

# *Finding Other Worlds with*

**SUPERWASP**

& **LCOGT.net**

*Rachel Street & Tim Lister*

- *The Story so far...*
- *Cosmic alignments*



- *Scanning the sky: SuperWASP*
- *Sorting the wheat from the chaff*
- *New Discoveries*

*Lynette Cook*

**UCSB**

# *The Story So Far...*

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1992: Surprise no. 1: First planets discovered around another star...a pulsar [1]

1995: Surprise no. 2: 51 Peg b, first planet orbiting a Solar-type star [Hot Jupiter]

*Migration theory, revisited*

1998: First two multi-planet systems found

1999: HD 209458 b, first transiting planet discovered [2,3]

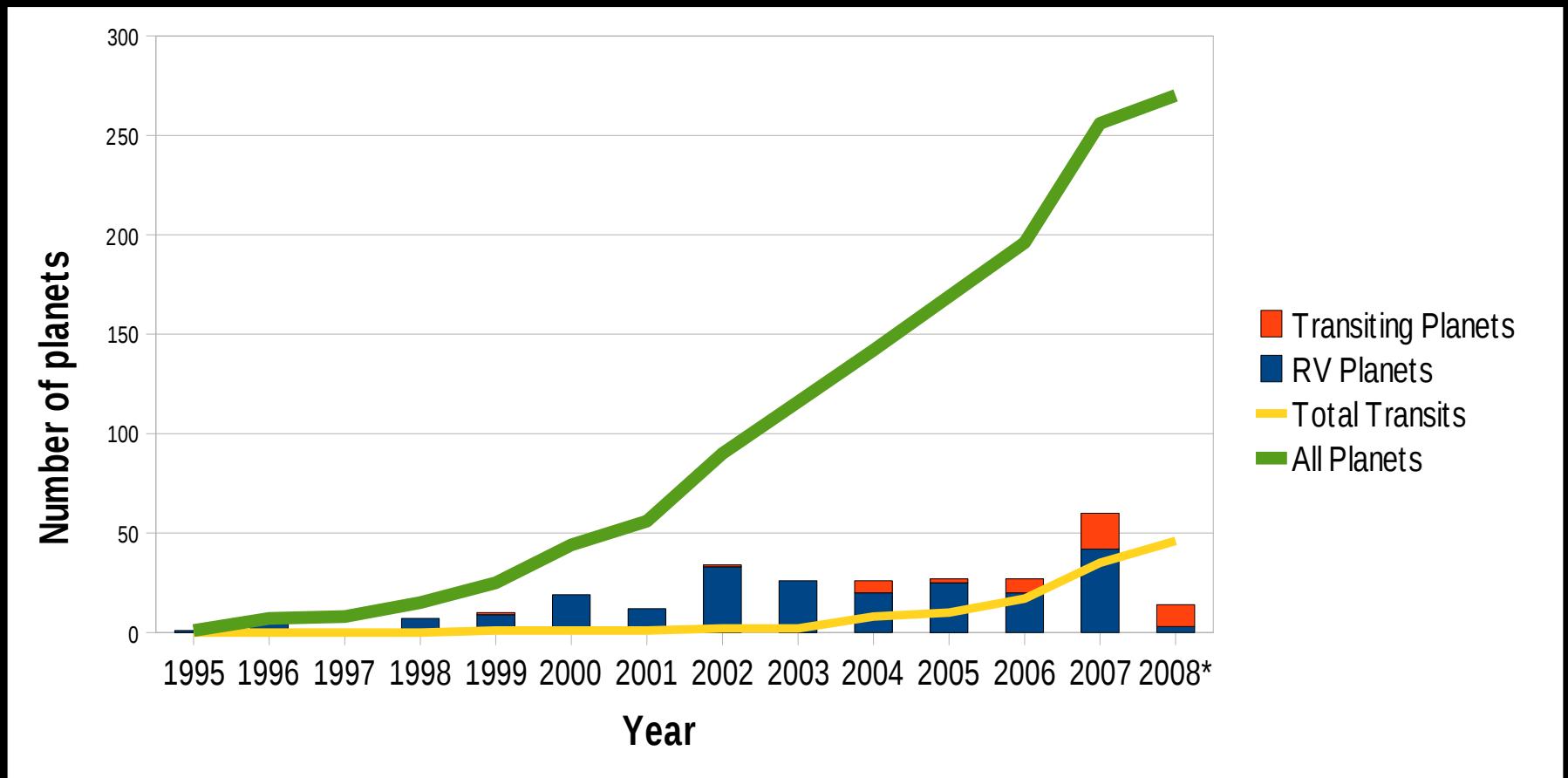


2001: > 50 known exoplanets

2003: > 100 known exoplanets

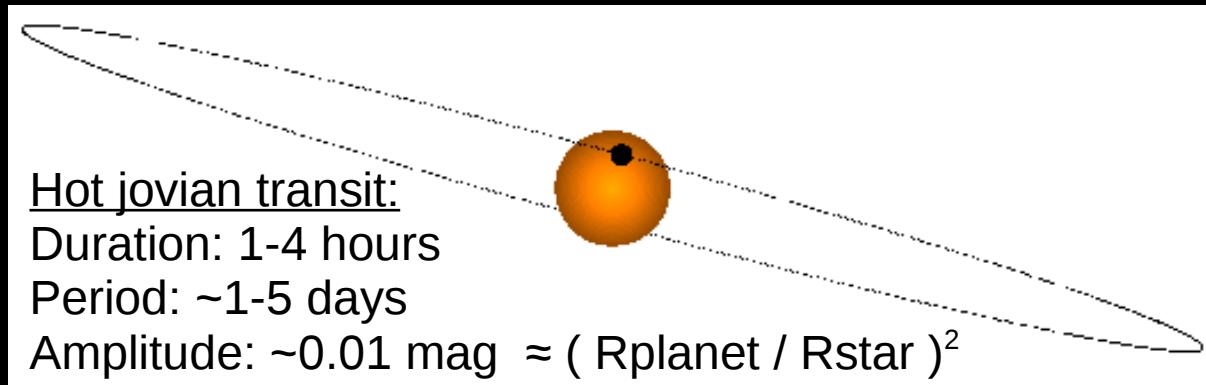
2007: > 200 known exoplanets

# *Planet discoveries*



2000: WASP0 prototype  
> much work <  
2004: SuperWASP-N operational  
2006: SuperWASP-S operational

