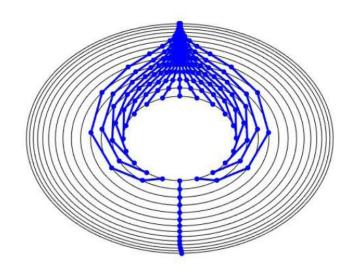


Black holes and overcooled vacuum in driven cavities

Ivar Martin



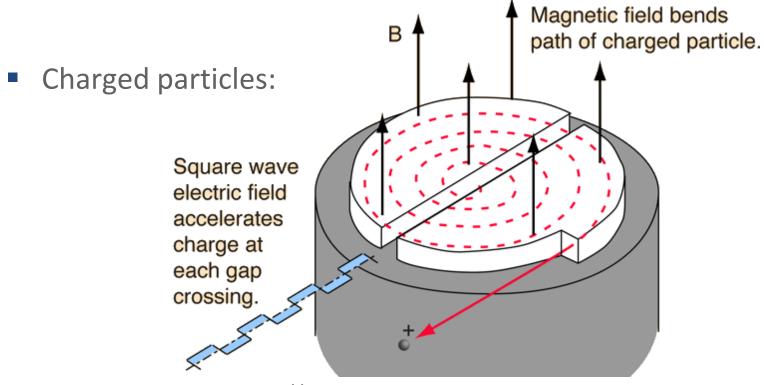
Floquet dynamics of classical and quantum cavity fields, arXiv:1809.02621



Outline

- Photon accelerator
- Maps, Fixed points, Chaos
- Solving wave equation by *Floquet map*
- Energy singularities
- Dynamical Casimir and "supercooled" vacuum
- Signal compression/decompression
- GR analogy

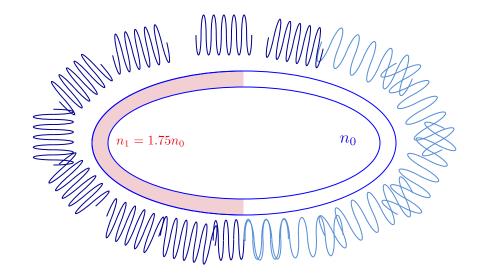
Cyclotron Accelerator



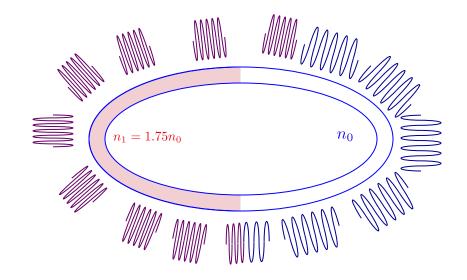
http://hyperphysics.phyastr.gsu.edu/hbase/magnetic/cyclot.html

Same for photons?

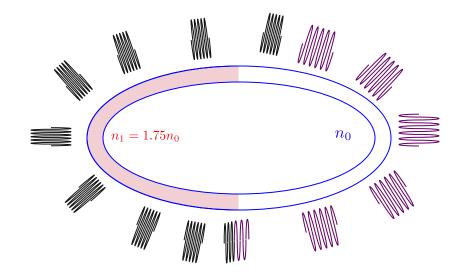
Photon accelerator 1



Photon accelerator 2

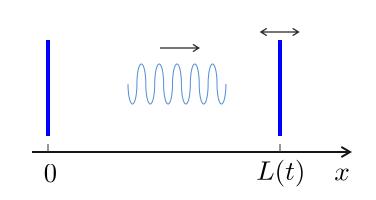


Photon accelerator 3



Alternative: Modulation of physical length

• Cavity with a moving mirror – dynamical Casimir effect

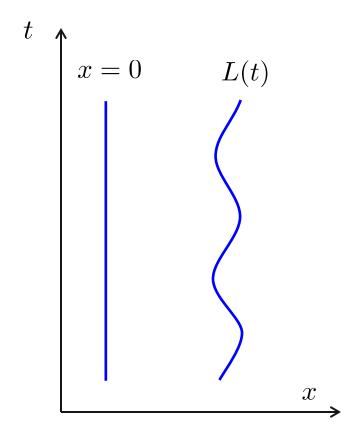


Wave equation

$$\frac{\partial^2 \mathcal{A}}{\partial t^2} - \frac{\partial^2 \mathcal{A}}{\partial x^2} = 0.$$

