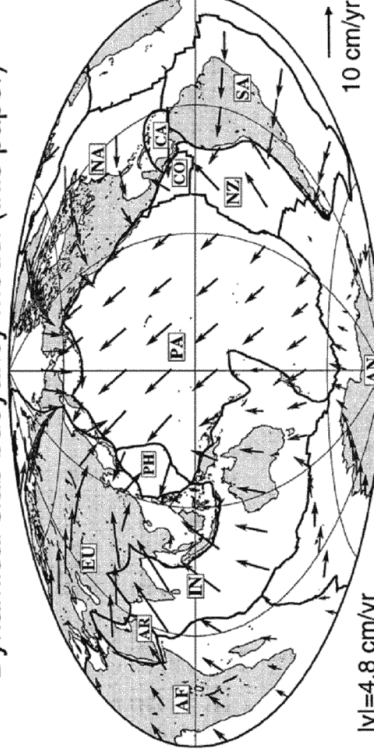


Hotspot reference frame model (Gordon & Jurdy, 1986)



a)

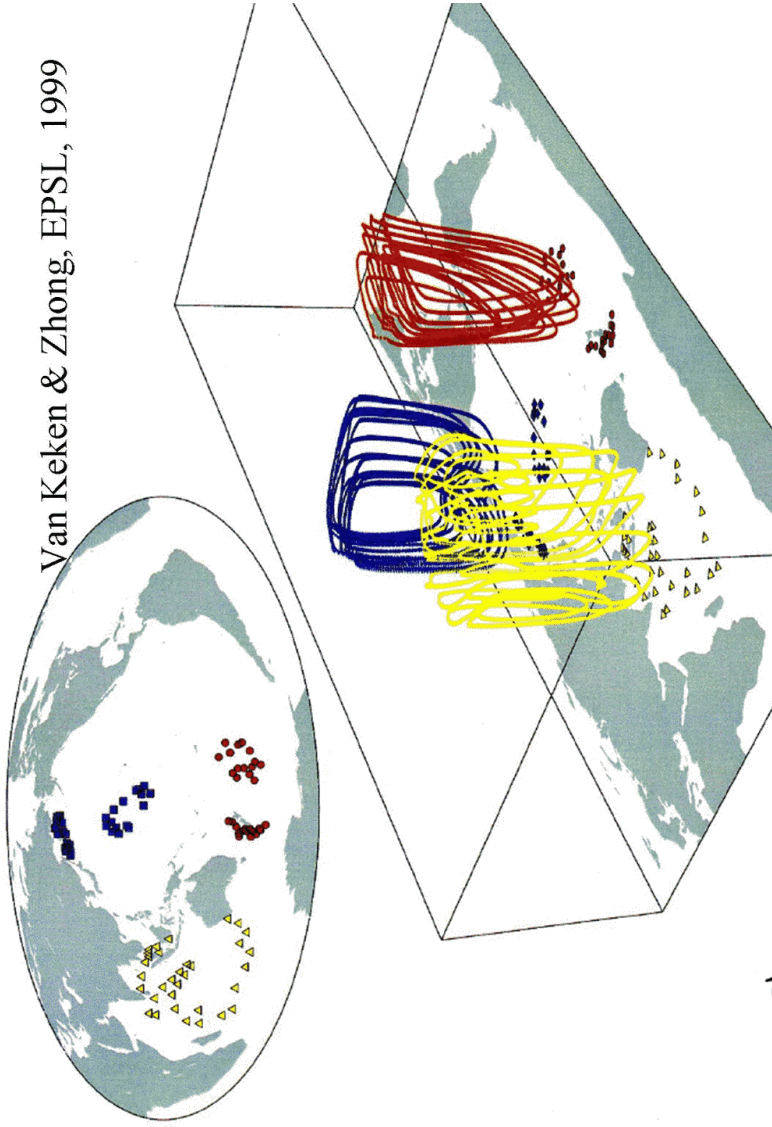
Dynamical slab buoyancy model (this paper)



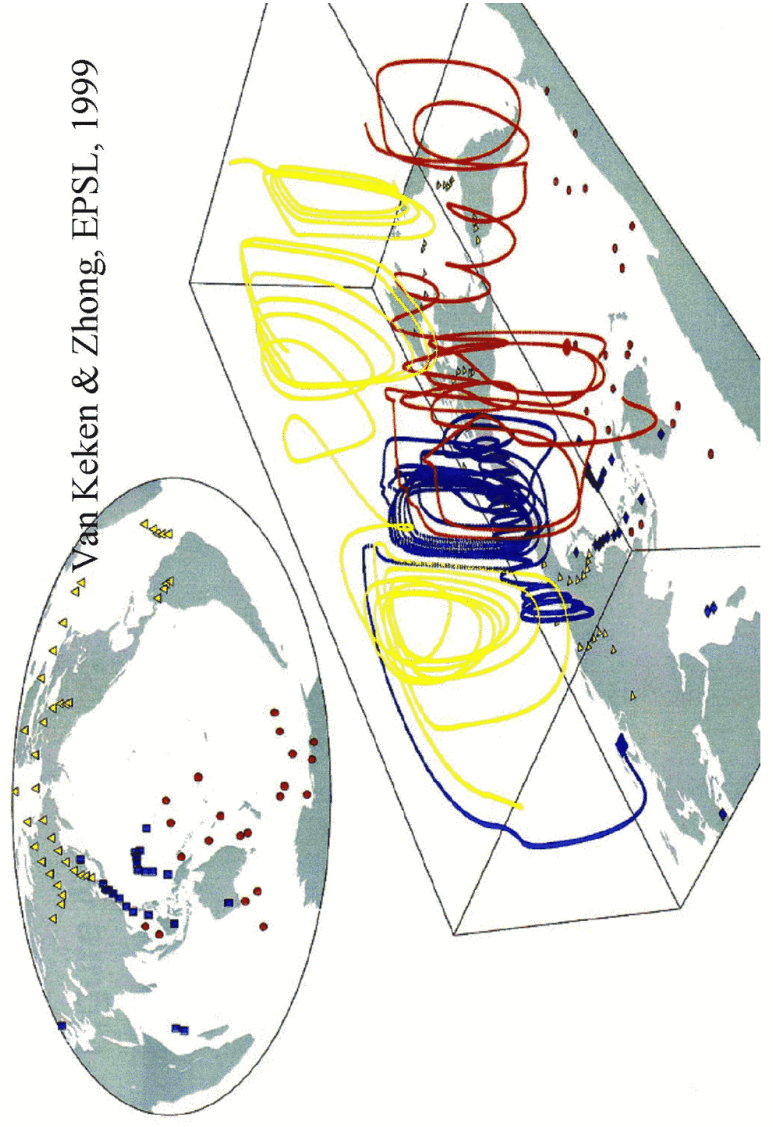
b)

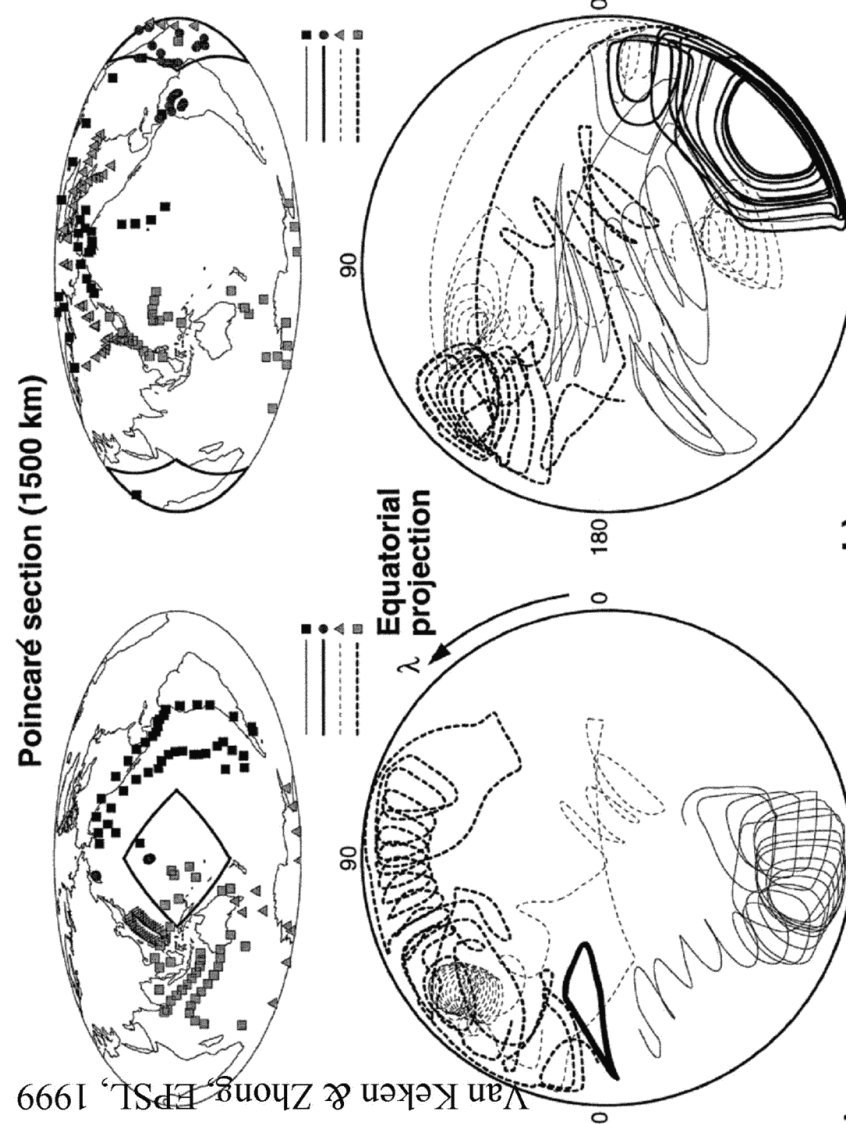
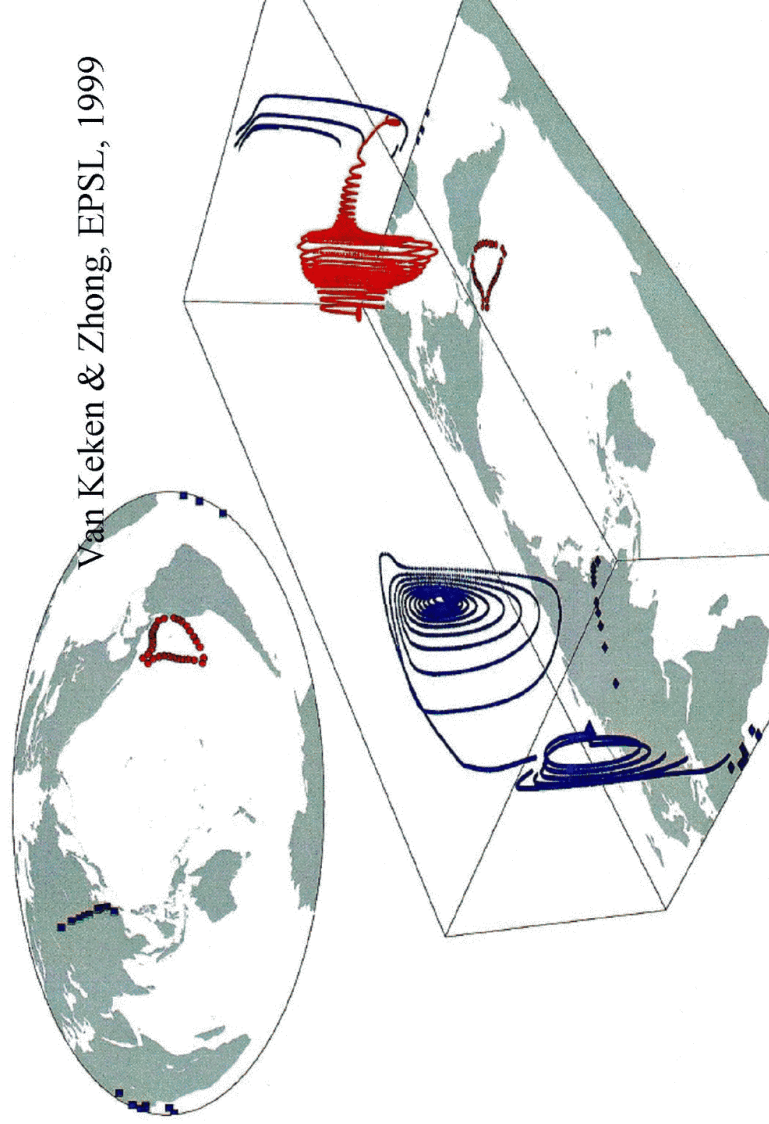
Van Keken & Zhong, EPSL, 1999