# *Cosmic Alignments*



→ Selection effects favour:  
Large planets  
Small stars  
Short period

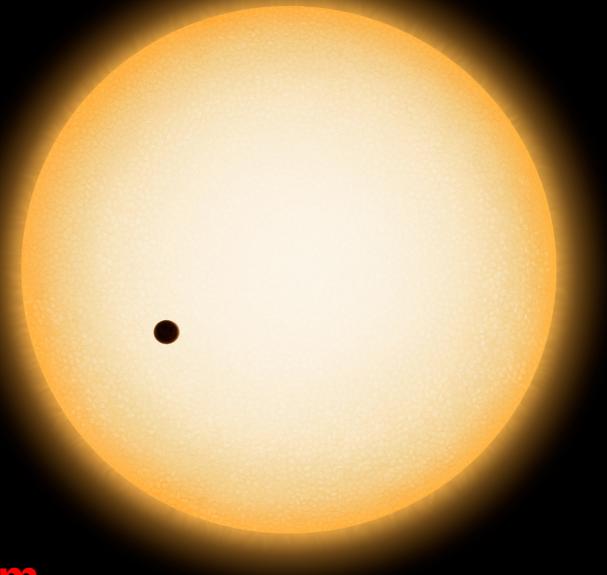
~10% of hot Jupiter class planets should transit

RV: ~1% of solar neighborhood stars harbor hot Jupiters  
~50% of stars are dwarfs

**Need to monitor very large numbers of stars!**

**Transit data + RV = complete characterization of system**

**Wide range of further study possible**



# *Scanning the skies:*



- Robotic, dedicated ultra-wide field photometric survey
- 8 cameras per station
  - 486.72 sq deg. total field of view
  - ~ 200,000 – 1,000,000 stars measured every exposure
- Two stations: La Palma (from 2004)  
South Africa (from 2006)
- Observe planets fields every 8mins  
All sky survey fields at least every day
- ~58 GB raw image data / station / night
- Photometric database inc. 24 million unique objects  
>113 billion datapoints
- Automated data reduction and transit search