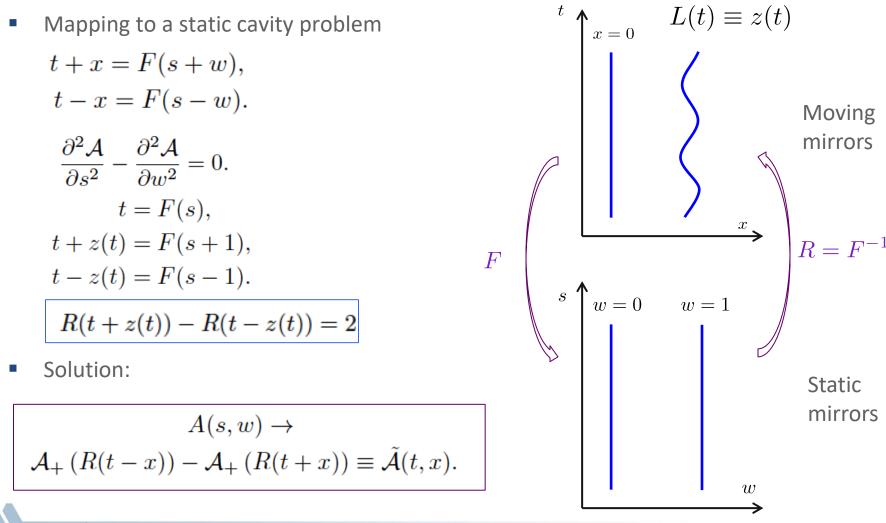
Boundary conditions

$$\mathcal{A}(t,0) = 0 \qquad \qquad \mathcal{A}(t,L(t)) = 0$$



Quantum Theory of the Electromagnetic Field in a Variable-Length One-Dimensional Cavity*[†]





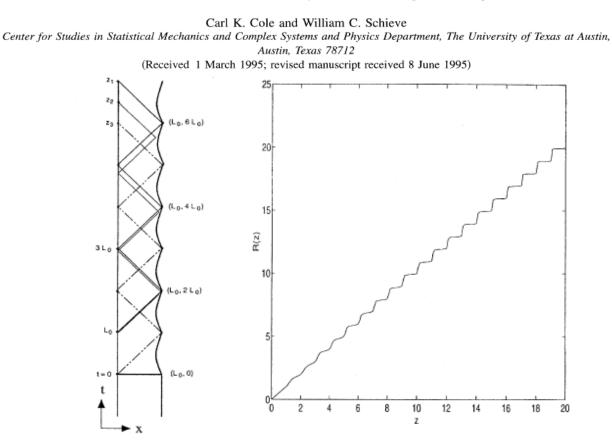
Need conformal map function; near resonance becomes singular!