Van Keken & Zhong, EPSL, 1999

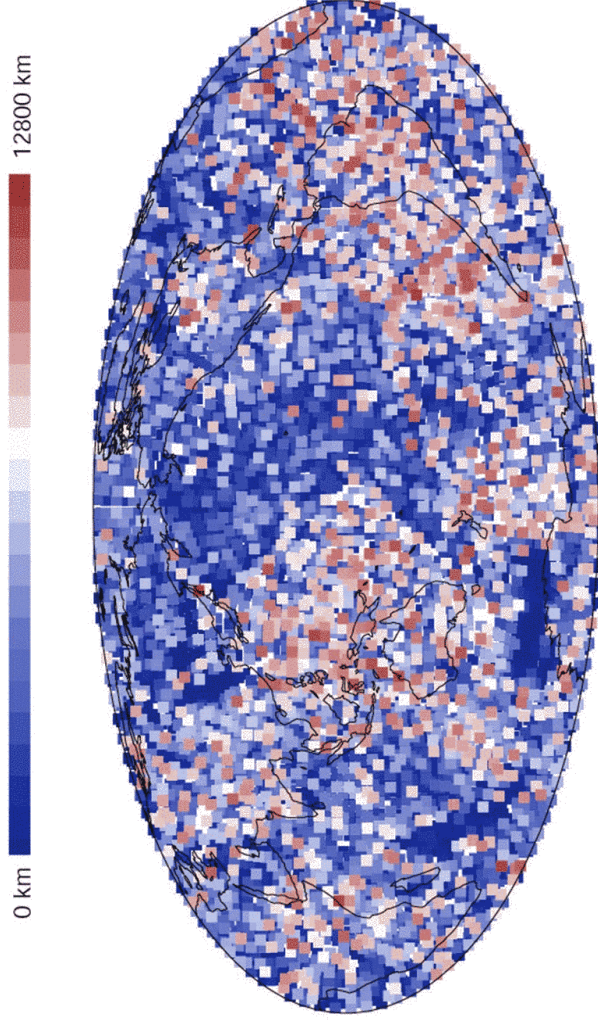


Van Keken & Zhong, EPSL, 1999

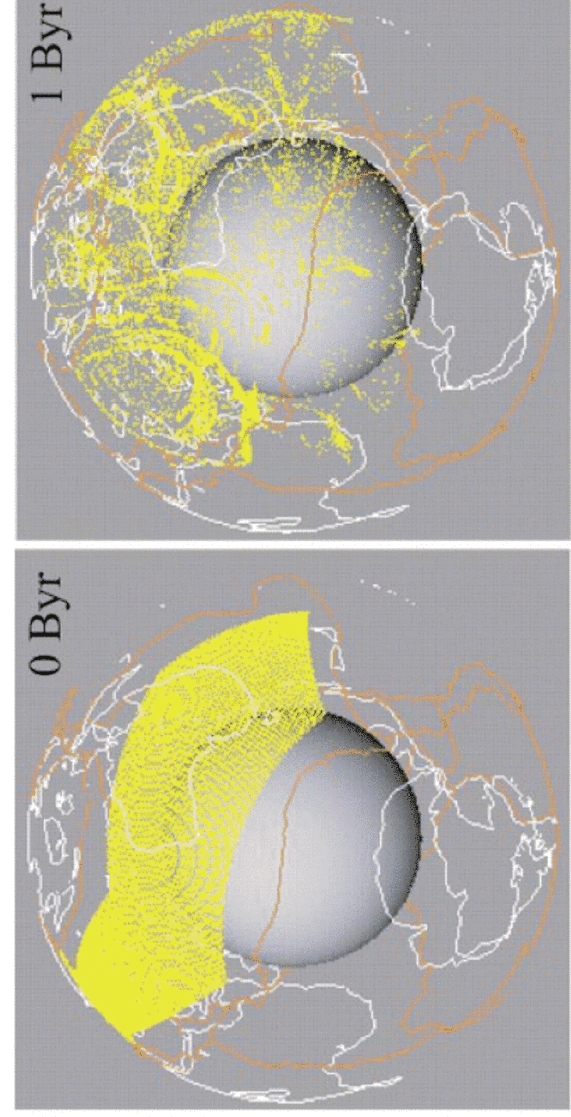




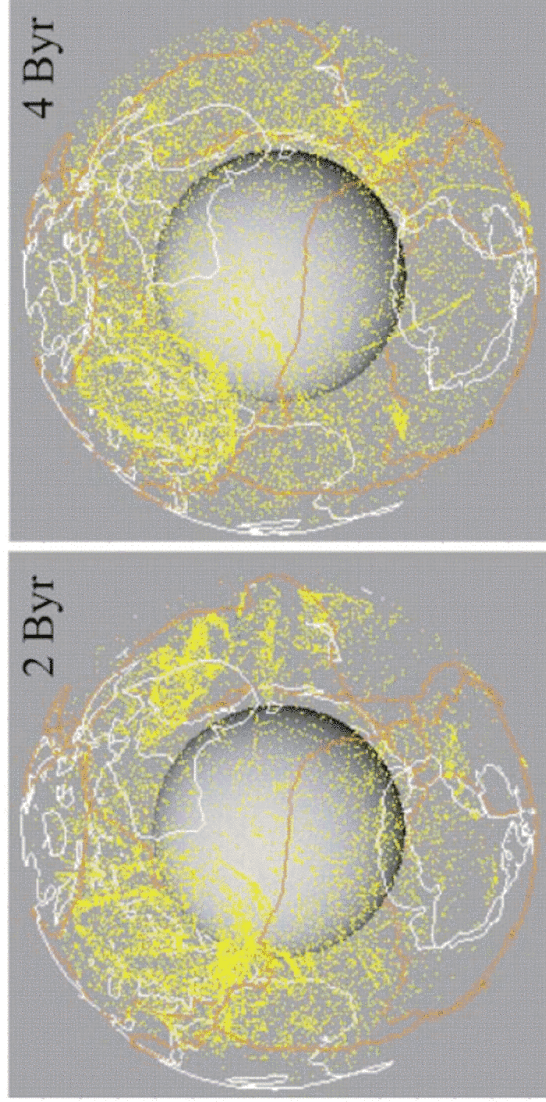




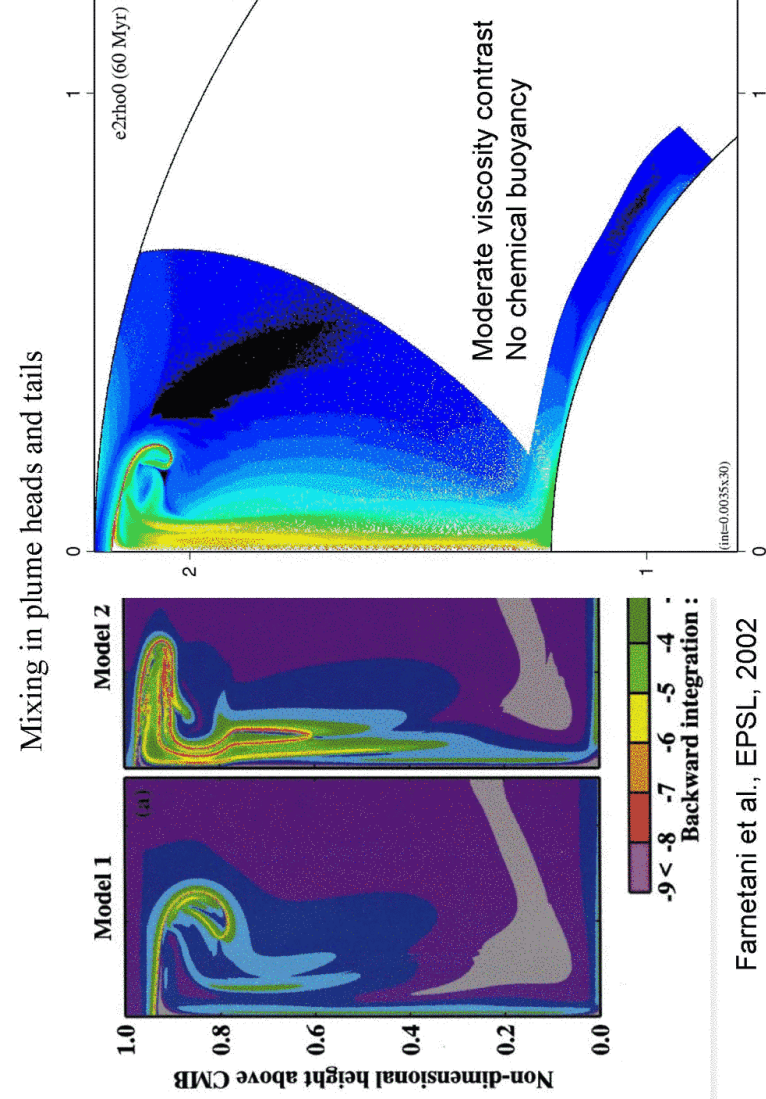
Van Keken & Zhong, EPSL, 1999



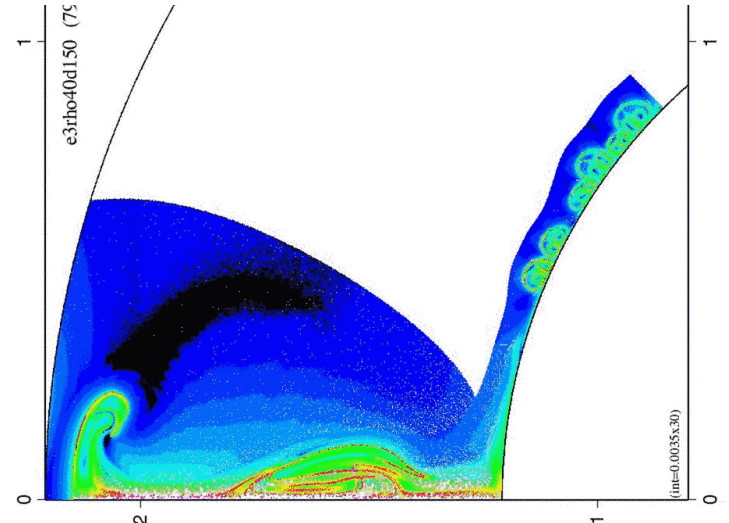
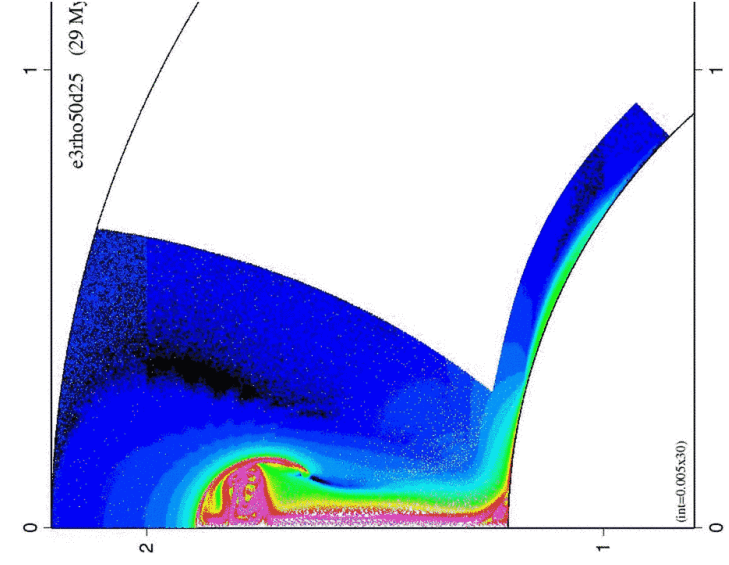
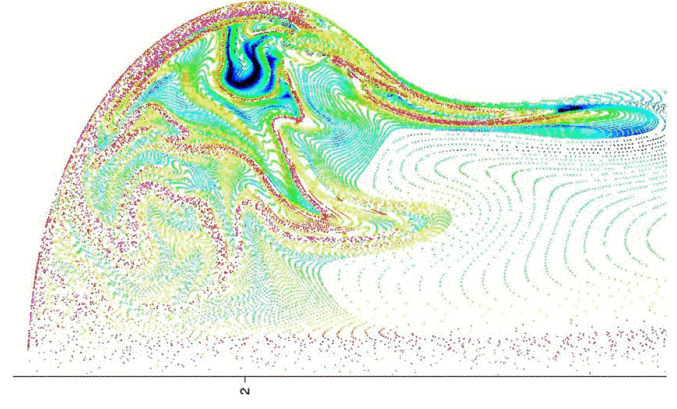
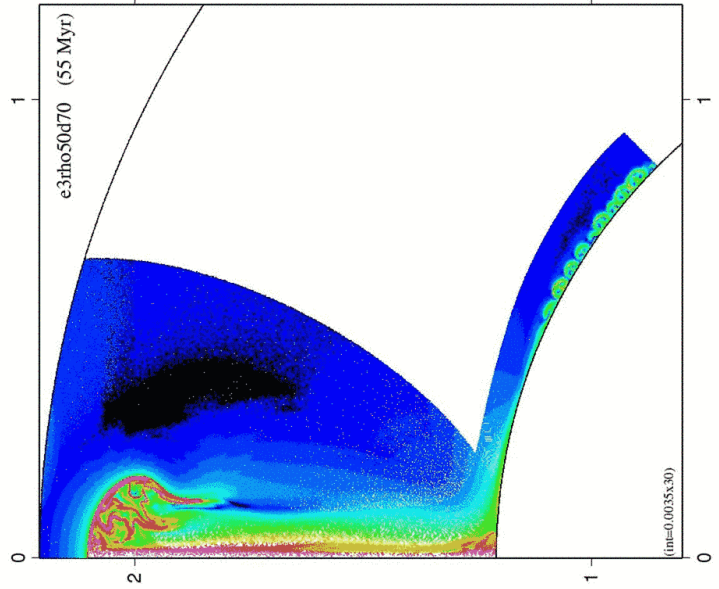
Van Keken & Zhong, EPSL, 1999