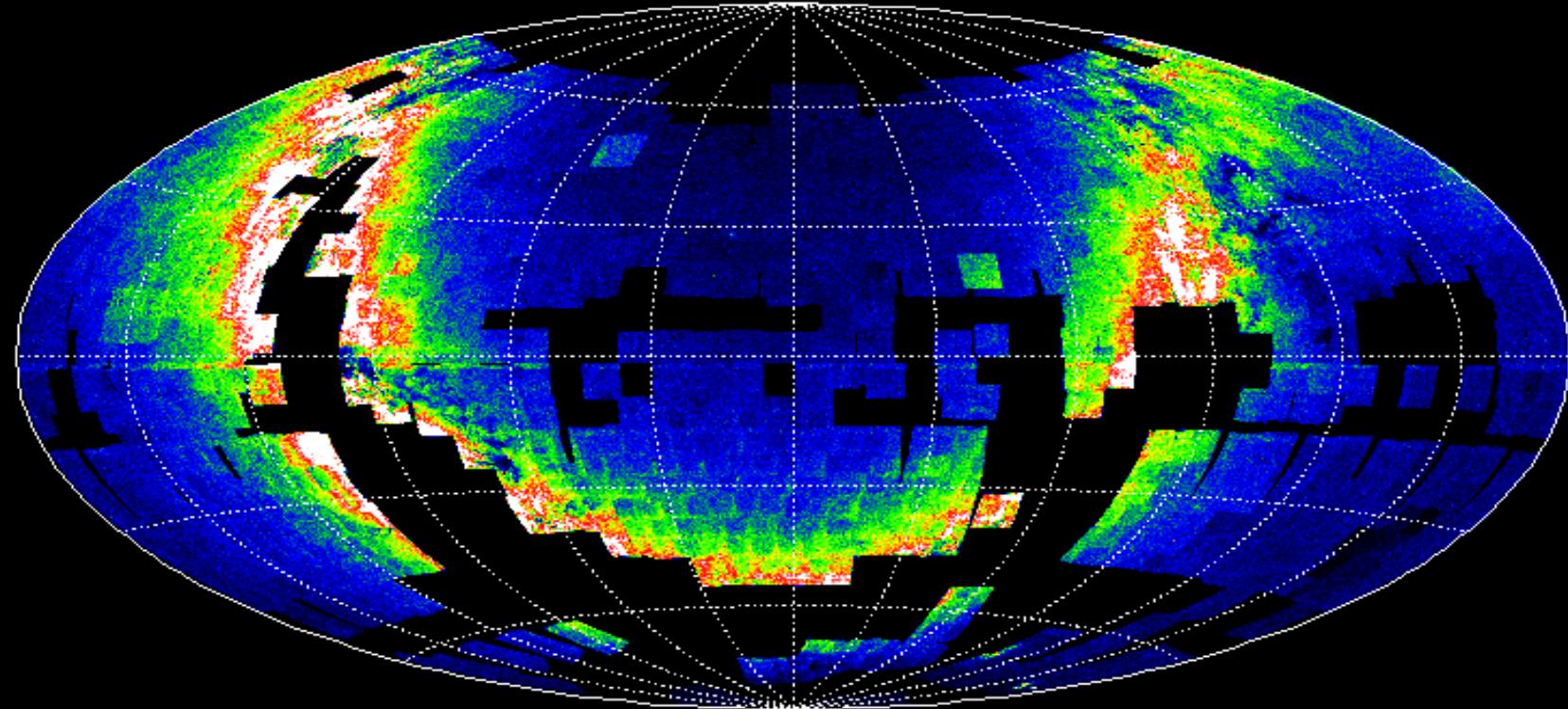
[www.superwasp.org](http://www.superwasp.org)

Pollacco, D. et al. (2006), Ap&SS, 304, p.253.