PHYSICAL REVIEW A

VOLUME 52, NUMBER 6

DECEMBER 1995

Radiation modes of a cavity with a moving boundary



Consequences of singularity: energy density peaks

VOLUME 73, NUMBER 14

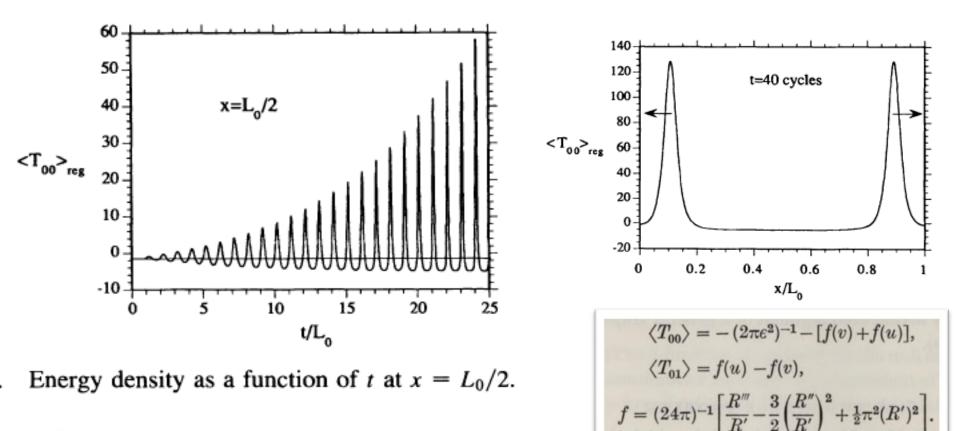
PHYSICAL REVIEW LETTERS

3 October 1994

Resonance Response of the Quantum Vacuum to an Oscillating Boundary

C. K. Law

Department of Physics and Astronomy, University of Rochester, Rochester, New York 14627 (Received 20 December 1993)



Outline

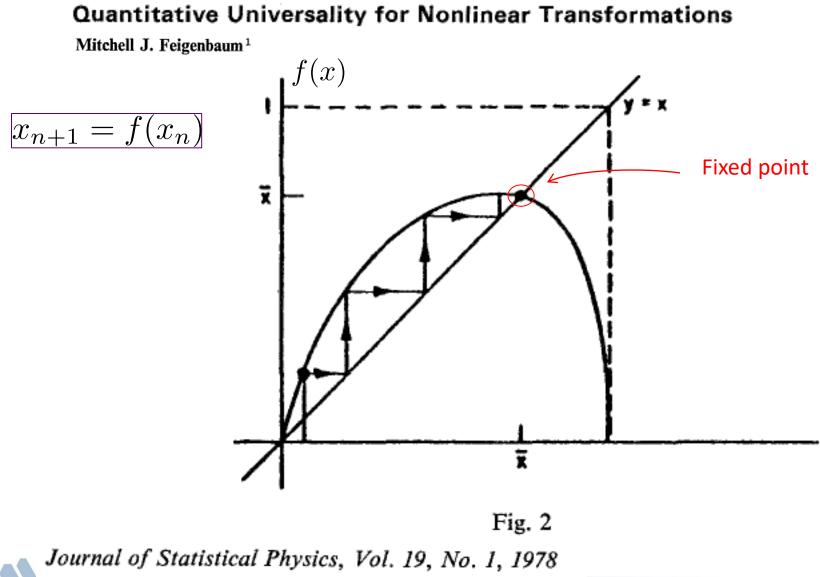
Photon accelerator

Maps, Fixed points, Chaos

- Solving wave equation by *Floquet map*
- Energy singularities
- Dynamical Casimir and "supercooled" vacuum
- Signal compression/decompression

GR analogy

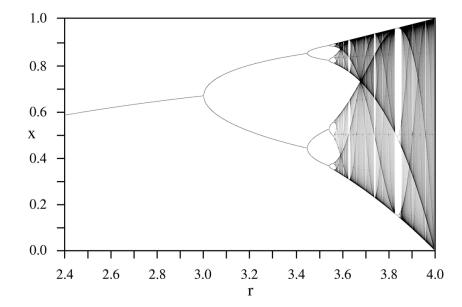
A different kind of map: dynamical systems



Bifurcations, chaos, fractals

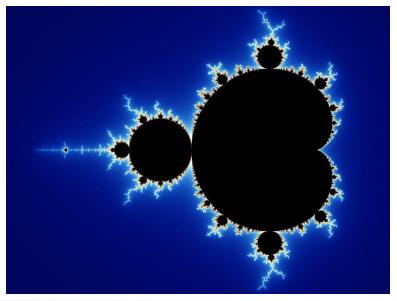
Real logistic map

$$x_{n+1} = rx_n(1-x_n).$$



 Complex logistic map (Mandelbrot set)