Van Keken & Zhong, EPSL, 1999

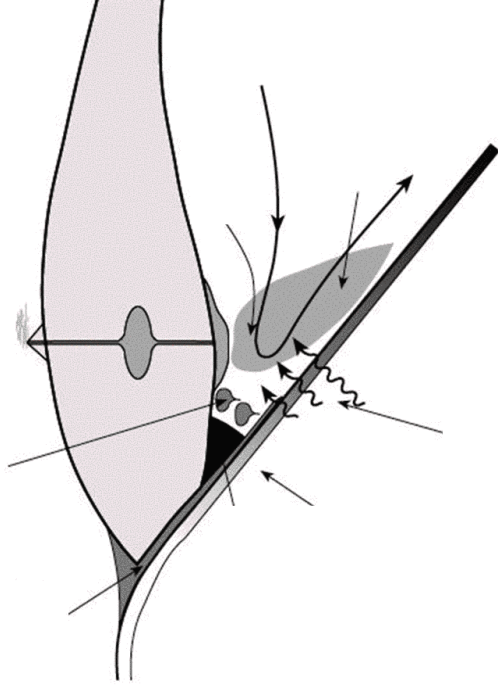






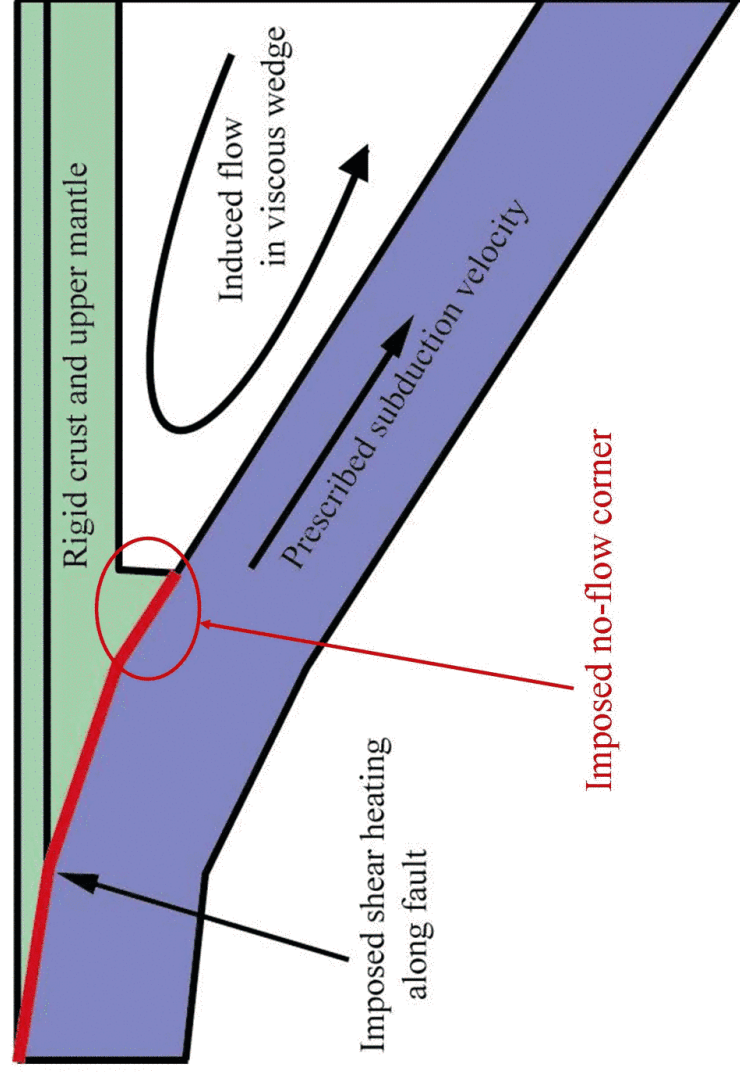
C en cidre, part deux

Controls on slab thermal structure:  
 age & speed of incoming slab  
 geometry  
 structure and dynamics of the wedge  
 shear heating along the seismogenic zone  
 phase changes