# *Co-ordinated Global Follow-Up Program*

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- Currently 363 'A' transit candidates widely distributed in RA & Dec
- Unavoidably high fraction of false positives
  - Need large-scale, efficient and flexible follow-up to pre-select candidates



# *Co-ordinated Global Follow-Up Program*

- LCOGT is a major contributor to the 2007-2008 follow-up campaign

## Photometry

48 hrs  
60 hrs  
175 hrs  
~60 hrs  
>150 hrs

IAC 80cm (Tenerife)  
Liverpool Telescope (La Palma)  
24-inch (Keele,UK)  
1m JGT (St. Andrews)  
Swiss 1.2m/*EulerCam*

>520 hrs  
23 hrs  
41 hrs

Tenagra 0.81m (Arizona)  
FTN (Hawai'i)  
FTS (Siding Spring)

## Spectroscopy

64 hrs  
16 hrs  
56 hrs  
40 hrs  
80-120 hrs

NOT/FIES (La Palma)  
TNG/SARG (La Palma)  
INT/IDS (La Palma)  
OHP-1.93m/*Sophie* (France)  
Swiss 1.2m/*Coralie* (ESO, La Silla)



# *The Scoreboard To Date...*

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## Tenagra

59 targets observed to date

11 rejected (identified EB neighbor blended in SW data)

7 no transit seen (probable ephemeris drift)

14 need more data, no conclusions yet (e.g. poor data)

5 low-mass eclipsing binaries

3 triple star systems

2 targets with data to be reduced

11 promising candidates need more data

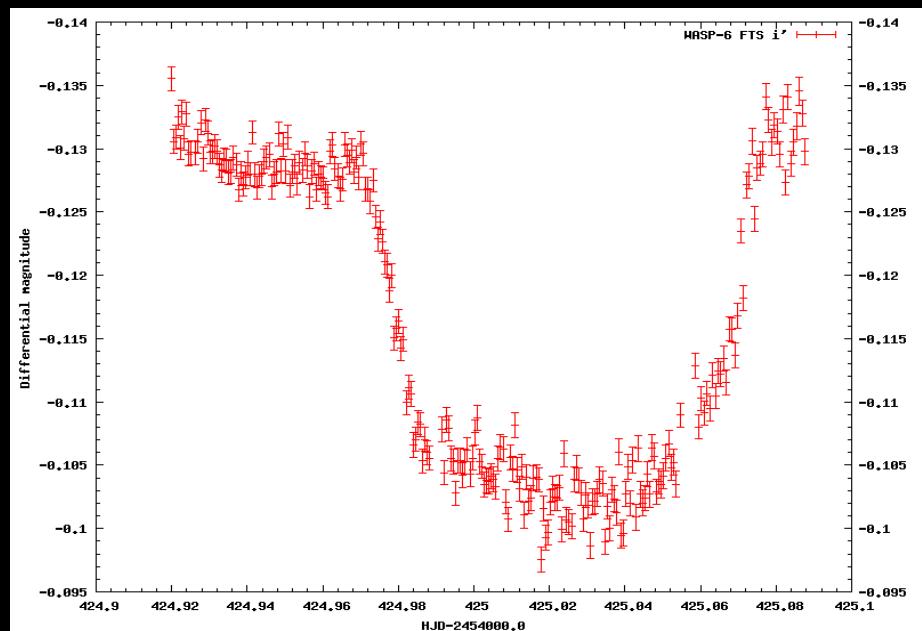
All 5 new northern-hemisphere confirmed planets