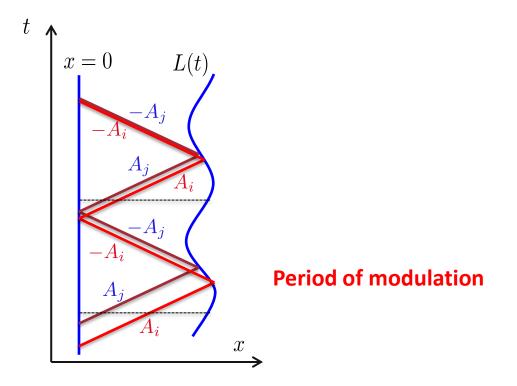
$$z_{n+1} = z_n^2 + c$$



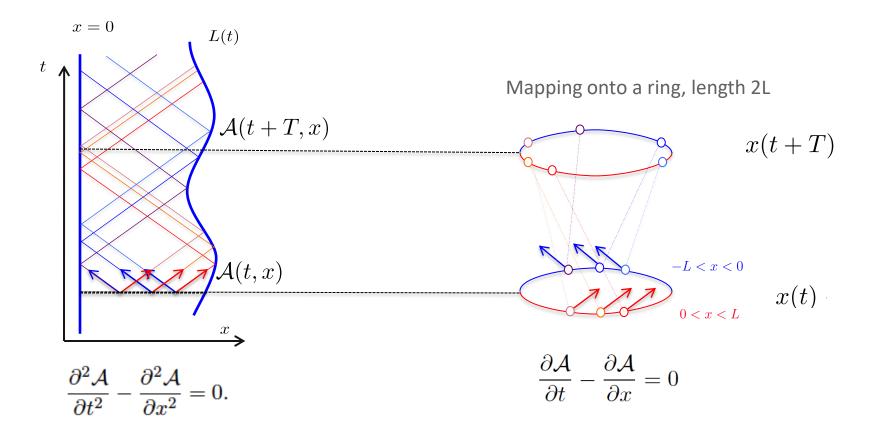
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Transport of A (vector potential) along characteristics



Cavity field problem as a map



Vector potential remains constant on characteristics (= light rays = null lines)

One period map

$$t_m - t_0 = \frac{L(t_m) - x_0}{c}$$

$$L(t_m) - x_0 + L(t_m) + x_1 = cT$$

$$x_{n+1} = f(x_n)$$

$$t_0 + T$$

$$t_m$$

$$t_0$$

$$t_m$$

$$t_0$$

$$L(t_m)$$

$$\mathcal{A}(x_0, t_0) = (\pm)\mathcal{A}(x_1, t_0 + T) = (\pm)\mathcal{A}(x_2, t_0 + 2T) = \dots$$

Weak modulation - explicit map

Approximation:

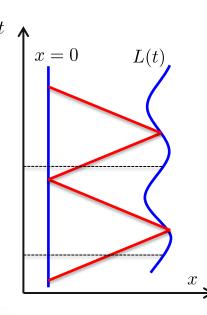
$$t_m - t_0 = \frac{L(t_m) - x_0}{c} \longrightarrow t_m \approx t_0 + \frac{L_0 - x_0}{c}$$
$$x_1 = x_0 + cT - 2L(t_m) \longrightarrow x_1 = x_0 + cT - 2L(\phi_0 - x_0)$$