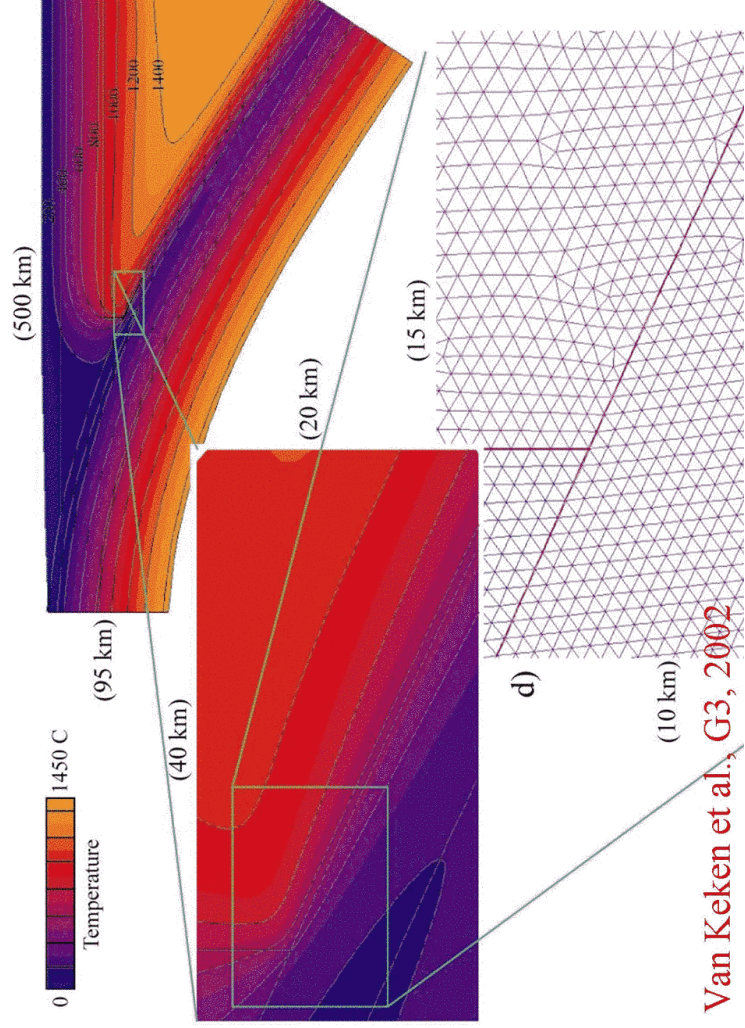
Van Keken, EPSL, 2003

II Kinematic/dynamic models



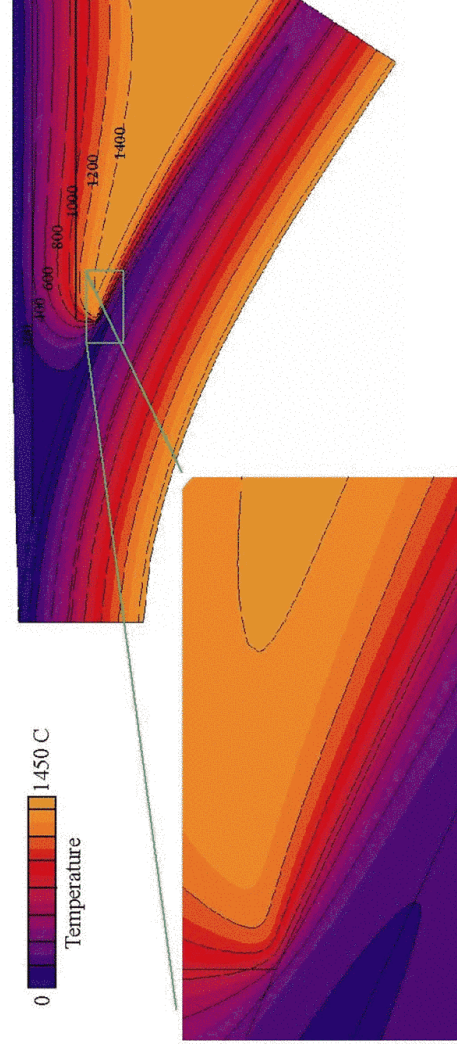


### Honshu; isoviscous

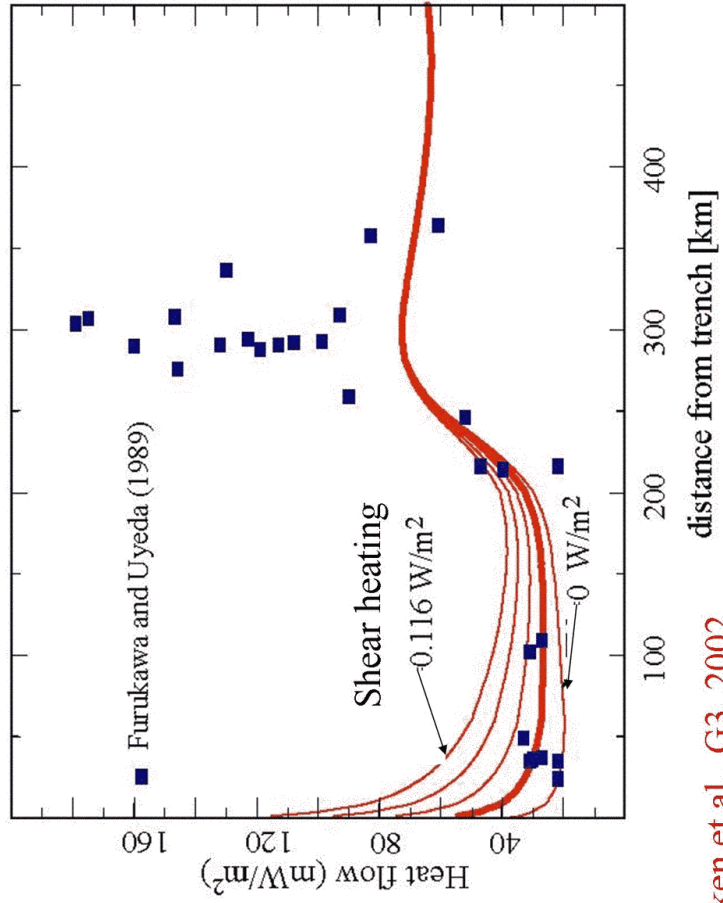


Van Keken et al., G3, 2002

### Honshu; Olivine rheology

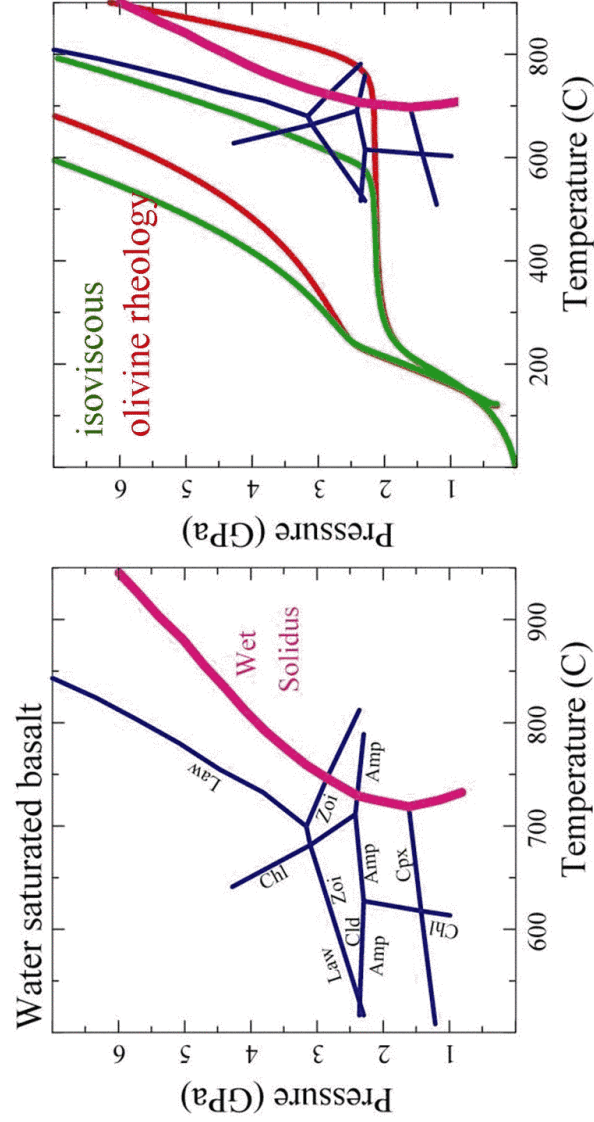


Van Keken et al., G3, 2002



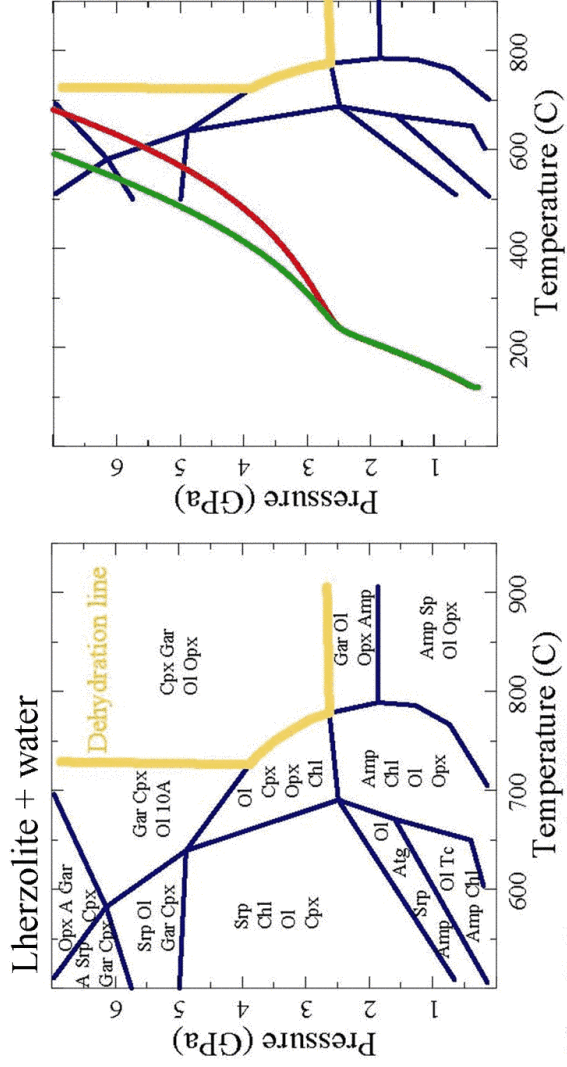
Van Keken et al., G3, 2002

### Temperature in subducted oceanic crust



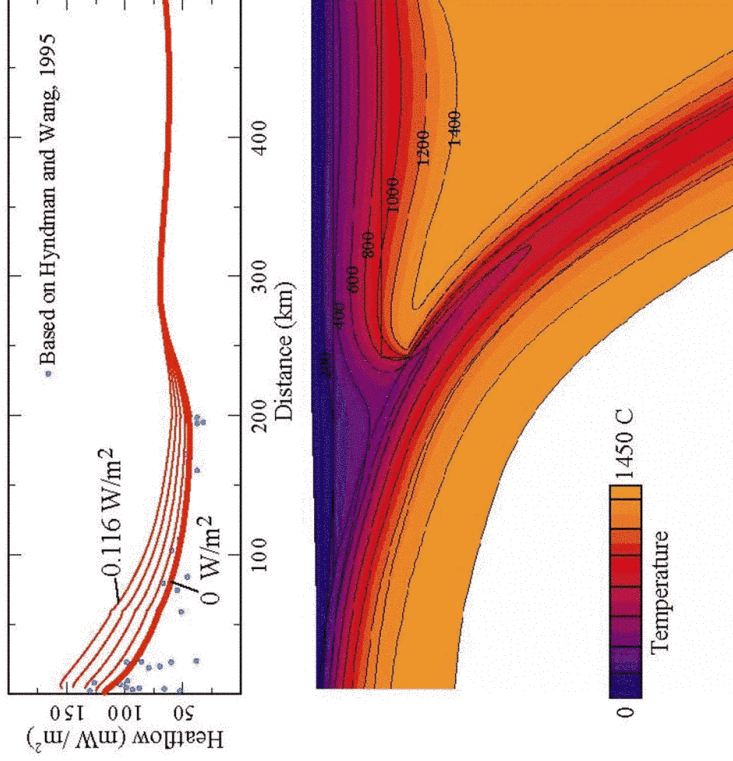
Van Keken et al., G3, 2002

Temperature in subducted oceanic upper mantle

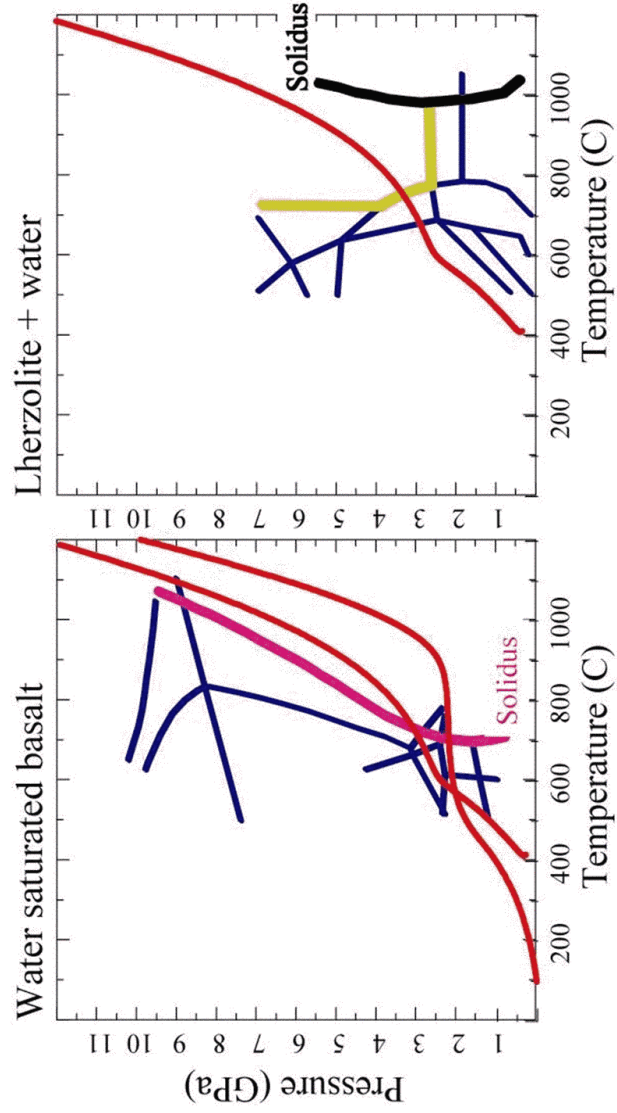


Van Keken et al., 2002

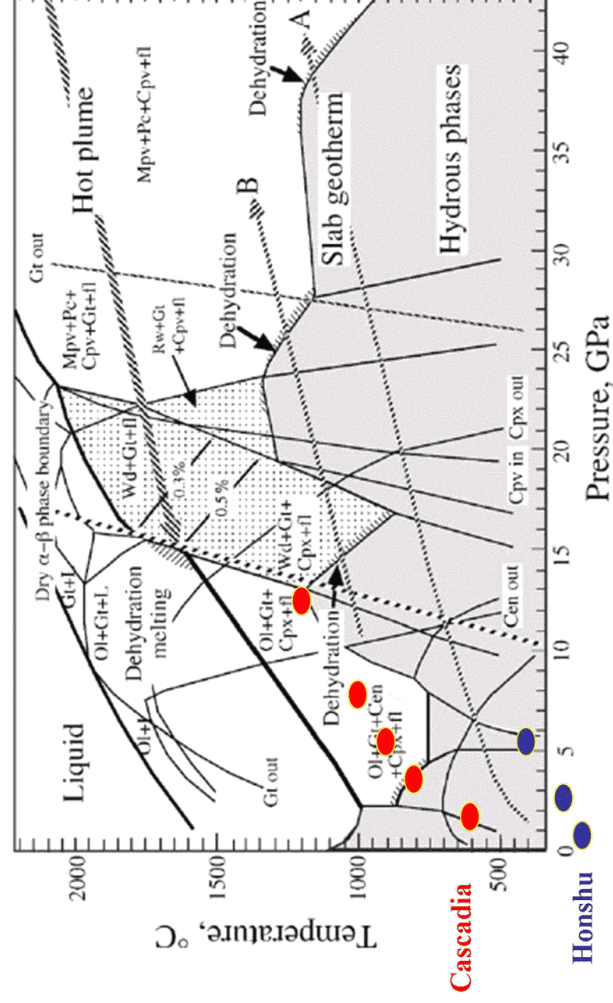
Cascadia; Olivine rheology







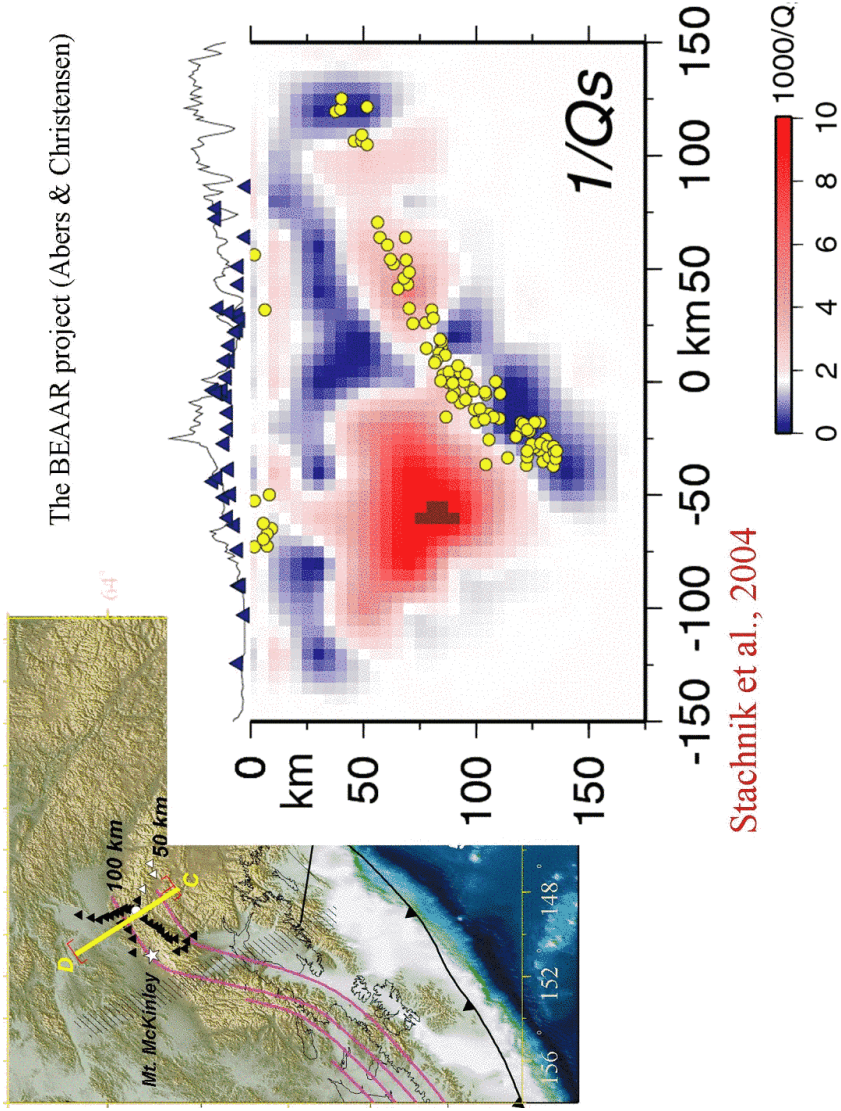
E. Ohtani et al. / *Physics of the Earth and Planetary Interiors* 143-144 (2004) 255-269