1 new southern-hemisphere confirmed planet

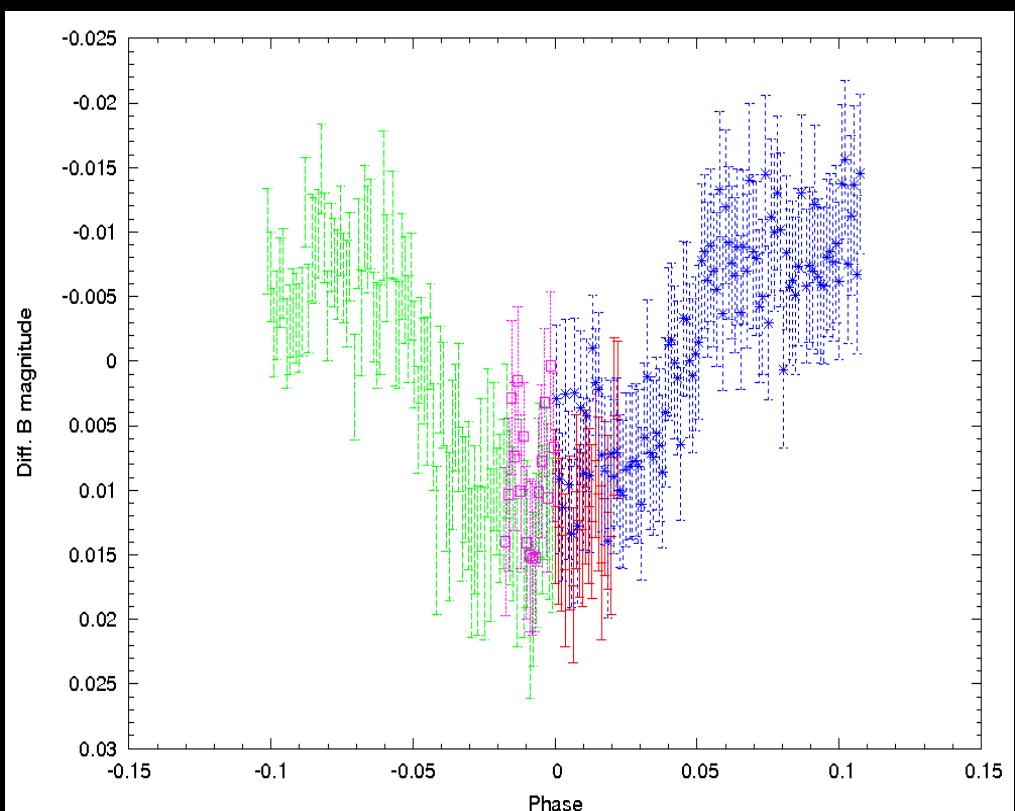
**FTN**      11 targets observed      inc. 2 confirmed planets