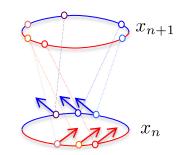
Single-step map

$$x_{n+1} = x_n + cT - 2L(\phi_0 - x_n)$$

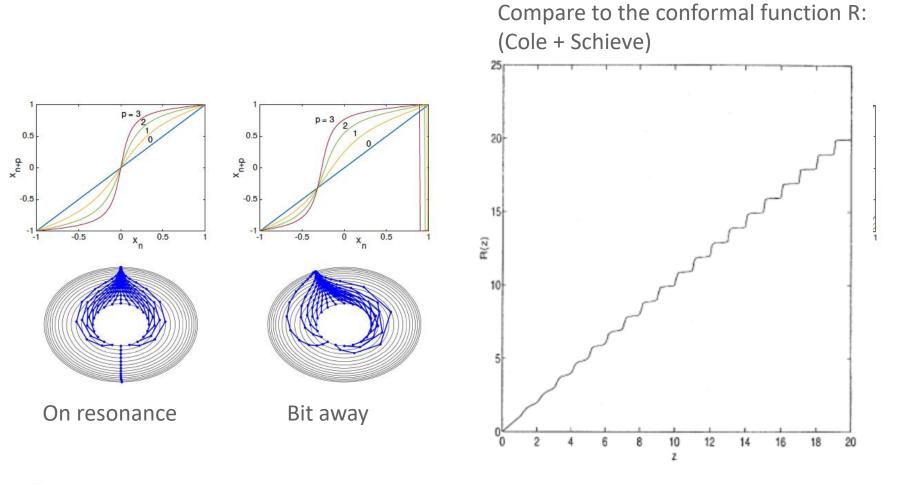
Multi-step map, continuous limit

$$\frac{dx}{dn} = cT - 2L(\phi_0 - x)$$

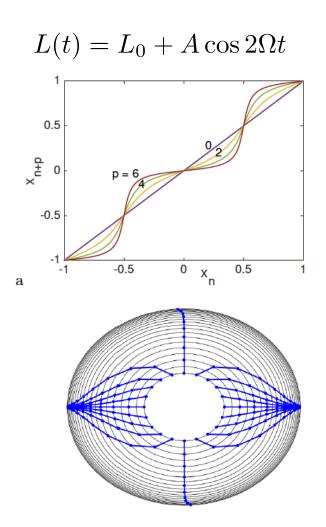




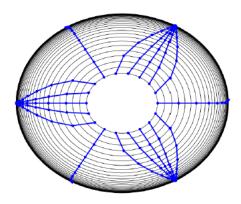
 $L(t) = L_0 + A\cos\Omega t$



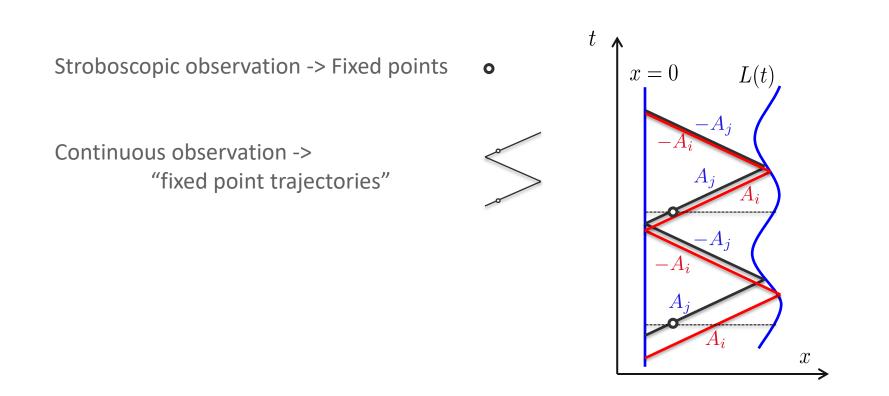
Driving near higher cavity resonances



 $L(t) = L_0 + A\cos 3\Omega t$



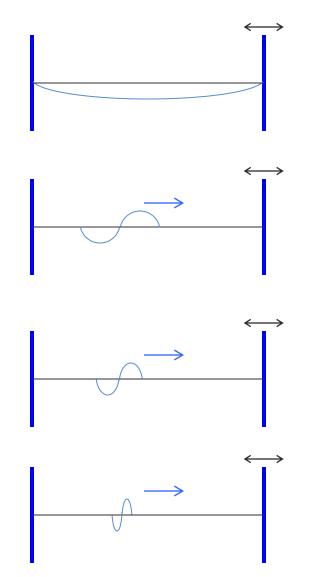
Back to continuous time



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Qualitative picture



$$k(t) \sim k_0 \left(\frac{c+v}{c-v}\right)^{t/T} \approx k_0 e^{\frac{2vt}{cT}}$$