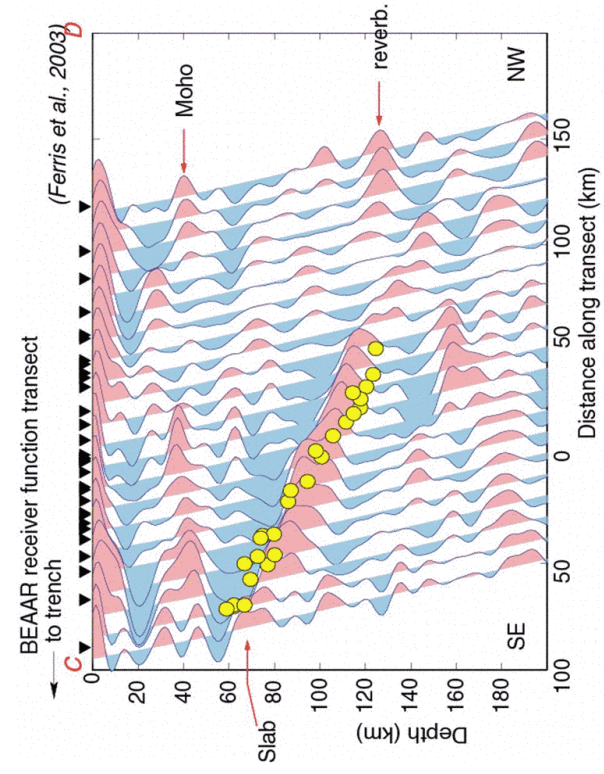
T along particle paths at 10 km below top of slab

Hirschmann and van Keken, in prep.

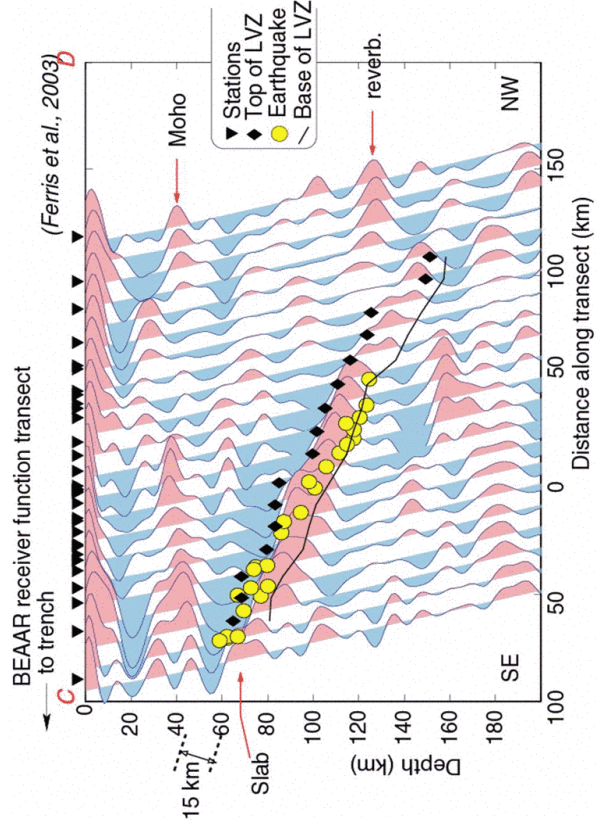




NW Back-azimuth: Processed Image

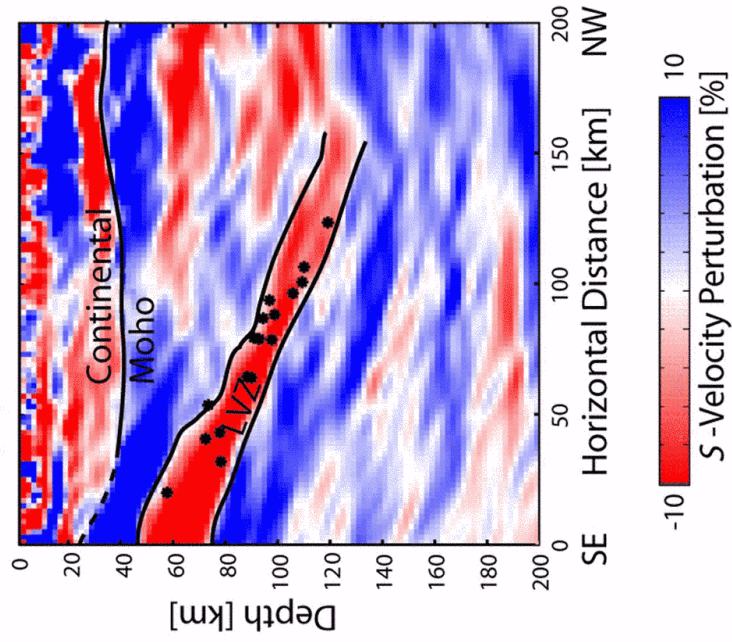


### Subducting Low-Velocity Zone



### Migration

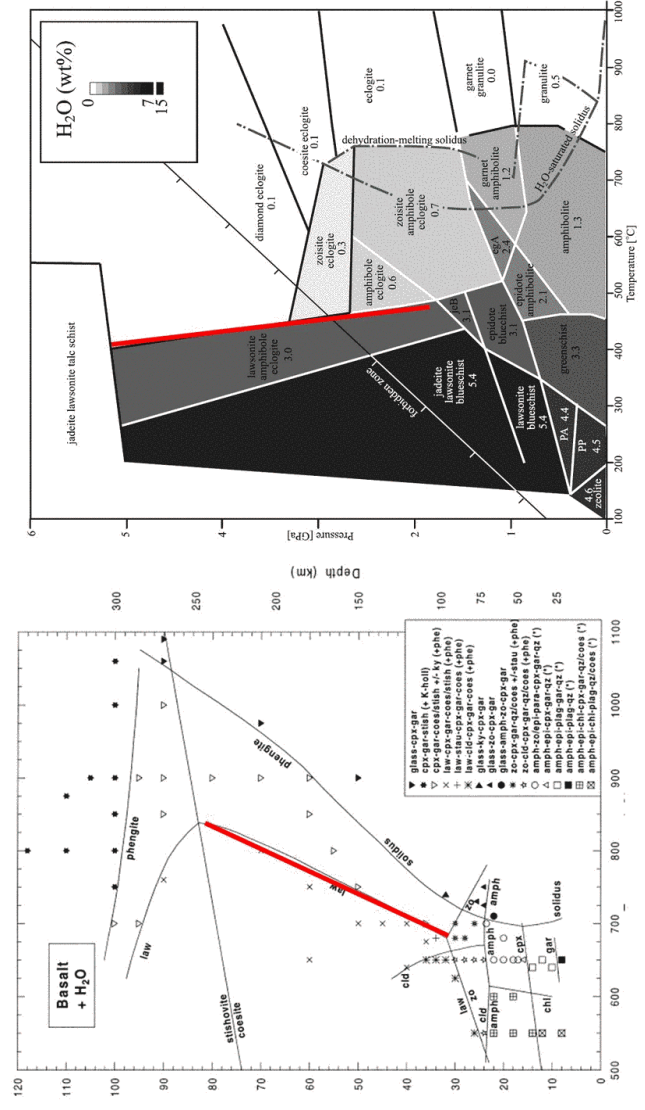
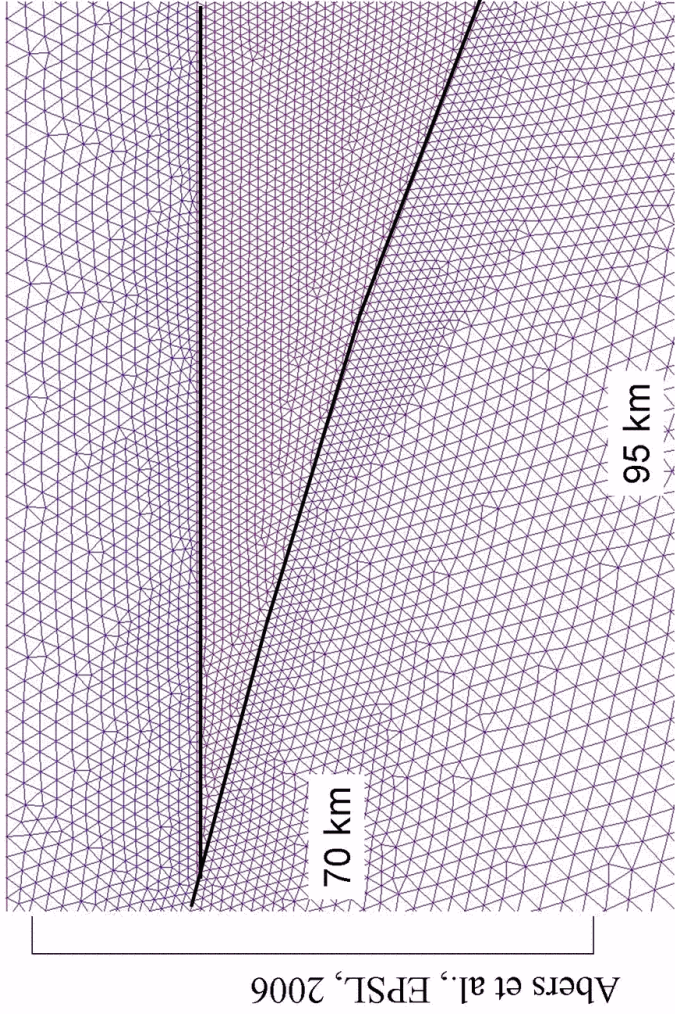
### $\delta\beta/\beta$ perturbations



(S. Rondenay)

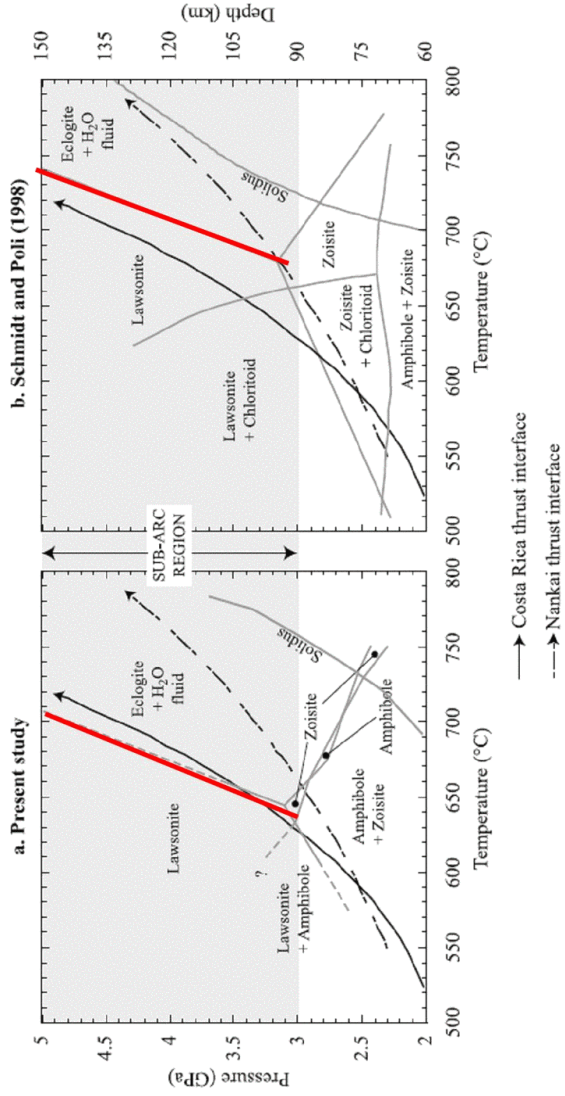


Thermal modeling of the Alaska subduction zone



Schmidt and Poli (1998)