**FTS**      12 targets observed      inc. 3 confirmed planets

# *Data on Top Candidates*

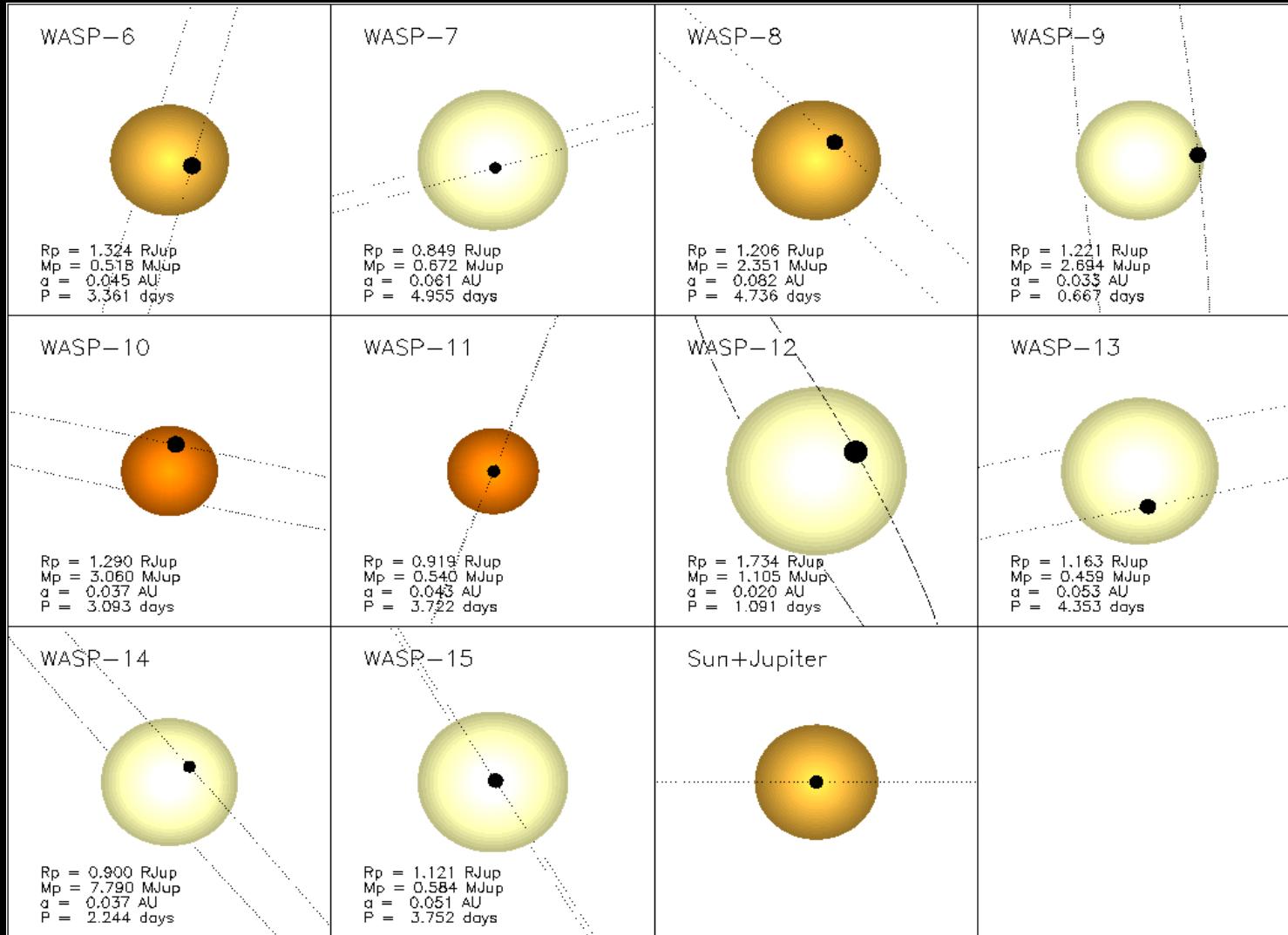


FTS SDSS-i band data, WASP 6  
2007 Nov 20

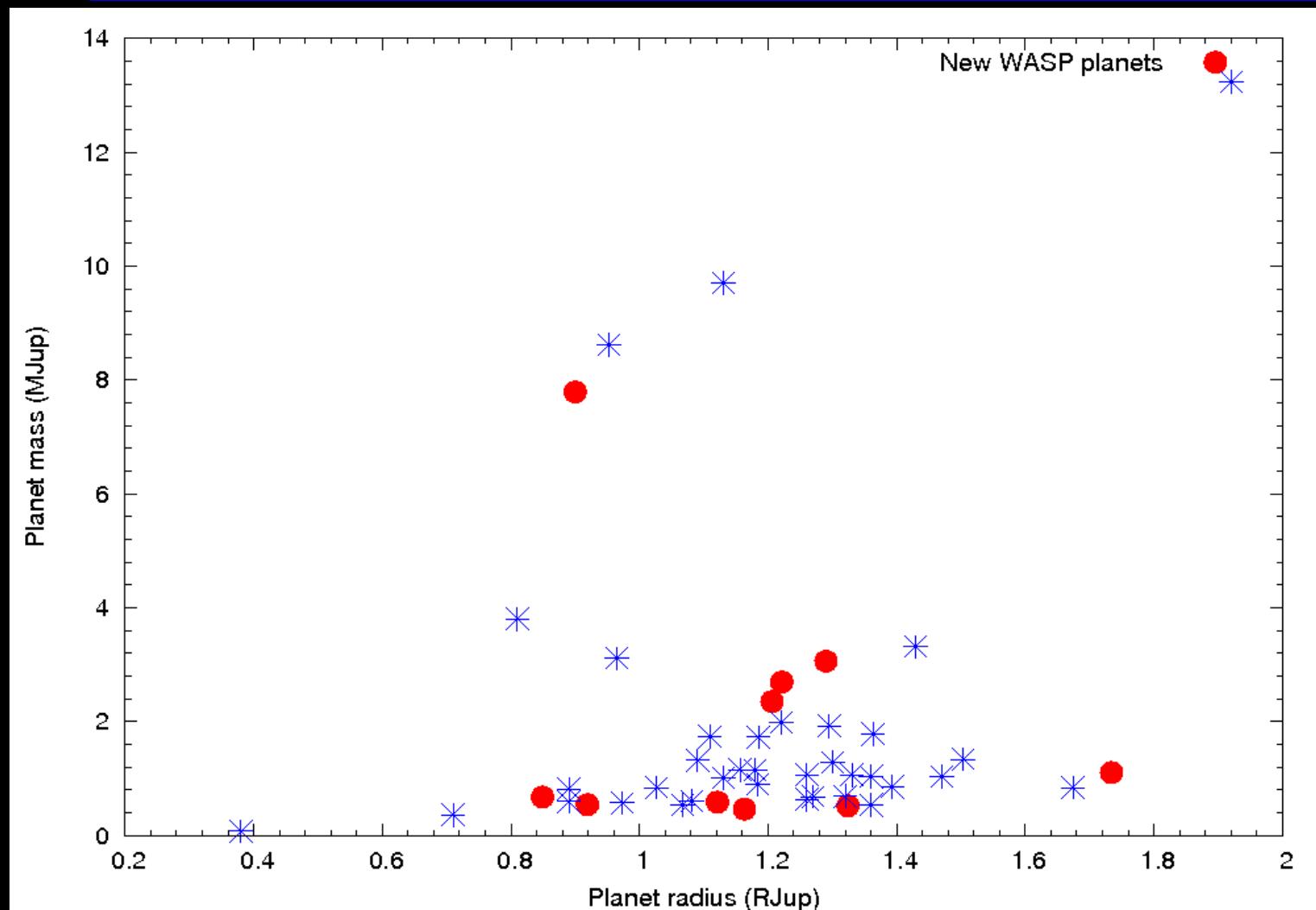


Tenagra B band data, WASP-12  
2008 Mar 27/28

# *The new WASP Planets*



# In Context



# *In Context*

## *“Bloated” planets*

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<u>Planet</u>	<u>Mass/MJ</u>	<u>Radius/RJ</u>	<u>Period/d</u>
TrES-4b	0.84	1.674	3.554
WASP-12b	1.105	1.735	1.091
HD 209458b	0.69	1.32	3.525

- *Burrows et al 2007:*  
Increased planetary atmospheric opacities, retain internal heat
- *Baraffe et al 2003:*  
Irradiation alone cannot explain large radii → extra heat source?
- *Mandushev et al. 2007 (TrES-4b):*  
→ Larger than predicted radius for its orbital separation & age

Burrows et al. 2007 ApJ 661, 502

Burrows et al. 2003, ApJ, 594, 545

Baraffe et al. 2003 A&A 402 701.

Fortney, J. et al. 2007, ApJ, 659, 1661

Mandushev et al. 2007, ApJ, 667, L195

# *In Context*      SuperGiant Planets?

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- Type-II migration (gap forming) more efficient for low-mass stars  
[Ida & Lin 2005 ApJ 626, 1045]

<u>Planet</u>	<u>Mass/MJ</u>	<u>Radius/RJ</u>	<u>Host type</u>
HAT-P-2b	8.62	0.951	F8
WASP-14b	7.790	0.9	F9

[Loeillet et al. 2008 A&A, 481, p.529]

→ Density  $\sim 12.5 \text{ g cm}^{-3}$

Gas giants  $\sim 1 \text{ g cm}^{-3}$

Low mass stars  $> 50 \text{ g cm}^{-3}$

- Different formation mechanism?

# *Future work*

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- New planet papers due out soon, inc. LCOGT data
- Analysis of low-mass binaries
- Improved identification techniques for false positives (blends, triples etc)
- New candidates still being identified from SW-N & SW-S
- 11 promising targets still in queue with some data already
- Continuous dataflow is prompting pipeline development

**...watch this space!**

# *Who cares about planets?*

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- Theories of planet formation revolutionized by exoplanet discoveries
  - Pulsar planets
  - Hot Jupiters
  - Planetary migration
- Distinctions between planets, brown dwarfs and stars
- Astrobiology
  - Looking for other habitable systems