v Mirror velocity at fixed point

Energy Density

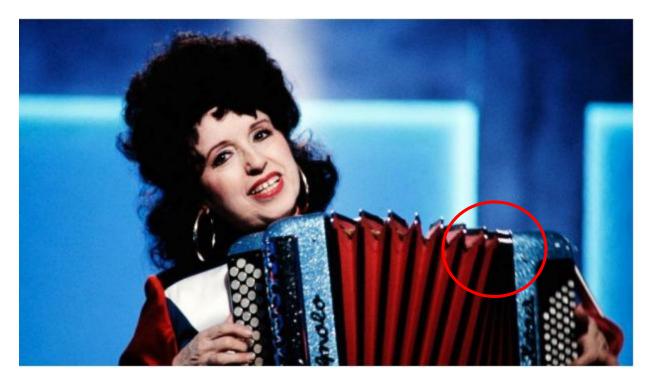
$$(\partial_x \mathcal{A})^2 + (\partial_t \mathcal{A})^2 \sim k^2 \mathcal{A}^2$$

Total Energy

 $[(\partial_x \mathcal{A})^2 + (\partial_t \mathcal{A})^2]/k \sim k \mathcal{A}^2$

Both grow exponentially with time

A known effect...



http://en.rfi.fr/20180612-frances-queen-accordeon-yvette-horner-dies/

Energy and energy density

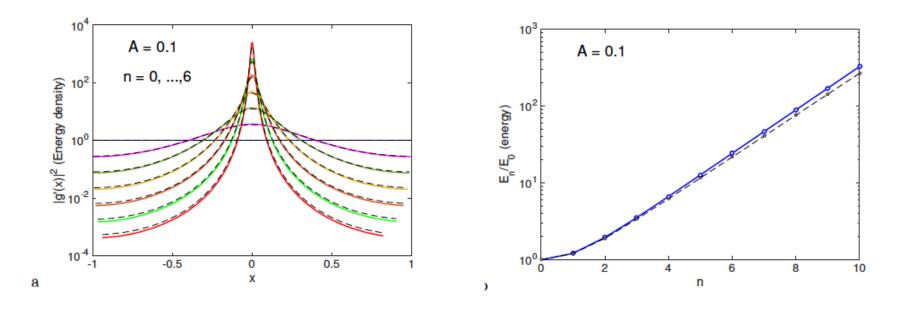
$$\begin{split} E(t) &= \int dx_t T_{00}(x_t, t) \\ &= \frac{1}{2} \int dx_t \left[\left(\frac{\partial \mathcal{A}(x_t, t)}{\partial x_t} \right)^2 + \left(\frac{\partial \mathcal{A}(x_t, t)}{\partial t} \right)^2 \right] \\ &= \int dx_t \left[\frac{\partial g^{(t)}(x_t)}{\partial x_t} \right]^2 \left[\mathcal{A}_0'(g^{(t)}(x_t)) \right]^2 \\ &= \int dx_0 \frac{\partial g^{(t)}(x_t)}{\partial x_t} \left[\mathcal{A}_0'(x_0) \right]^2 \\ &= \int dx_0 \frac{1}{\partial f^{(t)}(x_0) / \partial x_0} \left[\mathcal{A}_0'(x_0) \right]^2 \end{split}$$

$$g = f^{-1}$$

Weak harmonic modulation, $\, {\tilde A} = \pi A/L \ll 1 \,$

$$\tilde{x}_0 = \tilde{g}(\tilde{x}_n) = 2 \arctan(e^{-2\tilde{A}n} \tan \frac{\tilde{x}_n}{2})$$

Energy and energy density, numerical



Over time, energy density increases (exponentially) near fixed point trajectories, and **decreases (exponentially) everywhere else**.