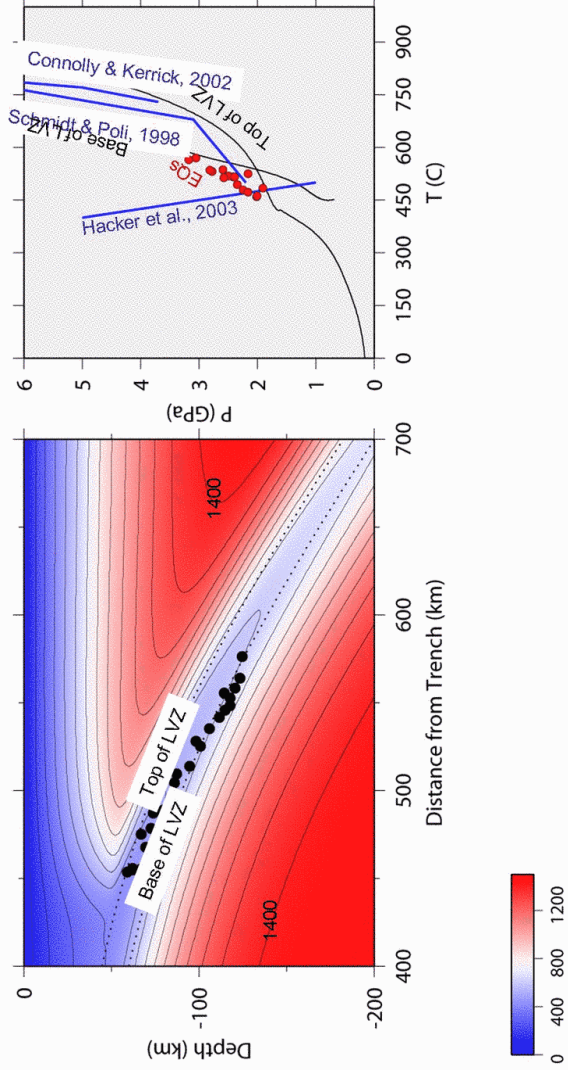
Hacker et al., JGR, 2003



Forneris & Holloway, EPSL, 2003

Base model: Isoviscous wedge  
 55 mm/yr

38 Myr half space cooling model ( $T_m = 1450\text{ C}$ )  
 16 mW/m<sup>2</sup> shear heating (10 MPa)

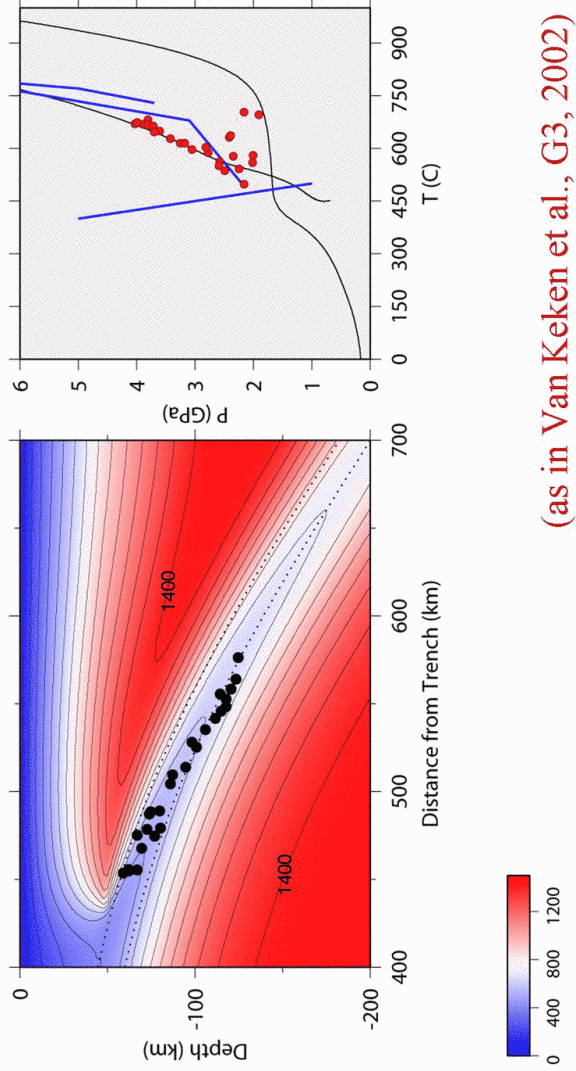


(Similar to Peacock & Gutscher 2003)



Better model: Olivine rheology for mantle wedge

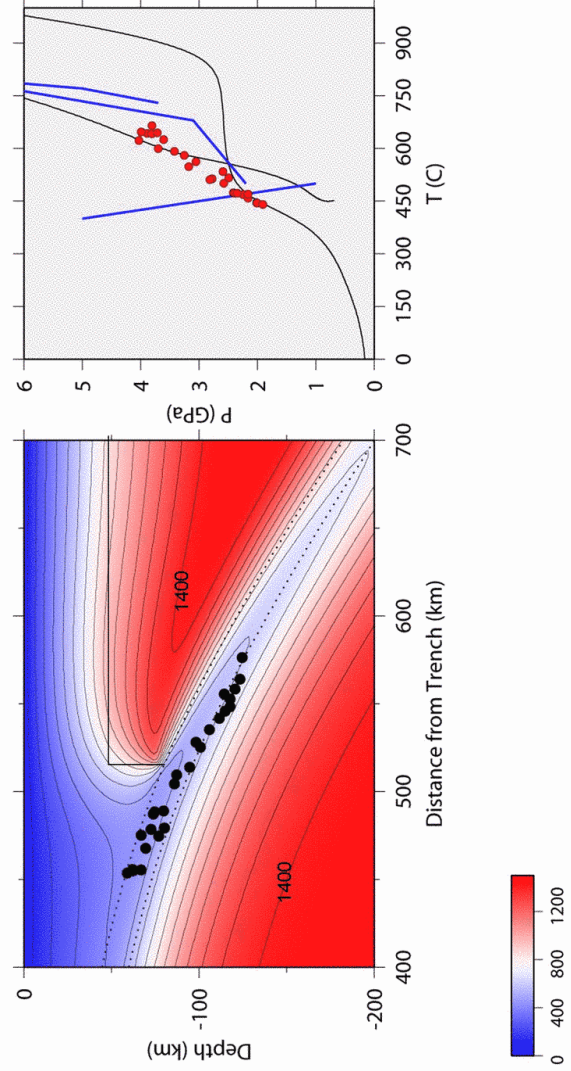
55 mm/yr  
38 Myr  
10 MPa



(as in Van Keken et al., G3, 2002)

Improvement: Olivine rheology with cold nose

55 mm/yr  
38 Myr  
10 MPa

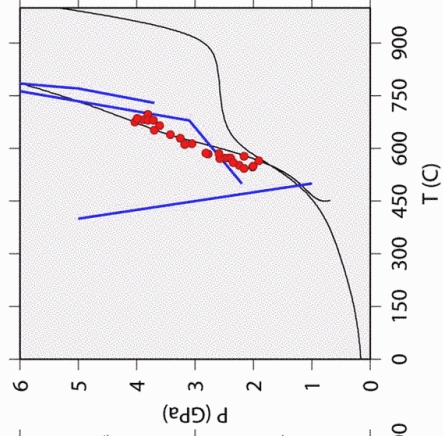
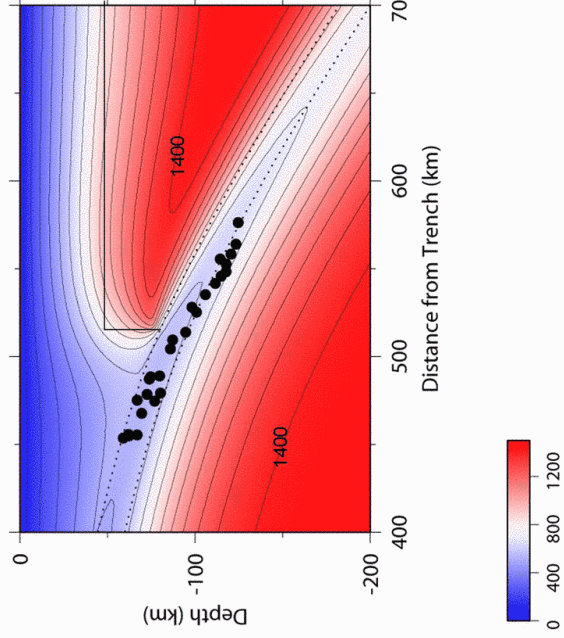


Olivine rheology with cold nose

55 mm/yr

38 Myr

30 MPa

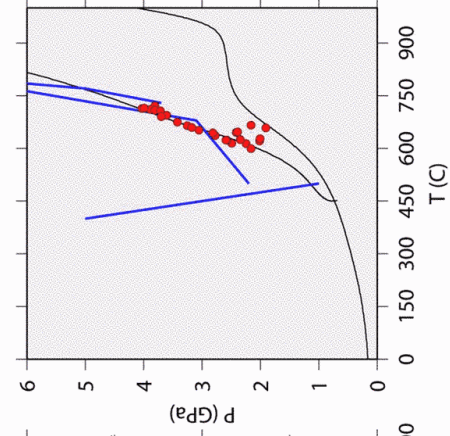
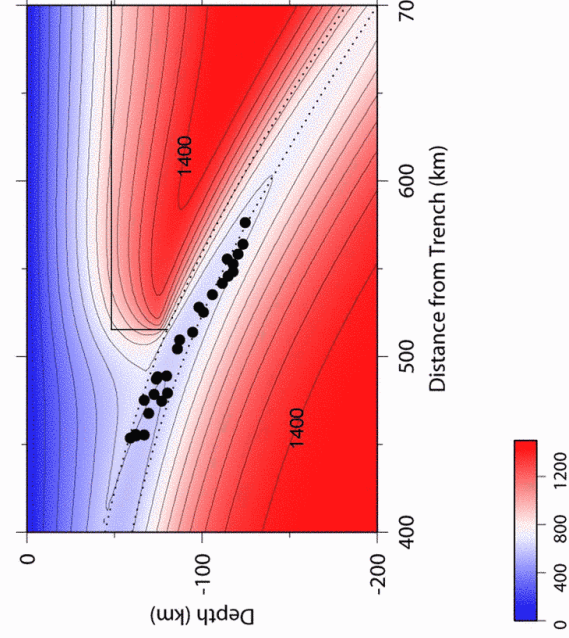


Olivine rheology with cold nose

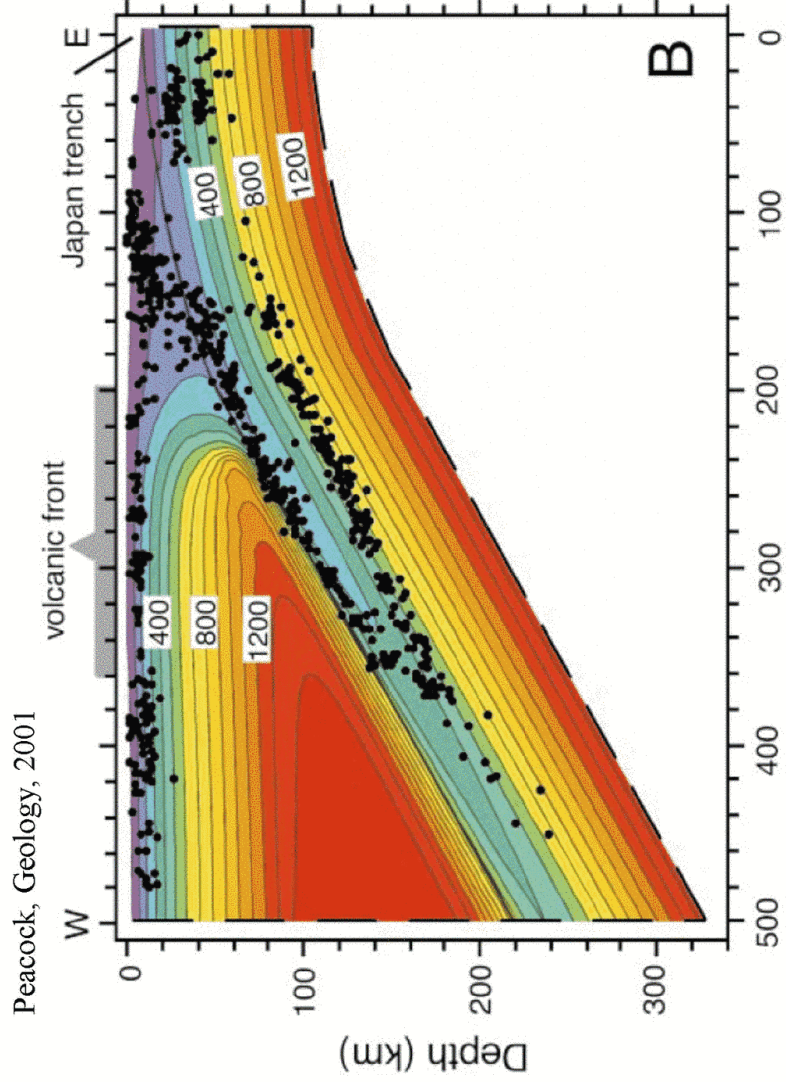
55 mm/yr

38 Myr

45 MPa



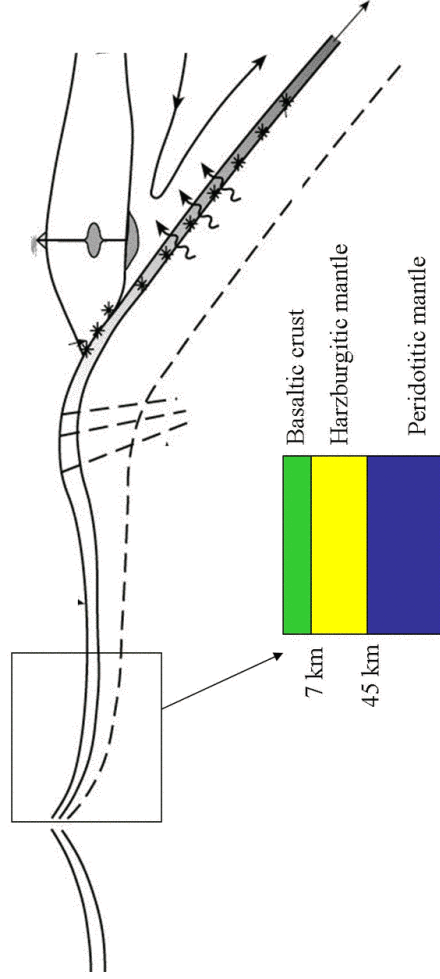




Penrose conference on start of plate tectonics, Lander, WY

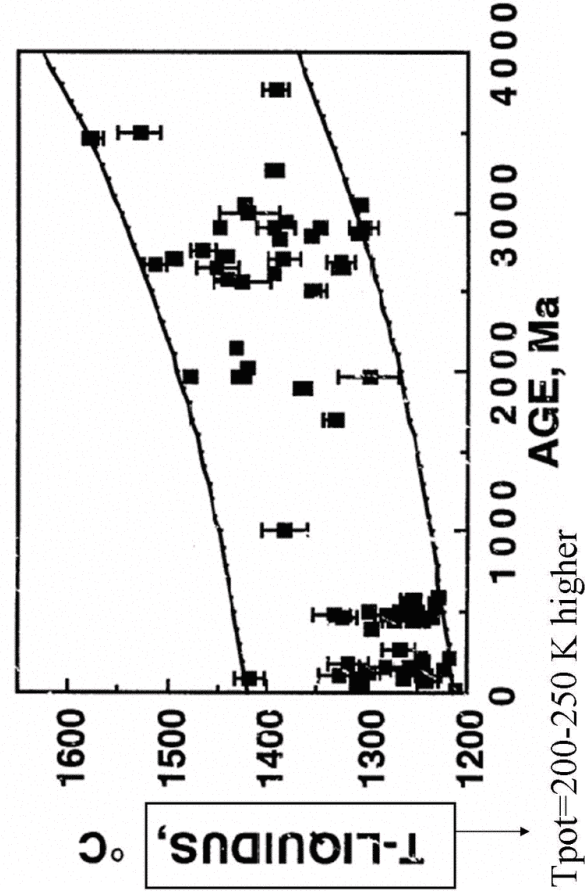


Compositional stratification



Van Keken, EPSL, 2003

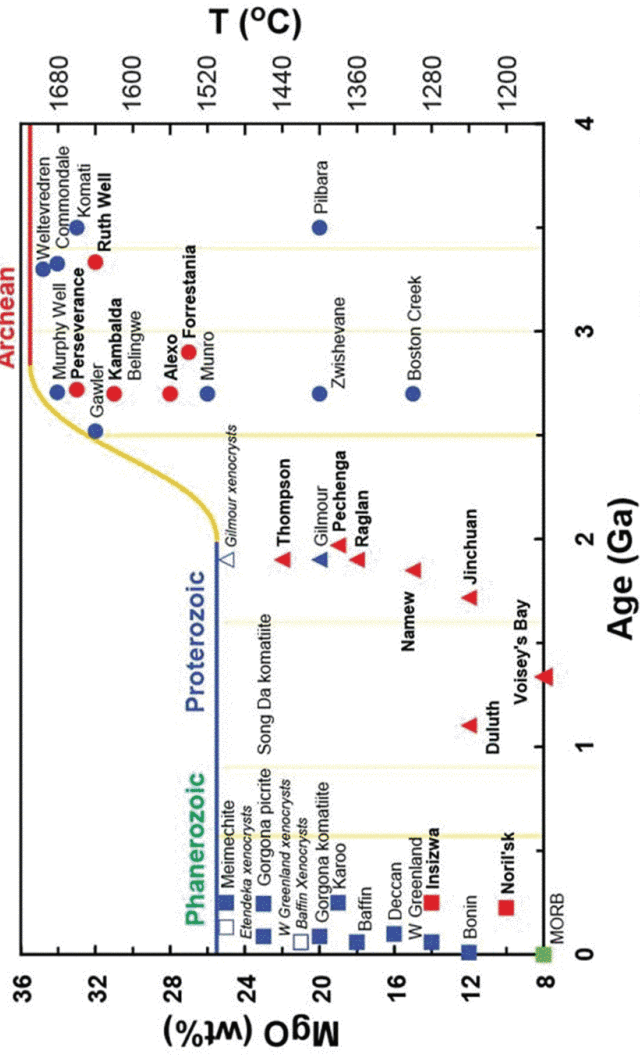
ALL MORB-LIKE SUITES



Abbott et al., JGR, 1994



# High-MgO Magmas vs. Time

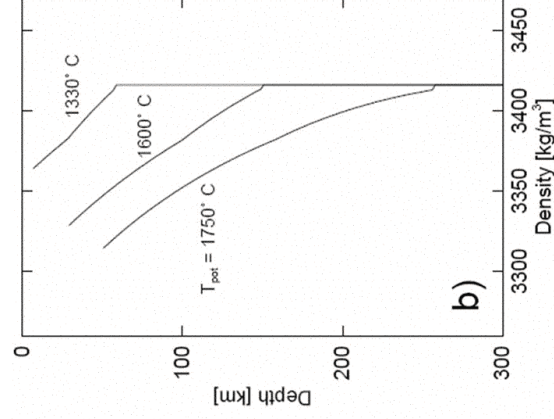


From Charles Lesher

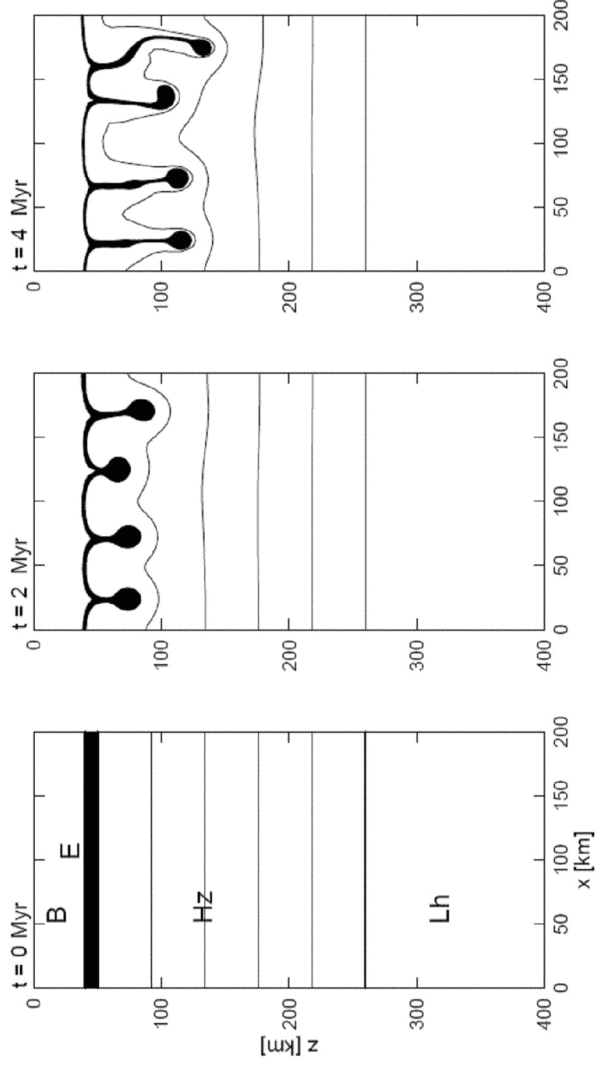
## Compositional stabilization of oceanic lithosphere

### Oceanic lithosphere

30-50 km	basalt	$\rho=3.0 \text{ g/cm}^3$ $\rho=3.315 \text{ g/cm}^3$
250 km	harzburgite	$\rho=3.416 \text{ g/cm}^3$
	lherzolite	$\rho=3.416 \text{ g/cm}^3$



Pre-plate tectonic cooling mechanism (?)