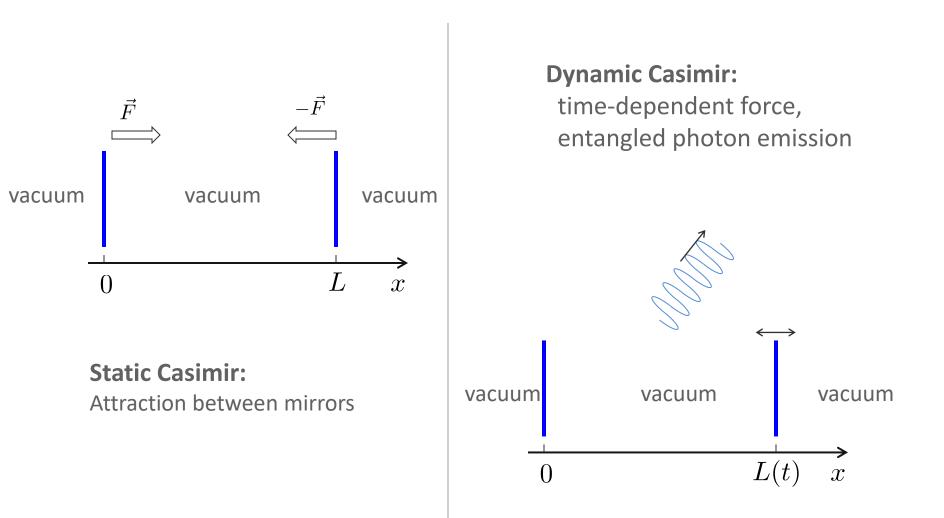
Uses:

- 1) pulsed energy source (similar to Mode-Locked laser)
- 2) Cavity cooling (Use a "dumb" Maxwell's demon)

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Casimir effects



Field quantization

[Moore 1970; Filling and Davies, Proc R Soc Lond A 1976; P. D. Nation et al, RMP 2012]

Conformal map function:

$$R(t+z(t))-R(t-z(t))=2$$

Mode function:

$$\phi_n(x,t) = (4\pi n)^{-1/2} [e^{-i\pi nR(t+x)} - e^{-i\pi nR(t-x)}]$$

Quantized field

$$\phi(x, t) = \sum_{n} a_n \phi_n(x, t) + a_n^{\dagger} \bar{\phi}_n(x, t)$$

Energy density

$$\begin{split} \langle T_{00} \rangle &= -\left(2\pi e^2\right)^{-1} - [f(v) + f(u)], \\ \langle T_{01} \rangle &= f(u) - f(v), \\ f &= (24\pi)^{-1} \bigg[\frac{R'''}{R'} - \frac{3}{2} \Big(\frac{R''}{R'} \Big)^2 + \frac{1}{2} \pi^2 (R')^2 \bigg]. \end{split}$$

Schwarzian

Static Casimir

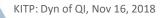
1D Casimir results

- Driving
 - Harmonic
 - Mirror oscillation amplitude $\,A\,$

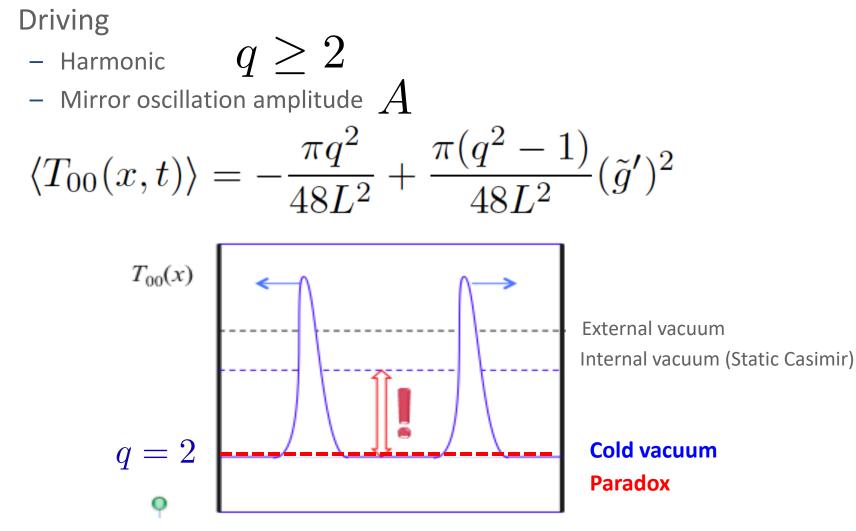
$$\langle T_{00}(x,t) \rangle = -\frac{\pi q^2}{48L^2} + \frac{\pi (q^2 - 1)}{48L^2} (\tilde{g}')^2$$
$$\frac{g'(x)}{1} = \frac{1}{\cosh(2qn\tilde{A}) + \sinh(2qn\tilde{A})\cos q\tilde{x}}$$

• Energy density for static (A = 0), and first hamonic ($q = 1, A \neq 0$) π

$$\langle T_{00}(x,t) \rangle = -\frac{\pi}{48L^2}$$



Casimir results, 2



Proc. R. Soc. Lond. A. 348, 393-414 (1976) Printed in Great Britain

Radiation from a moving mirror in two dimensional space-time: conformal anomaly

BY S. A. FULLING AND P. C. W. DAVIES

Department of Mathematics, University of London, King's College

(Communicated by R. Penrose, F.R.S. - Received 13 August 1975)

The energy-momentum tensor is calculated in the two dimensional quantum theory of a massless scalar field influenced by the motion of a perfectly reflecting boundary (mirror). This simple model system evidently can provide insight into more sophisticated processes, such as particle production in cosmological models and exploding black holes. In spite of the conformally static nature of the problem, the vacuum expectation value of the tensor for an arbitrary mirror trajectory exhibits a non-vanishing radiation flux (which may be readily computed). The expectation value of the instantaneous energy flux is negative when the proper acceleration of the mirror is increasing, but the total energy radiated during a bounded mirror motion is positive. A uniformly accelerating mirror does not radiate; however, our quantization does not coincide with the treatment of that system as a 'static universe'. The calculation of the expectation value requires a regularization procedure of covariant separation of points (in products of field operators) along time-like geodesics; more naïve methods do not yield the same answers. A striking example involving two

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Discrete time reversal: Compression/decompression

For a fixed stroboscopic sampling, it is possible to construct discrete timereversal protocol:

$$Z_{TR}(t) = \begin{cases} z(t), & t < t^* \\ 2L_0 - z(t), & t \ge t^*. \end{cases}$$

ь

Signal compression/decompression

L

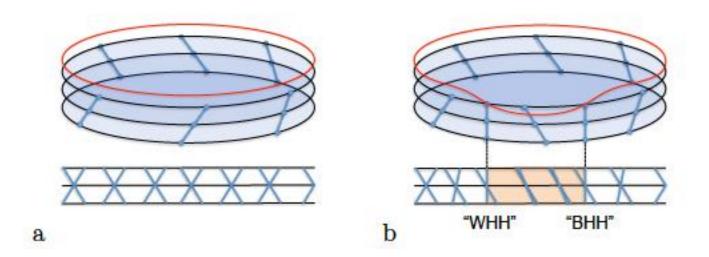
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"Stroboscopic GR"

For a fixed stroboscopic sampling, the light ray evolution mimics propagation in curved spaces

(stable/unstable) fixed point <-> (white/black) hole horizon



For details see Martin, arXiv:1809.02621

Summary

- Problem of periodically modulated cavities can be solved with dynamical maps
- The effect of maps is to concentrate energy around the fixed point trajectories
- Applications: cooling, compression/coding
- Qualitative picture of dynamical Casimir effect (vacuum squeezing)
- Connections to GR

Future

- Manipulation of individual photon states
- Understand overcooled vacuum (entanglement?)
- Application to cooling tech (can use phoNons instead of phoTons)
- Other systems (BEC, etc)

Thanks

- D. Orgad
- B. I. Halperin
- E. Demler
- V. Rosenhaus
- H. Kapteyn
- K. Schwab
- KITP