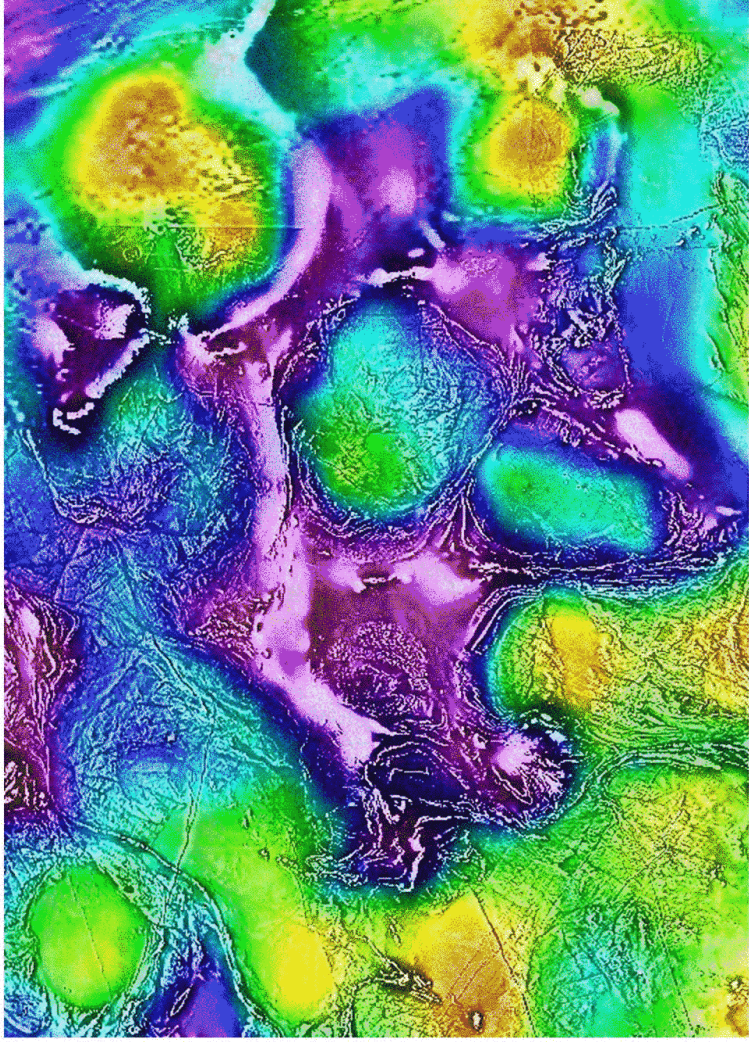


Sufficient to cool to near-present day  
Tpot within 1 Byr

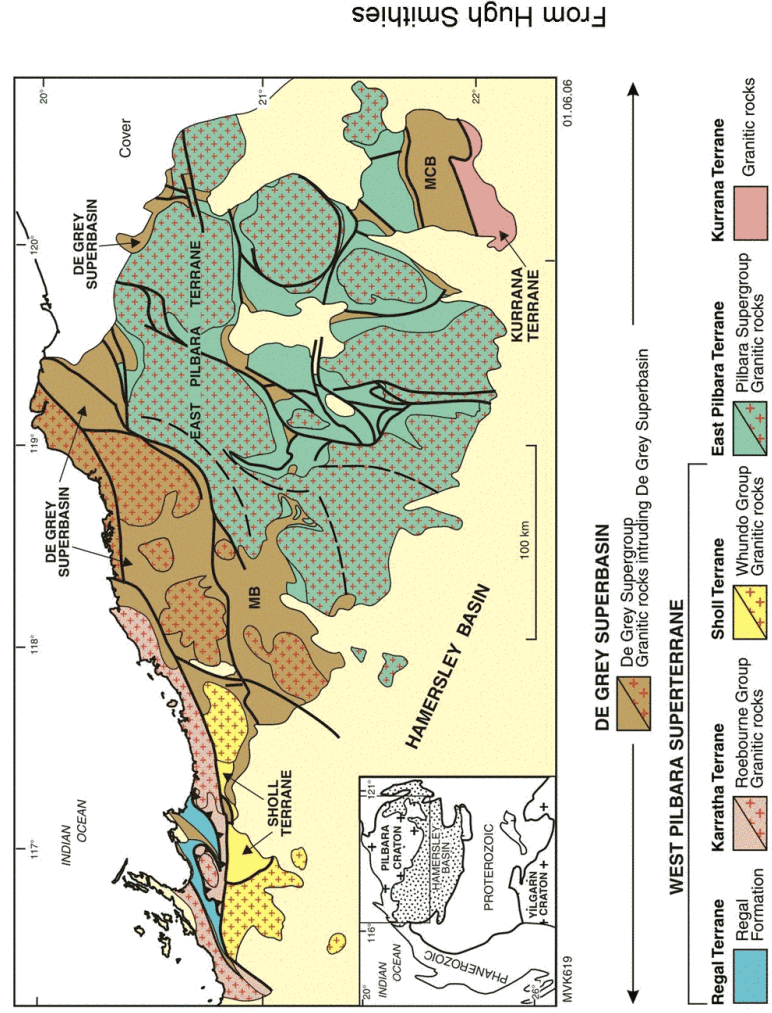
Vlaar et al., EPSL, 1994



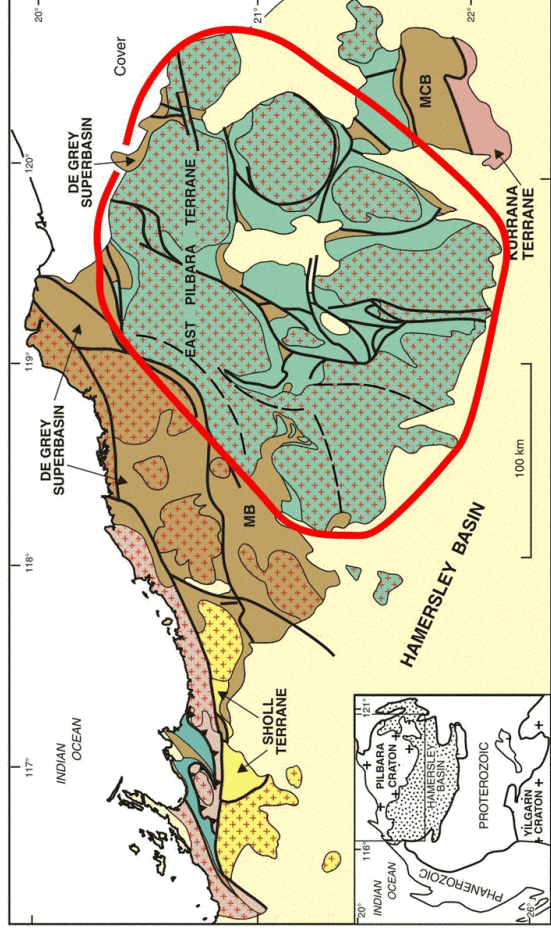




Courtesy of R. Blewett

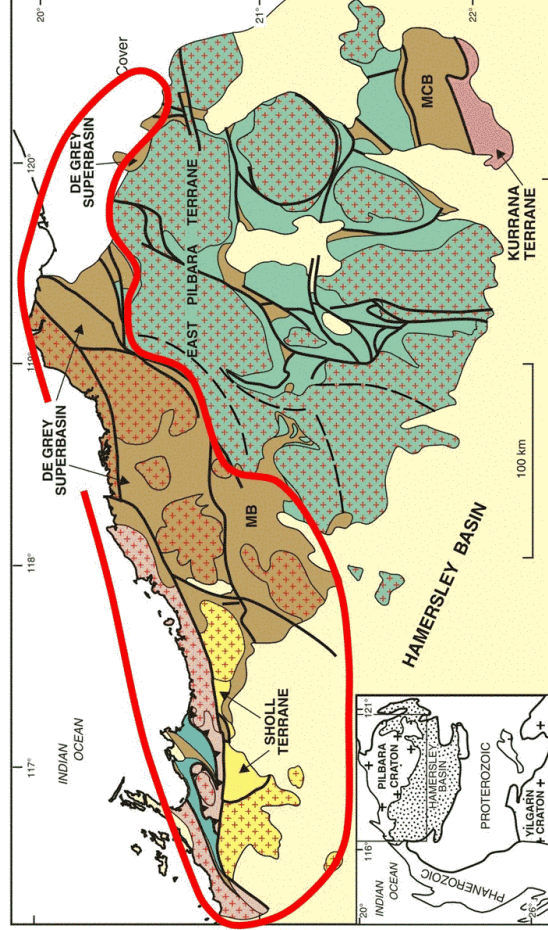






From Hugh Smithies

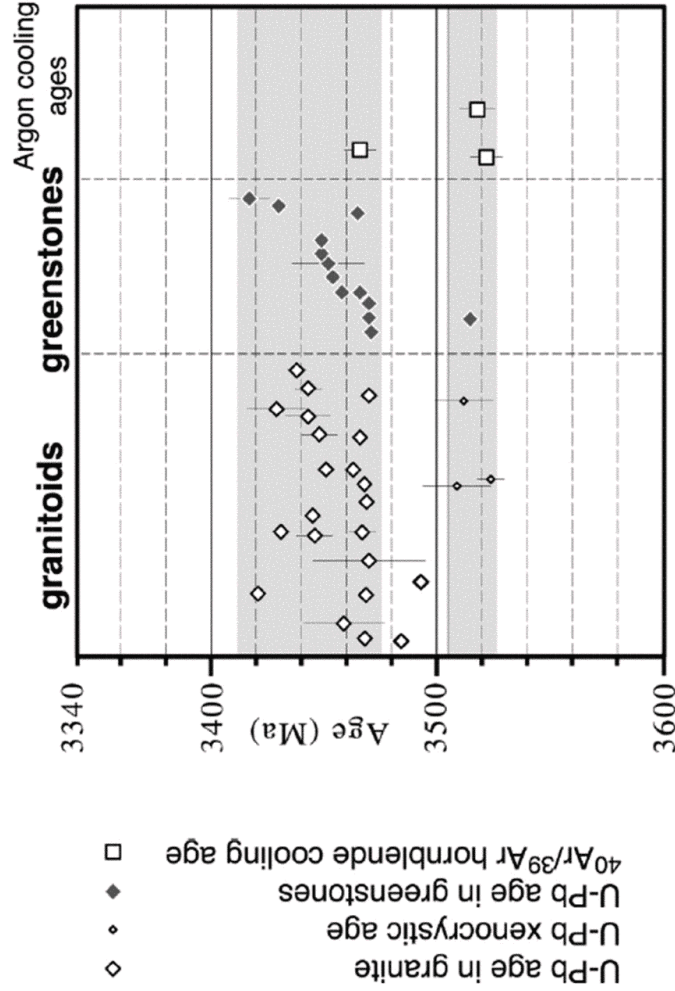
Old (3.53-3.2 Ga) east Pilbara nucleus - contains no clear evidence for modern-style plate tectonics. This thick crustal block most likely began as some form of oceanic plateau type crust.



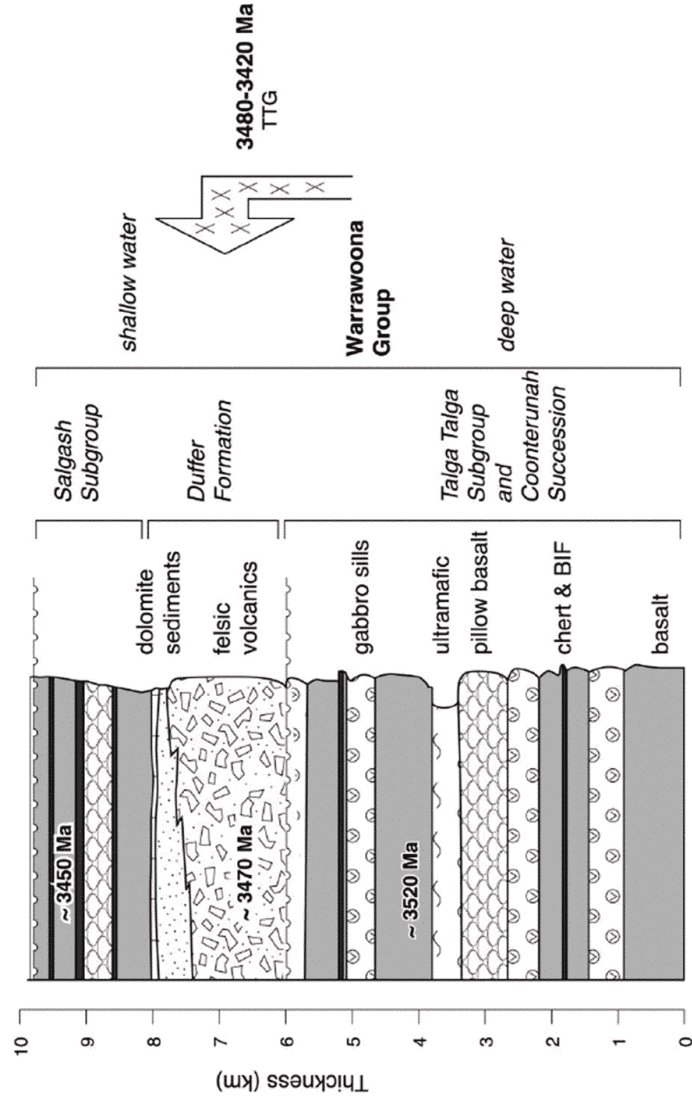
From Hugh Smithies

**The West Pilbara Superterrane**  
 A younger (3.3 - 3.05 Ga) amalgamation of terranes - contains extensive set of features that collectively present a compelling case for modern steep-style subduction at ~3.2 Ga. This represents accretion peripheral to the East Pilbara nucleus.

Zegers and Van Keken, *Geology*, 2001



Zegers and Van Keken, *Geology*, 2001



Zegers and Van Keken, *Geology*, 2001

